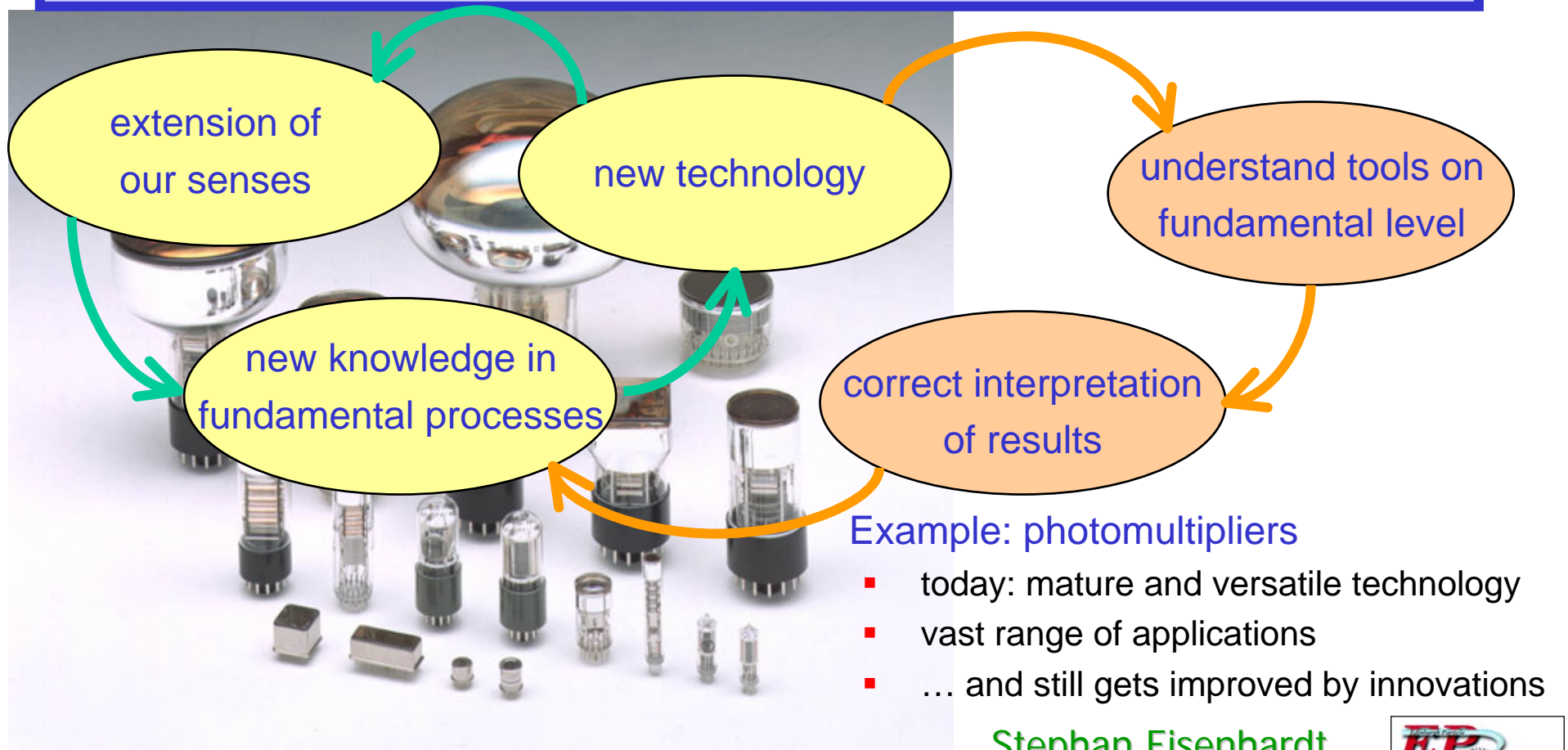


Photomultipliers: Eyes for your Experiment



Edinburgh, 23.07.2003

Stephan Eisenhardt
University of Edinburgh



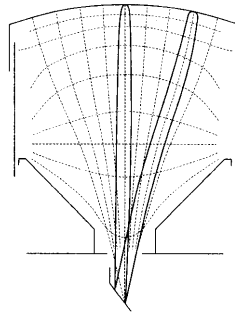
Basic Principle

- Photo emission from photo cathode

$$Q.E. = N_{p.e.}/N_{photons}$$

- Electron collection

- Focusing optics
- optimise efficiency
- minimise transient time spread



Equipotentials and trajectories in a fast input system

- Secondary emission from dynodes:

→ Electric potential

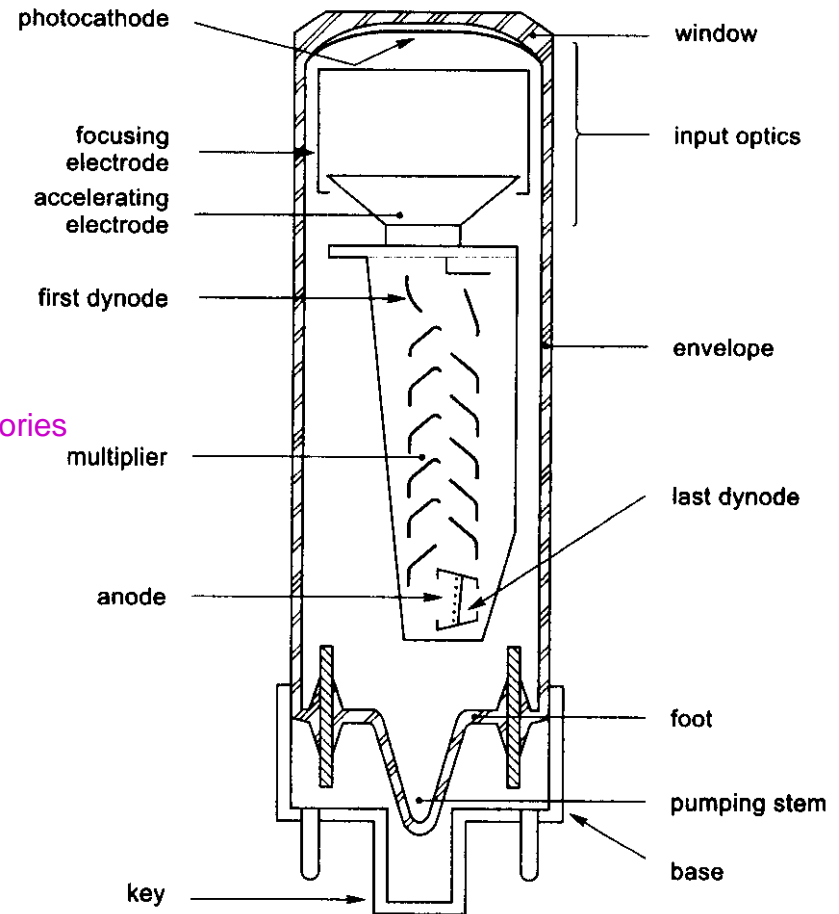
→ Electron multiplication

$$\text{dynode gain: } g(E) = 3 \dots 50$$

- Total gain: $G = \prod g_i$

e.g. 10 dynodes with $g=4$

$$\rightarrow G = 4^{10} \approx 10^6$$



Schematic of Photomultiplier Tube (Philips Photonic)

A Brief History

- 1887: photoelectric effect discovered by Hertz
 - 1902: first report on a secondary emissive surface by Austin et al.
 - 1905: Einstein: "Photoemission is a process in which photons are converted into free electrons."
 - 1913: Elster and Geiter produced a photoelectric tube
 - 1929: Koller and Campbell discovered compound photocathode (Ag-O-Cs; so-called S-1)
 - 1935: Iams et al. produced a triode photomultiplier tube (a photocathode combined with a single-stage dynode)
 - 1936: Zworykin et al. developed a photomultiplier tube having multiple dynode stages using an electric and a magnetic field
 - 1939: Zworykin and Rajchman developed an electrostatic-focusing type photomultiplier tube
 - 1949 & 1956: Morton improved photomultiplier tube structure
- commercial phase, but still many improvements to come

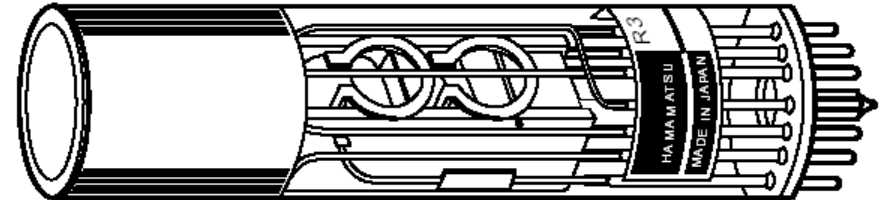


Extension of Vision

□ Human eye:



□ Photomultiplier:



□ Spectral sensitivity:

$400 < S(\lambda) < 750 \text{ nm}$

$110 < S(\lambda) < 1600 \text{ nm}$

□ Time resolution:

$\sim 50 \text{ ms}$

$50 \text{ ps} \dots 10 \text{ ns}$

□ Spatial resolution:

$\sim 100 \text{ lines/mm}$ (2500dpi)

$2 \text{ mm} \dots 50 \text{ cm}$

□ Intensity range:

$O(10^{16} \gamma/\text{mm}^2\text{s})$ (daylight)

$1 \gamma \dots \sim 10^8 \gamma/\text{mm}^2\text{s}$

single photon sensitivity
after adaptation

($\sim 1 \text{ mA}$ anode current for 1" tube)

□ Life time:

$O(10^5 \text{ C/mm}^2 \text{ 70 yrs})$

$O(10 \text{ C})$ for semi-transparent cathode

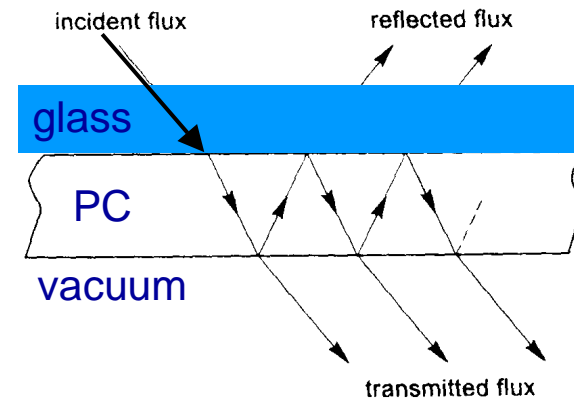
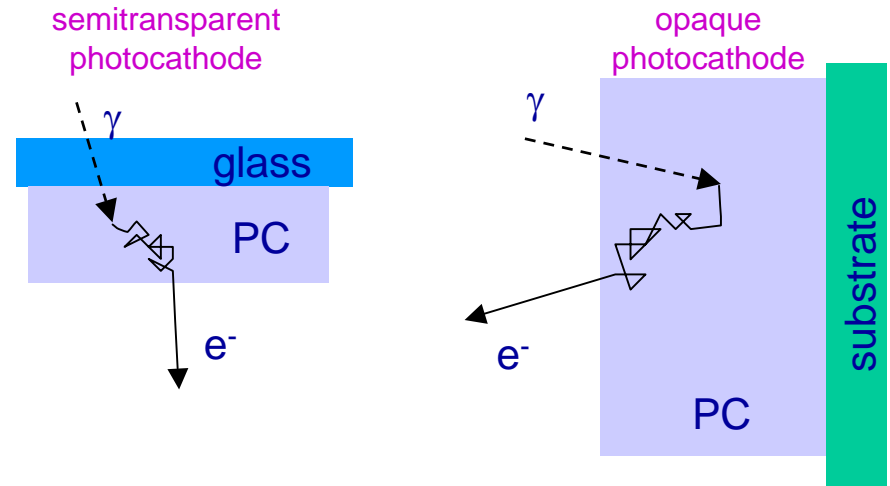
Photoemission

- 2-step process:
 - photo ionisation
 - escape of electron into vacuum

- Multi-reflection/interference due to high refractive index

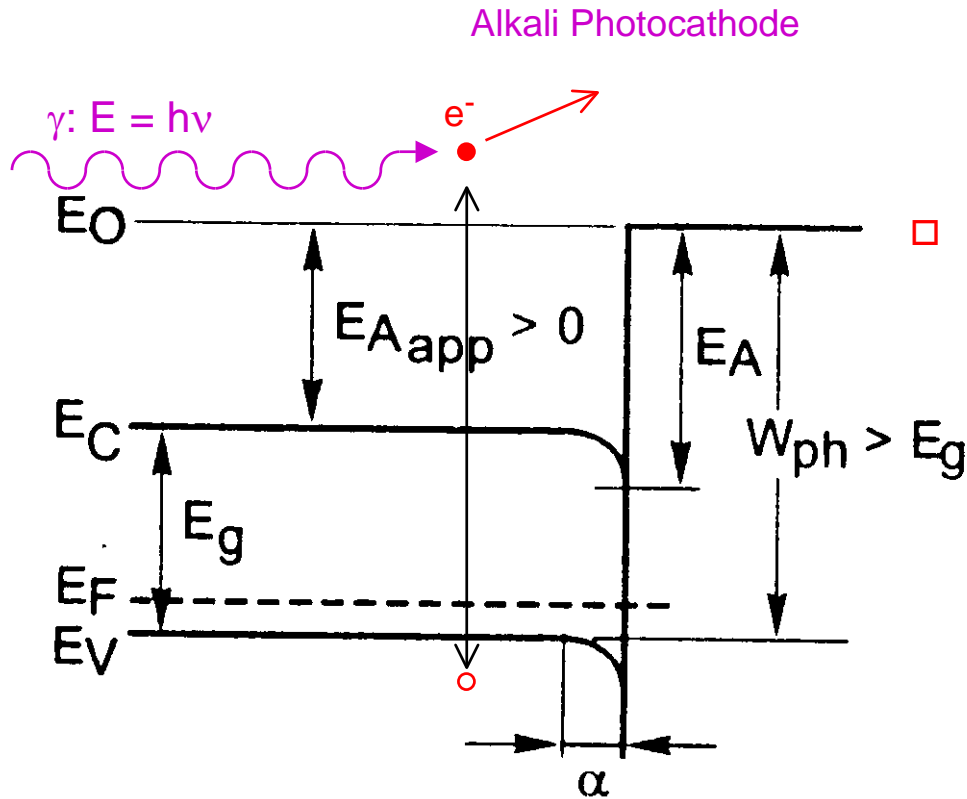
alkali: $n(\lambda = 442 \text{ nm}) = 2.7$

- Q.E. difficult to measure
- often only an effective detection efficiency determined:
 - + internal reflection from a metallic surface
 - + collection of the photoelectrons
 - + electronic threshold...



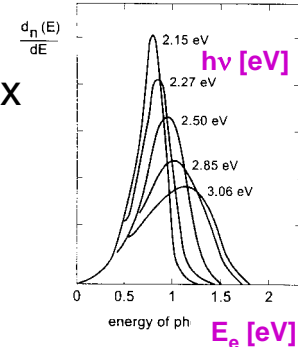
Spectral Response

- In the band model:



- Photo electric effect

- $h\nu > E_g + E_A + dE/dx$
- $E_e = h\nu - W - dE/dx$



- Spectral response:

- Quantum efficiency

$$QE(\lambda) = \frac{N_{p.e.}}{N_\lambda} = (1 - R) \frac{P_v}{\alpha} P_s \left(\frac{1}{1 + l/\alpha d} \right)$$

R = reflection coefficient

$\alpha = \gamma$ absorption coefficient

$l = e^-$ mean escape length

P_v = excitation probability

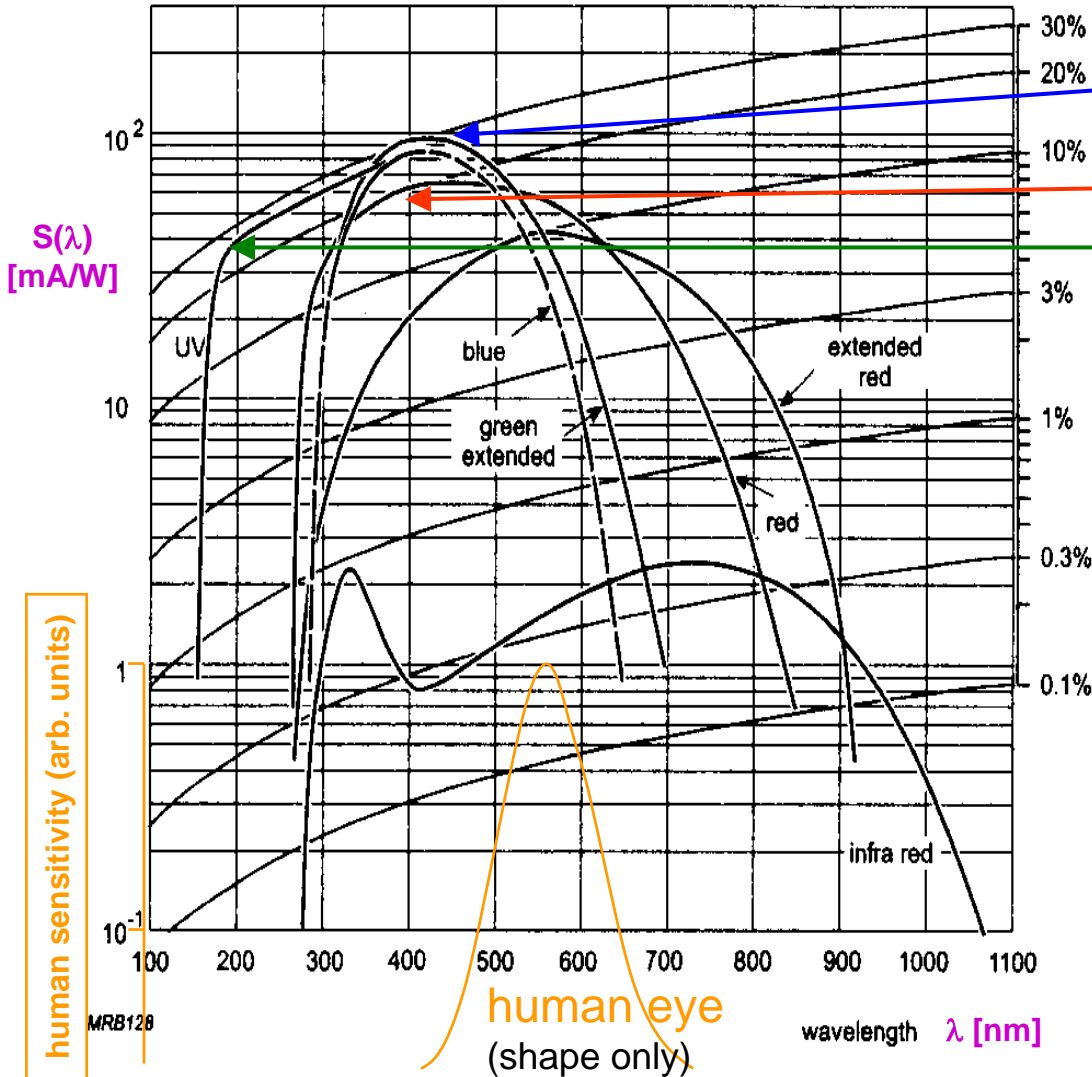
P_s = surface transition probability

- Cathode sensitivity

$$S(\lambda) = \frac{I_c}{P(\lambda)} = \frac{\lambda [\text{nm}] QE(\lambda)}{1240} [\text{A/W}]$$

Alkali Photocathodes

(Philips Photonic) semitransparent photocathodes

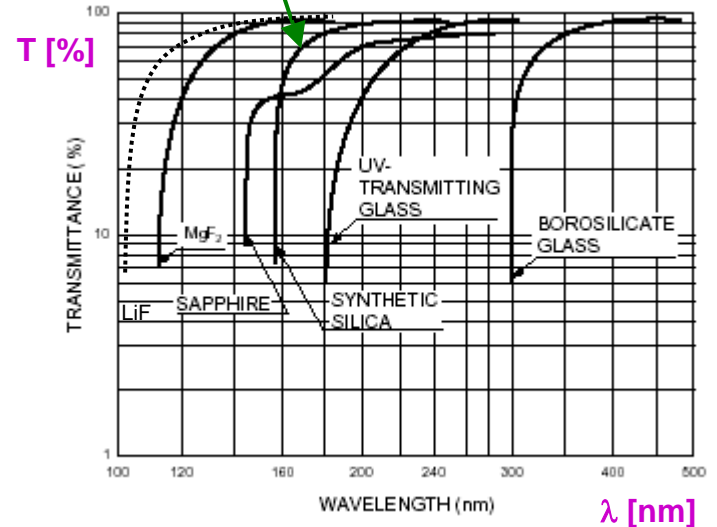


Q.E.

Bialkali : SbK_2Cs , SbRbCs

Multialkali : SbNa_2KCs

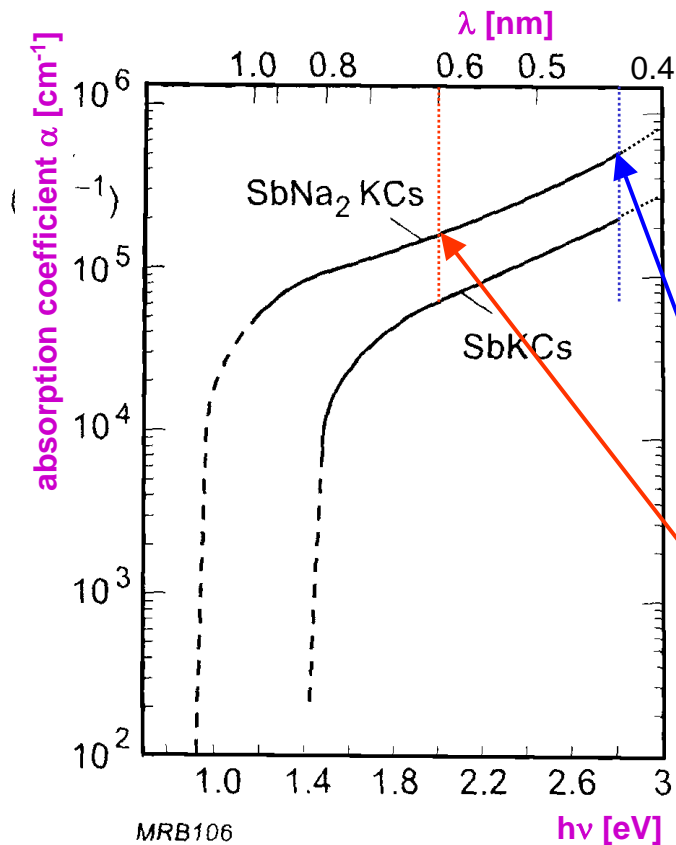
Solar blind : CsTe
(cut by quartz window)



transmittance of window materials

Photocathode Thickness

Blue light is stronger absorbed than red light!



$$\lambda_A = 1/\alpha$$

$$\alpha \approx 4 \cdot 10^5 \text{ cm}^{-1}$$

$$\lambda_A \approx 25 \text{ nm}$$

$$\alpha \approx 1.5 \cdot 10^5 \text{ cm}^{-1}$$

$$\lambda_A \approx 60 \text{ nm}$$

- semi-transparent cathodes
- best compromise for the thickness of the PC:

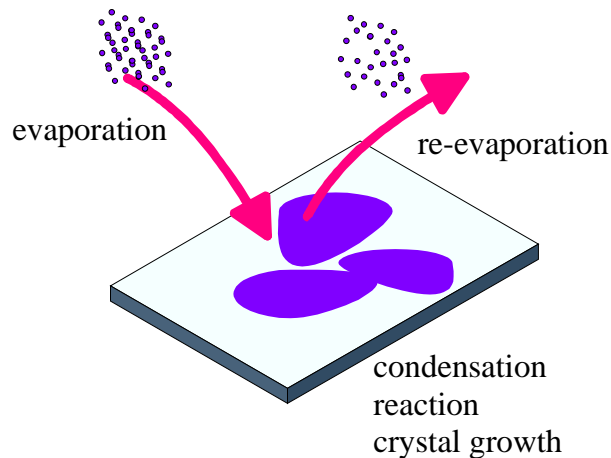
↗ photon absorption length $\lambda_A(E_{\text{ph}})$

↘ electron escape length $\lambda_E(E_e)$

- Q.E. of thick cathode:
 - red response ↗
 - blue response ↘
- Q.E. of thin cathode:
 - blue response ↗
 - red response ↘

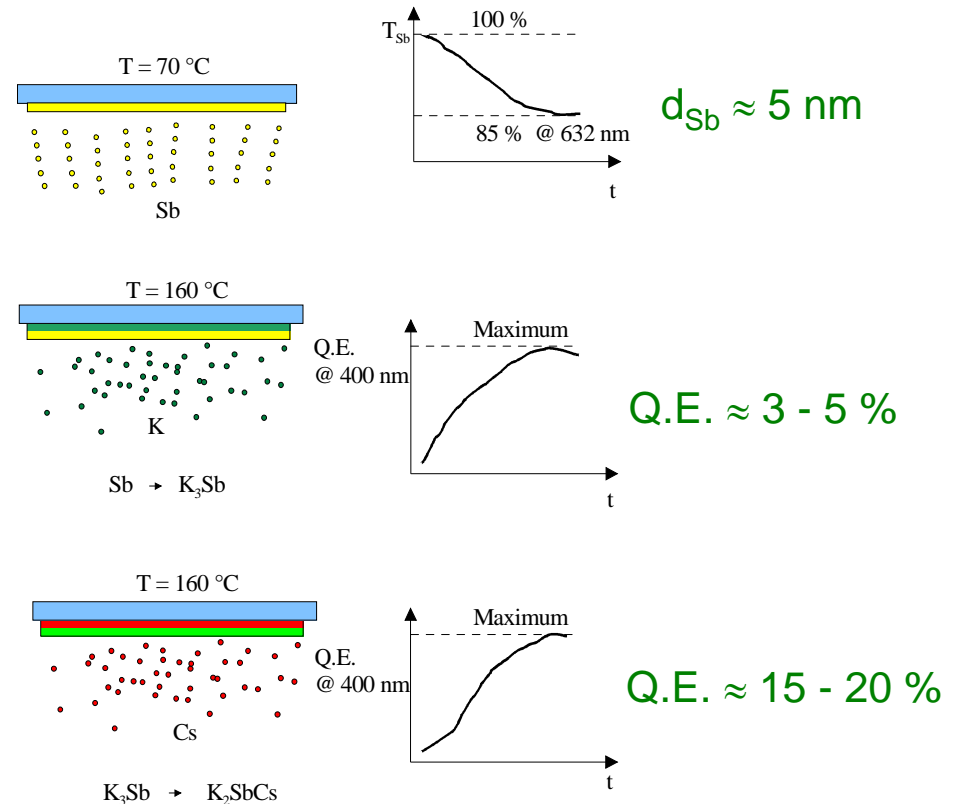
Alkali Photocathode Production

- evaporation of metals in high vacuum
 - $< 10^{-7}$ mbar
 - $< 10^{-9}$ mbar H_2O partial pressure
 - no other contaminants (CO , $C_xH_y...$)
 - bakeout of process chamber ($>150^\circ C$) and substrate ($>300^\circ C$)
- condensation of vapour and chemical reaction on entrance window



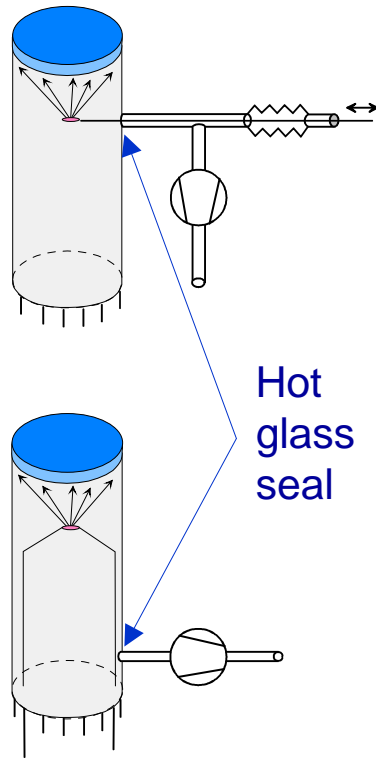
→ relatively simple technique

□ Example: SbK_2Cs
simplified sequential bialkali process

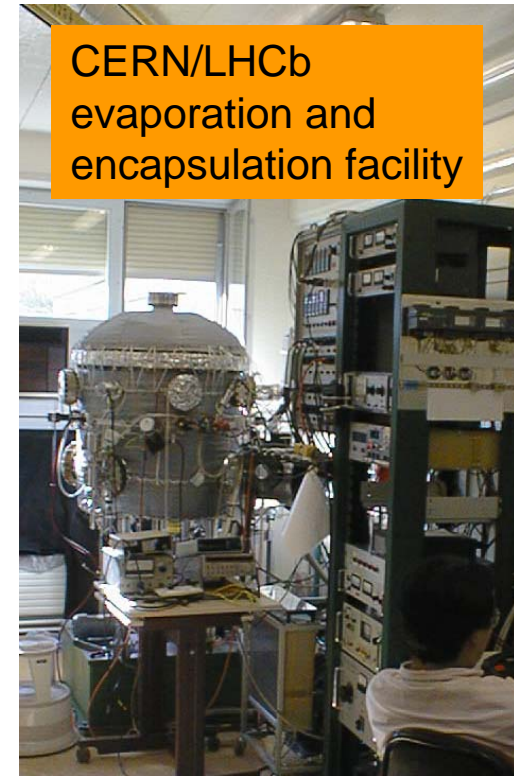
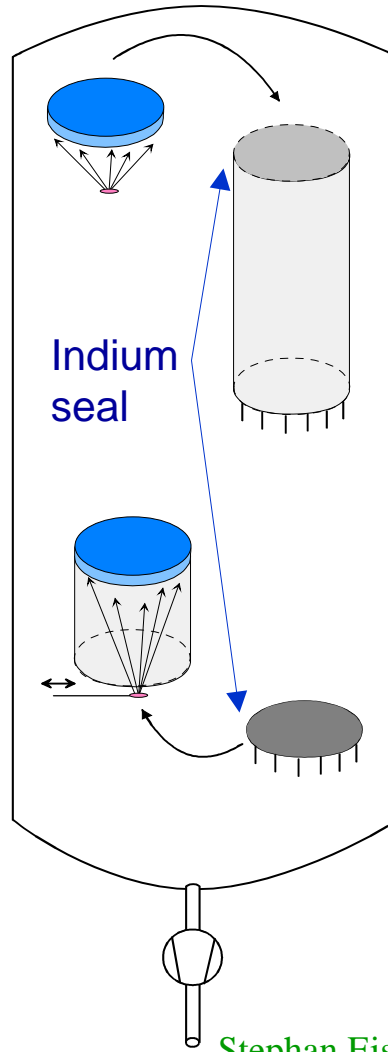


Phototube Fabrication

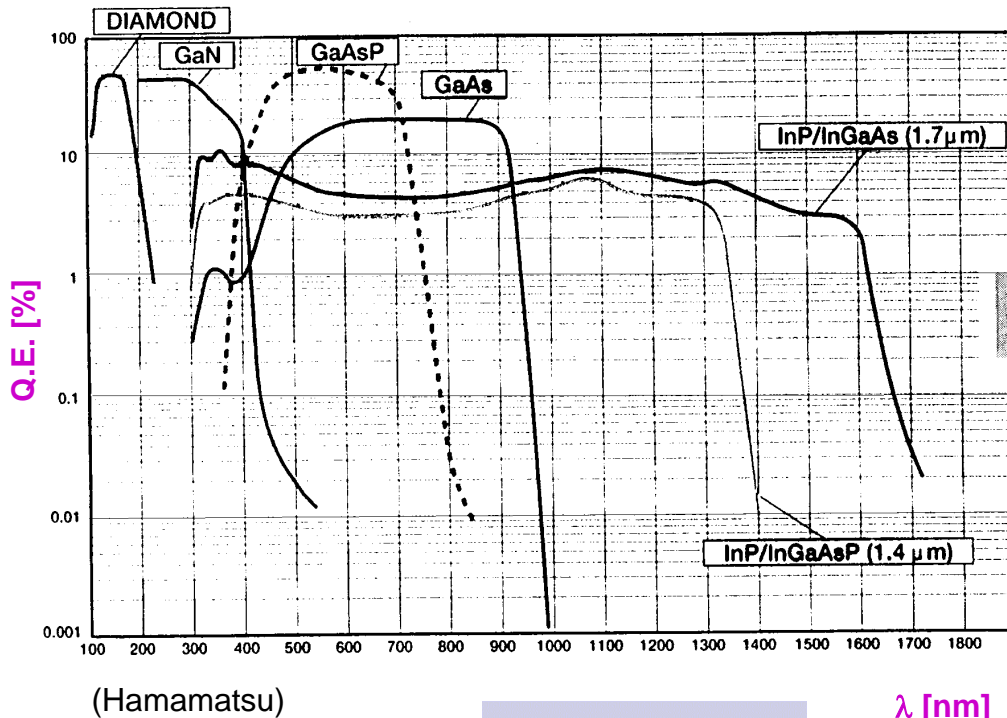
□ internal:



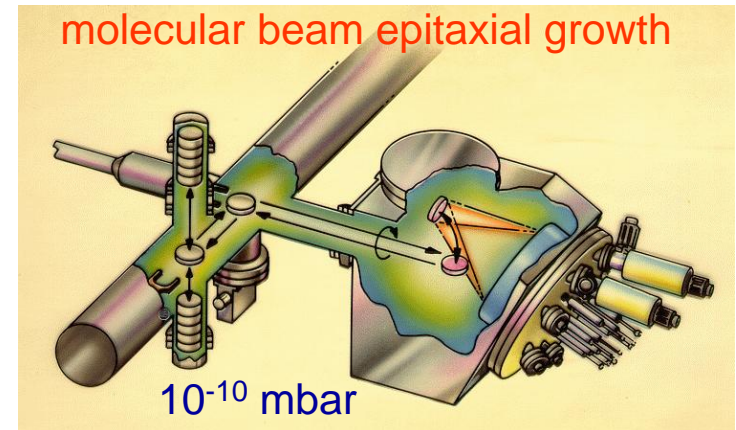
□ external:



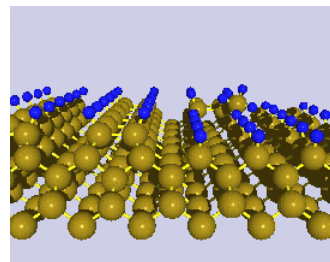
Semiconductor Photocathodes



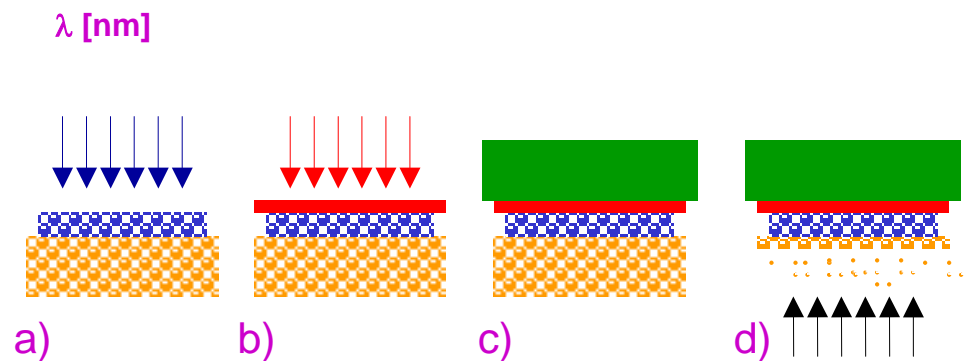
- ☐ 😊 high Q.E. and spectral width
- ☐ 😊 negative electron affinity
- ☐ ☹️ complex production



a) grow PC on crystalline substrate



- b) create interface layer
c) fuse to entrance window
d) etch substrate away



Secondary Emission

- alloy of alkali or earth-alkali and noble metal
- alkaline metal oxidises → insulating coating

- Large gain Gain $\delta = a(\Delta V)^k$ with $k = 0.7...0.8$
- Stability for large currents
- Low thermal noise

$$I = aT^2 \exp\left(\frac{-e\Phi}{kT}\right)$$

- Statistical process: Poisson distribution

$$P(\bar{n}, m) = \frac{\bar{n}^m e^{-m}}{m!}$$

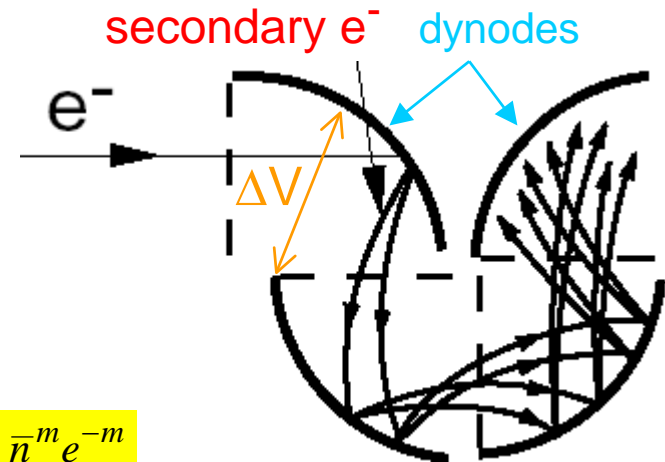
- Spread (RMS)
- Largest at 1st and 2nd dynode

$$\frac{\sigma_n}{\bar{n}} = \frac{\sqrt{\bar{n}}}{\bar{n}} = \frac{1}{\sqrt{\bar{n}}}$$

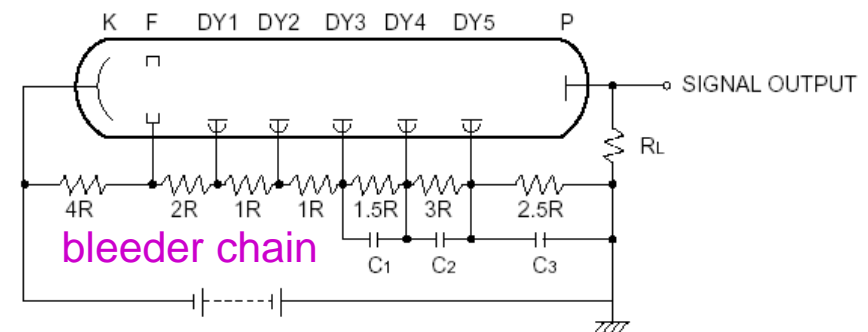
- typically 10 ... 14 dynode stages

- linearity limits:

- pulse: space charge
- DC: photo current \ll bleeder current



electron multiplication

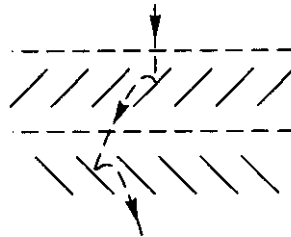


scheme of external circuit for dynode potentials

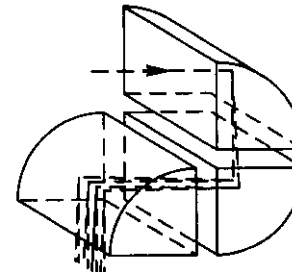
Classic Dynodes

- uniformity:
- photoelectron collection efficiency:

- simple design
- good for large PC \emptyset

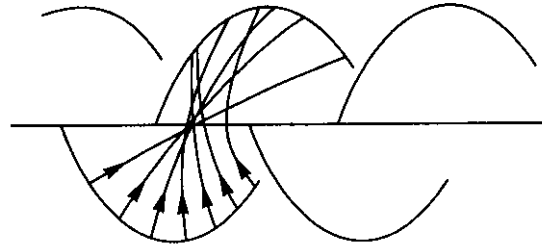


venetian blind 😊 😞

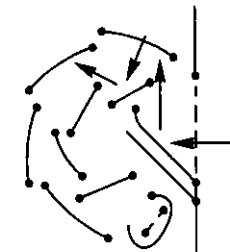


box and grid 😊 😊 😊

- excellent linearity
- good time resolution
- fast time response



linear focusing 😞 😊



circular cage 😞 😊

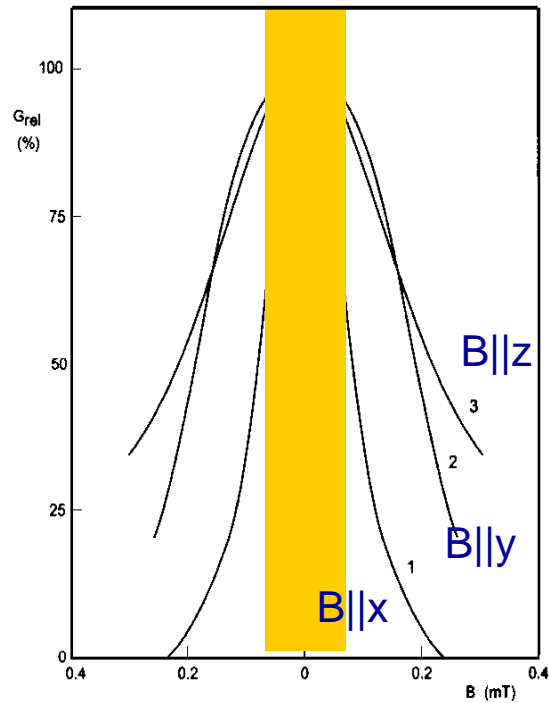
- best photoelectron collection efficiency
- good uniformity

- compact
- fast time response

sensitive to Earth B-field (30-60 μ T)!
no spatial resolution

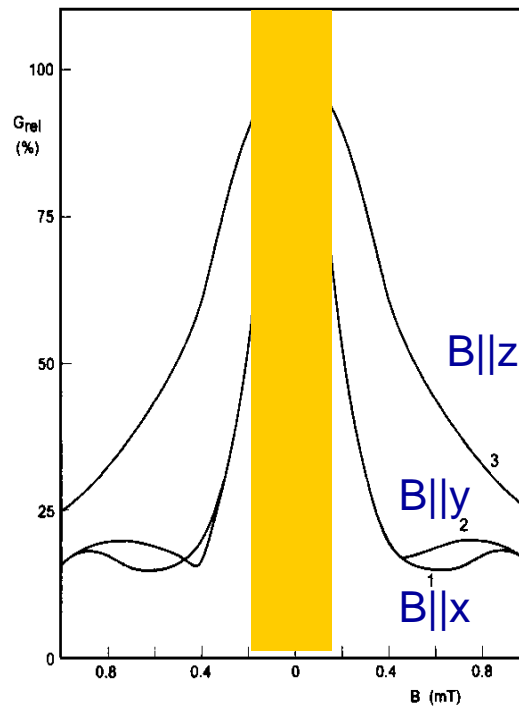
Sensitivity to Magnetic Fields

□ linear focusing

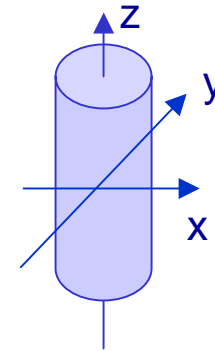


$\pm 50 \mu T$
 ± 0.5 Gauss

□ venetian blind



$\pm 200 \mu T$
 2 Gauss

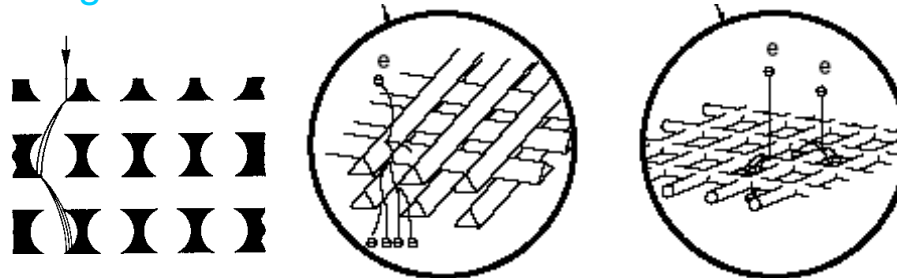


- Earth field: $30-60 \mu T$
- requires μ -metal shielding

Modern Dynodes

- ❑ B-field immunity up to 1.2T B-field
- ❑ spatial resolution via segmented anode

- ❑ good uniformity
- ❑ poor e⁻ collection efficiency!



foil 😊 😞 proximity mesh 😊 😞 ❑ excellent linearity

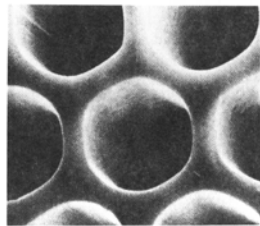
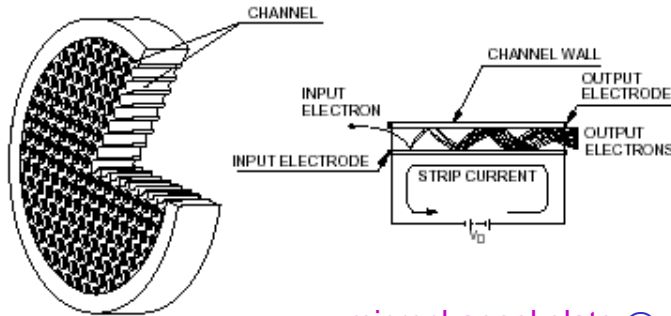
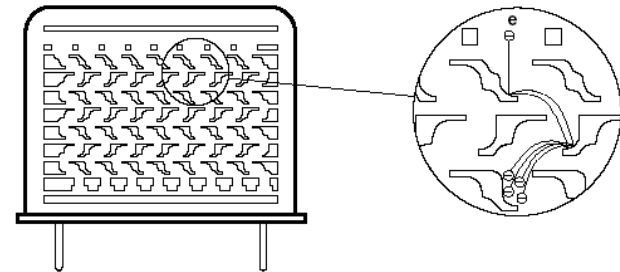


Fig. 5.6. Microphotograph of microchannel



microchannel plate 😊 😞

- ❑ excellent time resolution: transit time spread: 50ps

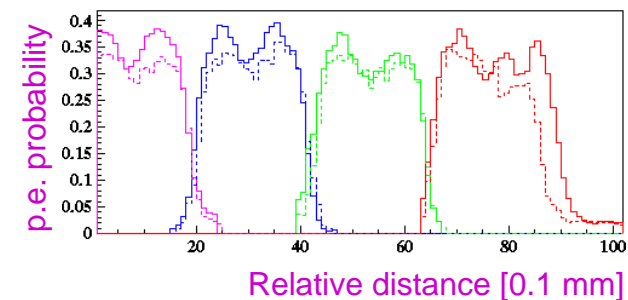
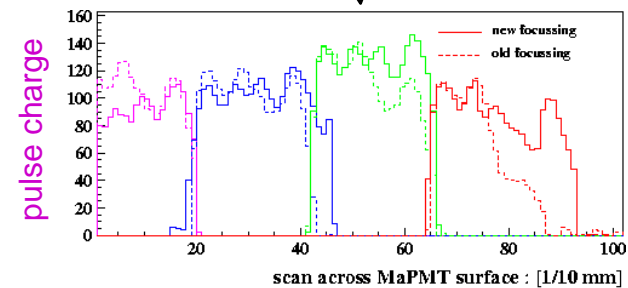
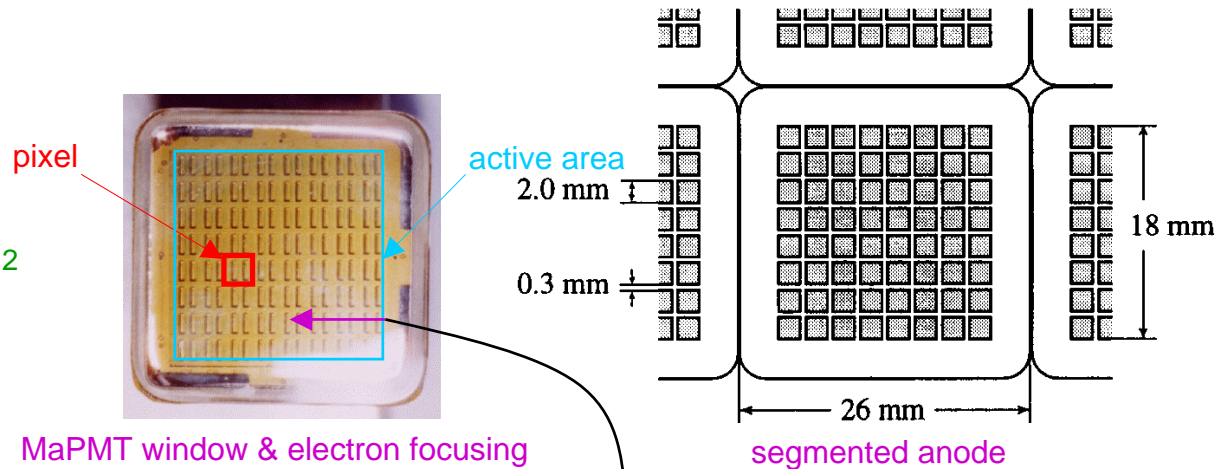


metal channel 😊 😊

- ❑ good e⁻ collection efficiency
- ❑ excellent time characteristics
- ❑ stable gain
- ❑ ...up to 20mT
- ❑ low cross-talk pixels

Multianode PMT

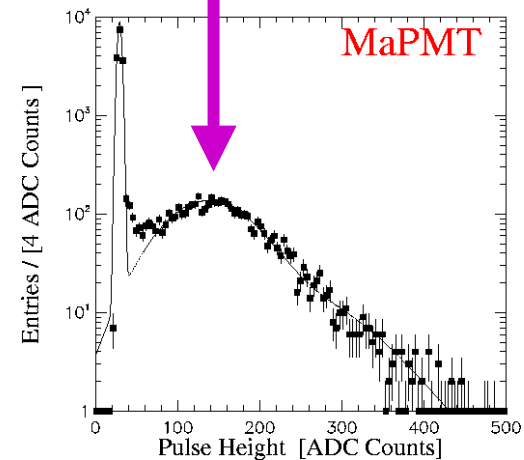
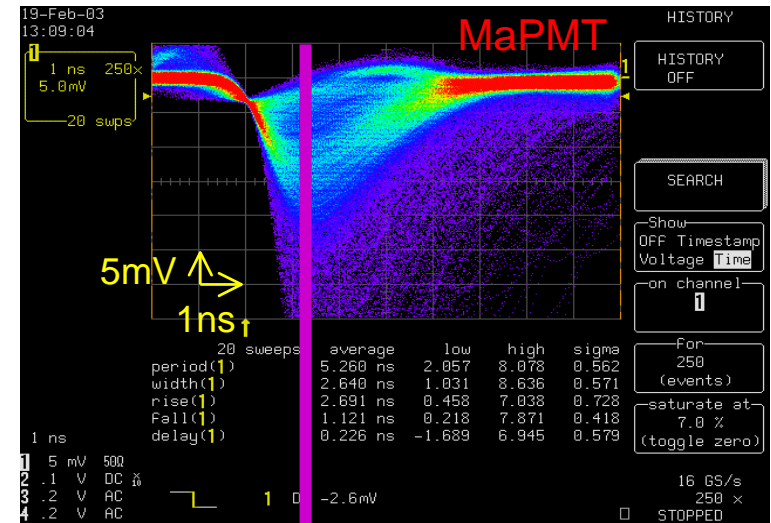
- Position sensitive PMT
 - 8x8 metal channel dynode chains in one vacuum envelope (26x26 mm²)
 - segmented anode: 2x2 mm²
 - active area fraction: 48%
- UV glass window
- Bialkali photo cathode
 - QE = 22...25% at $\lambda = 380$ nm
- Gain
 - $3 \cdot 10^5$ at 800 V
- Uniformity, Crosstalk
 - much improved
- Applications:
 - medical imaging
 - HERA-B, LHCb: Ring Imaging Cherenkov counters



PMT Characteristics

- **Fluctuations**
 - number of secondary electrons
 - Poisson distribution
- **Saturation**
 - space charge
 - large photon current
- **Non-linearity**
 - at high gains
- **Stability**
 - drift, temperature dependency
 - fatigue effects
 - ➔ Monitoring
- **Sensitive to magnetic fields**
 - Earth: 30-60 μT
 - ➔ requires μ -metal shielding

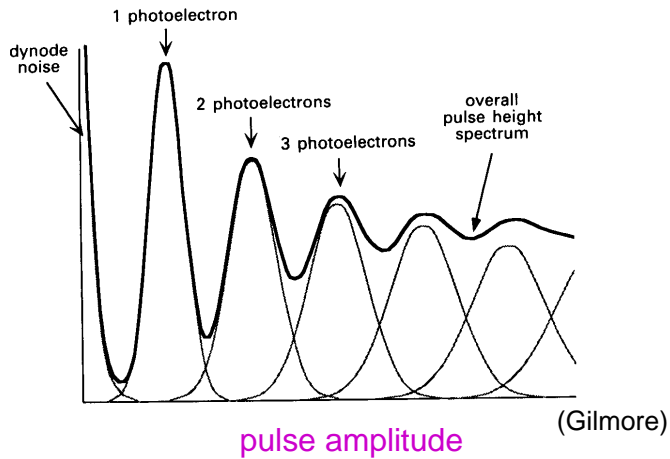
- single photon events to oscilloscope (50 Ω)



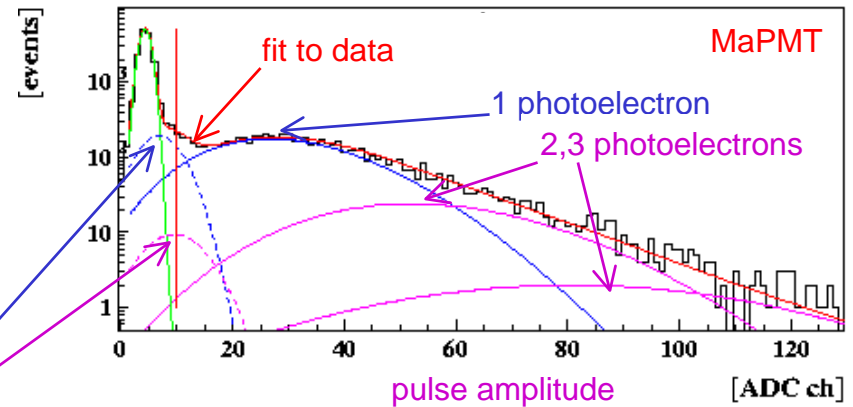
charge integration \rightarrow pulse height spectrum

Pulse Height Spectrum

- first dynode gain: $\delta = 25$
 - clear separation of $n\gamma$ signals
 - Gaussian shape
 - Poisson for γ probability



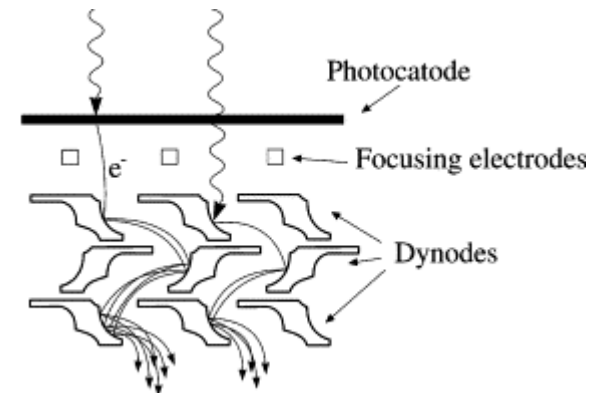
- first dynode gain: $\delta = 4$
 - broad signal distributions
 - Poisson shape for $n_{pe}=1,2$
 - significant signal loss below threshold cut



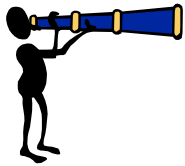
- Photo conversion possible at 1st dynode

- Gain reduced by δ_1
- ~~- explains data between noise and 1 p.e.~~

- contradiction when HV is changed!!
 - this data is contaminated with incompletely sampled charge pulses



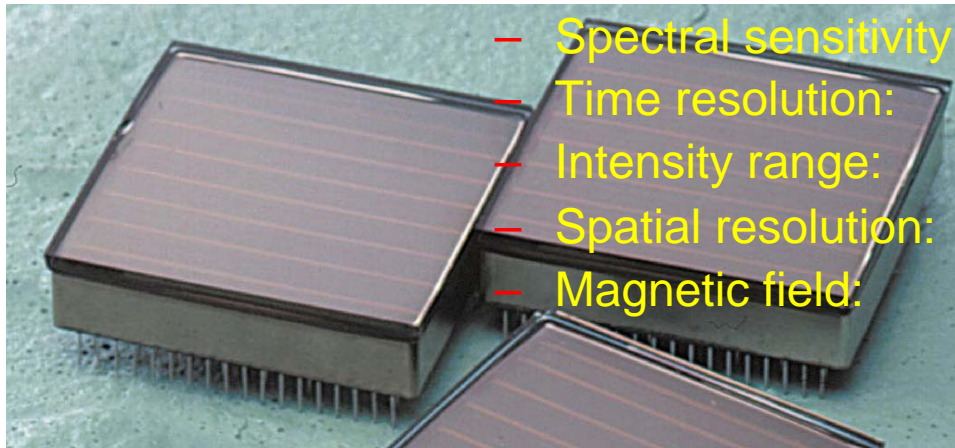
Conclusion



- know your tools
- don't fool yourself with immature conclusions
- always cross-check as far as possible



- photomultipliers are a mature and versatile technology
- suitable for a vast range of applications in light detection:



- Spectral sensitivity: $110 < S(\lambda) < 1600 \text{ nm}$
- Time resolution: down to 50ps
- Intensity range: $1\gamma \dots$ anode current $O(1 \text{ mA})$
- Spatial resolution: down to 2mm with low cross-talk
- Magnetic field: up to 1.2T

flat panel PMT: next generation MaPMT

- development goes on...



Super-Kamiokande: 20" tubes