

Super/Hyper-Kamiokande

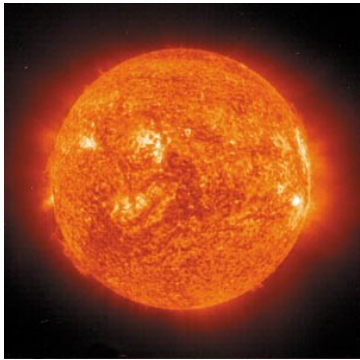
# T2K Latest Results & Hyper-Kamiokande Status.

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University of Edinburgh  
4 February 2014

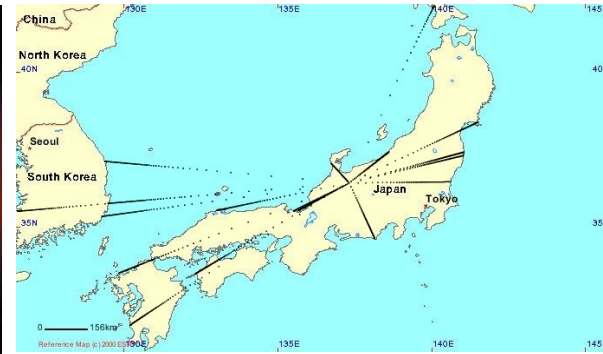


# Neutrino Oscillations: Established

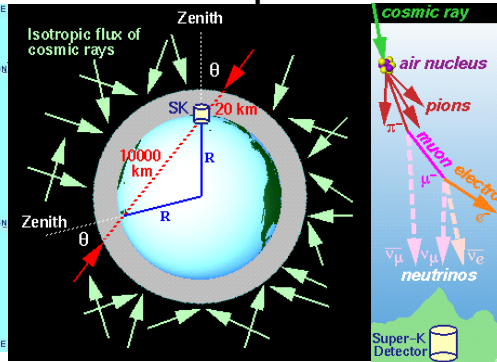
Sun



Reactors



Atmosphere



Accelerators



Homestake,  
SAGE, GALLEX,  
Super-K, SNO,  
Borexino

KamLAND, Chooz

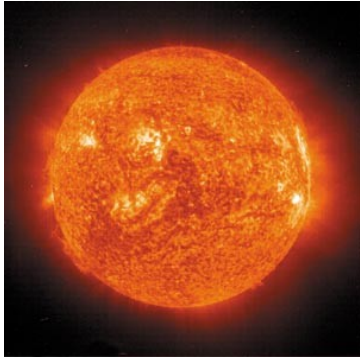
Super-Kamiokande

K2K, MINOS, T2K

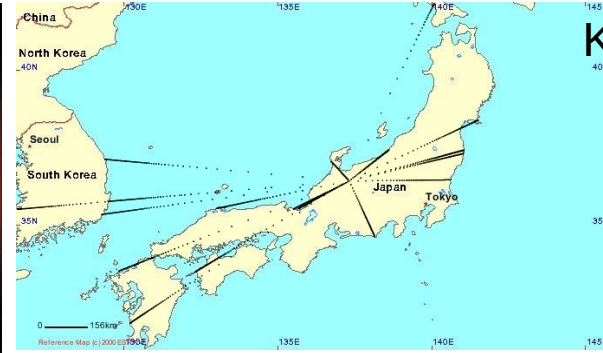
$\nu_{\mu} \rightarrow \nu_{\tau}$ or $\text{anti-}\nu_{\mu} \rightarrow \text{anti-}\nu_{\tau}$	atmospheric & beam experiments
$\nu_e \rightarrow \nu_{\mu, \tau}$	solar experiments
$\text{anti-}\nu_e \rightarrow \text{anti-}\nu_{\text{other}}$	reactor experiments
$(\text{anti-})\nu_{\mu} \rightarrow (\text{anti-})\nu_{\text{other}}$	atmospheric & beam experiments
$\nu_{\mu} \rightarrow \nu_e$	beam experiments

# Neutrino Oscillations: Established

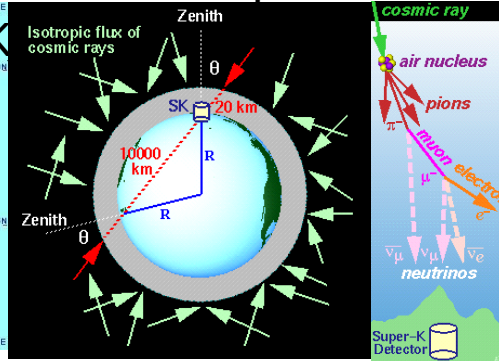
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Super-Kamiokande

PMNS (Pontecorvo, Maki, Nagakawa, Sakata) matrix for  $\nu$ :

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

Flavour eigenstates  
(coupling to the W)

Mass eigenstates  
(definite mass)

Unitary PMNS mixing matrix

# Open Questions in $\nu$ Oscillations (I)

$$U_{PMNS} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & +c_{23} & +s_{23} \\ 0 & -s_{23} & +c_{23} \end{pmatrix} \begin{pmatrix} +c_{13} & 0 & +s_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13} e^{i\delta} & 0 & +c_{13} \end{pmatrix} \begin{pmatrix} +c_{12} & +s_{12} & 0 \\ -s_{12} & +c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$\theta_{23}$ : "atm." mixing angle  
 ( $c_{ij} = \cos \theta_{ij}, s_{ij} = \sin \theta_{ij}$ )

Unknown:  $-i\delta$   
 $\theta_{13}$                        $\theta_{12}$ : "solar" mixing angle

Free parameters:  $\theta_{12}, \theta_{23}, \theta_{13}, \delta$

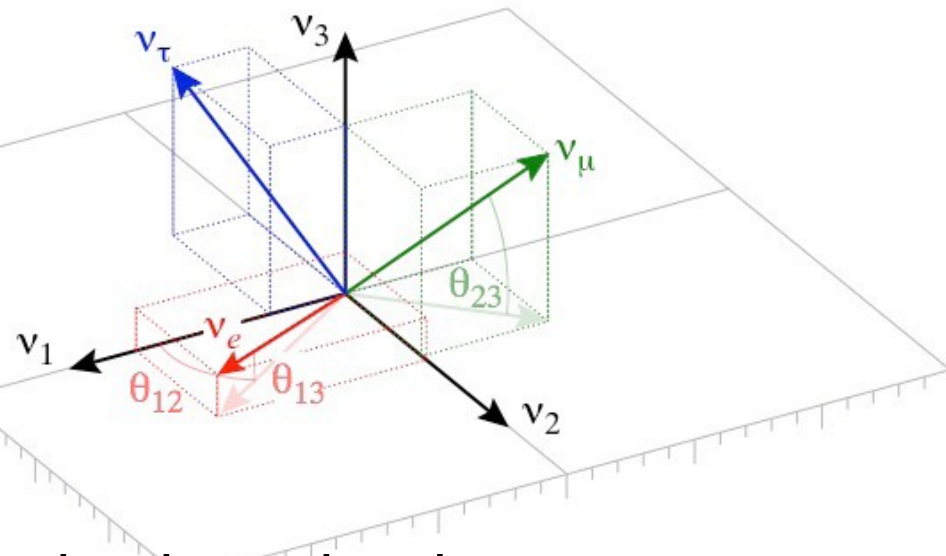
$$\theta_{12} = 34^\circ \pm 1^\circ$$

$$\theta_{23} = 40^\circ + 5^\circ / -2^\circ$$

How close to  $45^\circ$ ?

$$\theta_{13} = 9.1^\circ \pm 0.6^\circ$$

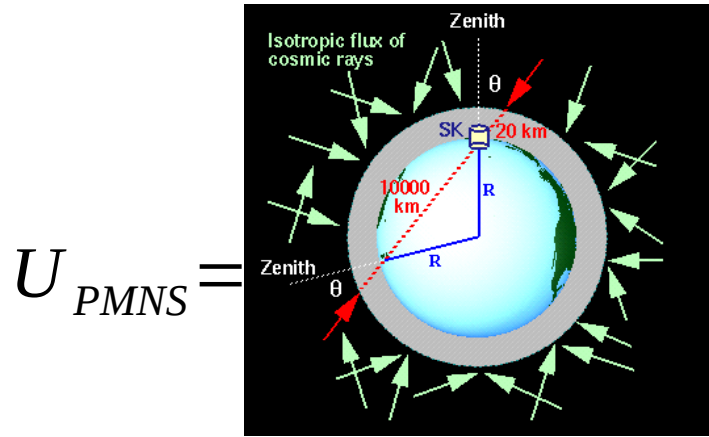
Reactor experiments



$\delta$  is the only unknown parameter

Key to understand the origin of matter-dominated Universe (Leptogenesis)

# Open Questions in $\nu$ Oscillations (I)



$$U_{PMNS} =$$

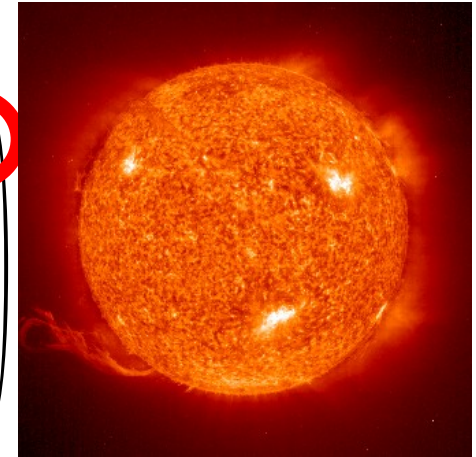
$$\begin{pmatrix} +c_{13} & 0 & +s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & +c_{13} \end{pmatrix}$$

$\theta_{23}$ : "atm." mixing angle  
 $(c_{ij} = \cos \theta_{ij}, s_{ij} = \sin \theta_{ij})$

Unknown!

$$\begin{pmatrix} 0 & +s_{13}e^{-i\delta} \\ 1 & 0 \\ 0 & +c_{13} \end{pmatrix}$$

$\theta_{13}$



$\theta_{12}$ : "solar" mixing angle

Free parameters:  $\theta_{12}, \theta_{23}, \theta_{13}, \delta$

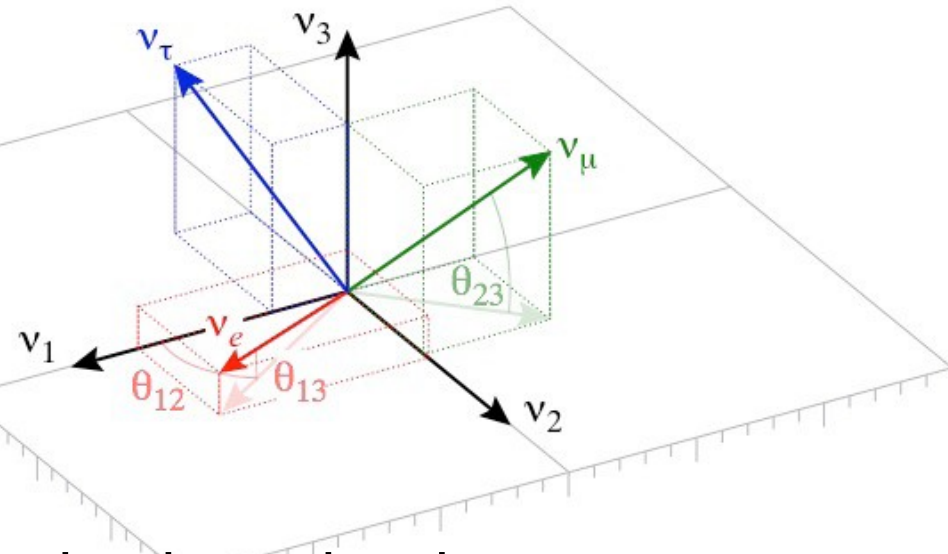
$$\theta_{12} = 34^\circ \pm 1^\circ$$

$$\theta_{23} = 40^\circ + 5^\circ / -2^\circ$$

How close to 45°?

$$\theta_{13} = 9.1^\circ \pm 0.6^\circ$$

Reactor experiments



$\delta$  is the only unknown parameter

Key to understand the origin of matter-dominated Universe (Leptogenesis)

# Neutrino Flavour Mixing

As neutrino propagate, the mass eigenstates interfere:

$$P_{\alpha\beta} = \delta_{\alpha\beta} - 4 \sum_{i>j} \text{Re}[U_{\beta i} U_{\alpha i}^* U_{\beta j}^* U_{\alpha j}] \sin^2 \left( \frac{\Delta m_{ij}^2 L}{4E} \right) + 2 \sum_{i>j} \text{Im}[U_{\beta i} U_{\alpha i}^* U_{\beta j}^* U_{\alpha j}] \sin \left( \frac{\Delta m_{ij}^2 L}{2E} \right)$$

Probability to observe  $\nu_\beta$  after starting with flavour  $\nu_\alpha$  depends on:

- Mixing matrix elements  $U_{\alpha\beta}$
- $L(\text{km})$ : Distance the neutrino has travelled
- $E(\text{GeV})$ : Energy of the neutrino
- $\Delta m_{ij}^2 (\text{eV}^2) = m_i^2 - m_j^2$  : mass splitting

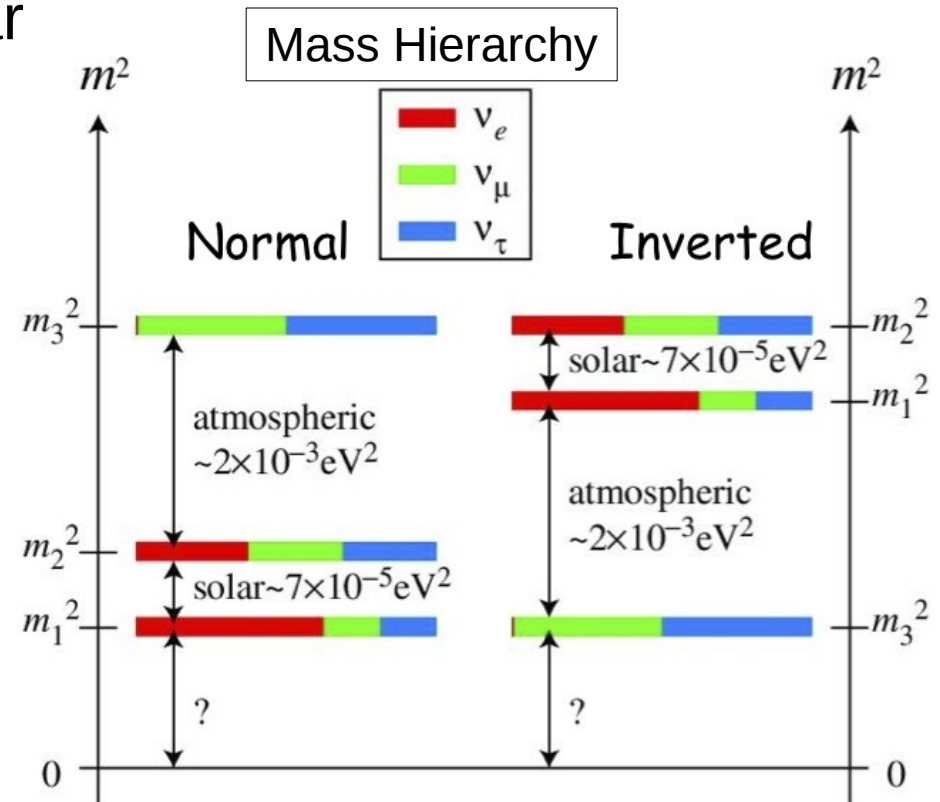
If neutrinos have no mass, or degenerate masses, no interference is possible.

# Open Questions in $\nu$ Oscillations (II)

Two observed mass splittings, determined from atmospheric and solar neutrino experiments, respectively:

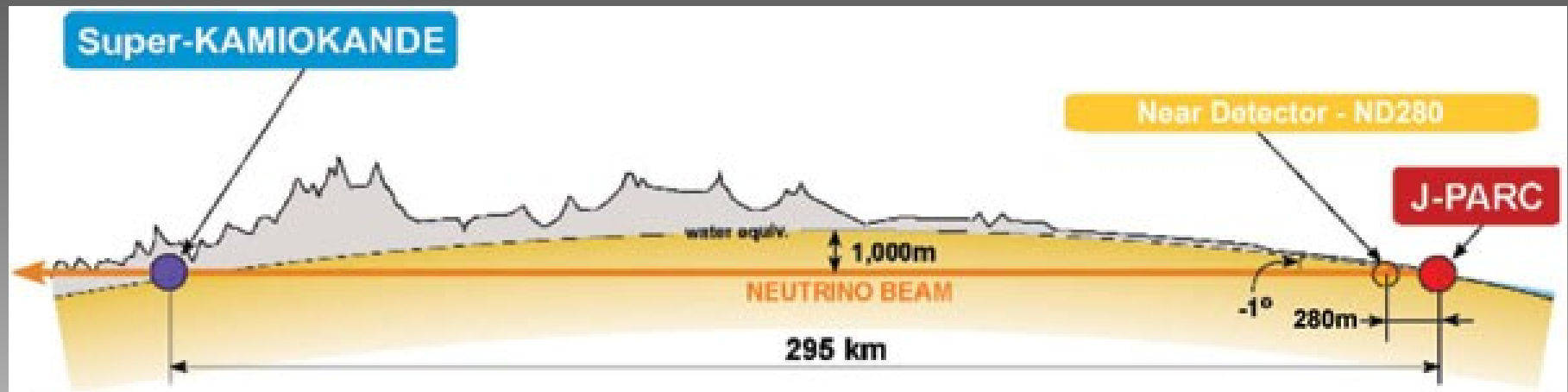
- $\Delta m^2$  (atmospheric) =  $|\Delta m^2_{32}| \sim 2.4 \times 10^{-3} \text{ eV}^2$
- $\Delta m^2$  (solar) =  $\Delta m^2_{21} \sim 7.6 \times 10^{-5} \text{ eV}^2$

The sign of  $|\Delta m^2_{32}|$ , or the “mass hierarchy” is still unknown:

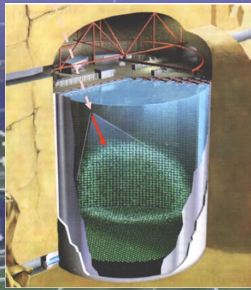


- *Normal* mass hierarchy is like quarks ( $m_1$  is lightest,  $\Delta m^2_{32} > 0$ )
- *Inverted* mass hierarchy has  $m_3$  lightest ( $\Delta m^2_{32} < 0$ )

# The T2K Experiment



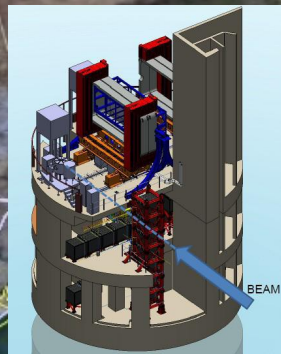




**Super-Kamiokande  
(Kamioka)**



**J-PARC  
(Tokai)**

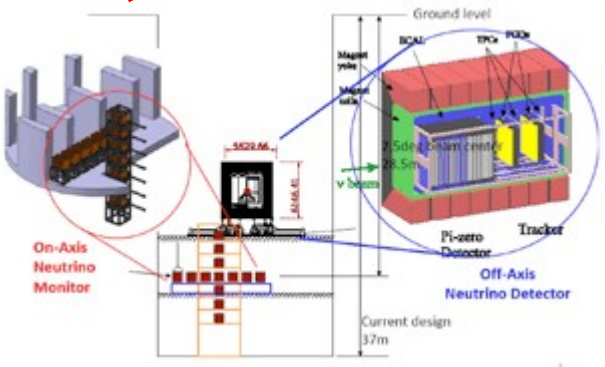


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# Japan Proton Accelerator Research Complex $\nu$ -Beamline

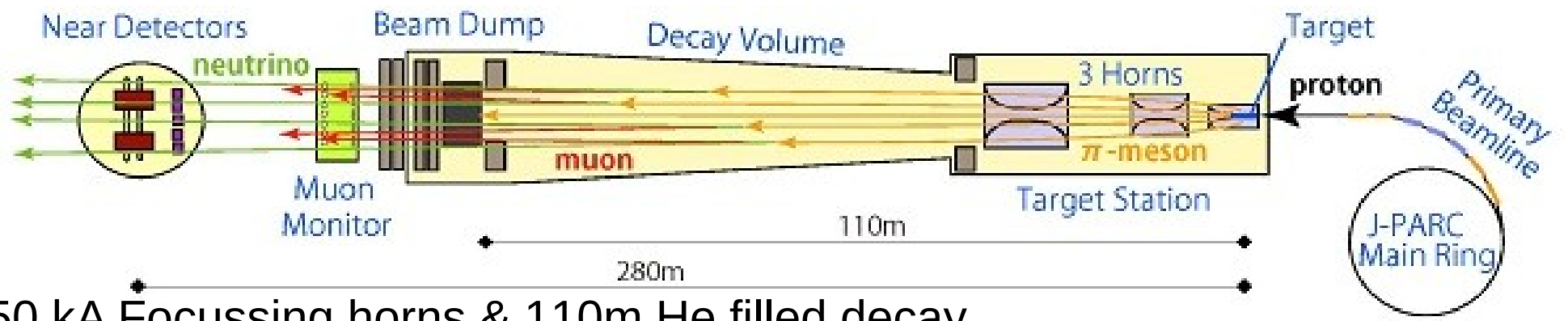


INGRID  
on-axis

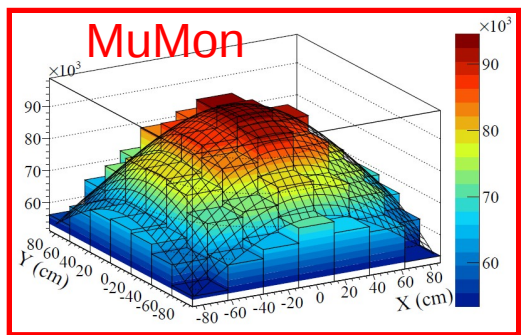


ND280  
off-axis

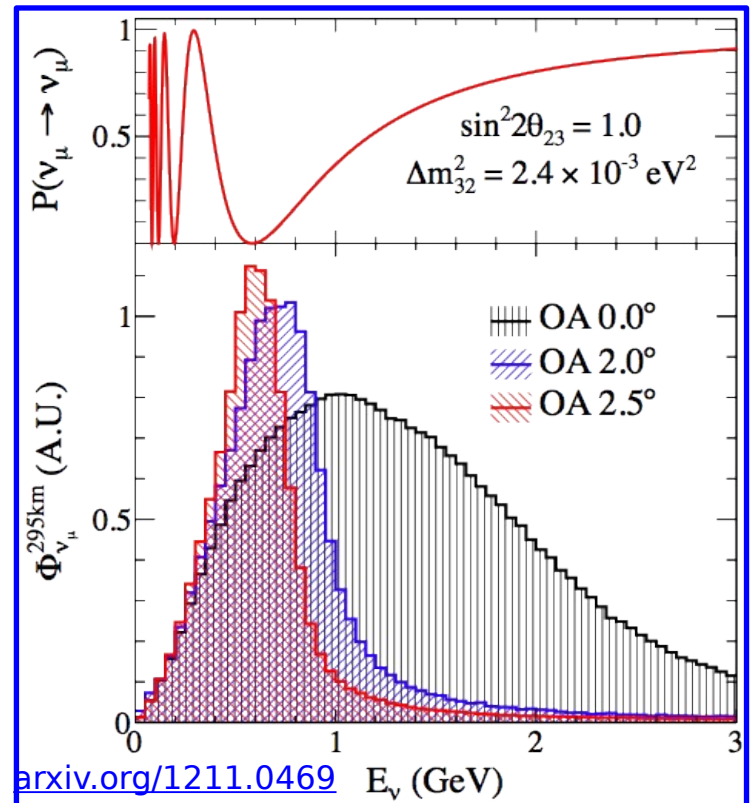
# Japan Proton Accelerator Research Complex $\nu$ -Beamline



- 250 kA Focussing horns & 110m He filled decay volume
- Series of beam monitors, **MuMon** muon monitor and **INGrid** near detector monitors beam centre
  - 2.5 $\sigma$  off-axis configuration
  - Reduces peak energy to oscillation maximum
  - Reduces spread of energies around peak.



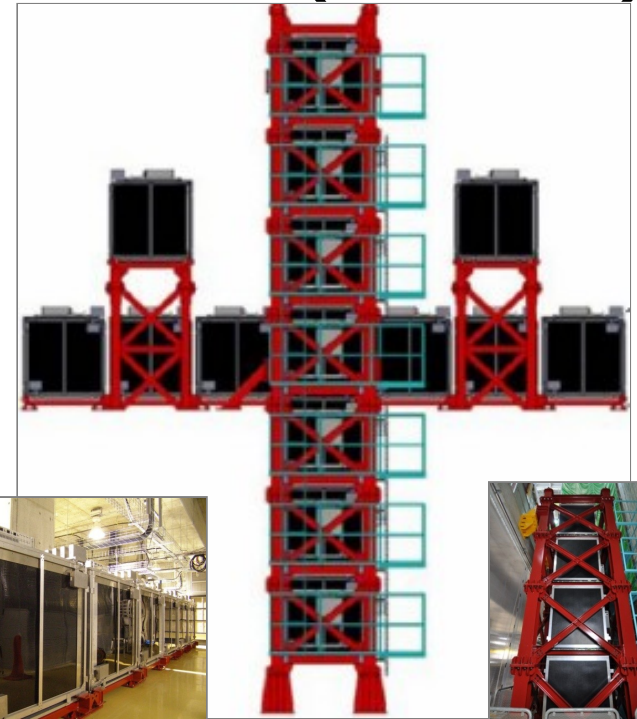
- Almost monochromatic beam at  $\sim 600$  MeV.
- Take advantage of Lorentz Boost and 2-body kinematics.
- Pure  $\nu_\mu$  beam with  $< 1\%$   $\nu_e$  contamination



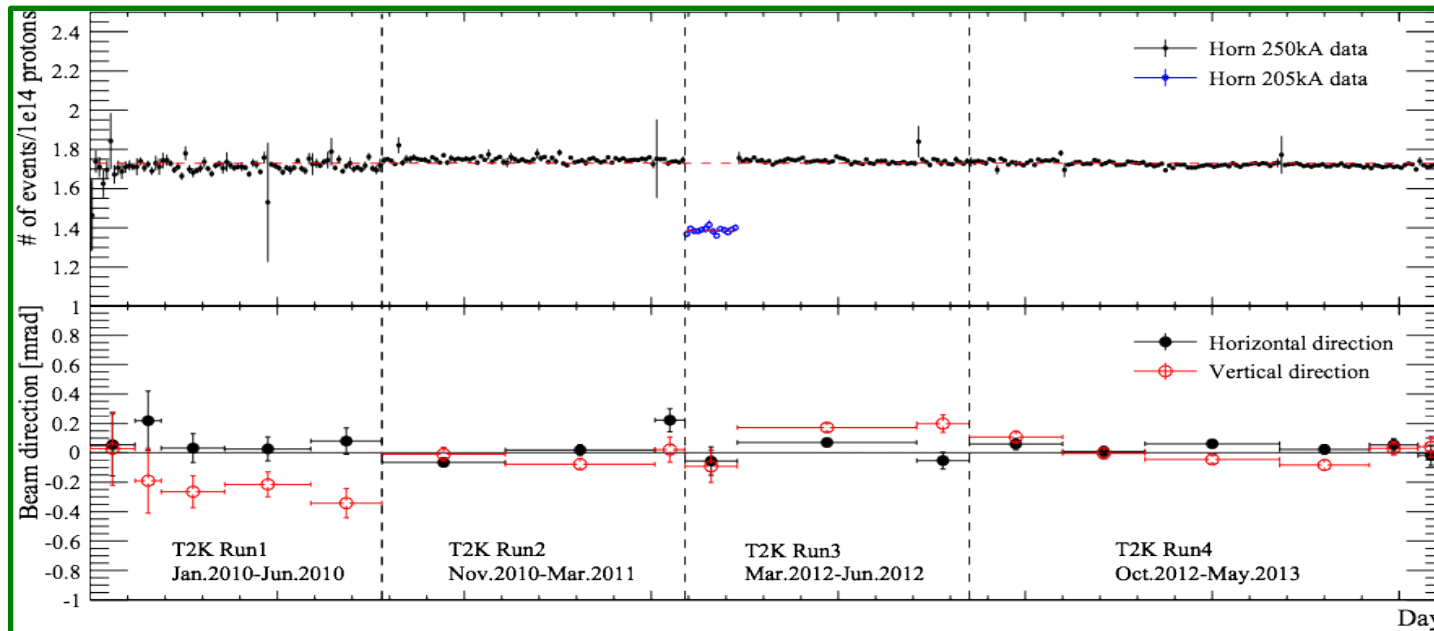
# INGRID on-axis near detector (280m)

INGRID (on-axis detector 280m from target):

- Designed to measure neutrino interactions & beam profile (beam intensity, direction & stability)
- Stability of beam direction requested  $< 1\text{mrad}$  to keep the peak energy at SK stable  $\delta E < 2\%$
- 7 + 7 modules (iron/scintillator planes sandwiches) in cross shape (central modules on-axis) + 3 extra modules.



Stability of  $\nu$  beam direction well within 1 mrad:



$\nu$  beam direction stability  $< 1\text{mrad}$

# ND280 off-axis near detector (280m)

- General purpose detector to measure:  $CC\nu_{\mu}$  events (normalization,  $E_{\nu}$  spectrum),
- $CC\nu_e$  events (background to  $\nu_e$  appearance), general neutrino interactions.

**2 FGDs Fine Grained Detectors:**  
Thin, wide scintillator planes.  
**Provides active target mass.**  
Optimized for p recoil detection.

**3 TPCs Time Projection Chambers:**  
momentum measurement of charged particles from FGD and P0D.  
PID via  $dE/dX$  measurement

**UA1/NOMAD magnet (0.2T),**  
Inner volume  $3.5 \times 3.6 \times 7 \text{ m}^3$

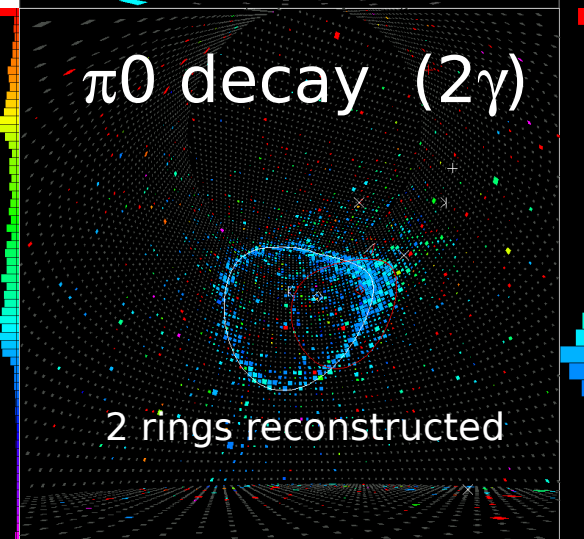
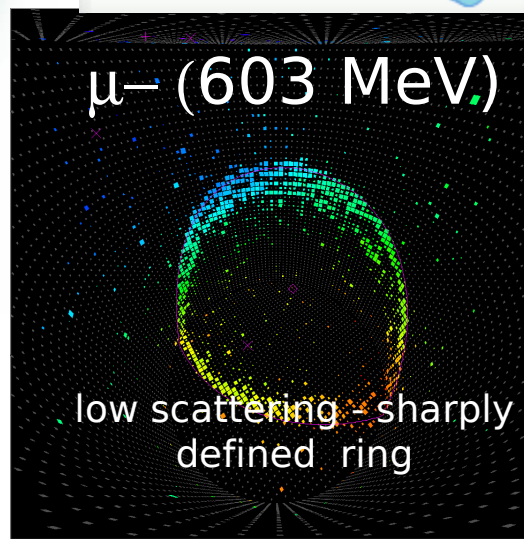
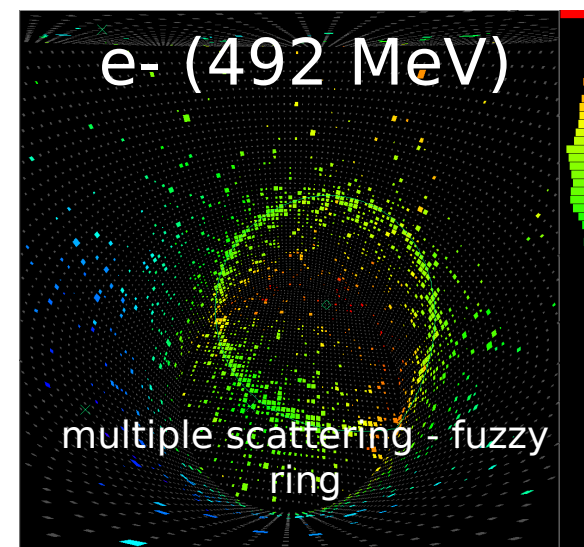
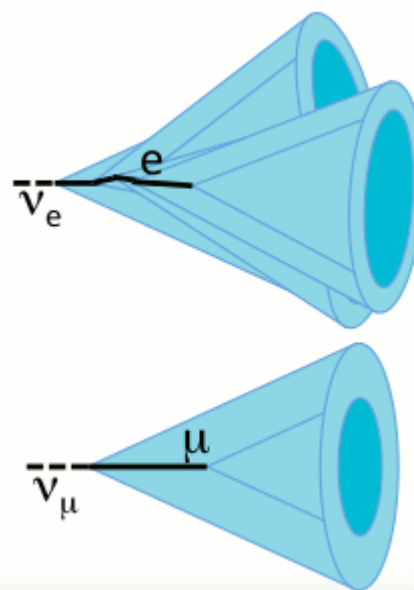
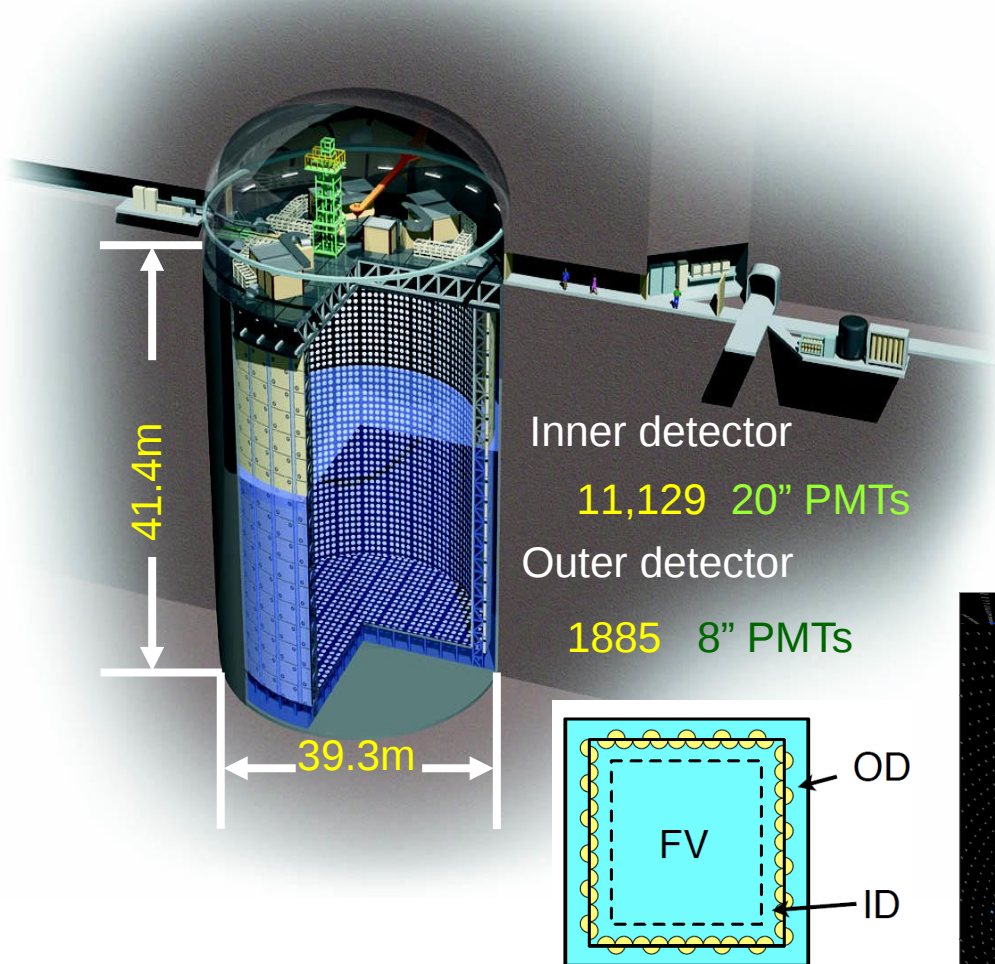
**SMRD Side Muon Range Detector**  
Scintillator planes in magnet yoke.  
Detector muons from inner detector.  
Momentum measurement.

**DSEcal, BarrelEcal, P0DEcal**  
**Electromagnetic calorimeter**  
Measure EM showers from inner detector

**P0D  $\pi^0$  detector**  
Scintillator planes interleaved with water and lead/brass layers  
Optimized for  $\gamma$  detection

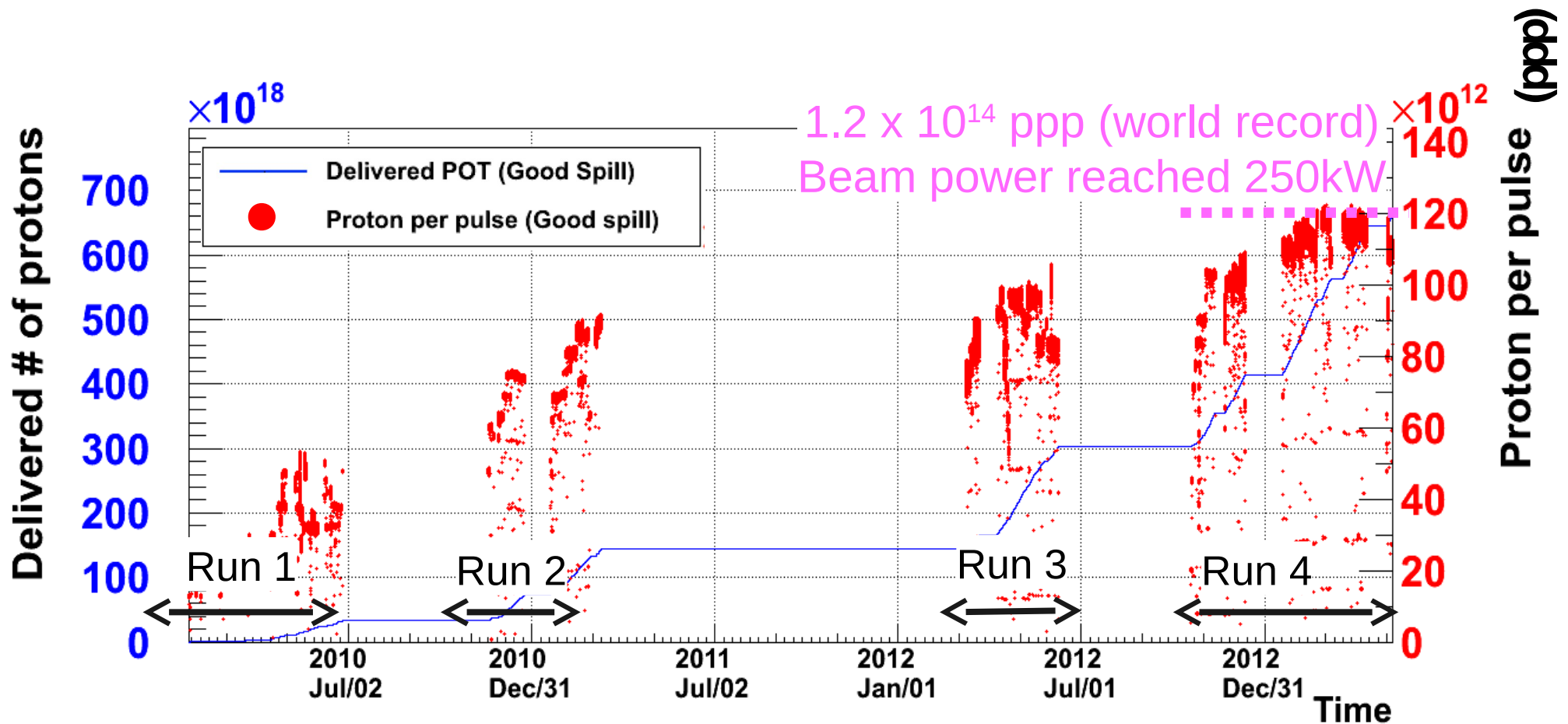
# T2K Far Detector: Super-Kamiokande

- 50 kton (22.5 kton fiducial) water Cherenkov detector
- Good reconstruction for T2K energy range
- Particle Identification (PID) based on shape of Cherenkov rings



Event displays show Monte Carlo

# T2K Dataset



- We collected  $6.63 \times 10^{20}$  protons on target (p.o.t.) so far
- Including  $0.21 \times 10^{20}$  p.o.t. with 205kA horn operation (13% flux reduction at peak) in Run3 (250kA horn current for nominal operation)

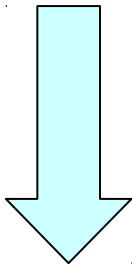
# (2013) Near Detector Constraint to SK

## Neutrino Flux Model:

- Data-driven: NA61/SHINE, beam monitor measurements

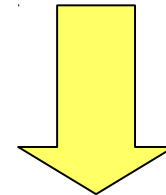
## Neutrino Cross Section Model (NEUT):

- Data-driven: External neutrino, electron, pion scattering data



NA61/Shine

- at the CERN SPS (North Area, H2 beam line)
- fixed target experiment on primary (ions) and secondary (ions, hadrons) beams



## Constraint from ND280 Data

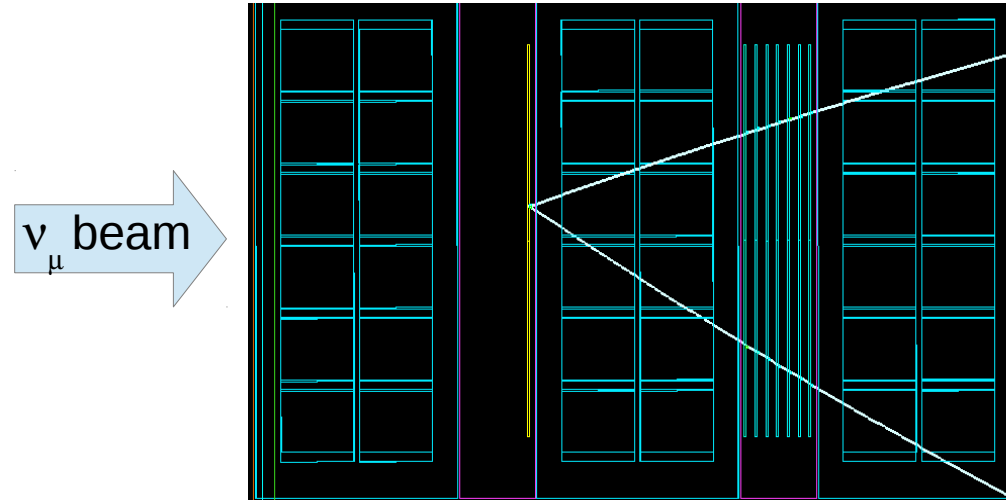
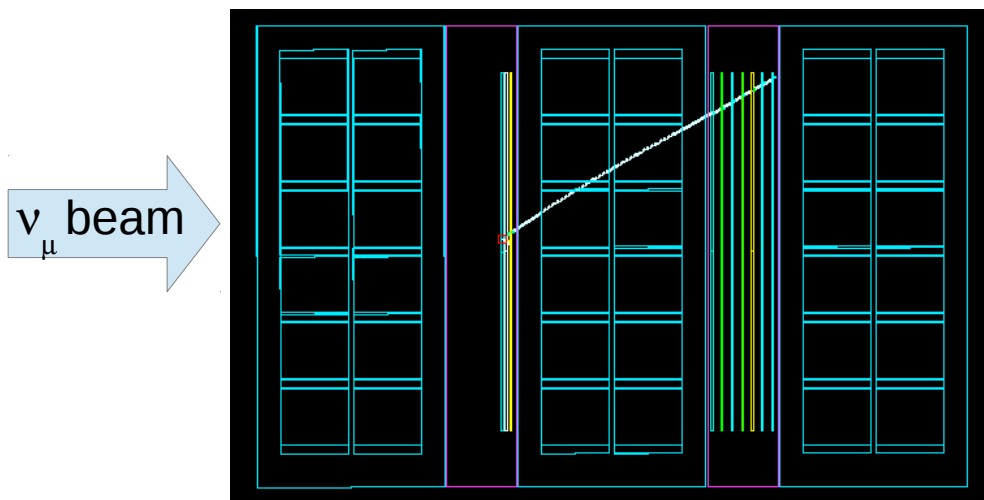
- Input: CC interactions with 0, 1 or multiple pions
- Fit to data constrains flux, and cross section parameters
- Constrained SK flux parameters and subset of cross section parameters are used to predict SK event rates



# ND280 Event Categories

Charged current (CC) with  $0\pi$

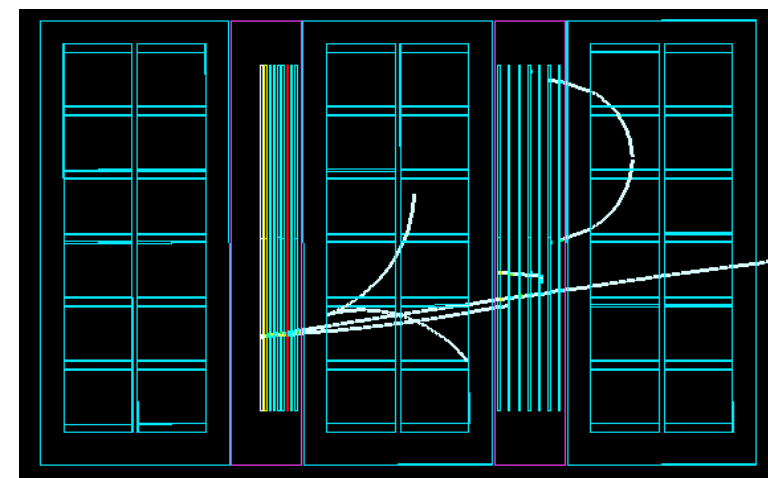
CC  $1\pi^+$



CC Other ( $\geq 1\pi$  - or  $\pi^0$ , or  $> 1\pi^+$ )

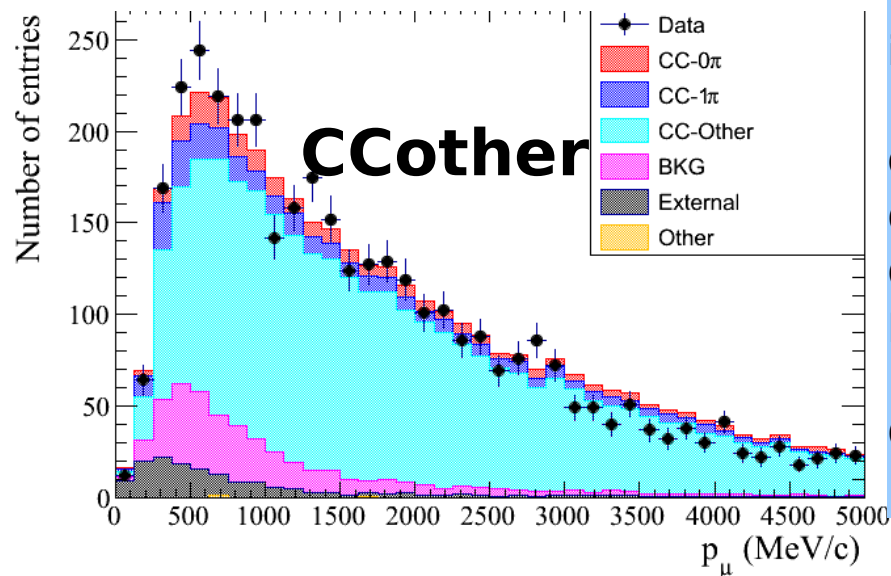
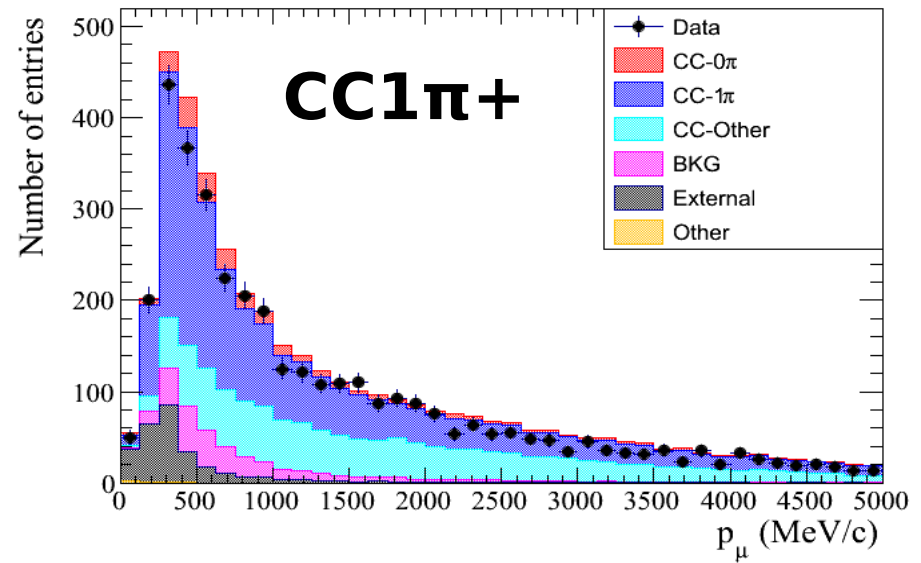
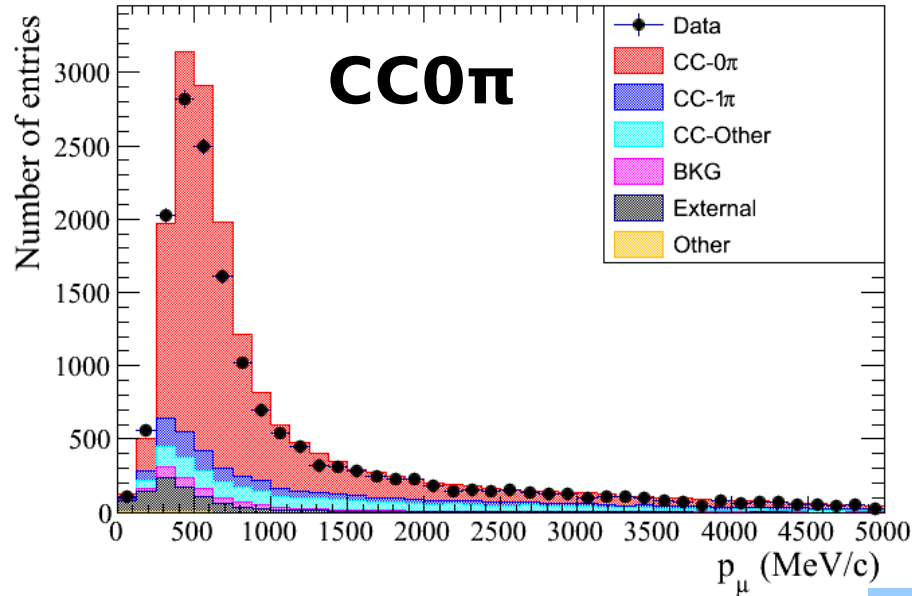
- $\pi^0$  candidates have identified electrons in the TPC

• Disappearance analysis joins CC  $1\pi^+$  and CC other together



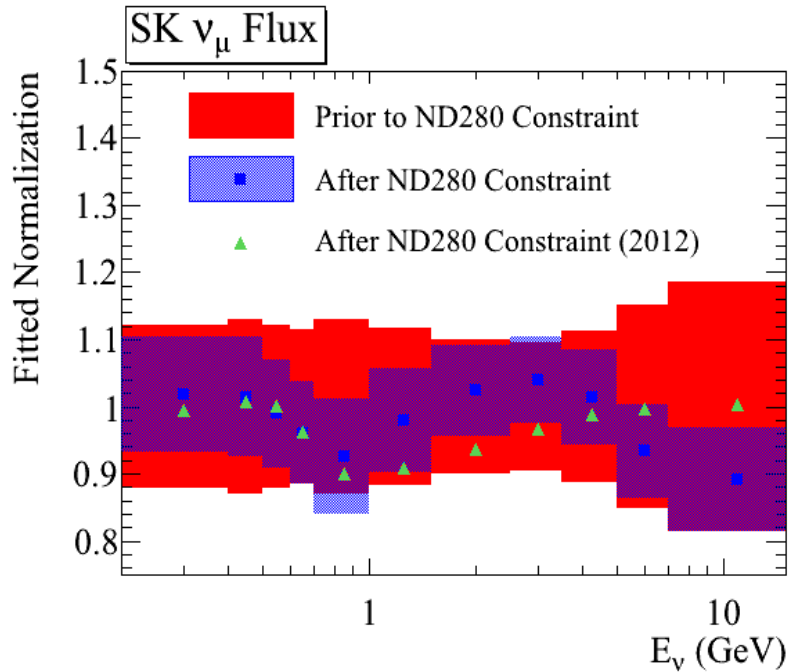
The event displays show only the tracking system at ND280

# Muon Momentum in ND280



True identification of interaction	CC0π sample	CC1π sample	CCOther sample
CC0π	72.6%	6.4%	5.8%
CC1π	8.6%	49.4%	7.8%
CCOther	11.4%	31%	73.8%
Bkg(NC+anti-nu)	2.3%	6.8%	8.7%
Out of FGD1 Fid Vol	5.1%	6.5%	3.9%

# Flux and X-Sections after Constraint



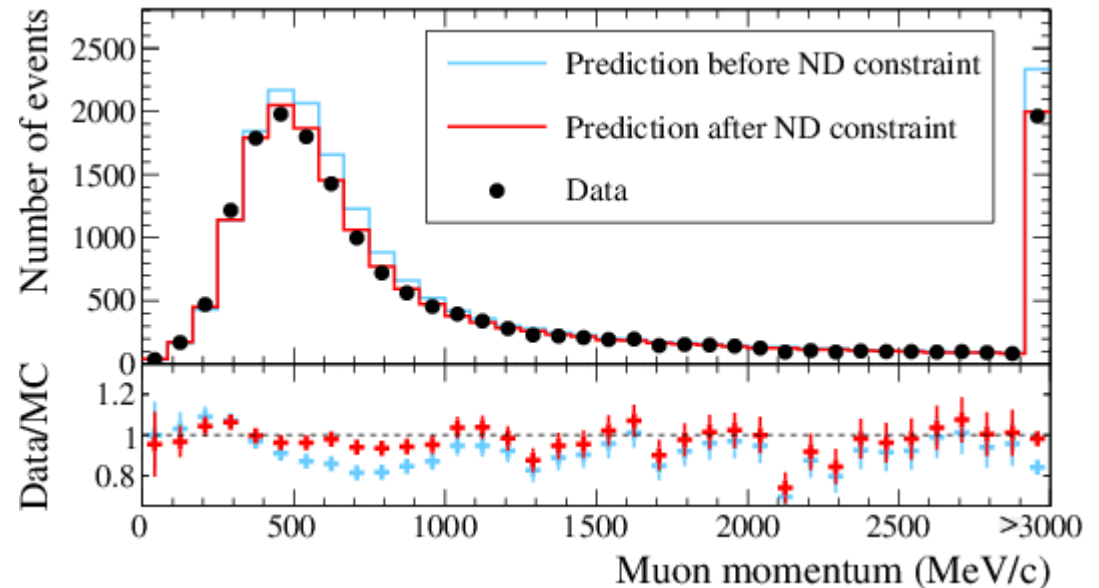
Parameter	Prior to ND280 Constraint	After ND280 Constraint
$M_A^{QE}$ (GeV)	$1.21 \pm 0.45$	$1.22 \pm 0.07$
CCQE Norm.*	$1.00 \pm 0.11$	$0.96 \pm 0.08$
$M_A^{RES}$ (GeV)	$1.41 \pm 0.22$	$0.96 \pm 0.06$
CC1 $\pi$ Norm.**	$1.15 \pm 0.32$	$1.22 \pm 0.16$

\*For  $E_\nu < 1.5$  GeV

\*\*For  $E_\nu < 2.5$  GeV

• ND280 constraint reduces both flux and cross-section model uncertainties individually

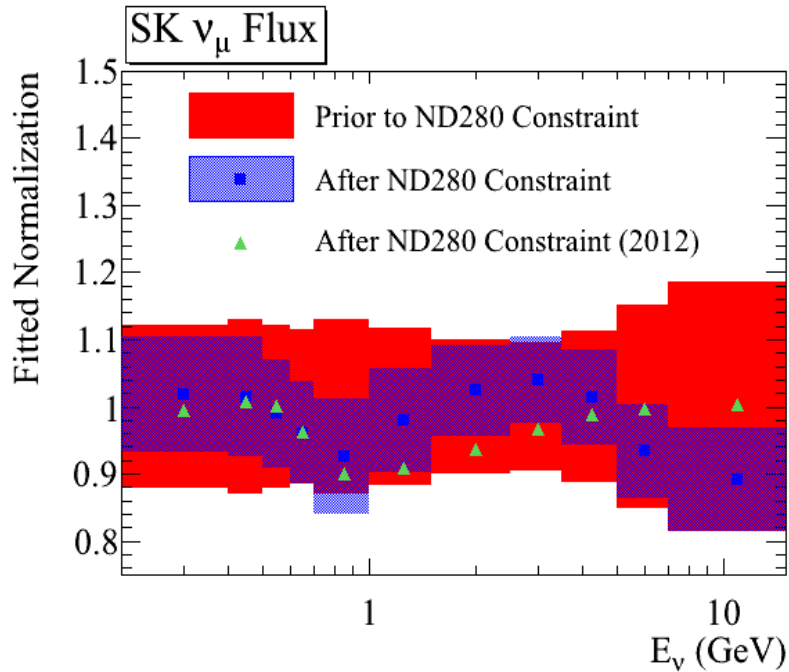
➤ Note reductions on the “ $M_A$ ” parameters which set  $Q^2$  shape of these events



$M_A^{QE}$  = mass parameter in the axial dipole form factor for QE interactions

$M_A^{RES}$  = mass parameter in the axial dipole form factor for resonant pion interactions

# Flux and X-Sections after Constraint

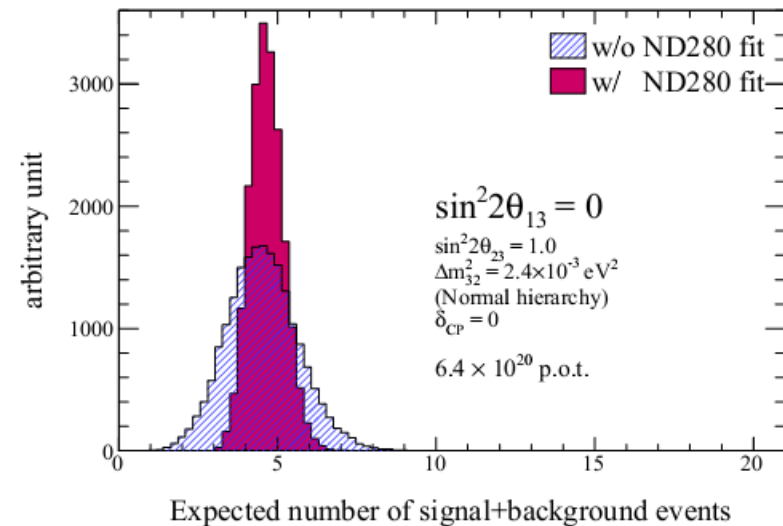
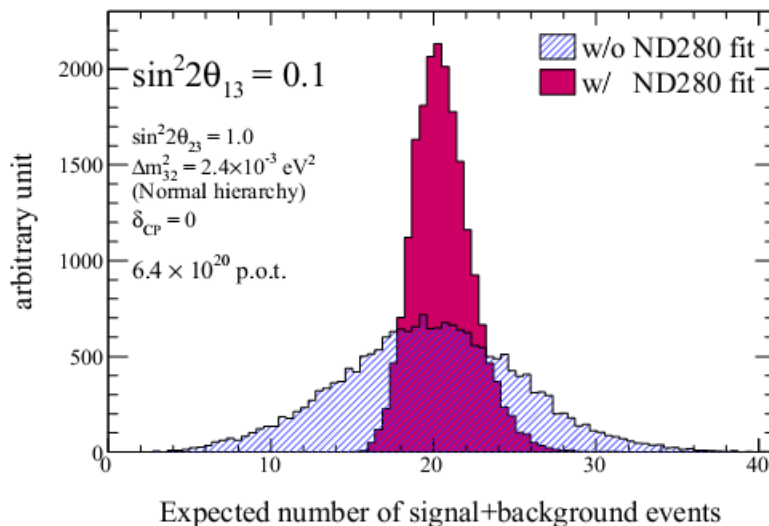


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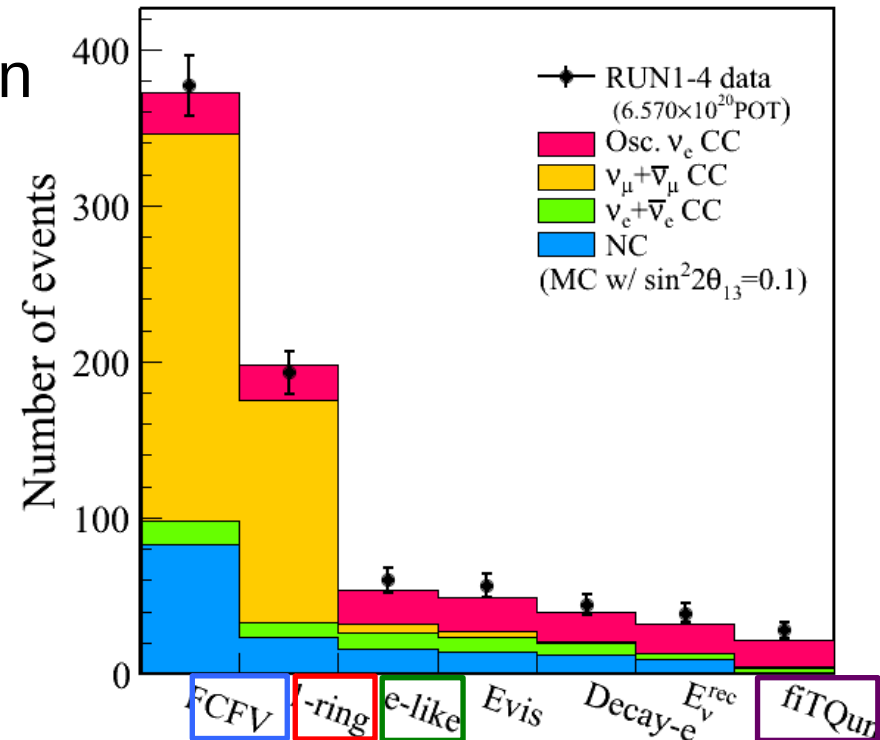
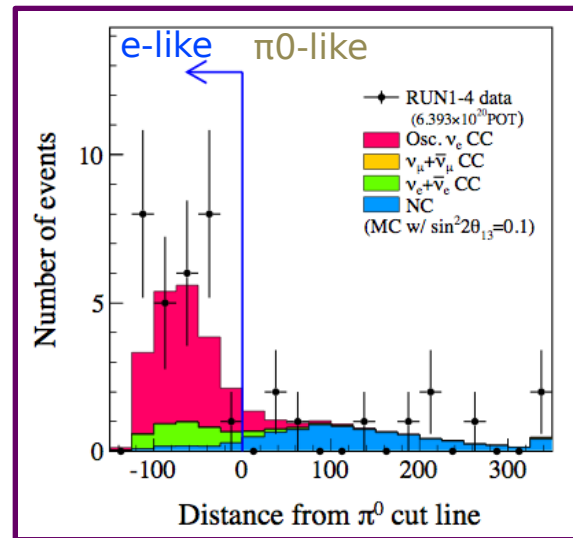
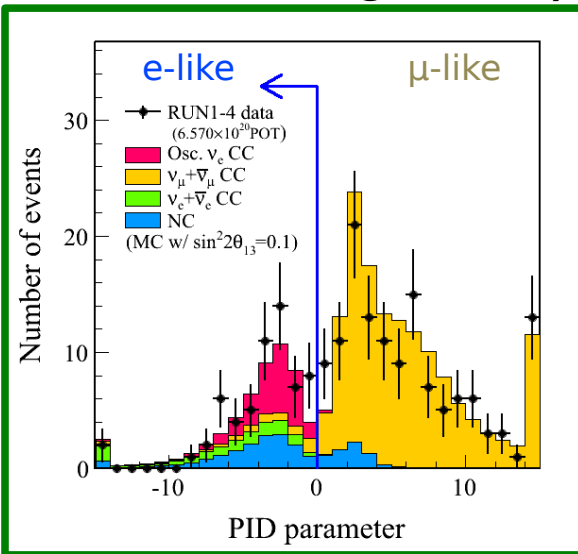
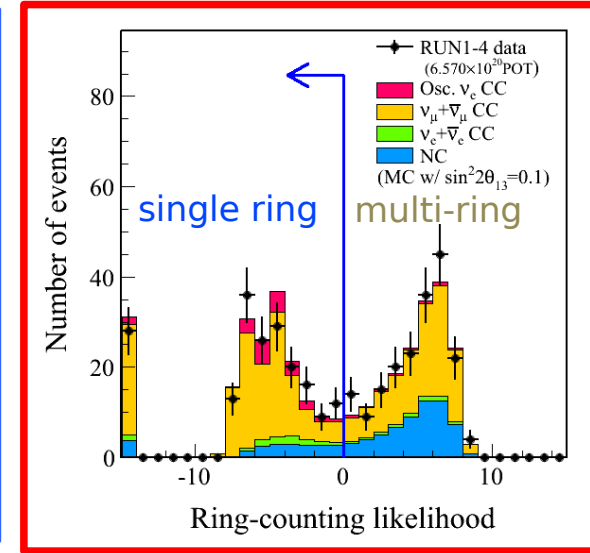
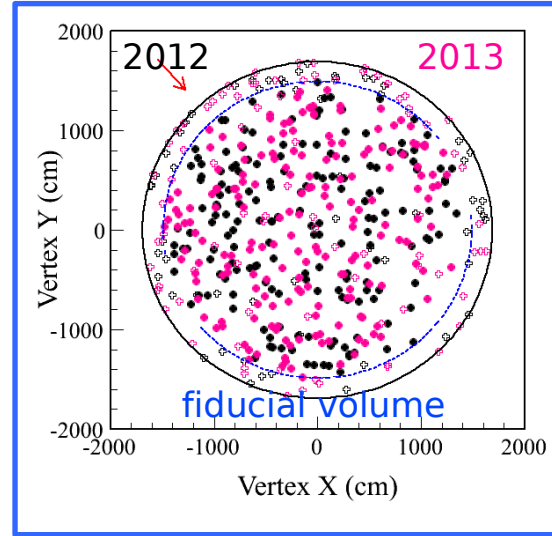
## Far detector prediction uncertainties after ND280 constraint



# $\nu_{\mu} \rightarrow \nu_e$ Event Selection

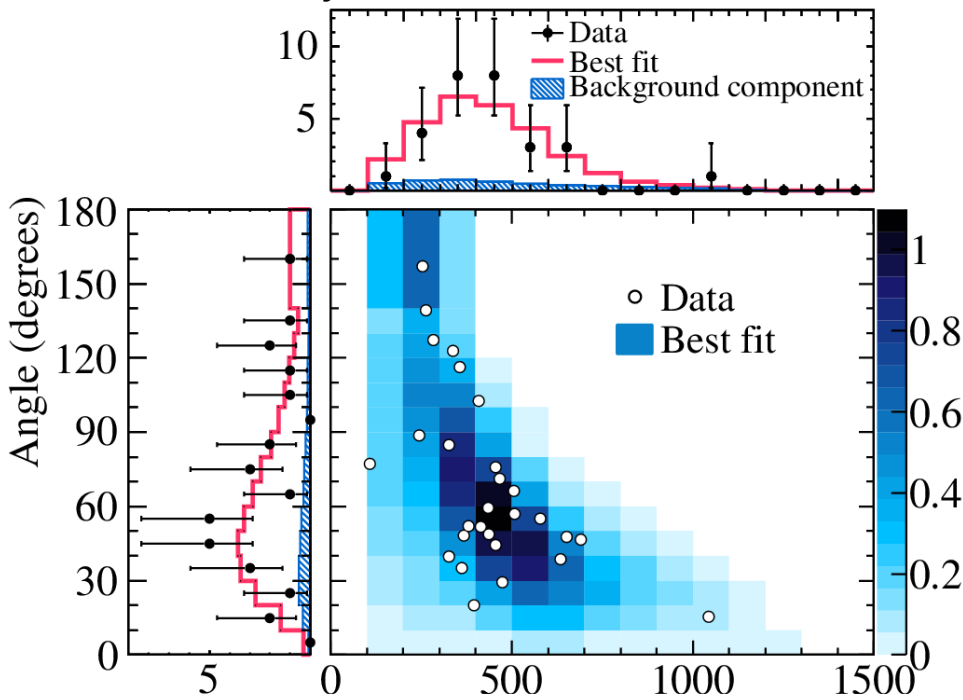
## Event selection:

- Fully contained in fid. volume
- Only one reconstructed ring
- Ring is electron-like
- Visible energy  $> 100\text{MeV}$
- No Michel Electrons
- Reconstructed energy  $< 1.25\text{ GeV}$
- New SK reconstruction ( $\sim 30\%$  reduction in  $\pi^0$  background)

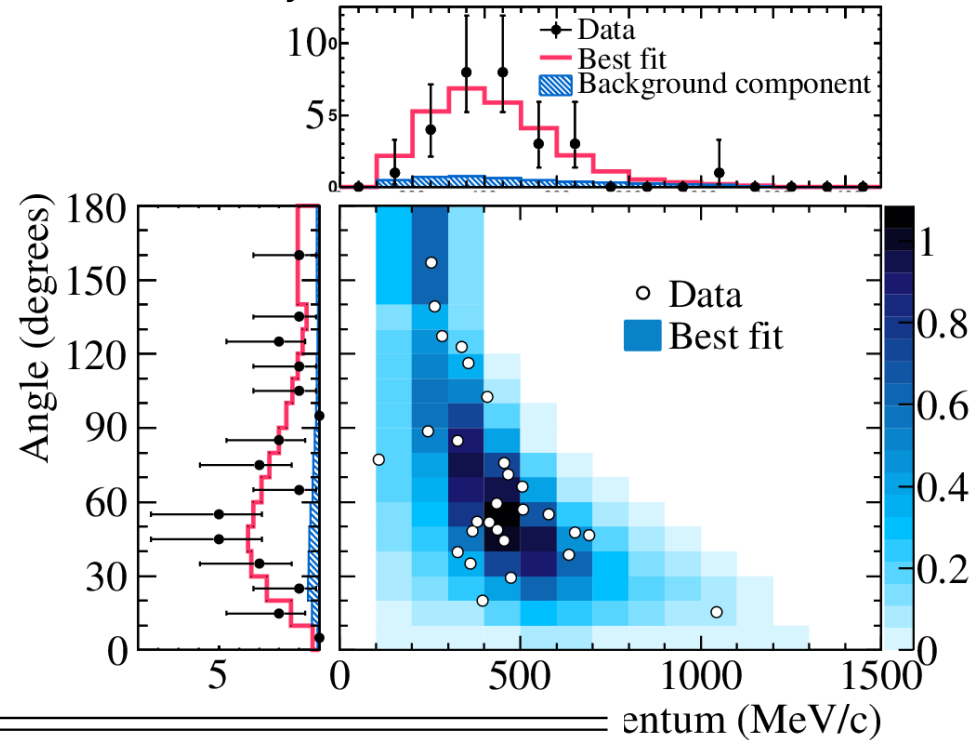


# Results

Normal Hierarchy



Inverted Hierarchy



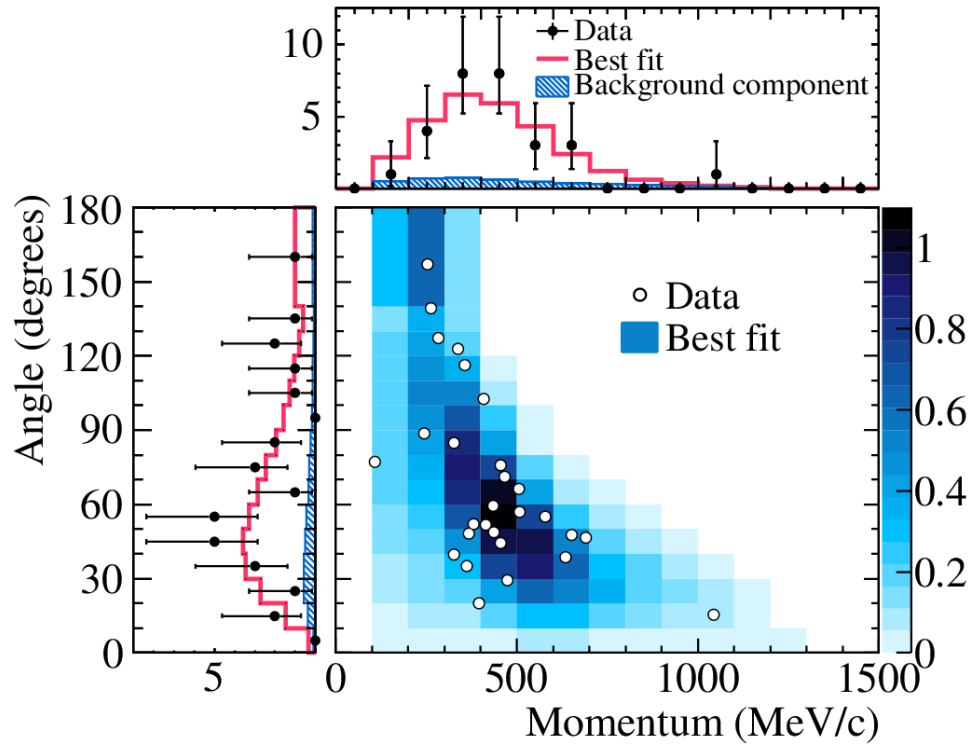
The predicted number of events

Event category	$\sin^2 2\theta_{13} = 0.0$	$\sin^2 2\theta_{13} = 0.1$
Total	4.92	21.56
$\nu_e$ signal	0.40	17.30
$\nu_e$ background	3.37	3.12
$\nu_\mu$ background	0.94	0.94
$\bar{\nu}_\mu$ background	0.05	0.05
$\bar{\nu}_e$ background	0.16	0.15

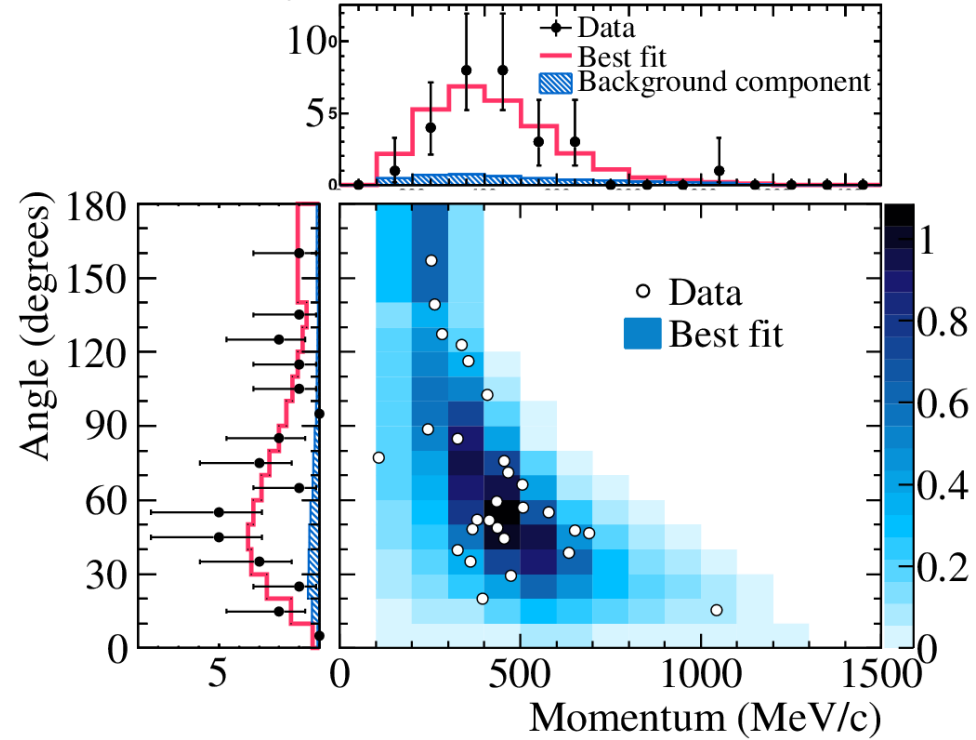
Data: 28

# Results

Normal Hierarchy



Inverted Hierarchy



Error source	$\sin^2 2\theta_{13} = 0$		$\sin^2 2\theta_{13} = 0.1$	
	w/o ND280 fit	w/ ND280 fit	w/o ND280 fit	w/ ND280 fit
BANFF	21.7	4.8	25.9	2.9
$\nu$ int. (other than BANFF)	6.8	6.8	7.5	7.5
SK+FSI	7.3	7.3	3.5	3.5
Total	24.0	11.1	27.2	8.8
2012 analysis	21.0	13.0	24.2	9.9

# T2K $\nu_e$ Appearance Fit Results

- Fit using  $(p, \theta)$  distributions
- Best fit value at 68% CL for  $\delta_{\text{CP}} = 0$

## normal hierarchy:

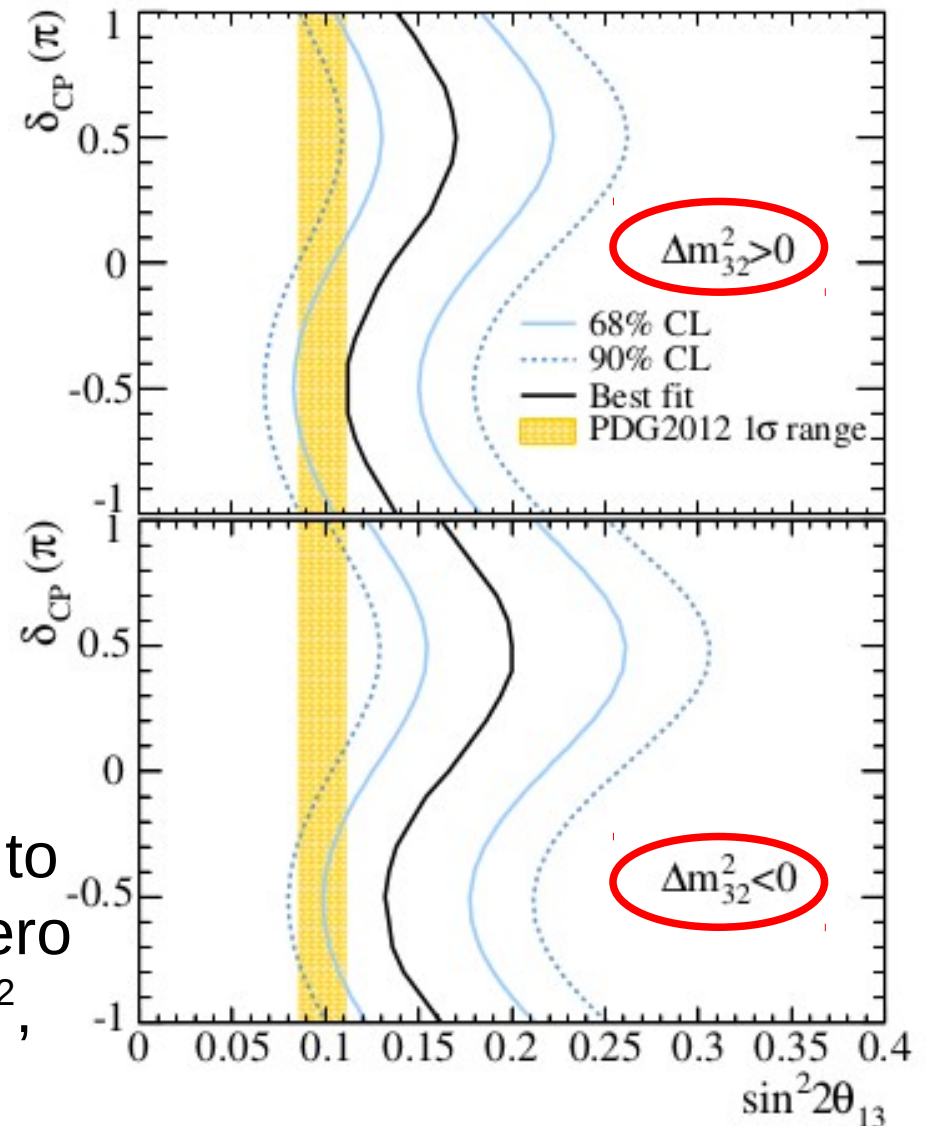
$$\sin^2 2\theta_{13} = 0.136^{+0.044}_{-0.033}$$

## inverted hierarchy:

$$\sin^2 2\theta_{13} = 0.166^{+0.051}_{-0.042}$$

- Comparing the best  $p$ - $\theta$  fit likelihood to null hypothesis gives a  $7.4\sigma$  for non-zero  $\theta_{13}$ , for  $\sin^2\theta_{23}=0.5$ ,  $|\Delta m_{32}^2|=2.4\times 10^{-3}\text{eV}^2$ ,  $\delta_{\text{CP}}=0$ , and normal mass hierarchy.

- First ever observation ( $>5\sigma$ ) of an explicit  $\nu_e$  appearance channel



NB: These are 1D contours for values of  $\delta_{\text{CP}}$ , not 2D contours in  $\delta_{\text{CP}}-\theta_{13}$  space

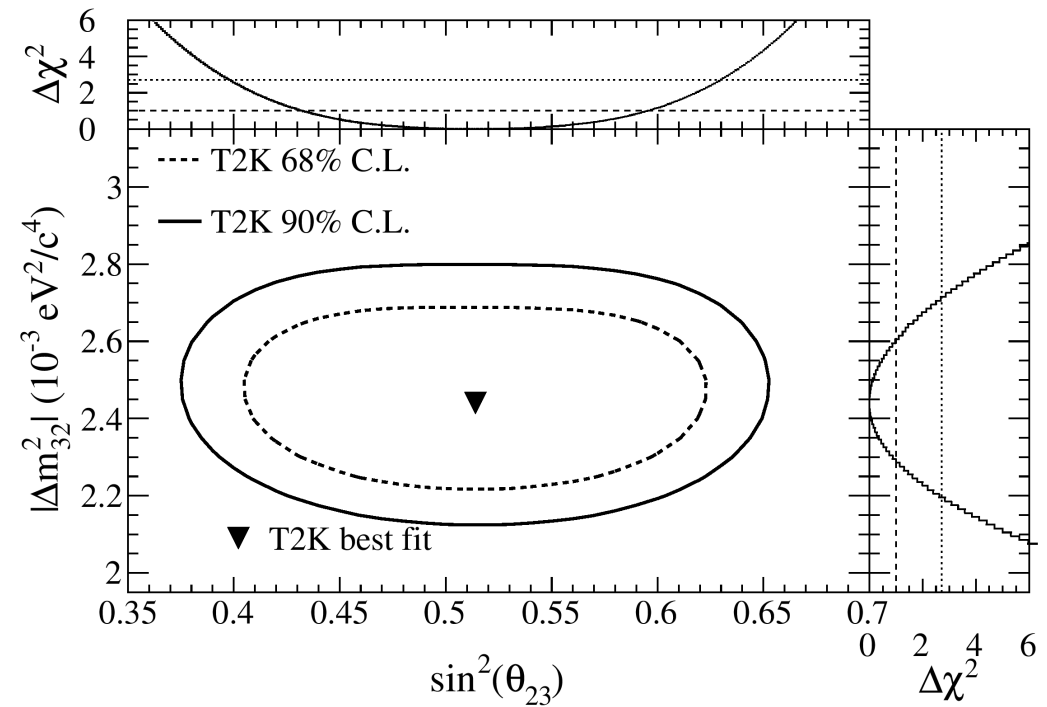
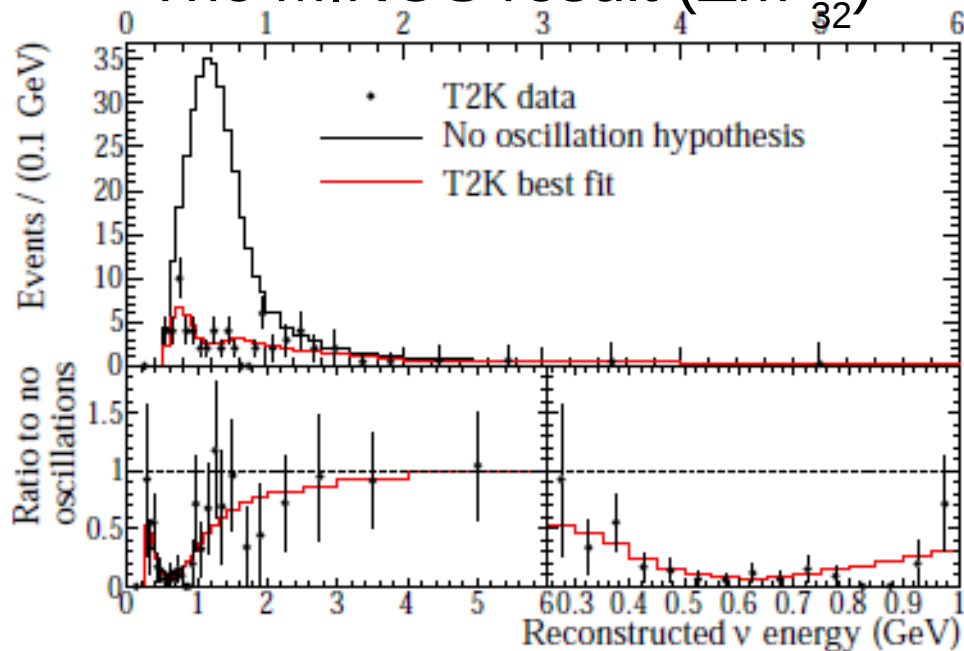


# $\nu_{\mu} \rightarrow \nu_{\mu}$ T2K Result

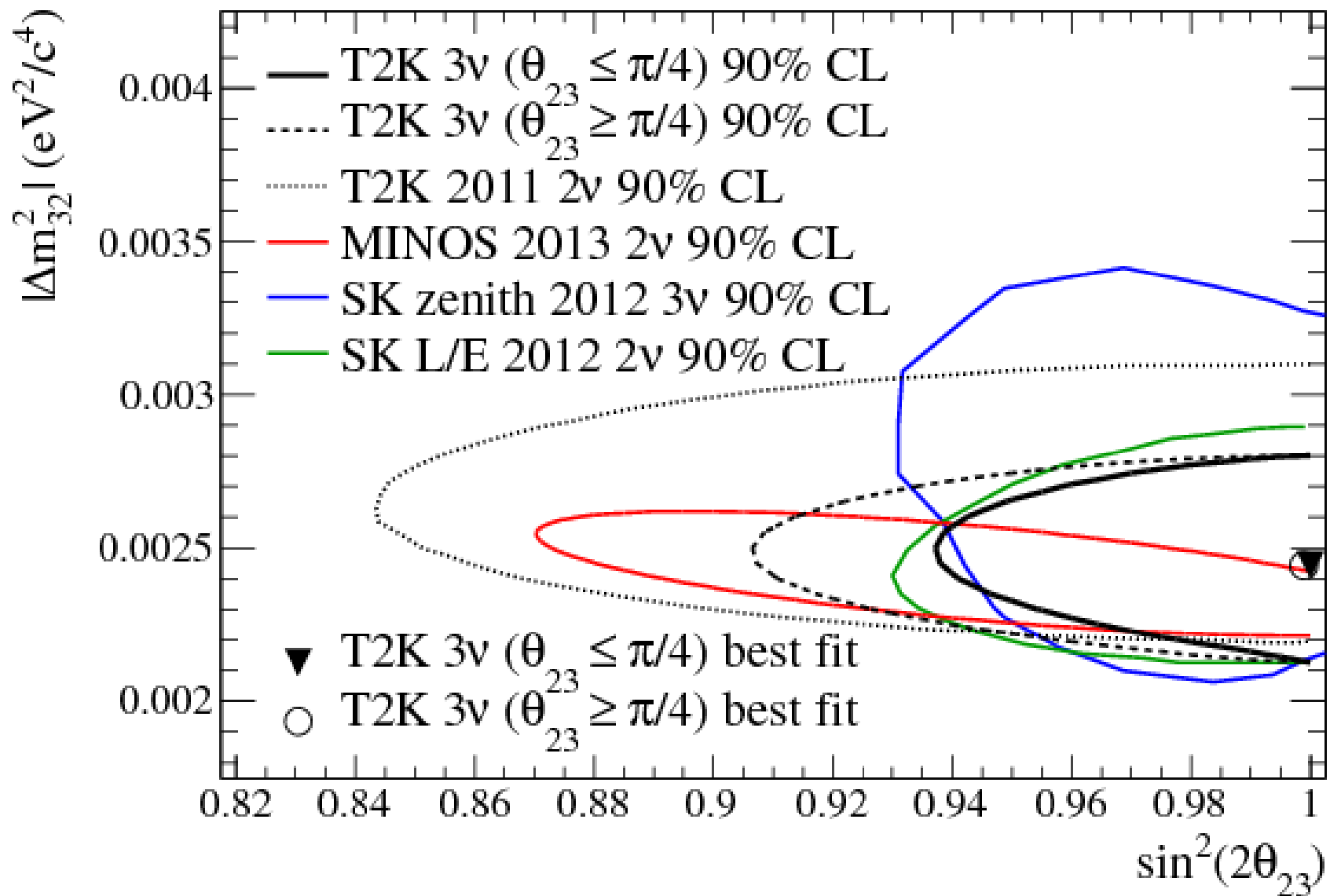
- Best-fit oscillation parameter values:

$$\sin^2 \theta_{23} = 0.514 \pm 0.082 \quad \left| \Delta m_{32}^2 \right| = 2.44_{-0.15}^{+0.17} \times 10^{-3} \text{ eV}^2/c^2$$

- Events: 58 (observed), 57.92 (predicted),  $205 \pm 17$  (no oscillation)
- Data prefers 2nd  $\theta_{23}$  octant
- $1\sigma$  confidence intervals are consistent with:
  - Maximal mixing ( $\sin^2 \theta_{23}$ )
  - The MINOS result ( $\Delta m_{32}^2$ )

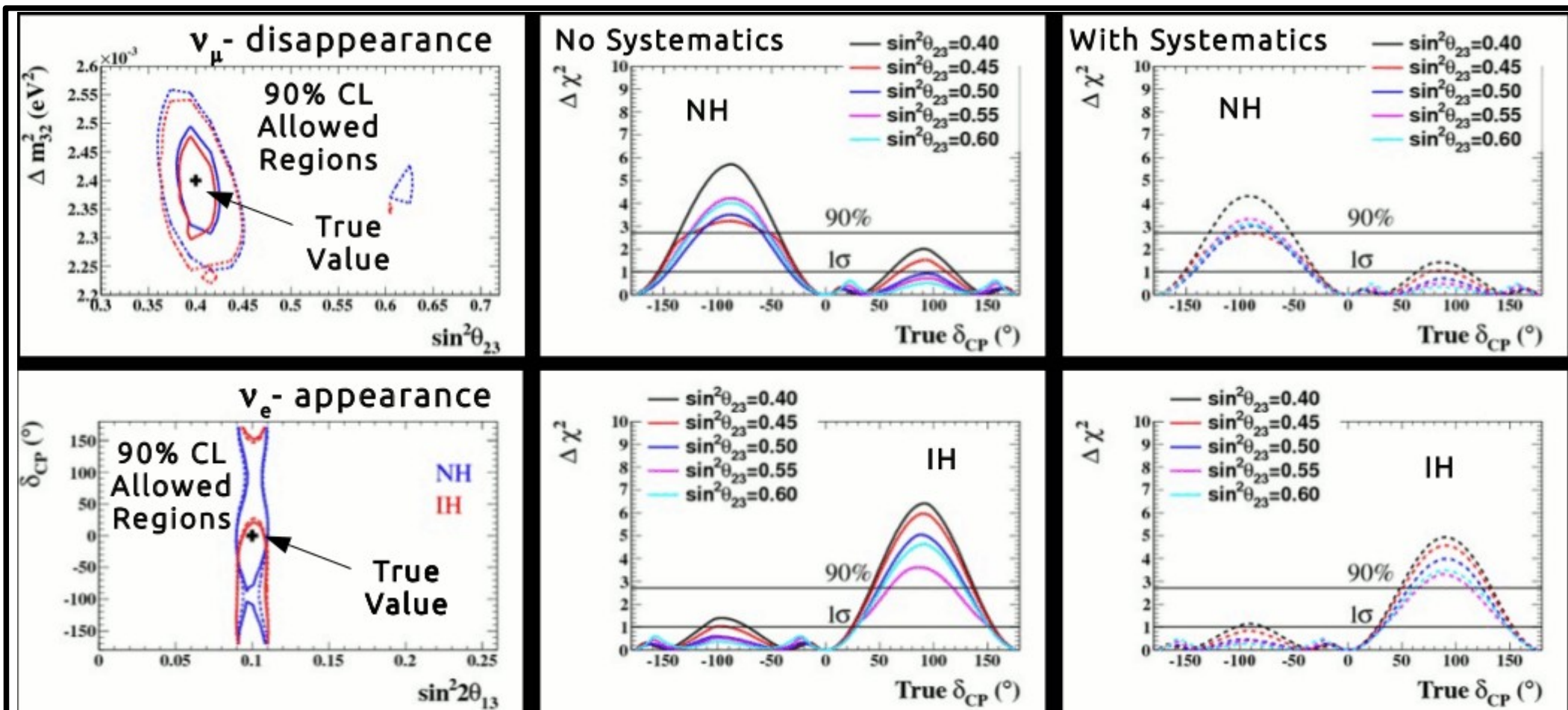


# Results

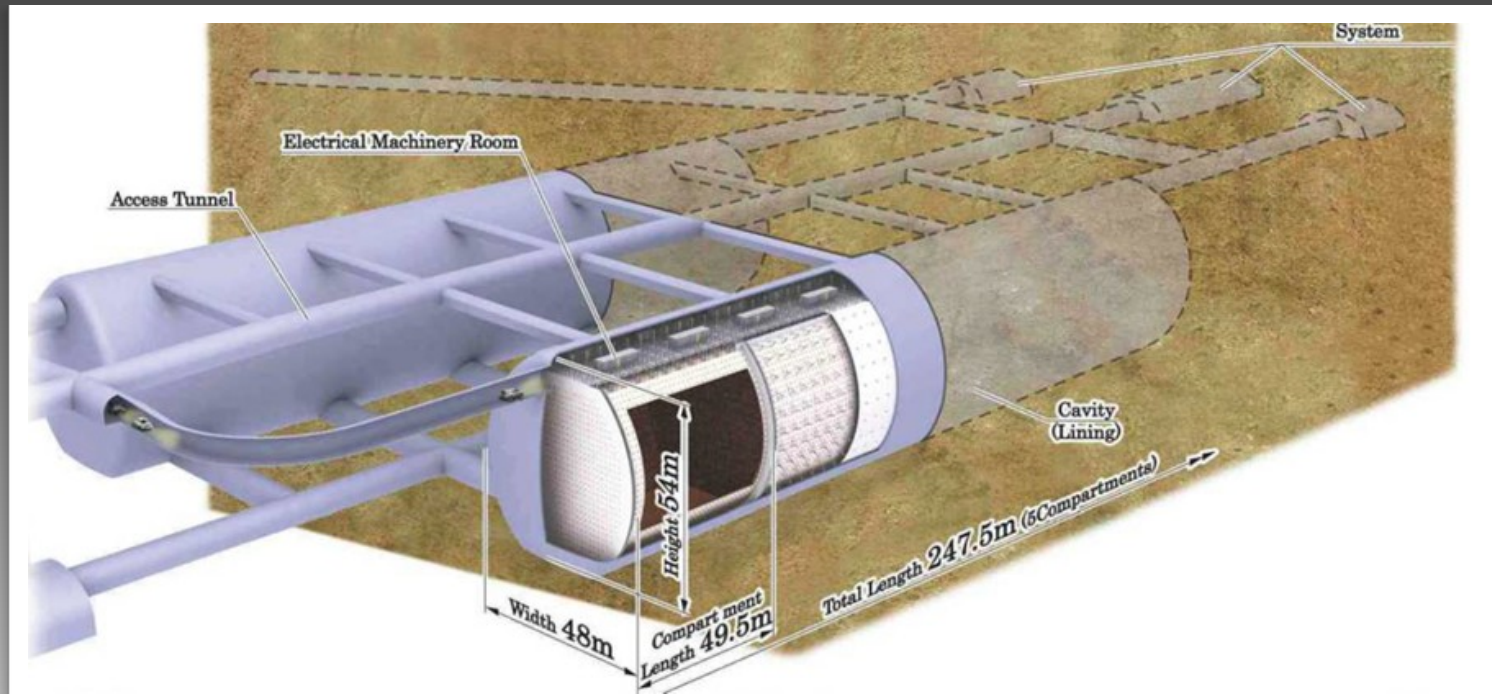


# T2K and J-PARC Run Plans

- Ongoing sensitivity studies to determine future potential and running conditions (neutrino vs anti-neutrinos)
- Combined studies with NovA for neutrino mass hierarchy ongoing.
- Assuming 50% neutrino and 50% antineutrino for T2K and combination with Daya Bay:



# Hyper-Kamiokande



# The Hyper-Kamiokande Project

- Four International Open Meetings (2012-2014) @ IPMU, Japan.
- Formed international working groups.

August 21-23, 2012

<http://indico.ipmu.jp/indico/conferenceDisplay.py?confId=7>



January 14-15, 2013

<http://indico.ipmu.jp/indico/conferenceDisplay.py?confId=10>



June 21-22, 2013

<http://indico.ipmu.jp/indico/conferenceDisplay.py?confId=23>



January 27-28, 2014

<http://indico.ipmu.jp/indico/conferenceDisplay.py?confId=29>



Please look at the slides from the meetings for detailed information on topics

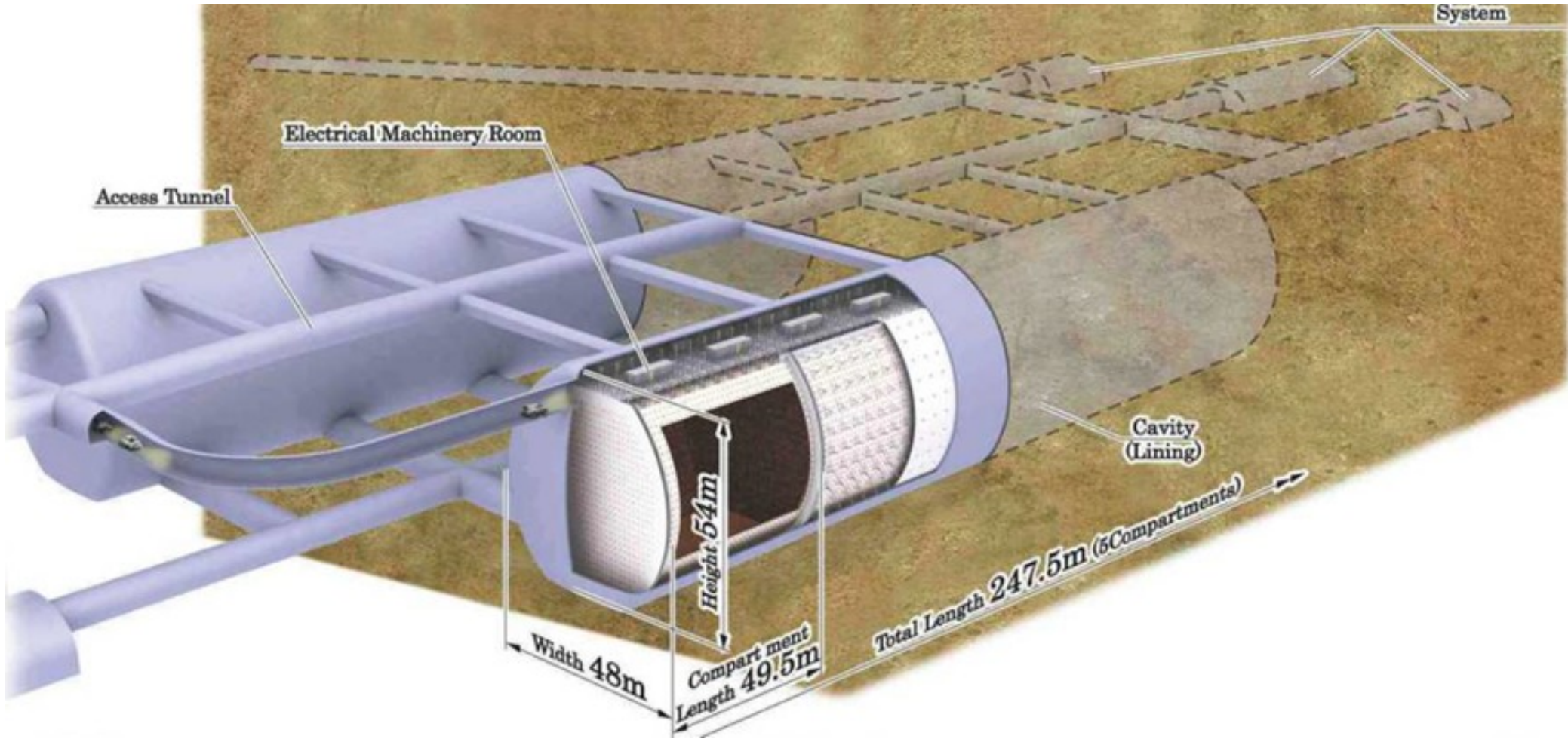
# The Hyper-Kamiokande Project

- First European Open Meeting (18 Dec. 2013, London, QMUL): <http://indico.cern.ch/e/HKEUOpenMeeting>
- More than 40 participants from 9 Countries.



- Discussed common issues. One more open meeting this year.
- Created mailing list <[hyper-kamiokande-eu@qmul.ac.uk](mailto:hyper-kamiokande-eu@qmul.ac.uk)>

# Hyper-Kamiokande Overview



25 x Super-Kamiokande 31

# Hyper-Kamiokande Overview

- **Water Cherenkov**, proved technology & scalability:
  - Excellent PID at sub-GeV region >99%
  - Large mass → statistics always critical for any measurements.

Total Volume	0.99 Megaton
Inner Volume	0.74 Mton
Fiducial Volume	0.56 Mton (0.056 Mton × 10 compartments)
Outer Volume	0.2 Megaton
Photo-sensors	<ul style="list-style-type: none"><li>• 99,000 20"Φ PMTs for Inner Detector (ID) (20% photo-coverage)</li><li>• 25,000 8"Φ PMTs for Outer Detector (OD)</li></ul>
Tanks	<ul style="list-style-type: none"><li>• 2 tanks, with egg-shape cross section 48m (w) × 50m (t) × 250 m (l)</li><li>• 5 optically separated compartments per tank</li></ul>

25 x Super-Kamiokande



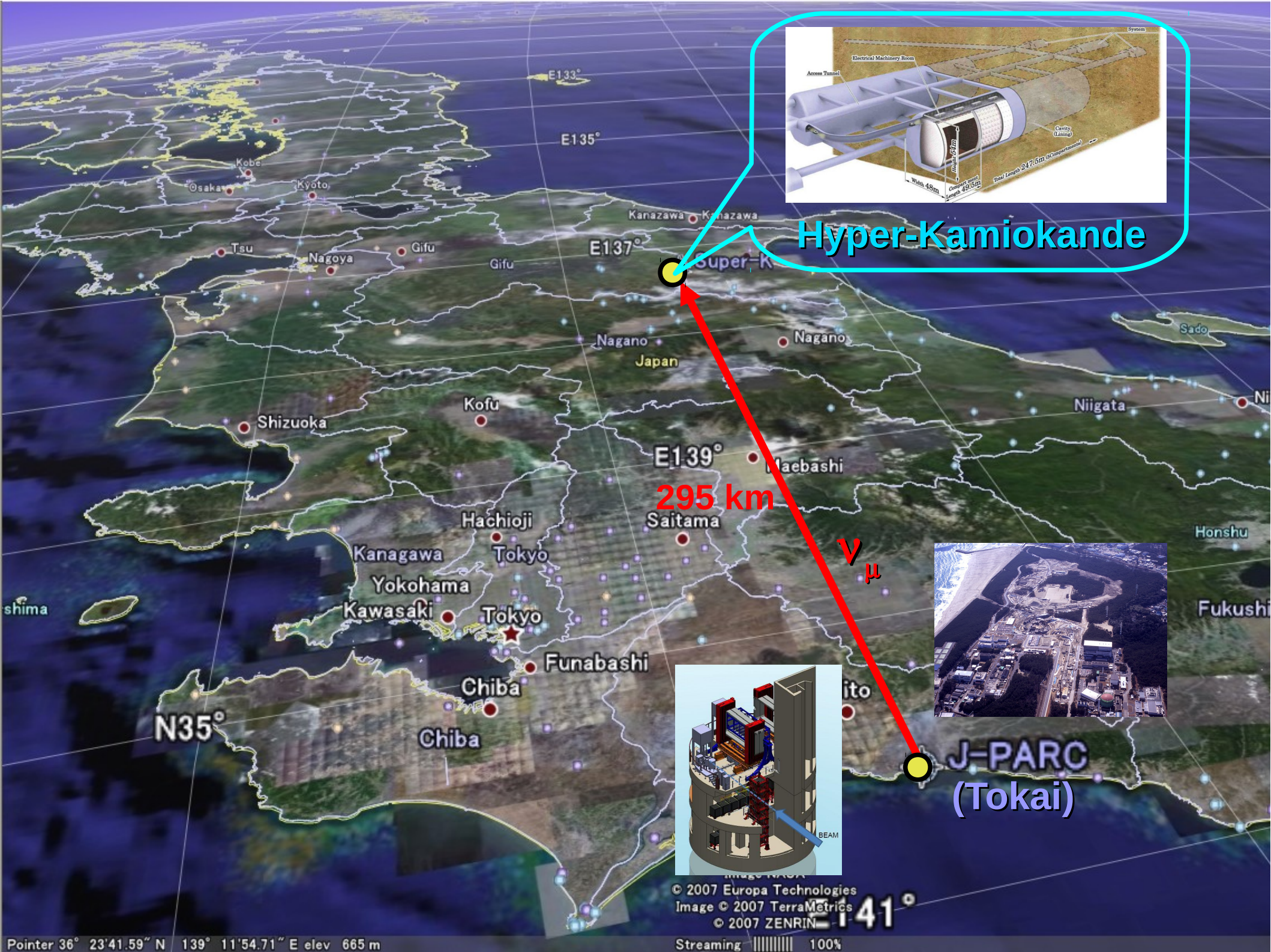
# Physics Topics



CAVEAT (Letter of Intent, Hyper-K WG arXiv:1109.3262 [hep-ex])

- 5% overall systematic error
- 3y:7y  $\nu$ -beam: $\nu$ -beam sharing
- No new SK reconstruction used ( $\Rightarrow$  higher  $\pi^0$  background)
- No new near detector

New updated results expected by Summer 2014 (LoI to J-PARC).



# Hyper-Kamiokande

295 km

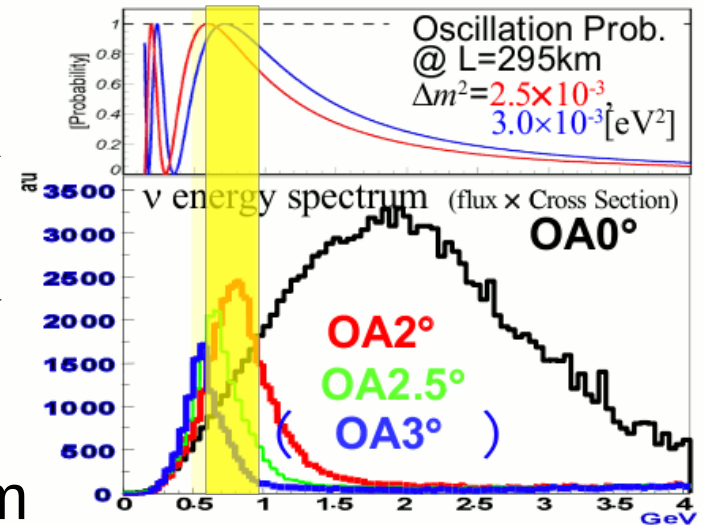
$\nu_\mu$

J-PARC (Tokai)

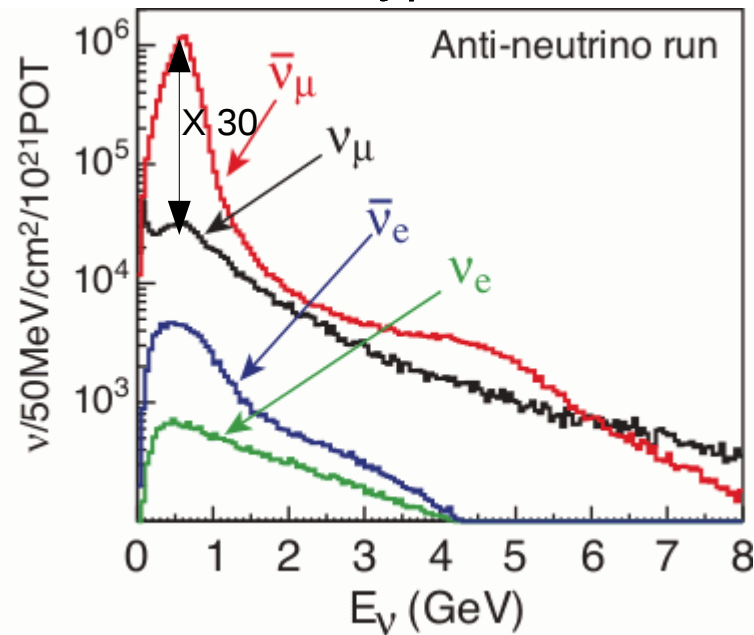
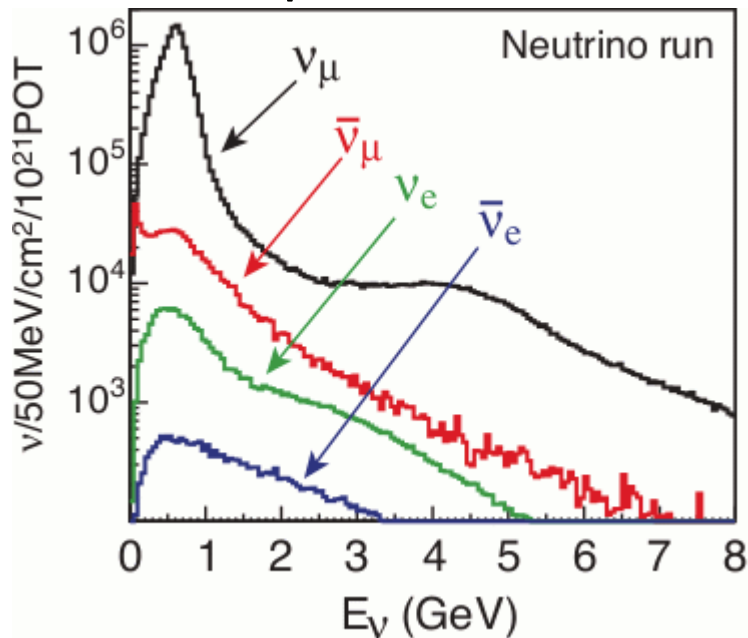
© 2007 Europa Technologies  
Image © 2007 TerraMetrics  
© 2007 ZENRIN

# Tokai-2-Hyper-Kamiokande

- Natural extension of the technique being proven by the success of T2K:
  - Off-axis narrow band beam: suppress background from high energy component ( $\nu_\tau$  negligible)
  - $E_\nu \sim 0.6$  GeV: peaked at oscillation maximum



Expected unoscillated neutrino flux at Hyper-K



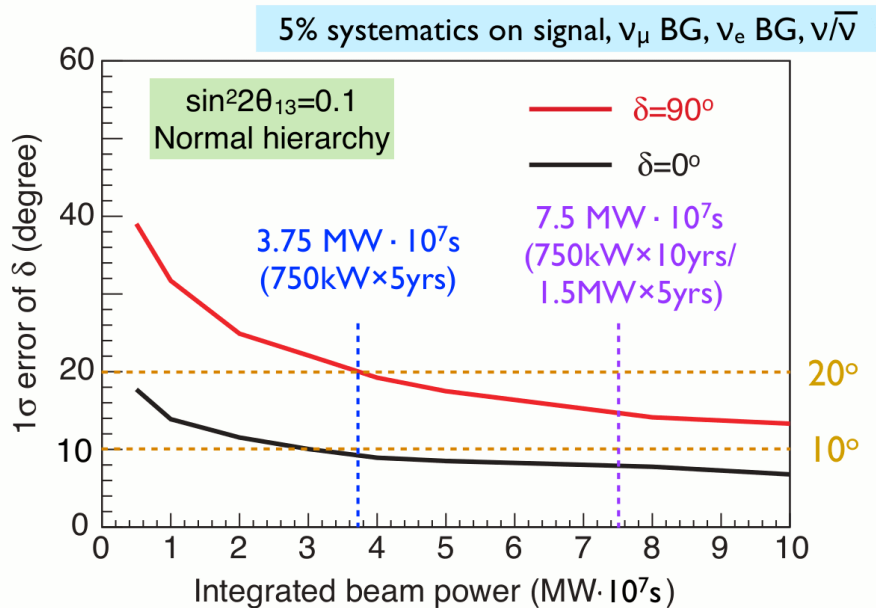
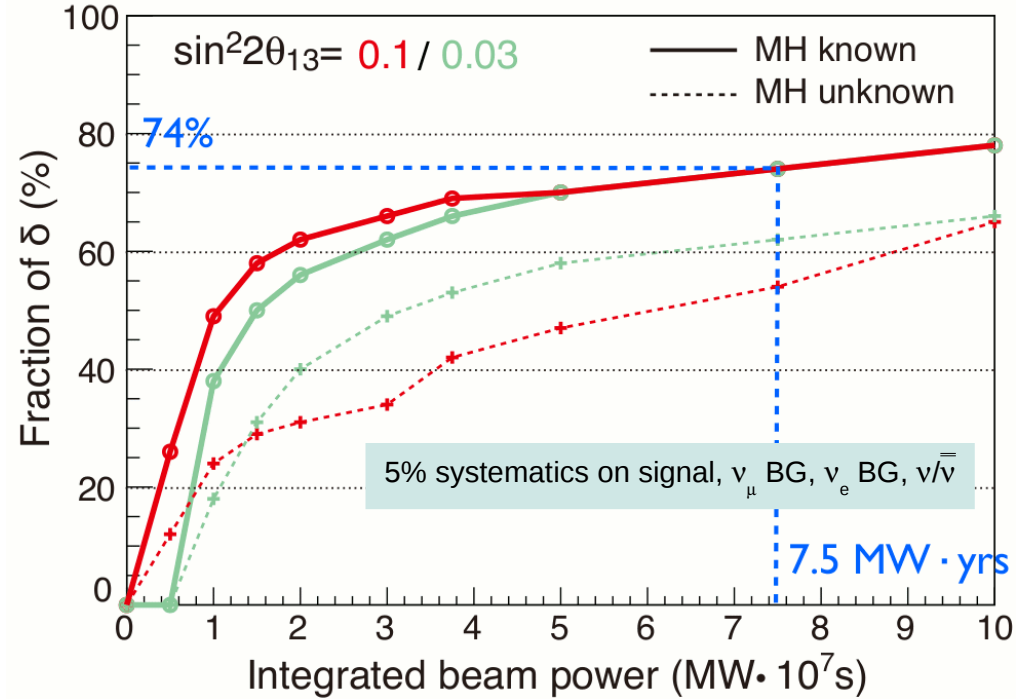
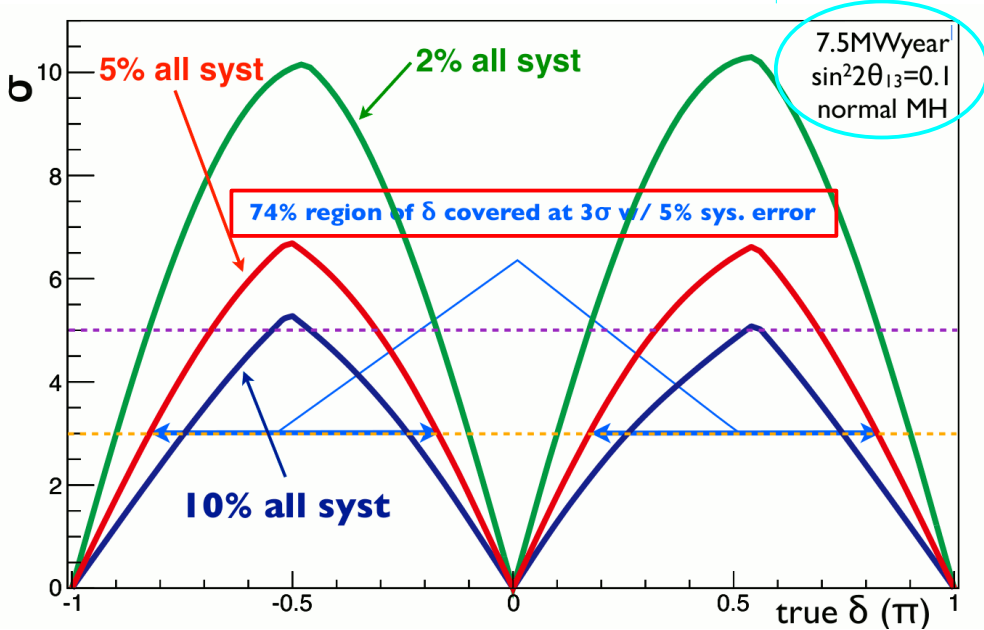
Beam sharing between neutrinos and anti-neutrinos

$\nu$ -mode:  $\bar{\nu}$ -mode  
 3y : 7y

# Expected Sensitivity to CP Violation

**CPV discovery sensitivity** w/ mass hierarchy known.

**Fractional region of  $\delta$  (%)** for which the CPV ( $\sin \delta \neq 0$ ) significance is  $> 3\sigma$



**$\delta$  coverage:**

CPV  $> 3\sigma$  ( $5\sigma$ ) for 74%(55%) of  $\delta$

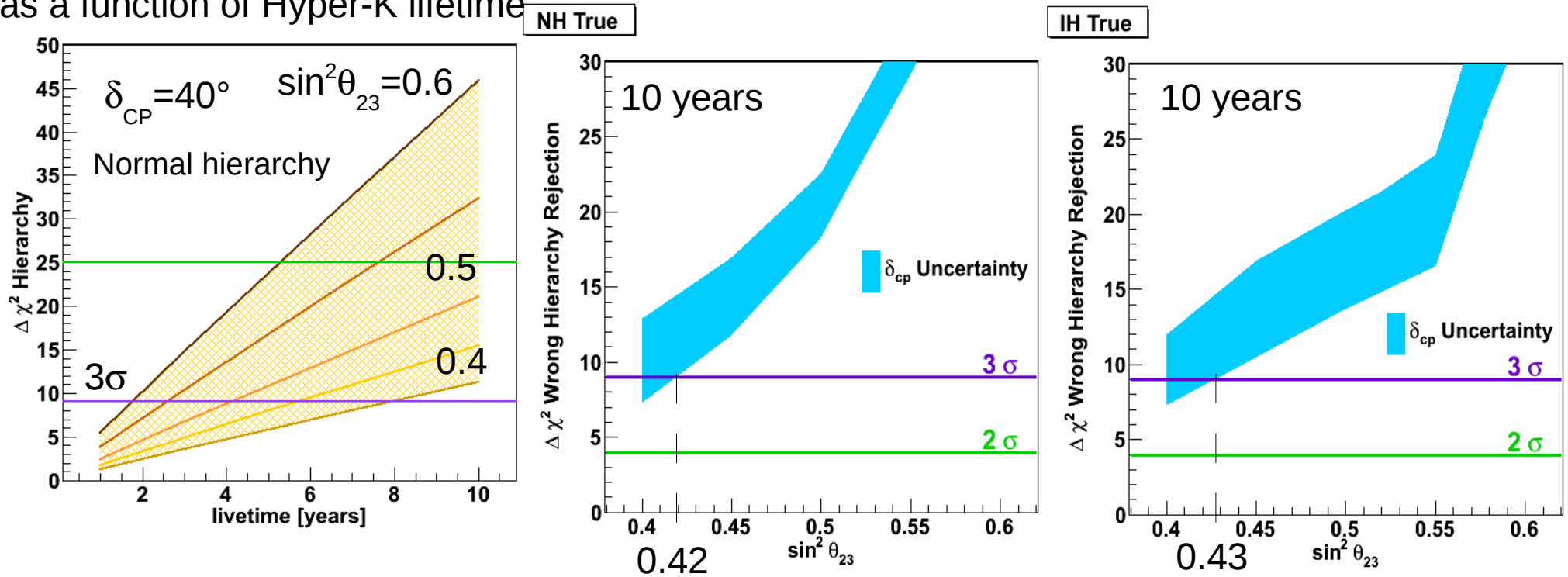
**$1\sigma$  uncertainty of  $\delta$  as a function of the beam power:**

$< 20^\circ$  ( $10^\circ$ ) for  $\delta = 90^\circ$  ( $0^\circ$ )

Modest dependence on  $\theta_{13}$

# Mass Hierarchy Sensitivity

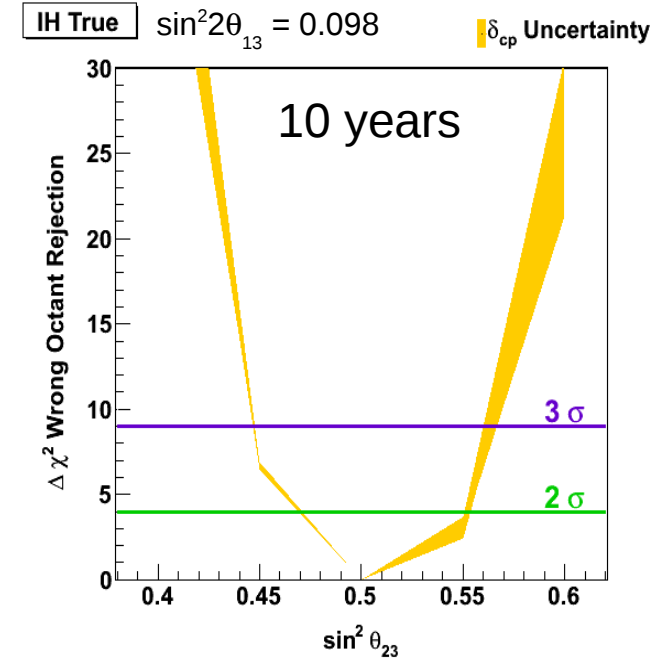
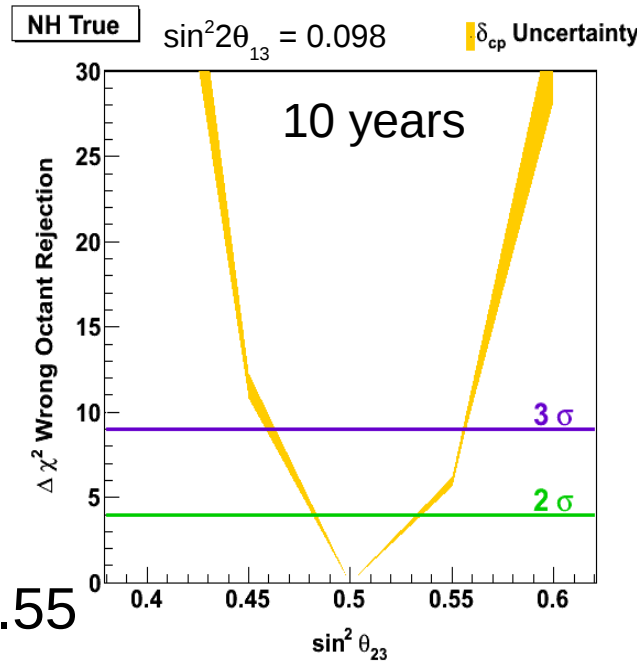
Significance for MH determination as a function of Hyper-K lifetime



- Sensitivity mainly depends on  $\theta_{23}$ ,  $\delta$  and slightly on the MH itself.
- $3\sigma$  mass hierarchy determination for  $\sin^2\theta_{23} > 0.42$  (0.43) for normal (inverted) hierarchy for 10y data taking.
- Caveat: the  $\Delta\chi^2$  method to determine the number of  $\sigma$ 's is used.

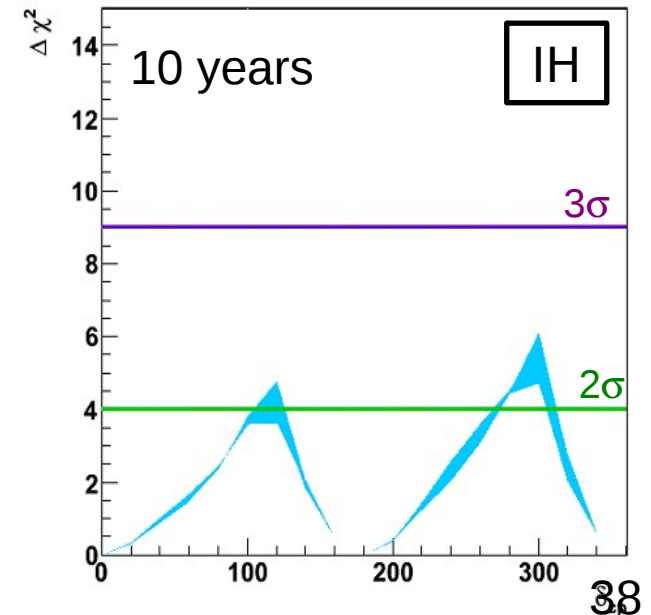
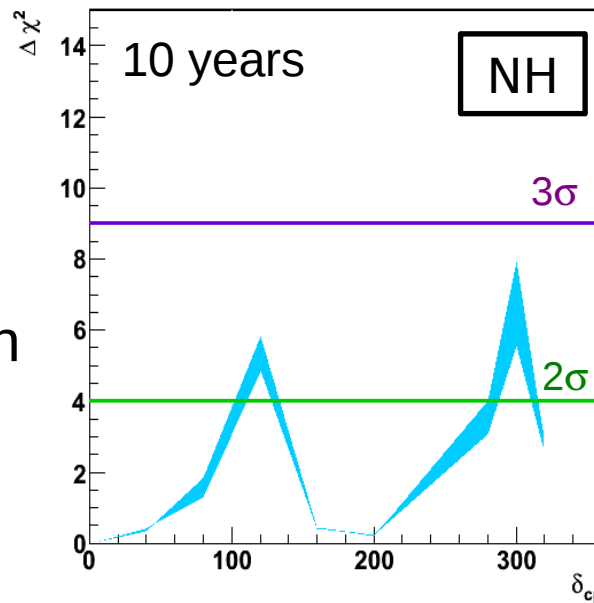
# Sensitivity for $\theta_{23}$ Octant and CPV

- $\theta_{23}$  octant sensitivity.
- Thickness of the band corresponds to the uncertainty from  $\delta_{CP}$ .
- We can expect discrimination between  $\sin^2\theta_{23}$  0.4-0.6, w/ limited discrimination b/w 0.45-0.55

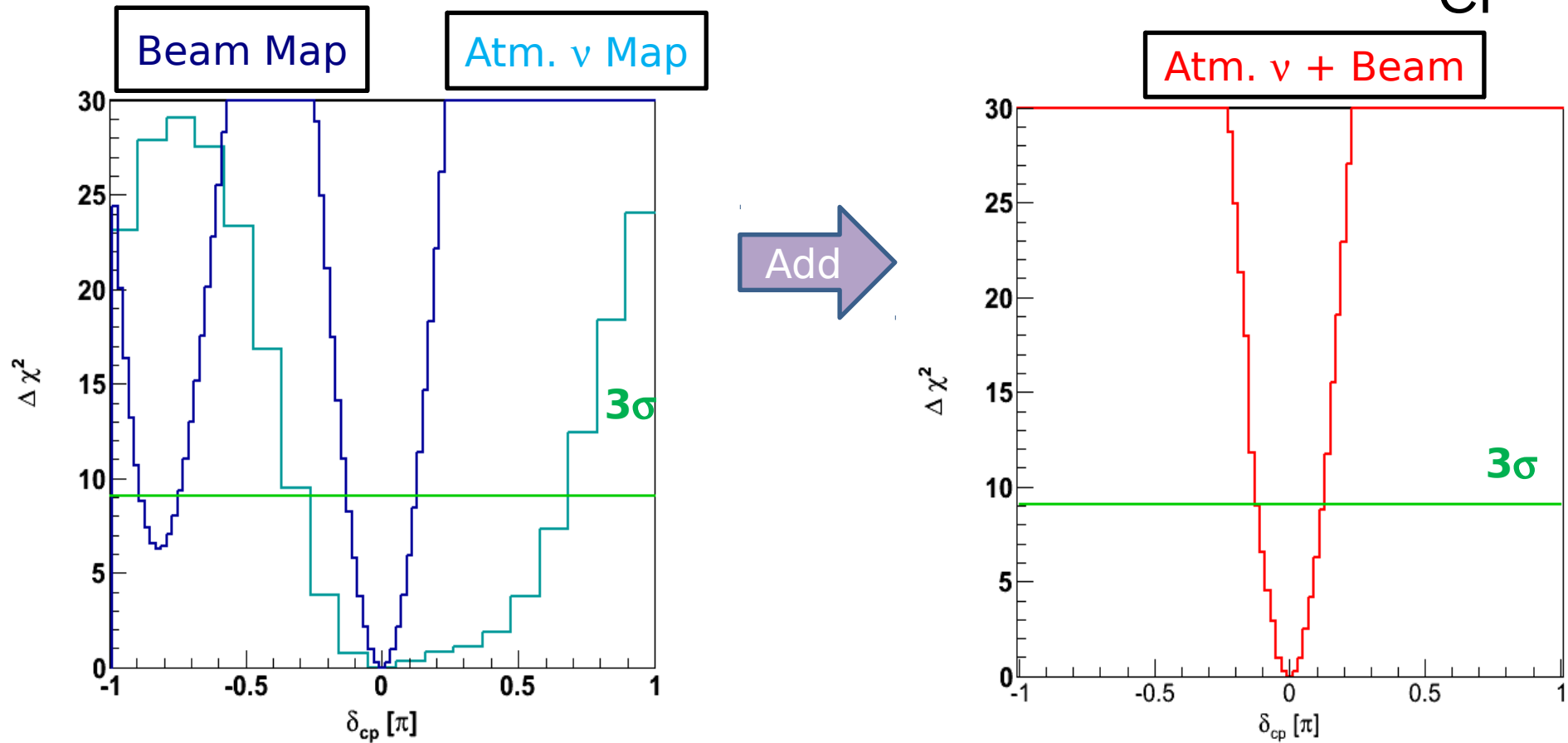


$\theta_{13}$  is fixed :  $\sin^2 2\theta_{13} = 0.099$

- Excluded  $3\sigma$   $\delta_{CP}$  fraction.
- Thickness of the band corresponds to the uncertainty from  $\sin^2 2\theta_{23}$
- Sensitivity to CP-violation is limited under both hierarchy assumptions.



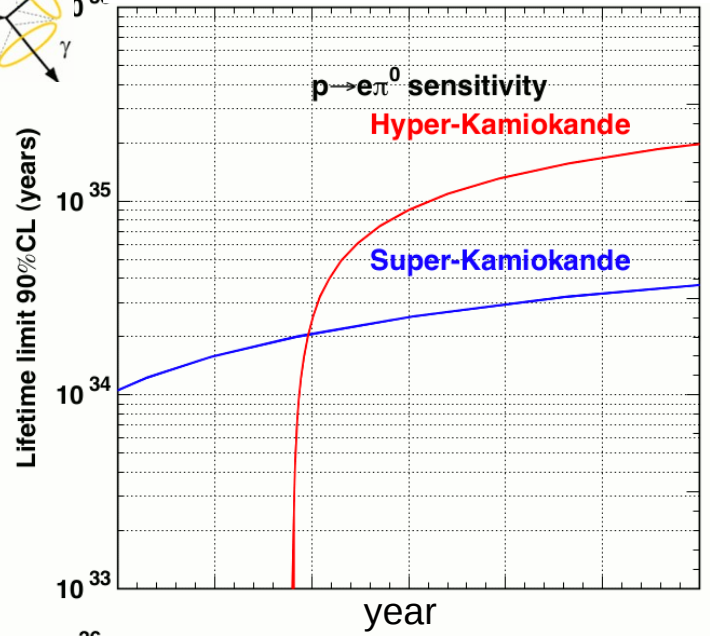
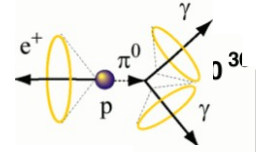
# Beam + Atmospheric $\nu$ : Allowed $\delta_{CP}$



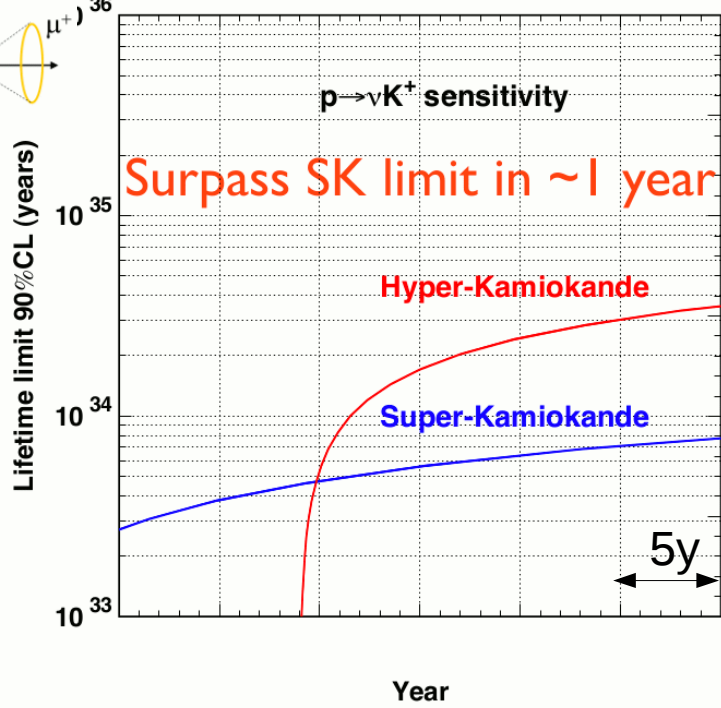
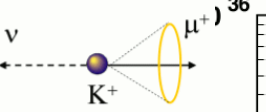
- Hierarchy is unknown, but NH is true.
- True  $\delta_{CP} = 0.0$ ;  $\sin^2 2\theta_{13} = 0.10$ ; Maximal mixing  $\sin^2 2\theta_{23} = 1.0$
- Degenerate solution exists at  $3\sigma$  in the beam only case.

# Proton Decay Sensitivities

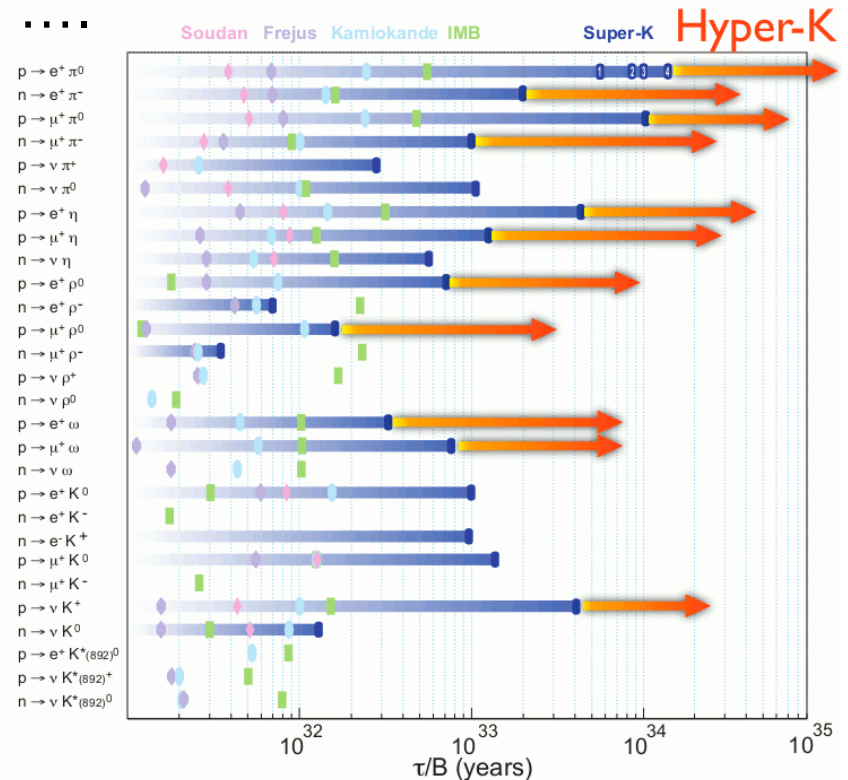
Surpass SK limit in ~1 year



Br = 63.5%



- 10 times better sensitivity than Super-K
- Hyper-K surpasses SK limits in ~1y
  - $p \rightarrow e^+ \pi^0$  :  $1.3 \times 10^{35}$  y at 90%CL
  - $p \rightarrow \bar{\nu} K^+$  :  $2.5 \times 10^{34}$  y at 90%CL
  - Many other modes:
    - $P(n \rightarrow e, \mu) + (\pi, \rho, \omega, \eta)$ ;  $10^{14} - 10^{35}$
    - $K^0$  modes
    - $\nu \pi^0, \nu \pi^+$
    - .....



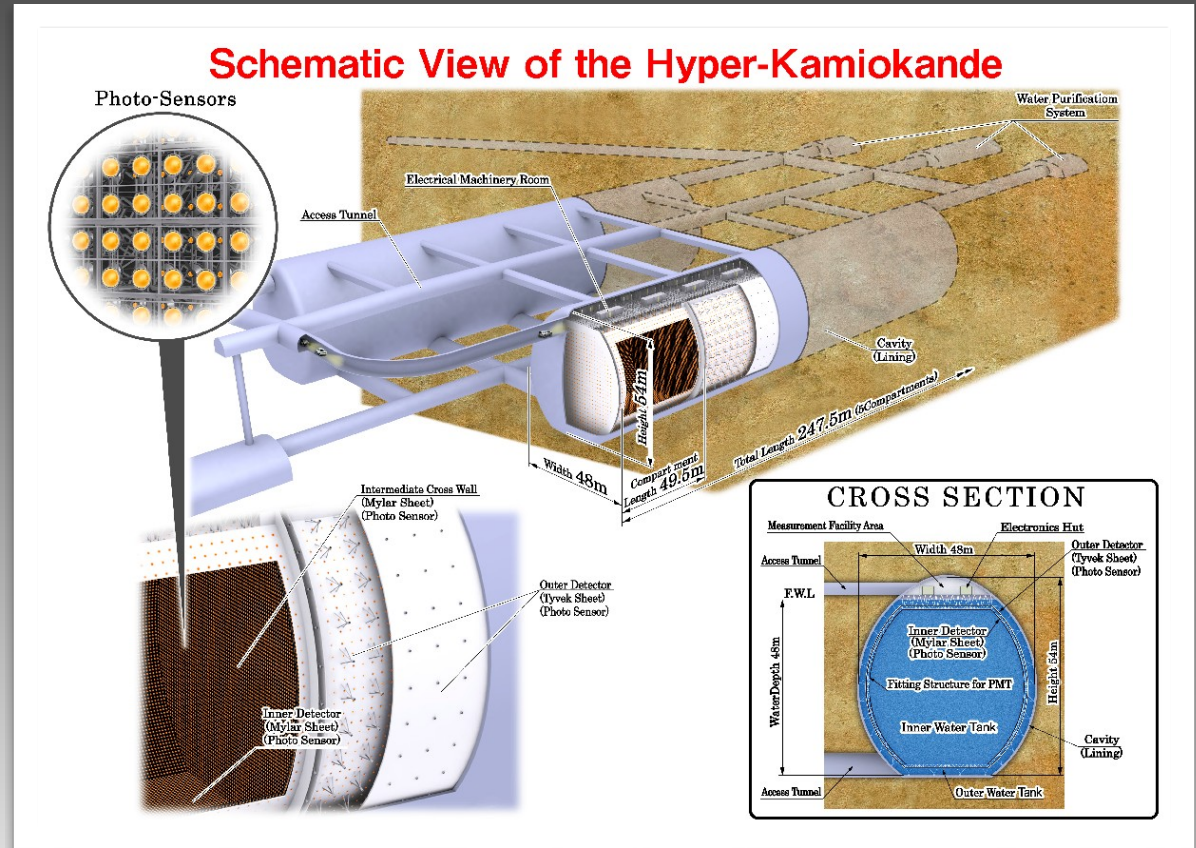


# “Other” Physics Topics at Hyper-K

More physics topics than the ones described can be investigated by Hyper-Kamiokande:

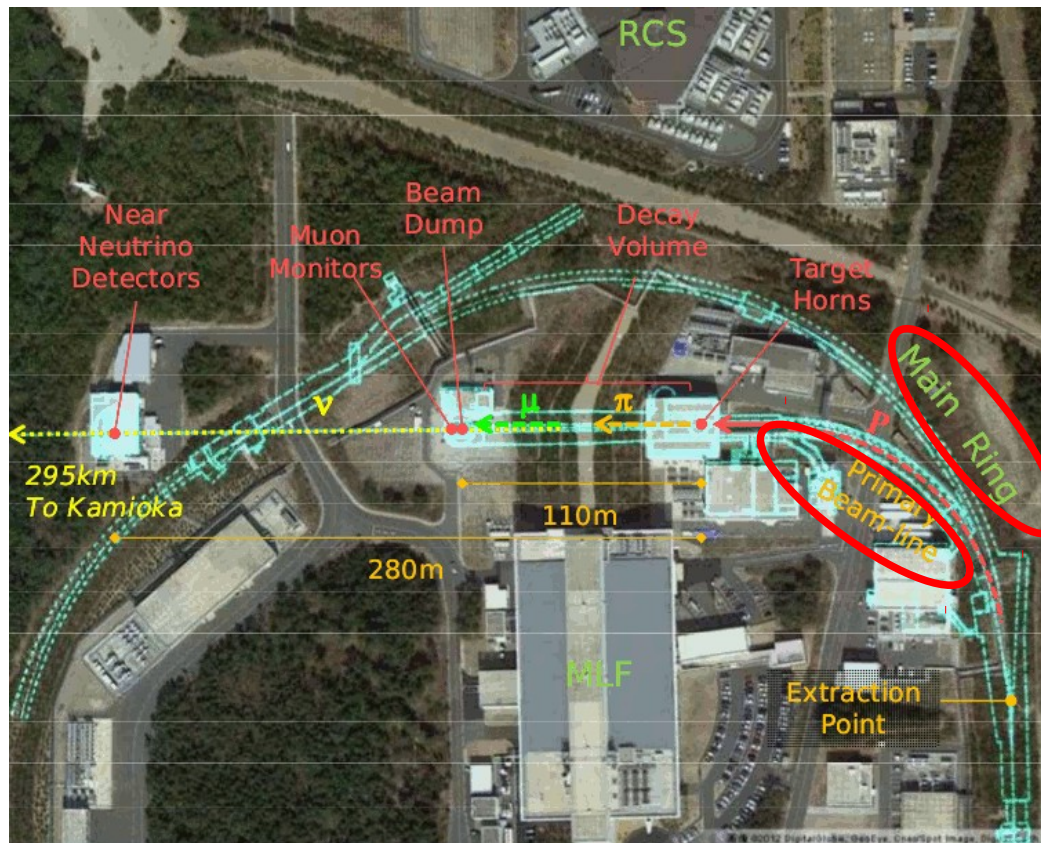
- Solar Neutrinos: 200  $\nu$ 's / day from Sun  $\rightarrow$  day/night asymmetry of the solar neutrinos flux can be precisely measured at HK.
- Solar flares can be discovered at Hyper-K (important information about particle acceleration at work in solar flares)
- Astrophysical neutrinos:
  - 200k  $\nu$ 's from Supernova at Galactic center (10kpc)  
 $\rightarrow$  time variation & energy can be measured with high statistics
- Indirect dark matter search, excellent capabilities at low mass region.
- Neutrino geophysics: neutrino radiography w/ atmospheric neutrinos for surveying the internal structure of the Earth.

# Experiment Status



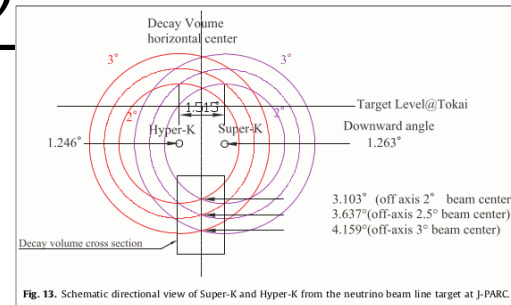
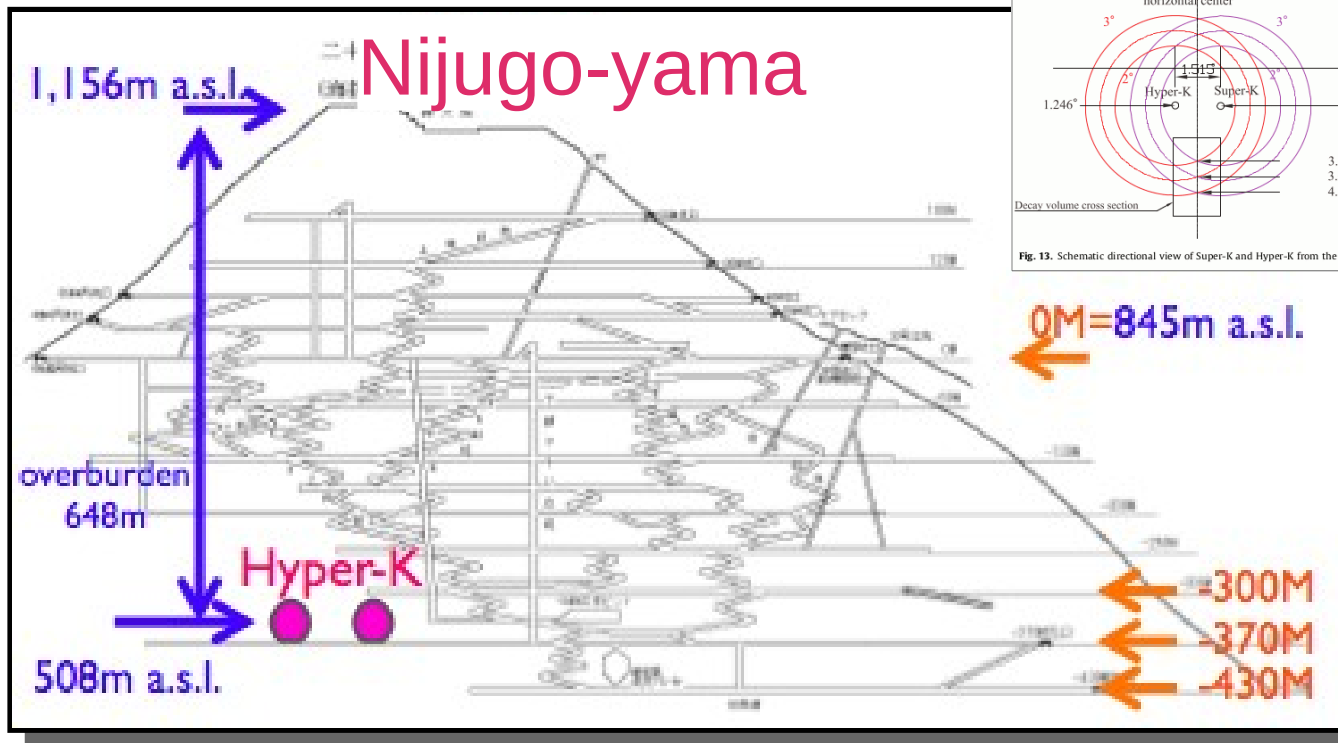
# Beam for Tokai-2-Hyper-Kamiokande

- Next upgrade (intermediate plan) towards a 750kW operation.
- UK/RAL contributions for T2K and planned for Hyper-K.
- The upgrade will concern:
  - Upgrade plan for J-PARC accelerators.
  - Upgrade plan for the neutrino beam-line to accept a 750MW beam.



# Candidate Site: Tochibora Mine

- Located under “Nijugo-yama” (Mt. 25), ~8km south from Super-K
- Identical baseline (295km) and off-axis angle ( $2.5^\circ$ ) to T2K
- Overburden ~650m (~1755 m.w.e.)



- The candidate site vicinity used for mining. Many existing tunnels and shafts.
- Historically many surveys have been done in wide area and at several levels/depths, especially mapping the location of faults.
- Confirmed that the HK cavern can be constructed w/ existing techniques.

# Geological Survey at Mozumi Mine

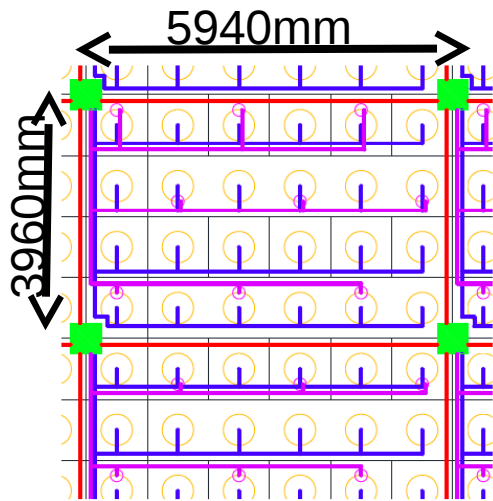
- Geological survey at the Mozumi mine, already used for Super-K, recently started, to have a deeper cavern ( $\sim 800\text{m}$ ).
- First rock mass characterization has been done: rock quality at Mozumi-site is comparable with Tochibora-site.
- More tests under way to complete the geological survey.
- Note: Tochibura and Mozumi are on the opposite sides of the beam, but same off-axis angle ( $2.5^\circ$ ).



# Design Work...

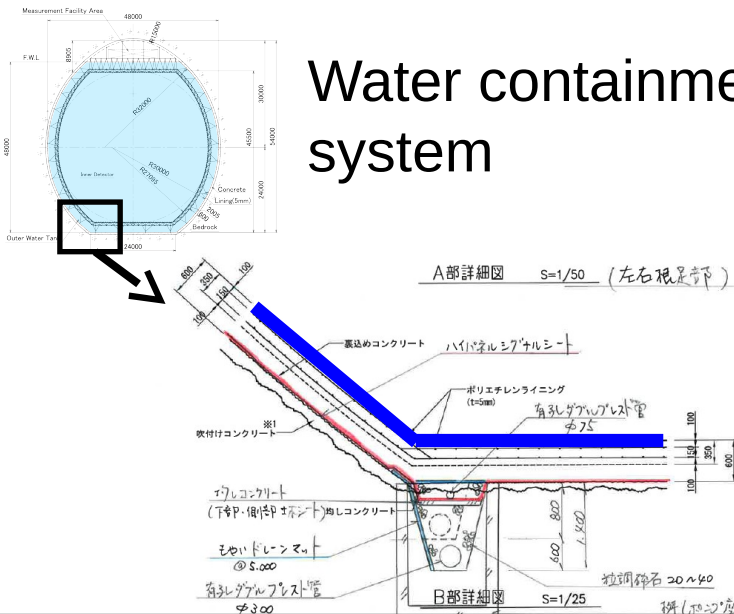
All major part of HK tank has been designed, water containment system, photosensors support, layout of water pipes, front-end electronics, cables, calibration holes, plug manholes, ... etc.

## Electronics & cable layout

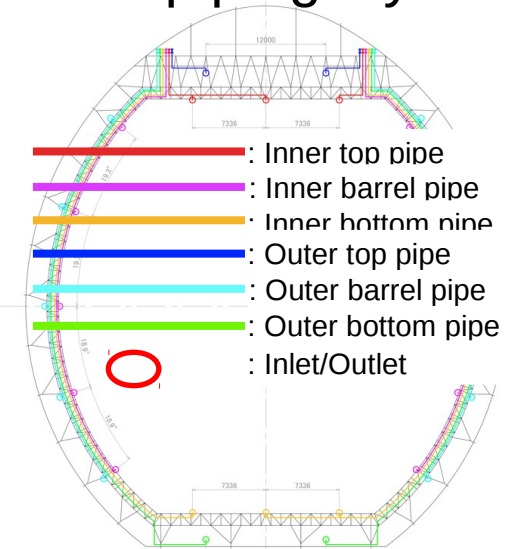


- : Support structure
- : Cable for inner PMT
- : Cable for outer PMT
- : Network/Power cable
- : Hub / Front End Electronics
- : Inner photo-sensor (20")
- : Outer photo-sensor (8")

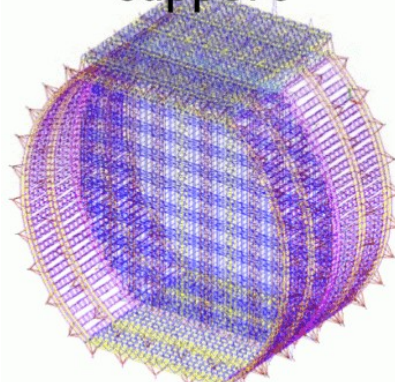
## Water containment system



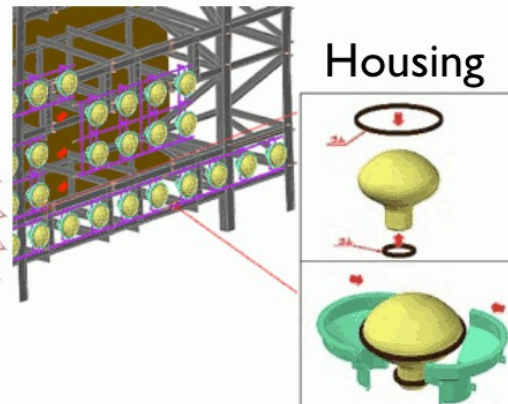
## Water piping layout



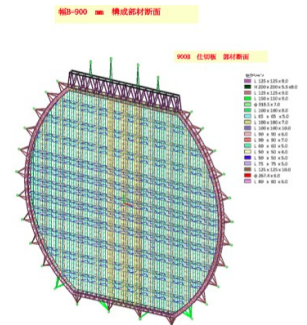
## Photo-sensor support



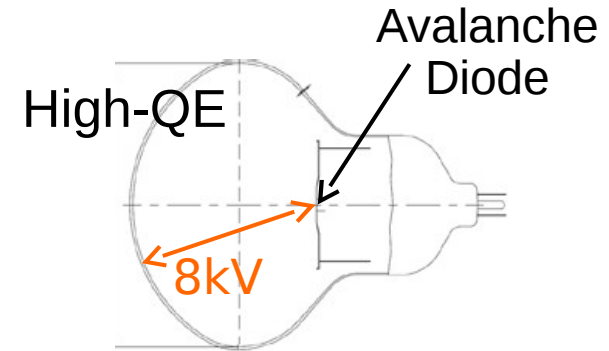
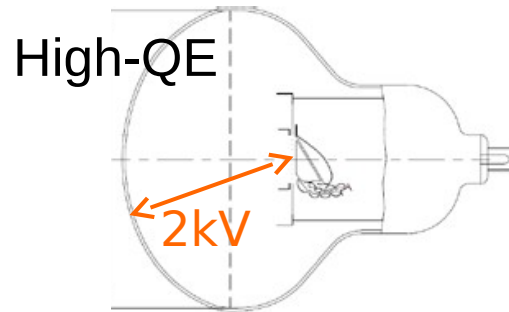
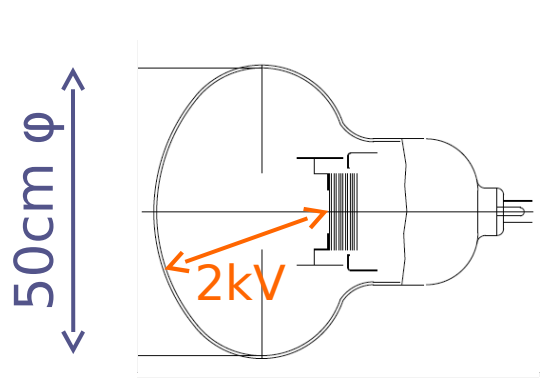
## Mounting Photo-sensor Housing



## Separation wall



# Photosensors Candidates



**20" PMT**  
(Venetian-Blind dynode)

- Super-K ID PMTs
- Used for ~20 years
  - Guaranteed
- Complex production
  - Expensive

**20" Improved PMT**  
(Box&Line dynode)

- Under development
- Better performance
- Same technology
  - Lower risk

**20" HPD**  
(Hybrid Photodetector)

- Under development
- Far better performance
- Simple structure
  - Lower cost
- New technology
  - Higher risk

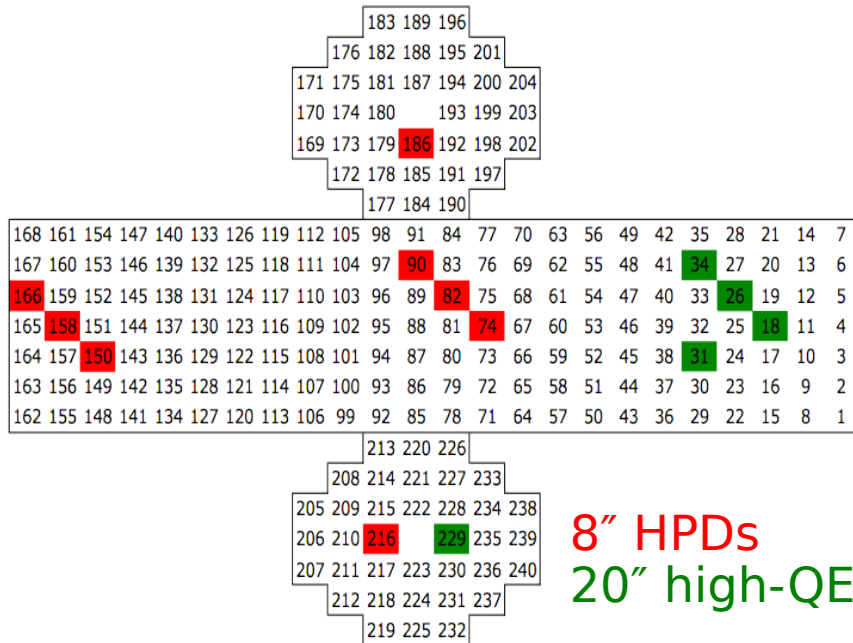
Lower Risk



Higher Performance

# Tests in a Water Cherenkov Detector

- EGADS detector : a 200 ton scale model of Super-K
  - To demonstrate the safety and effectiveness of “SK + Gadolinium”
  - 240 inward-facing photodetectors
  - Electronics : ATMs (used in SK-1,2,3), to be upgraded to QBEEs (SK4)
- Eight 8" HPDs and five 20" high-QE PMTs were mounted
  - Other 227 photodetectors are R3600, and can be used as references for the new photodetector evaluation



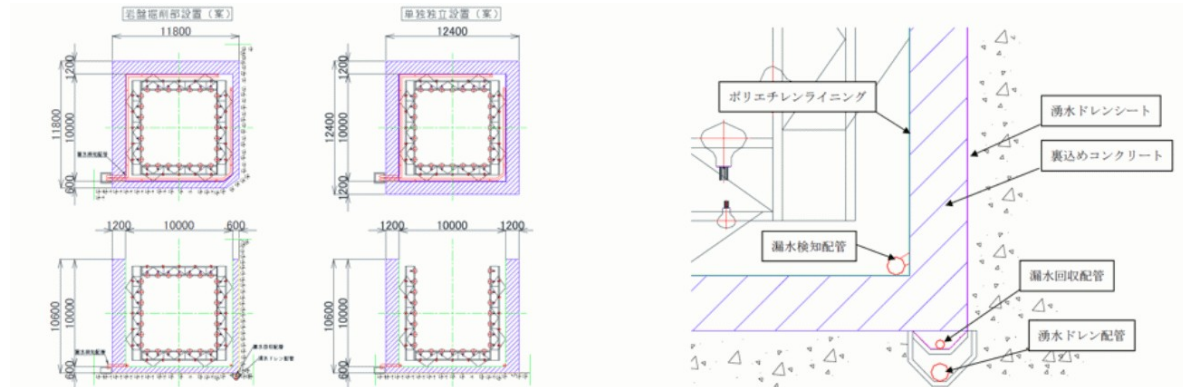
8" HPDs  
20" high-QE PMTs





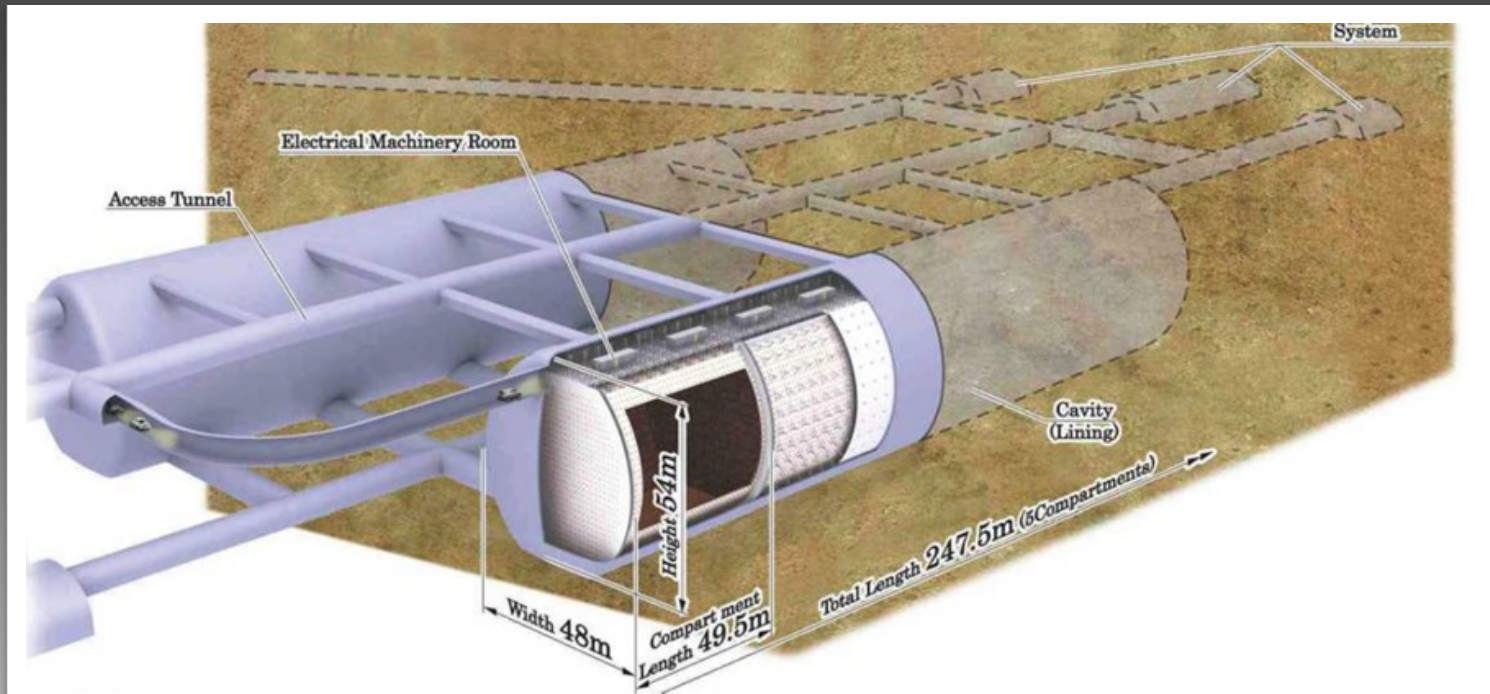
# 1kton WC Prototype

- Prototype (1kton,  $\sim 10 \times 10 \times 10 \text{ m}^3$ ) for R&D test approved in Japan as Grant-in-Aid:  $\sim \text{USD } 1.7\text{M}/5 \text{ years}$  (2013-17).
- It's one of the 20 proposals selected each year from all areas in science.
- Main feasibility studies:
  - Photosensor and corresponding support structure
  - Liners
  - Leak water collection detection
  - DAQ
  - Electronics
  - Calibration system
  - ....



- Location site (J-PARC, KEK, Kamioka) being discussed.

# UK Interests



# Summary of Current Areas of Interest

## Beam

contributed to T2K. Already intense work towards a  $\sim$ MW beam.

## ND280 upgrade

neutrino interaction measurements

## Intermediate Detector

work started towards the optimization of the design and requirements for an intermediate detector at  $\sim$ 2km.

## DAQ & Electronics

very successful performance in T2K. New ideas being investigated.

## Calibration

huge expertise (mainly SNO, SNO+). Starting to work on it.

## Photosensors

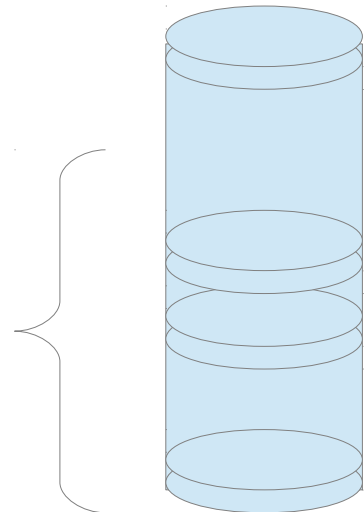
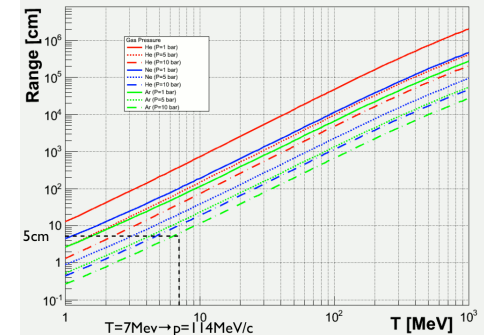
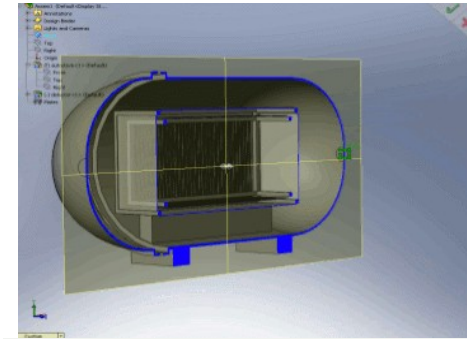
LAPPDs for Intermediate detector.

## Software/Computing

working on computing model.

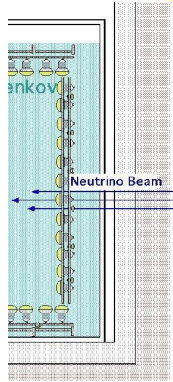
# ND280 Upgrade for T2K (T2HK)

- Several studies being performed for upgrade → beneficial for T2HK as well, if keeping ND280.
- Water target with vertex information
  - water based scintillator for P0D/FGD
  - high pressure TPC, fiber tracker:
    - × Access the low energy debris and help in the study of neutrino-nucleon interactions
    - × close to Oxygen, much lower energy threshold
  - water detector at the basement (B2) of ND280
- Enhancement on side/backward going tracks
  - Trip-t electronics upgrade or better calibration
- Neutrino-nucleon cross section for model input
  - D2O and CH2 targets for FGD/P0D
- New “water-column” detector at ~1km
  - Minimize dependence on neutrino interactions by sampling the beam at several off-axis angles.

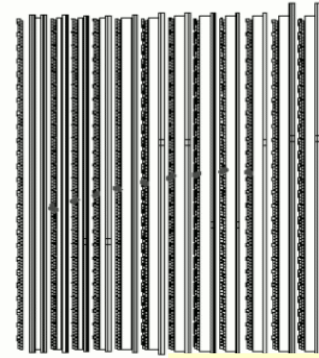


# Intermediate Detector

PMTs facing outwards in the veto region to reject particles that enter or are produced outside the detector

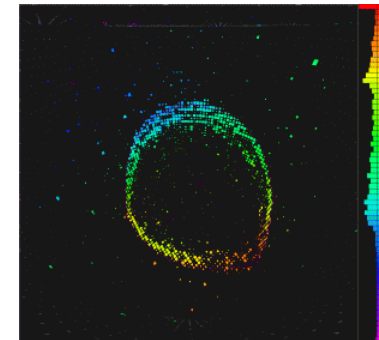
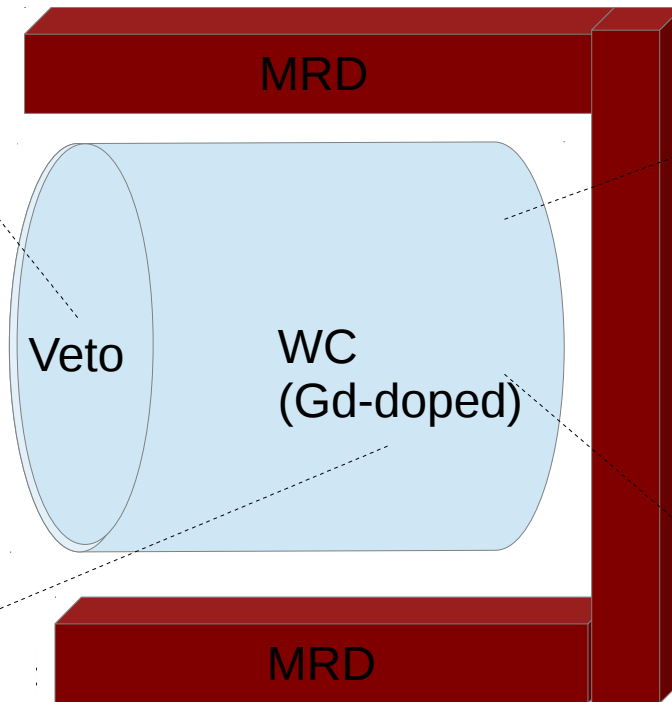


To measure energy, angle and production point of muon from CC interactions

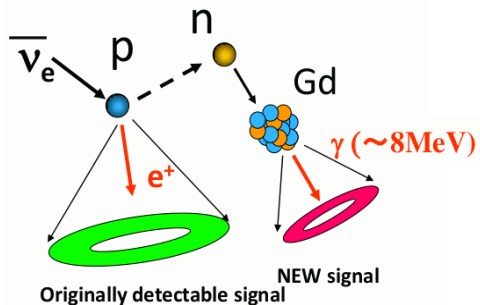


Water-Cherenkov target

$\nu$  beam



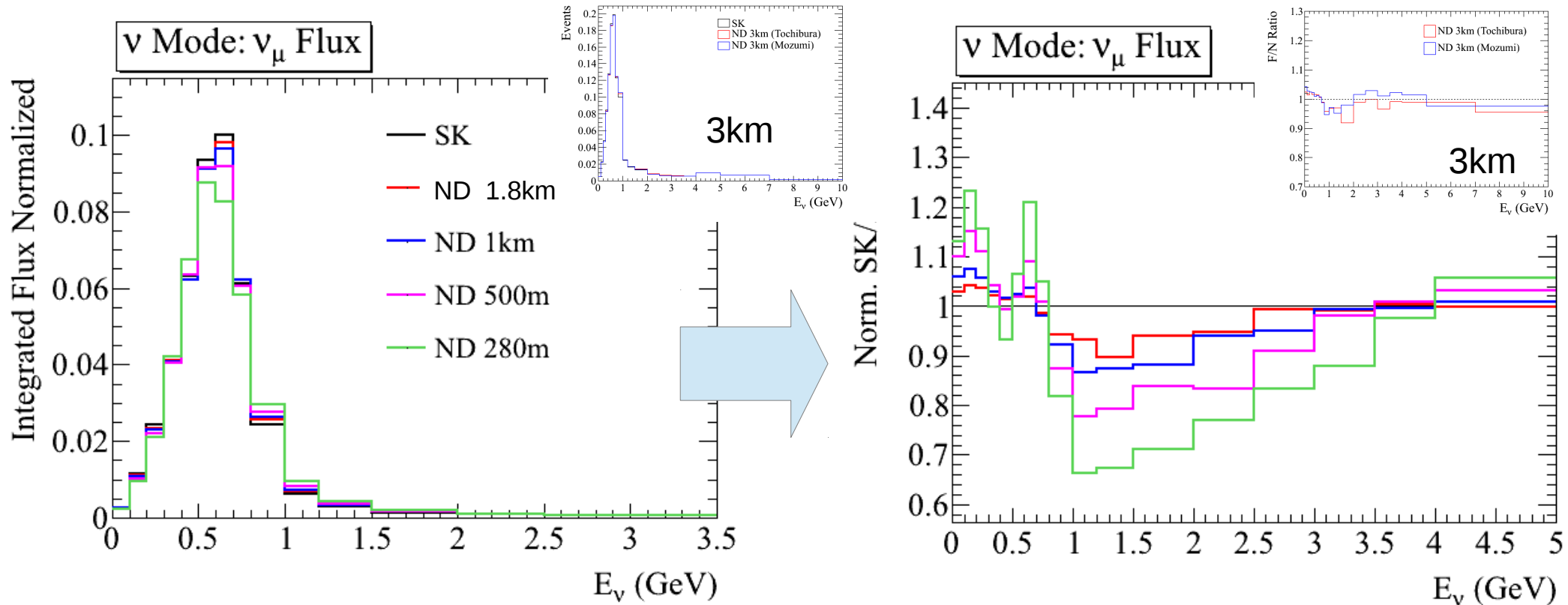
Aim to use LAPPDs (w/ HPD) Large Area Picosecond Photo-Detector (micro-channel plate photosensor)



Aiming for Gd doped water.



# Intermediate Detector



- At 280m: neutrino source not point-like, spectral differences with respect to SK
- Neutrino spectra at SK and 3km are almost the same: ~same beam → energy spectrum
- To improve our current precision we need to improve our errors on the flux predictions

# Detector Sites

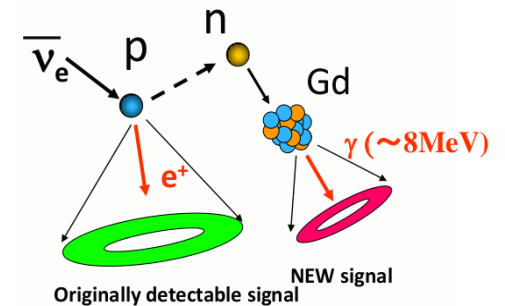


- 🌍 Site optimization:

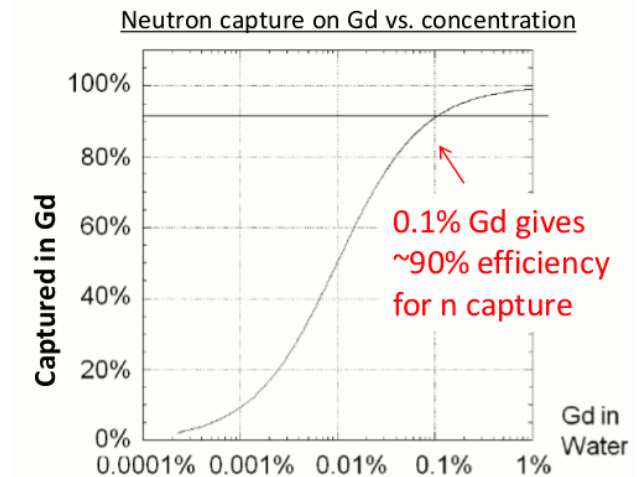
- compromise between physics and land availability.
- 🌍 A site at ~2-3km will see similar spectrum to Hyper-K.
- 🌍 Ongoing work to study cross section errors, pile-up, size, location, etc.
- 🌍 In the past, already performed investigations at 1.8 GeV.

# Gd-doping

- Aim to dope the experiment with Gd.
- For 0.1% concentration of Gd 90% of neutrons are captured on Gd rather than thermalized in water.



- We recently started to explore this idea. Within SK there are already studies performed with EGADS/GADZOOKS! at the Kamioka mine



## • CP-related goals:

- Use neutron tagging to reconstruct neutrino energy
- Use to separate CC versus NC interactions. In neutrino mode, neutron multiplicity is expected to be lower for CC interactions.
- Separation neutrinos versus anti-neutrino. More neutrons for anti-neutrinos. It's important to measure the neutrino component in an anti-neutrino beam.



# Gd-doping

## • Xsection physics:

- Study nuclear models in details. A predicted effect of two-body currents is a high nucleon multiplicity in the final states.

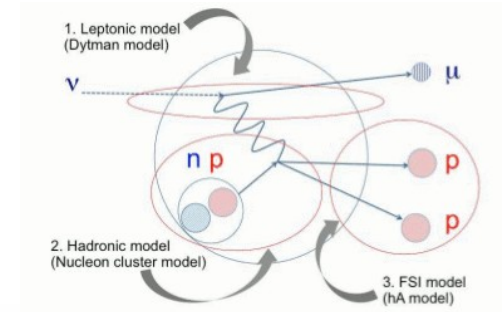


FIGURE 1. Basic strategy of modeling MEC in GENIE.

T.Katori  
arXiv:1304.6014

TABLE 1. Comparison of MEC models in neutrino interaction generators.

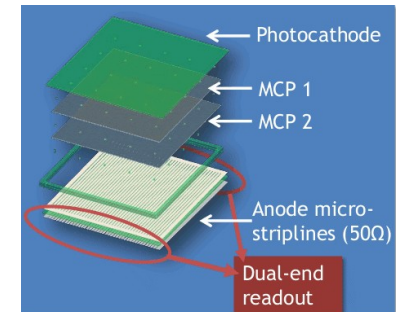
	GENIE	NuWro	GiBUU
Leptonic model	Dytman model	TEM, np-nh model, and Valencia model	Transverse projector
Hadronic model	nucleon cluster	nucleon cluster	phase space density
initial nucleon momentum	Fermi sea	Fermi sea	Fermi sea
initial nucleon momentum correlation	none	none	none
initial nucleon spatial correlation	none	none	2 nucleons are generated at the same location
initial nucleon pair	n-p:n-n=1:4	n-p:n-n=9:1	n-p:n-n=12:5
FSI model	isospin ansatz hA model	short range correlation cascade model	statistical average BUU transport

## • “Other” physics:

- Neutron interaction rate, relevant as a background for proton decay (currently will be investigated by ANNIE, but this detector has a later timescale and larger)
- SN neutrinos → lower statistics (~500 evts?) than HK, but it is relevant to get an SN alarm in coincidence with HK.

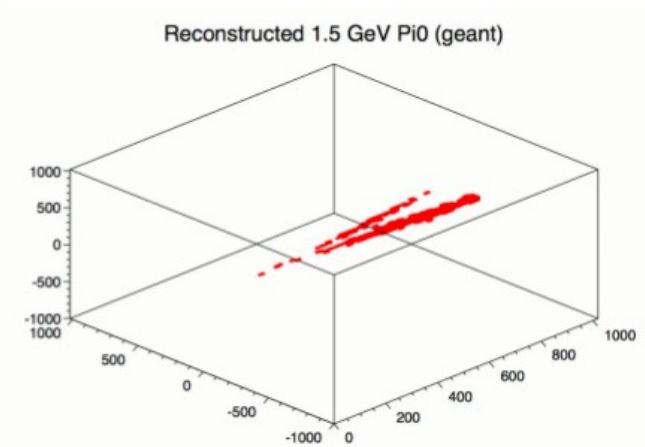
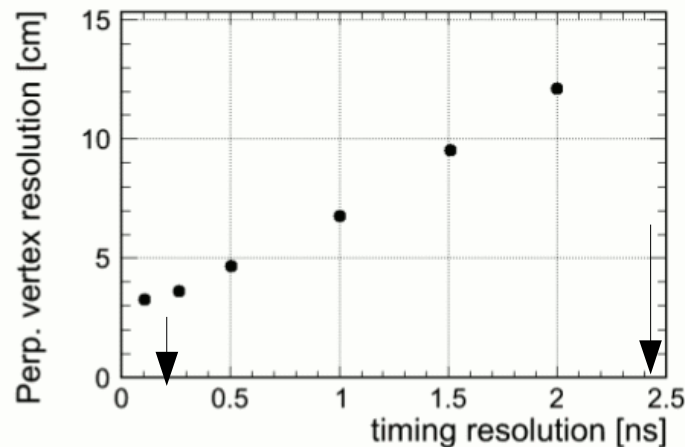
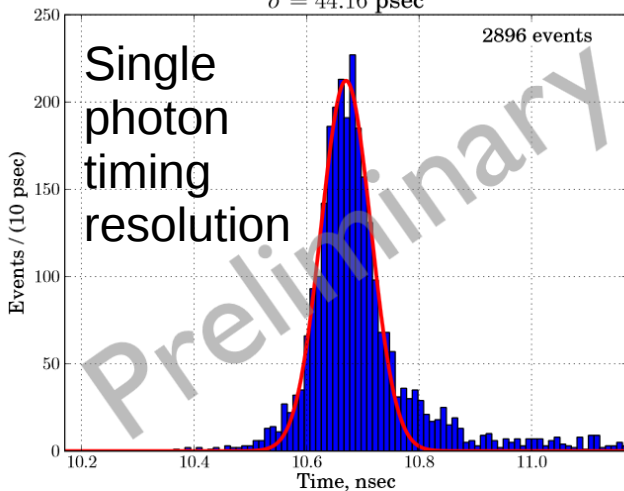
# LAPPDs

- Investigating the option of using the LAPPDs (Large Area Picosecond Photo-Detectors).
- LAPPD, developed at Argonne National Lab, is based on a micro-channel plate.
- Improved timing resolution, currently limited by PMT transit time spread (2-5ns per photon).
- LAPPDs show the benefit of excellent single timing resolution of  $\sim 50\text{ps}$ 
  - Improved vertex resolution
  - Improved spacial resolution

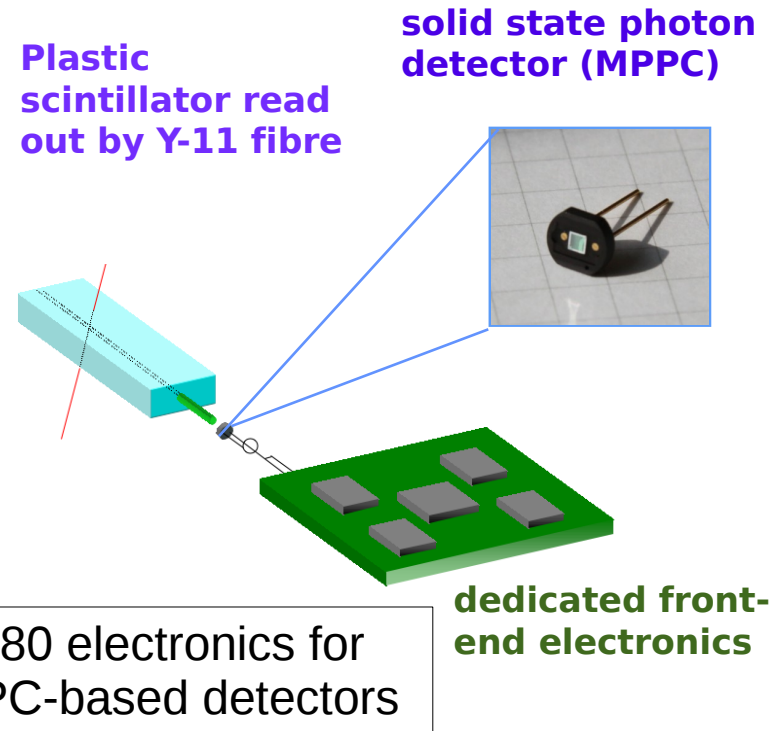
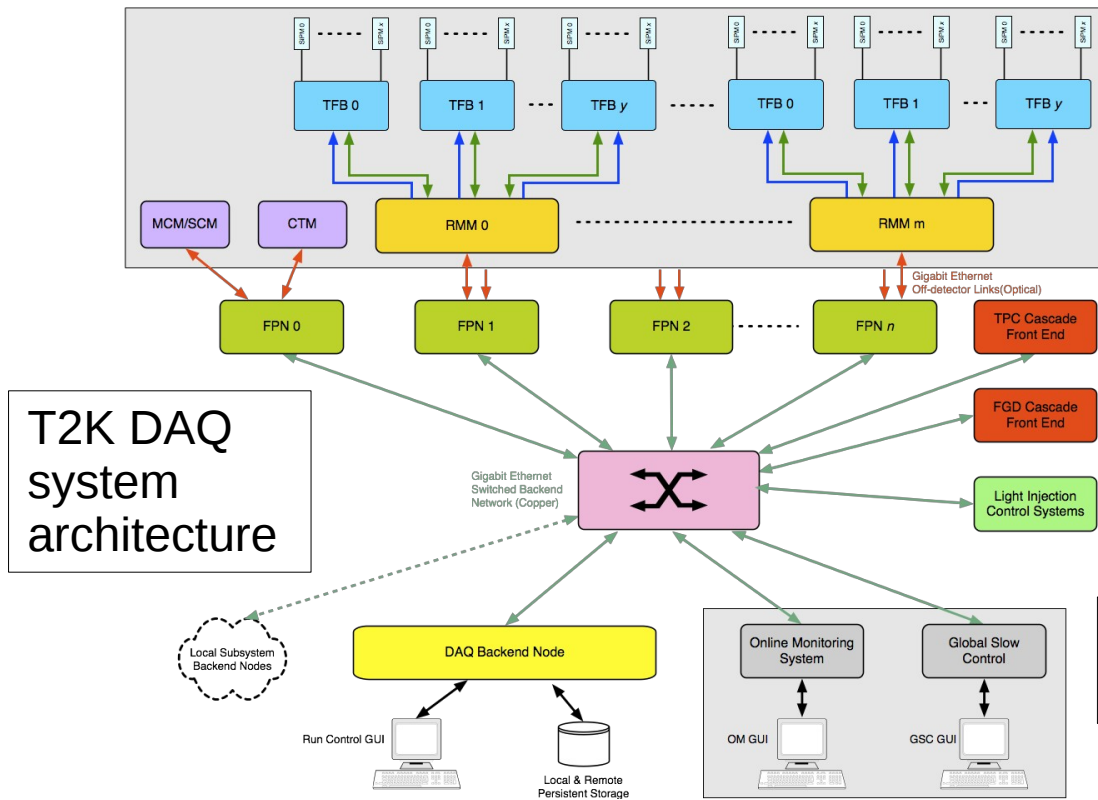


T.Xin, I. Anghel, M. Wetstein, M. Sanchez

$\sigma = 44.16 \text{ psec}$



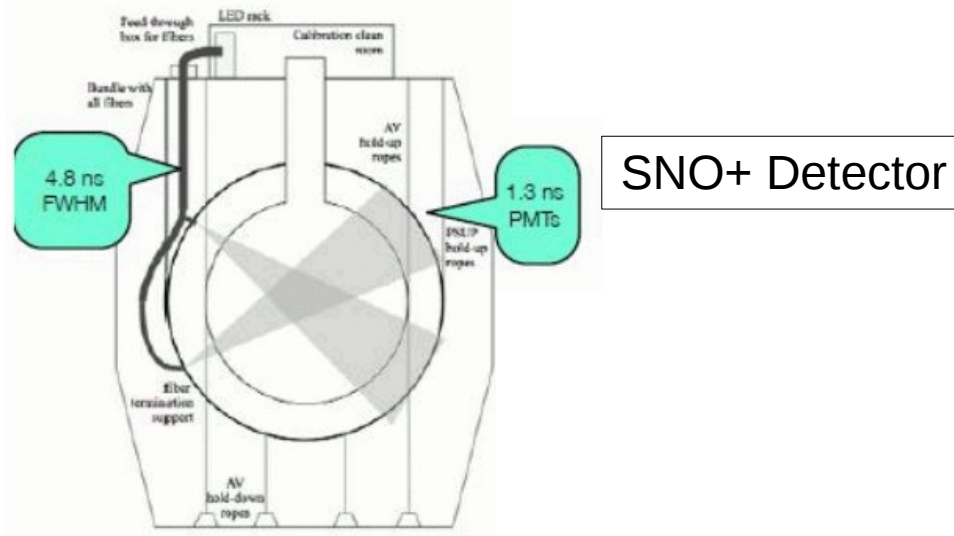
# DAQ & Electronics (for Hyper-K and Intern. Det.)



- DAQ & electronics are running stably since the start of T2K.
- Based on the previous experience, we are devising the DAQ for HK and Intern. Det. – new ideas/design being discussed.
- We are also interested in the electronics based on both T2K and other experiment expertise.

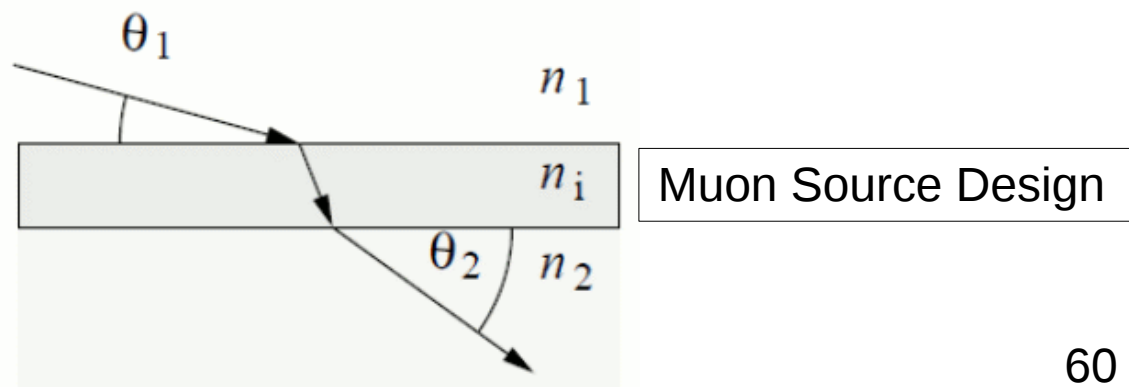
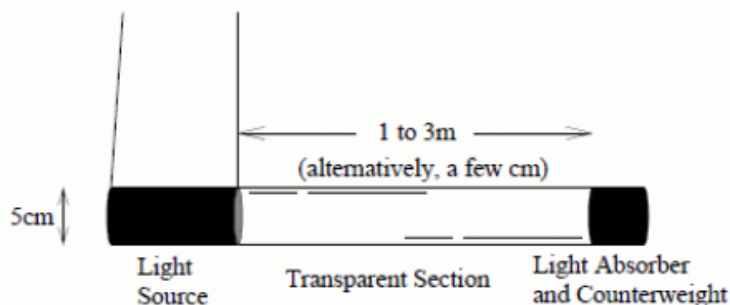
# Calibration Strategy (for Hyper-K and Interm. Det.)

- Exploit current expertise (e.g. SNO+, ANTARES..)



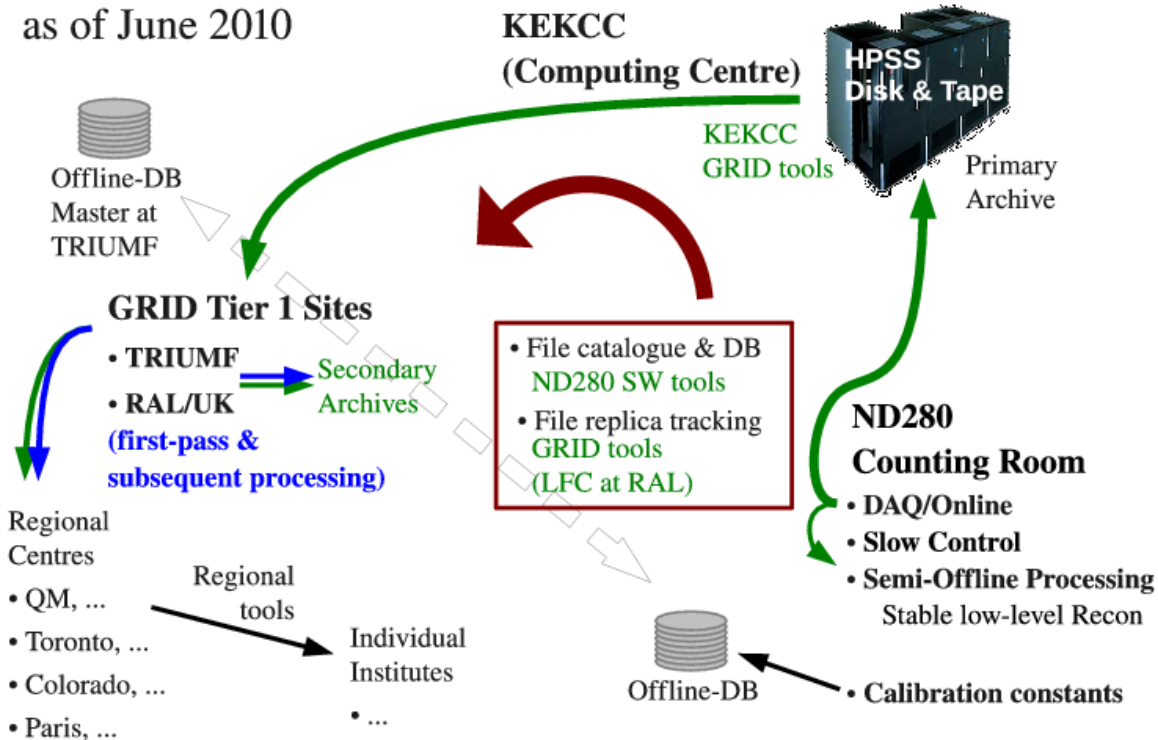
- Some initial work:

- Development of updated LED drivers for HK
- A source to simulate muons and test reconstruction

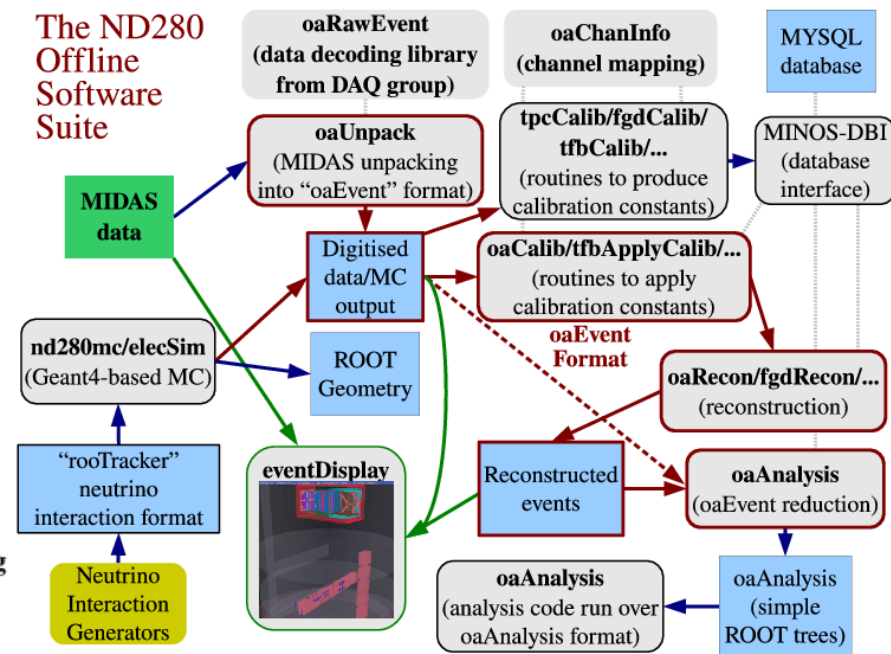


# Software/Computing

ND280 Data Distribution Implementation  
as of June 2010

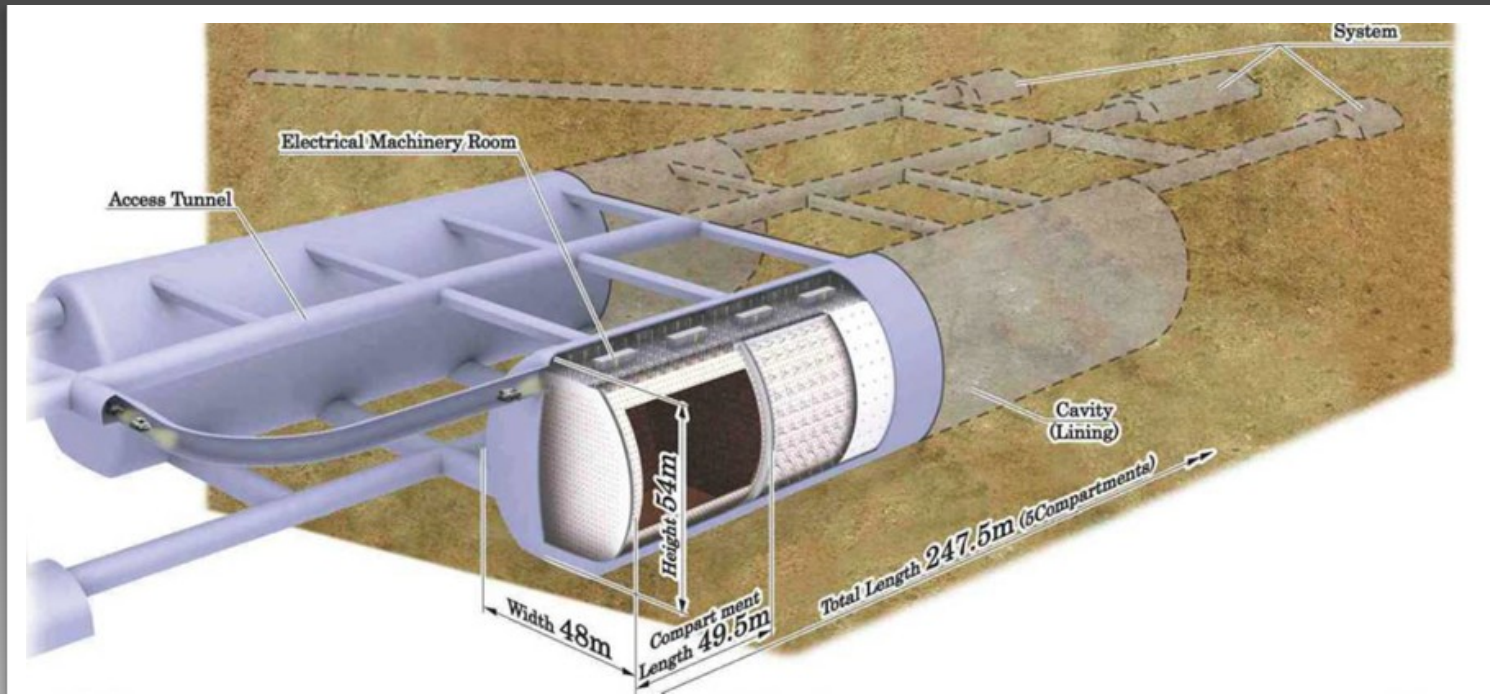


The ND280  
Offline  
Software  
Suite



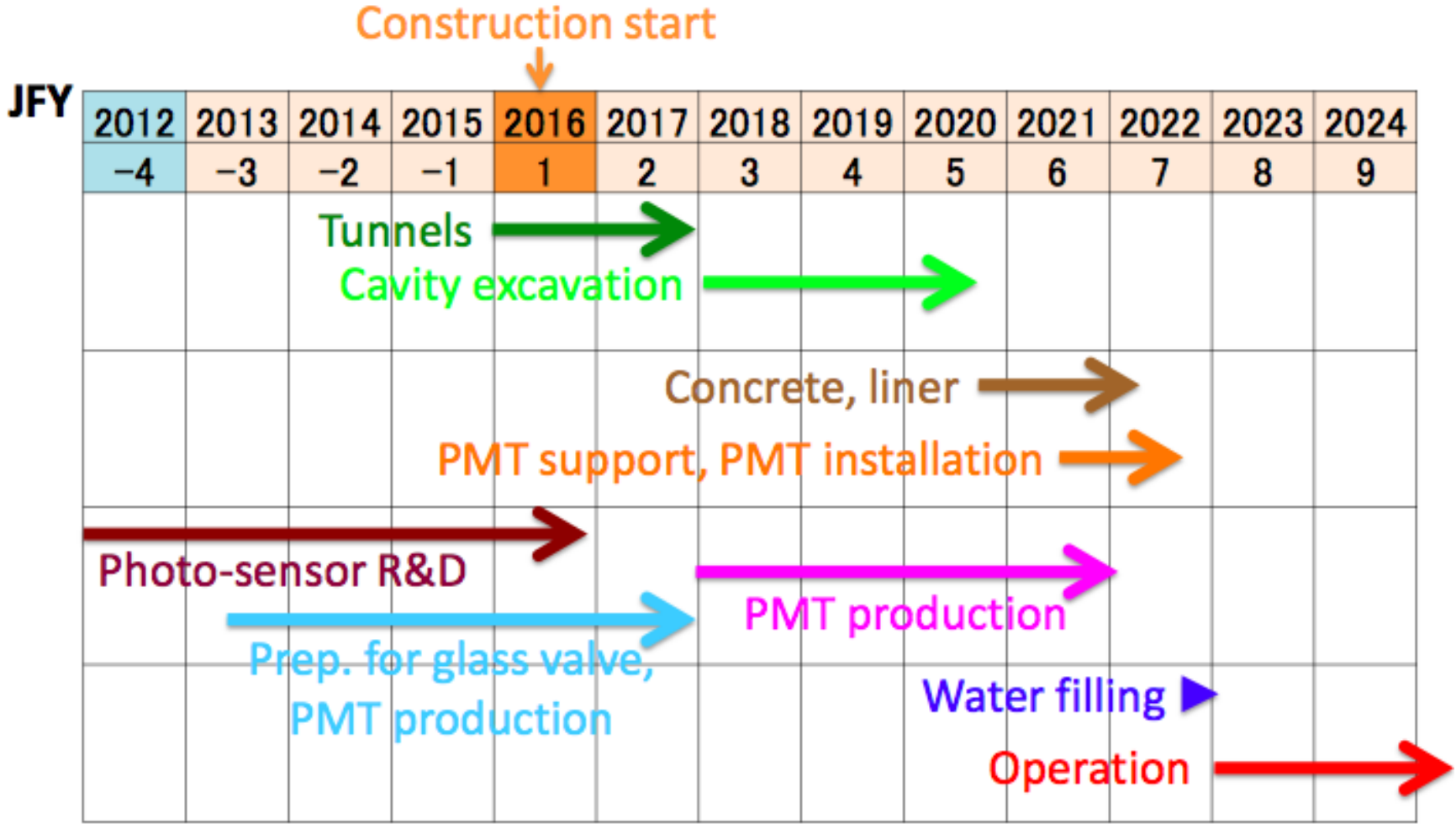
- Central role in software/computing for T2K.
- Already working on Hyper-K computing model.
- Currently producing Hyper-K simulated events.

# Schedule & Summary



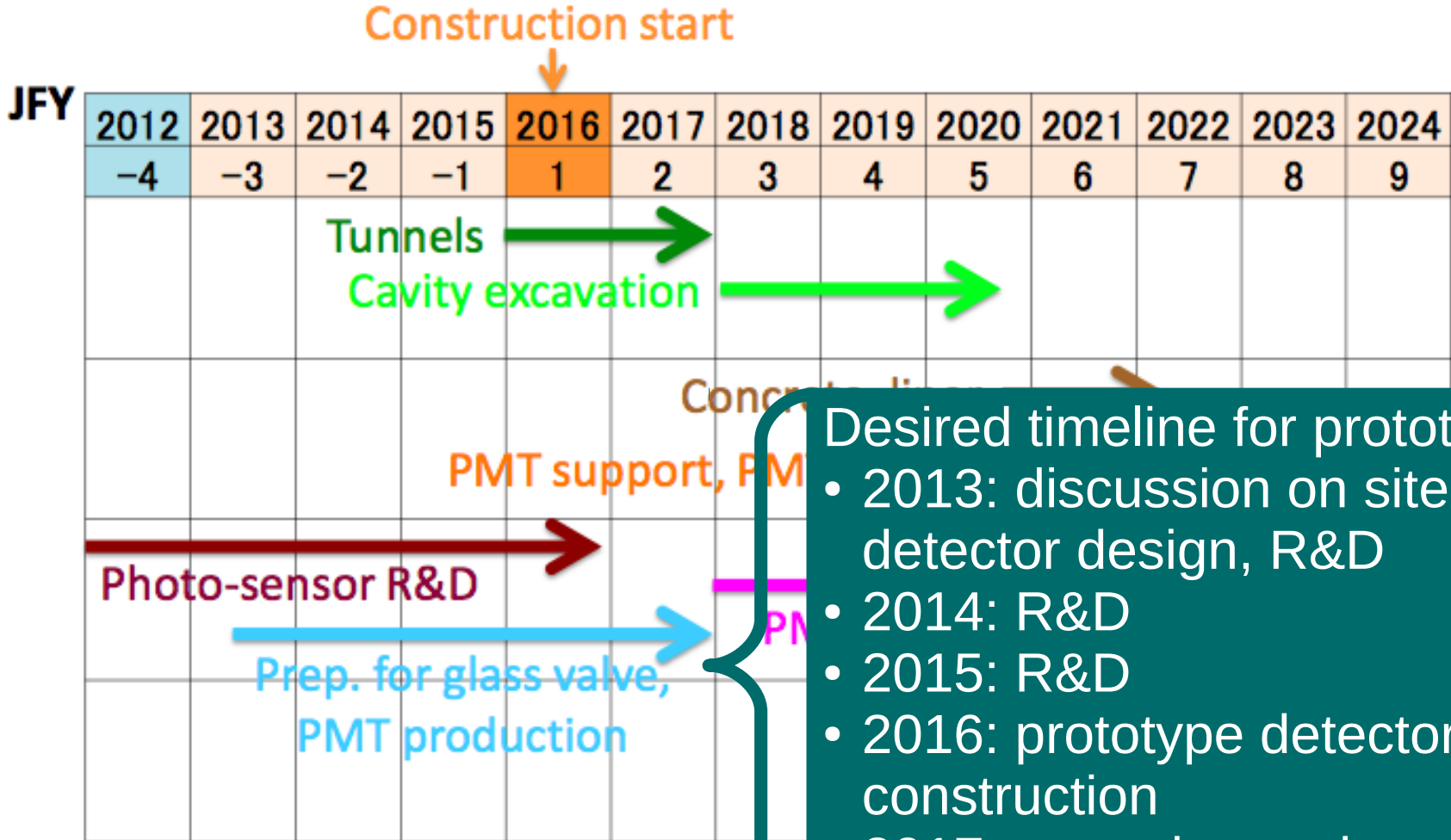
# Overall Project Schedule

- Overall HK construction: ~7 years
- Assuming full funding starting in 2016.



# Overall Project Schedule

- Overall HK construction: ~7 years
- Assuming full funding starting in 2016.



Desired timeline for prototype

- 2013: discussion on site, detector design, R&D
- 2014: R&D
- 2015: R&D
- 2016: prototype detector construction
- 2017: operation and conclusion on components



# Approval Status in Japan

- Approval of the experiment happens in just one phase that allocates the total funds for the experiment, in a given timeline.
- Community consensus crucial → bottom-up approach.
- R&D budget (w/ WC proto-type) for Hyper-K approved in 2013.
- Recommended by HEP community as one of the two major large scale projects: [http://www.jahep.org/office/doc/201202\\_hecsubc\\_report.pdf](http://www.jahep.org/office/doc/201202_hecsubc_report.pdf)
- KEK Roadmap includes Hyper-K:  
<http://kds.kek.jp/getFile.py/access?sessionId=1&resId=0&materialId=0&confId=11728>
- Cosmic Ray community endorses HK as large scale project.
- › Science Council of Japan master plan:
  - › Proposal submitted, expecting outcome soon.
- MEXT:
  - › Based on the SCJ master plan the MEXT will update the roadmap of the big projects. We should prepare report to MEXT in 2014.
- Lol to submit to J-PARC for T2HK in April 2014.

# Approval Status outside Japan

## 🌐EU:

- Statement-of-interest approved in the UK (2013). Proposal to STFC to be submitted in May 2014. Awarded “bridging” money to fill the gap up to the proposal approved.
- Hyper-Kamiokande strongly supported in other T2K Countries. New non-T2K Countries interested.

## 🌐Canada:

- Proposal to Canadian Foundation per innovation under preparation to submit around June 2014.
- Green light from TRIUMF to proceed.

## 🌐US:

- Under discussion in P5.
- Historically strong commitment to Super-K, K2K, T2K, and generally experiments in Japan (e.g. KamLAND, KamLAND-Zen, ...)

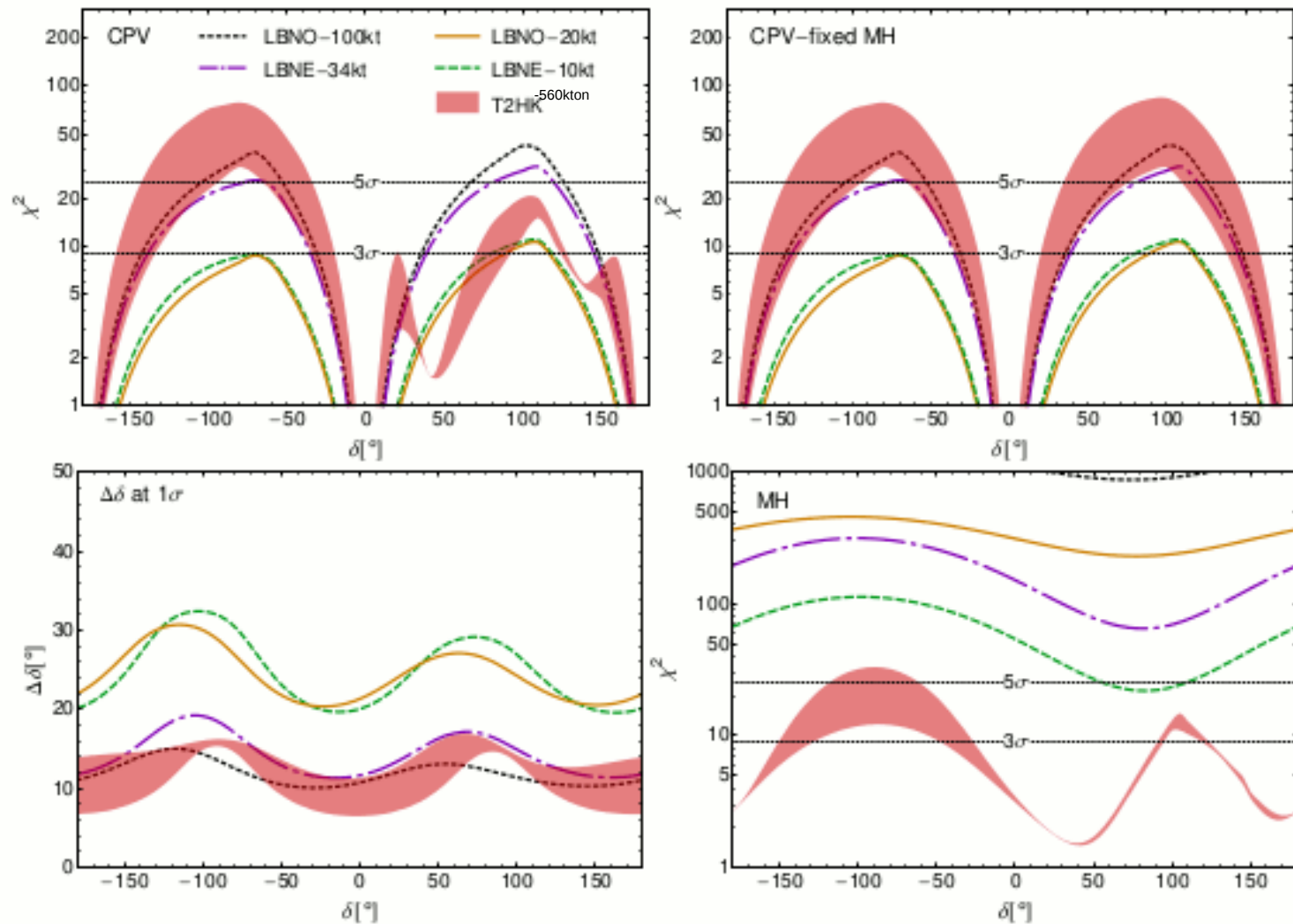
# Overall Cost Estimate

Total	800M USD	
Cavern	300M USD	
Tank & structure	200M USD	
Photo-sensors	200M USD	High QE HPD
Near Detector	30M USD	@Tokai

- Costs estimated based on the current design and including a new near detector.
- Proportional sharing of costs between the interested Countries is expected.

# Summary

- T2K measured non-zero  $\theta_{13}$  with  $7\sigma$  significance by observation of  $\nu_{\mu} \rightarrow \nu_e$ 
  - Major analyses improvements: ND280 selection and SK reconstruction
- Also measurement of  $\nu_{\mu} \rightarrow \nu_{\mu}$  which favors maximal mixing
  - Run4 will be added soon
- Hyper-K covers wide range of physics:
  - Neutrino oscillation with beam- $\nu$  & atmospheric- $\nu$ 
    - ➔ Main goal: CP violation
  - Nucleon decay search and astrophysical neutrinos
- R&D started in all areas and progressing
- Japan HEP community: HK at highest priority. Strongly growing international community.



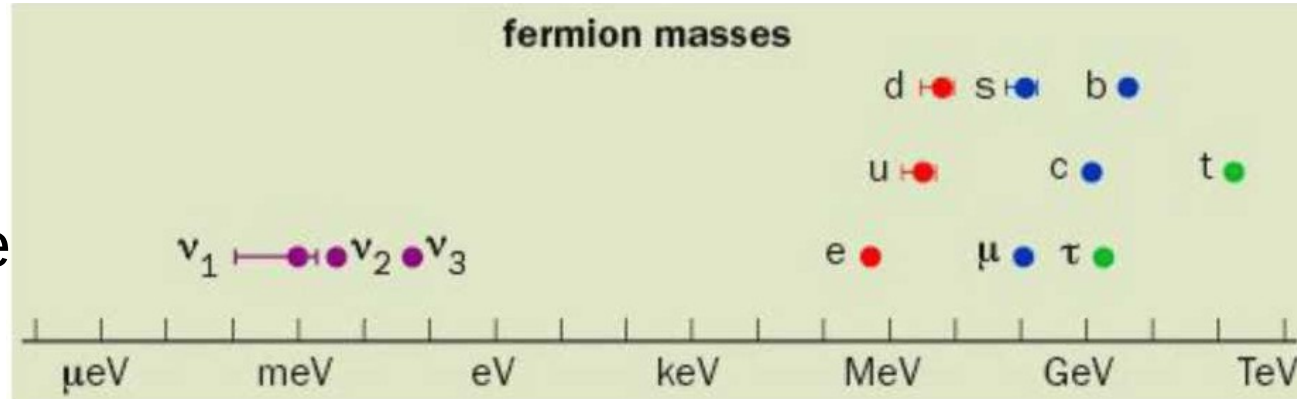
Input to strategy group for the update of the EU strategy

[http://cds.cern.ch/record/1628377/files/Briefing\\_book.pdf](http://cds.cern.ch/record/1628377/files/Briefing_book.pdf)(2013). Beam power: LBNE 700kW 10years, T2HK 1.66 MW 5 years, LBNO 800kW 10 years).

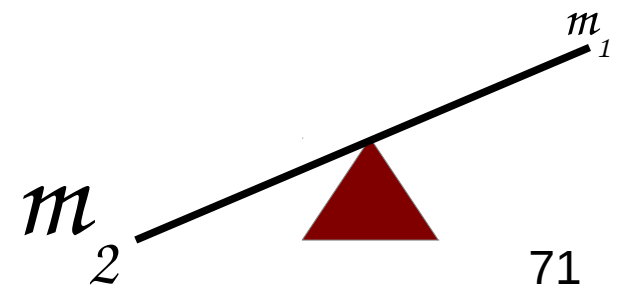
# Backup Slides

# What we Know about Neutrinos

- Three flavours:  $\nu_e, \nu_\mu, \nu_\tau$
- Neutral
- Interact via the weak force
- Abundant
- Massive



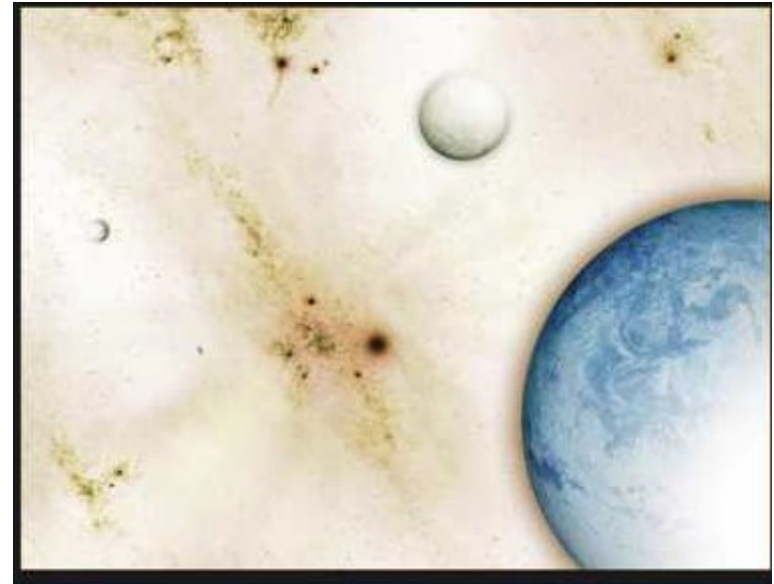
- In the Lagrangian, charged leptons have Dirac mass terms which couple left and right chiral components:  $-m_D (\bar{\nu}_L \nu_R + \bar{\nu}_R \nu_L)$
- Neutrinos are massless in the SM, only interact with the left handed left force
- The see-saw mechanism explains the lightness of the neutrino mass by adding a very heavy right handed neutrino,  $N_R$



# $\nu$ and the Matter-Antimatter Asymmetry



$\neq$

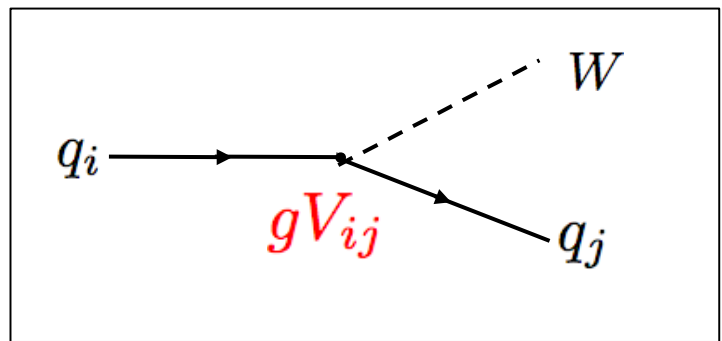


- Leptogenesis: CP violation in the decay of  $N_R$  can create a lepton number asymmetry which can be converted to a baryon number asymmetry (M. Fukugita and Y. Yanagita, PLB 174, 45 (1986)).
- For the matter-antimatter asymmetry to be generated, we require: non-thermal equilibrium, CP Violation, and baryon number violation (A.D. Sakharov, JEPT Lett. 5, 24 (1967)).



# Neutrino Oscillations

- Similar mechanism as in the quark oscillation (CKM matrix).
- Free parameters: 3 angles, 1 phase



CKM matrix

$$\begin{bmatrix} |d'\rangle \\ |s'\rangle \\ |b'\rangle \end{bmatrix} = \begin{bmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{bmatrix} \begin{bmatrix} |d\rangle \\ |s\rangle \\ |b\rangle \end{bmatrix} .$$

- PMNS (Pontecorvo, Maki, Nagakawa, Sakata) matrix for  $\nu$ :

Flavour eigenstates  
(coupling to the W)

Mass eigenstates  
(definite mass)

# Neutrino Oscillation Revisited

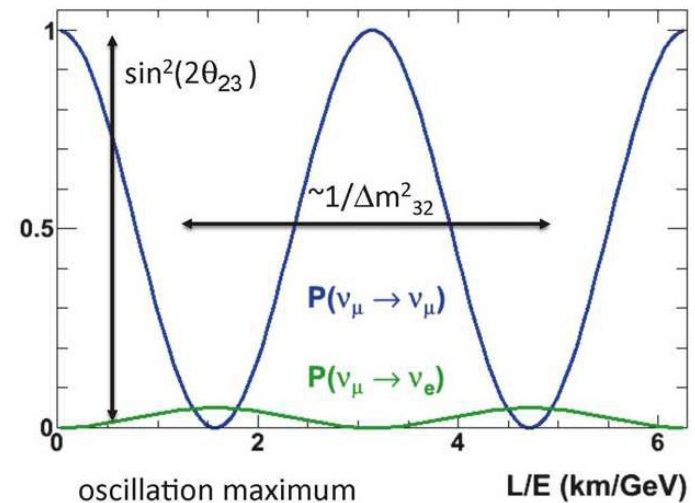
$$P_{\alpha\beta} = \delta_{\alpha\beta} - 4 \sum_{i>j} \text{Re}[U_{\beta i} U_{\alpha i}^* U_{\beta j}^* U_{\alpha j}] \sin^2 \left( \frac{\Delta m_{ij}^2 L}{4E} \right) + 2 \sum_{i>j} \text{Im}[U_{\beta i} U_{\alpha i}^* U_{\beta j}^* U_{\alpha j}] \sin \left( \frac{\Delta m_{ij}^2 L}{2E} \right)$$

- $\Delta m_{32}^2 \gg \Delta m_{21}^2$ , producing high frequency and low frequency oscillation terms in the probability formula.
- If choose L,E such that  $\sin^2(\Delta m_{32}^2 L/E) \sim 1$ , then  $\Delta m_{21}^2$  terms will be small.
- $\nu_\mu$  “disappear” in  $\nu_e, \nu_\tau$ :

$$P(\nu_\mu \rightarrow \nu_\mu) \cong 1 - \sin^2 2\theta_{23} \sin^2 \left( \frac{\Delta m_{32}^2 L}{4E} \right)$$

- A small fraction of  $\nu_e$  will appear:

$$P(\nu_\mu \rightarrow \nu_e) \cong \sin^2 2\theta_{13} \sin^2 \theta_{23} \sin^2 \left( \frac{\Delta m_{31}^2 L}{4E} \right)$$

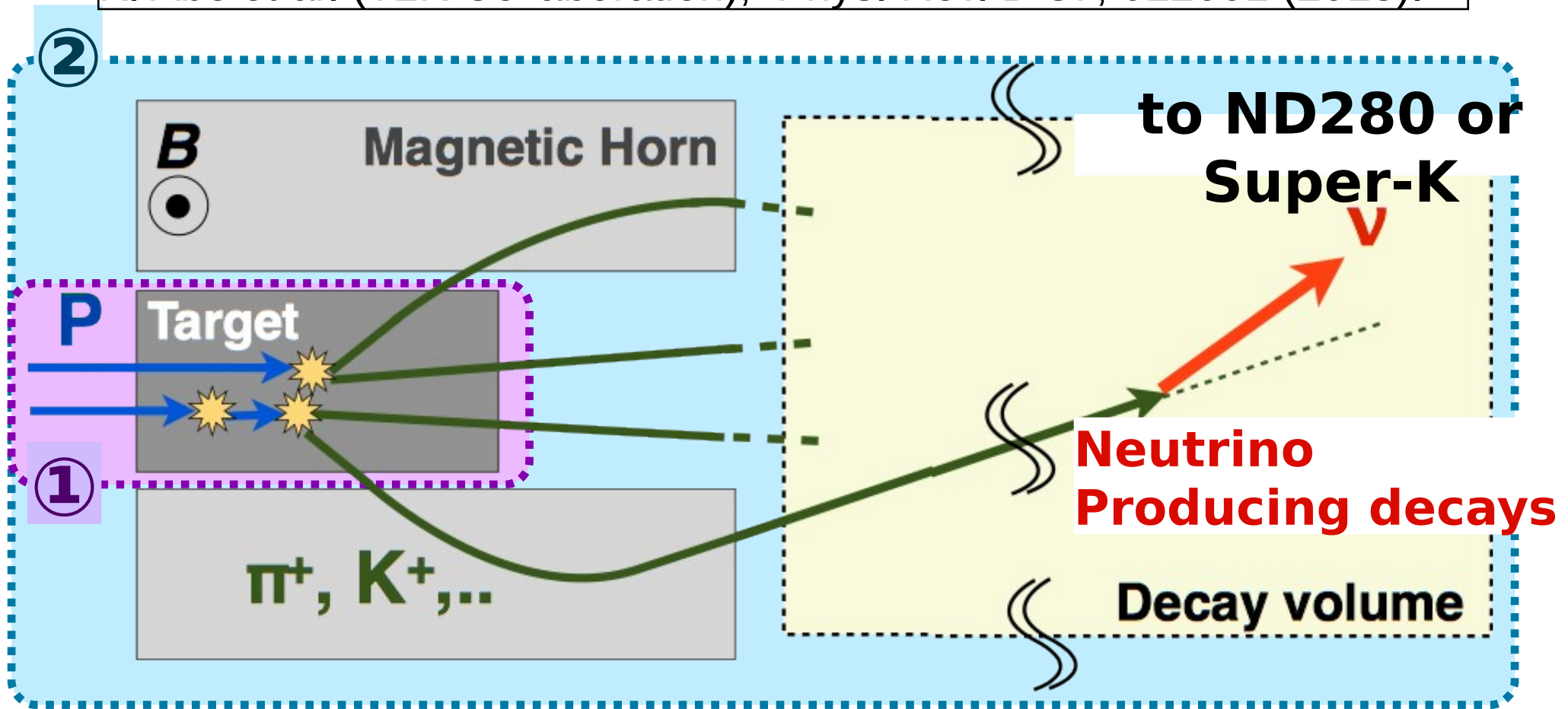


$$\Delta m_{31}^2 \sim \Delta m_{32}^2$$

Only leading order terms shown

# Simulating neutrino flux

K. Abe *et al.* (T2K Collaboration), Phys. Rev. D 87, 012001 (2013).



1. p interaction inside the carbon target with FLUKA2008.3d
  2. Tracking through horn fields and decay volume using GEANT3 with GCALOR
- Calculate neutrino producing decays  
Estimate the flux at the near/far detector

# External Data and Flux

- Hadro-production simulated with FLUKA2008.3d, weighted so that interactions match external data [1]
  - NA61/SHINE (CERN) [2][3], Eichten *et al.* [4], and Allaby *et al.* [5]

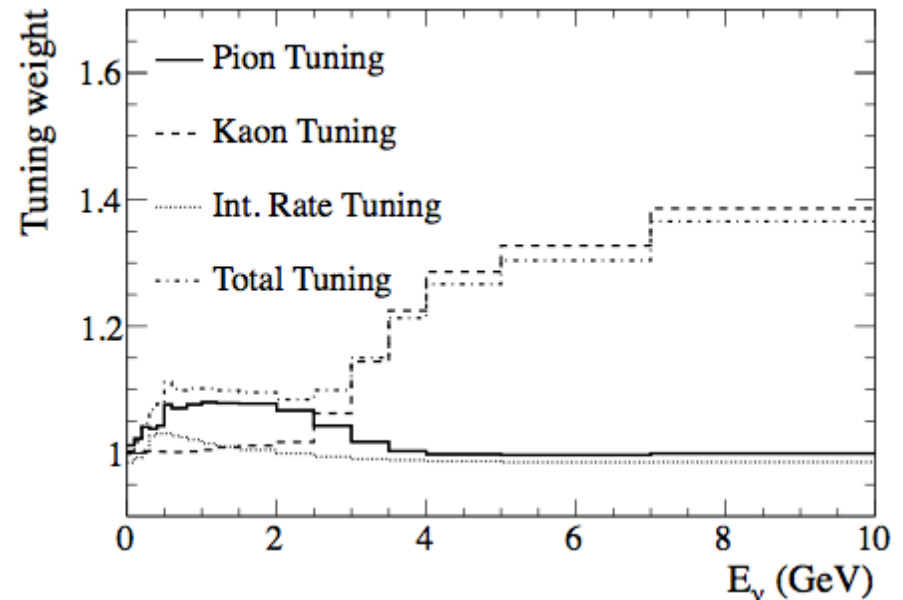
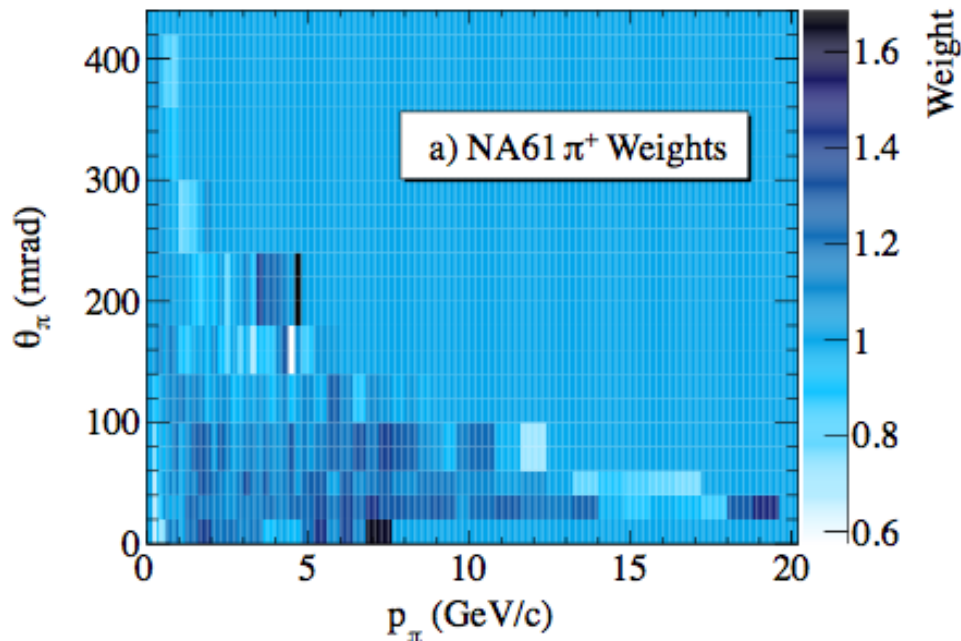
[1] K. Abe *et al.* (T2K Collaboration), Phys. Rev. D 87, 012001 (2013).

[2] N. Abgrall *et al.* (NA61/SHINE Collaboration), Phys. Rev. C 84, 034604 (2011)

[3] N. Abgrall *et al.* (NA61/SHINE Collaboration), Phys. Rev. C 85, 035210 (2012)

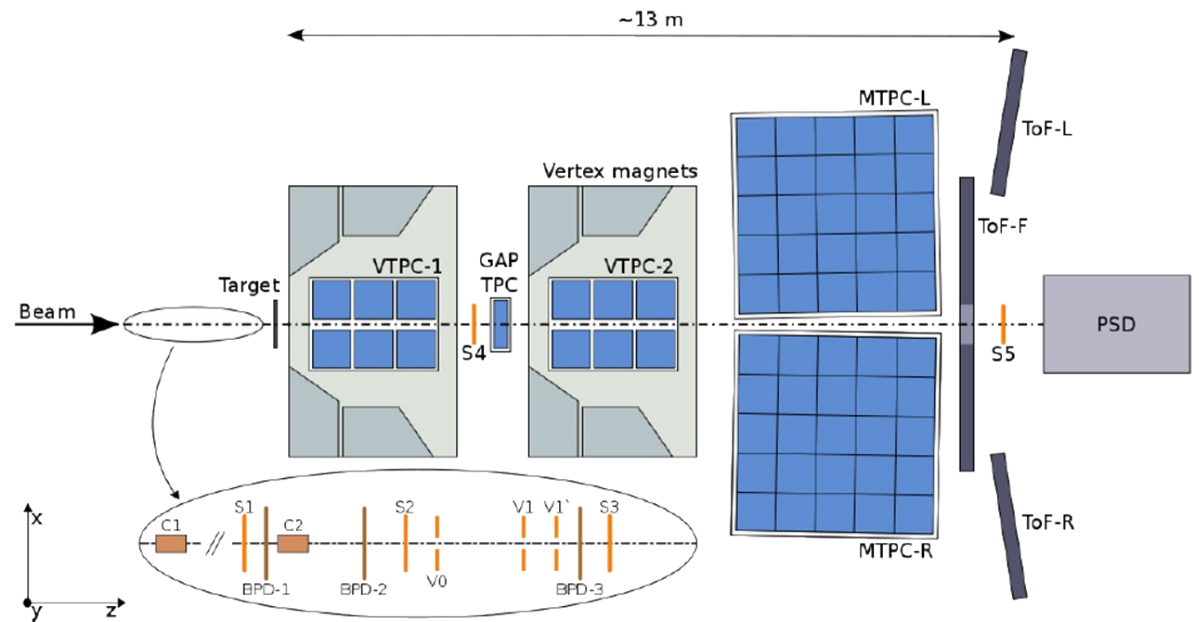
[4] T. Eichten *et al.*, Nucl. Phys. B 44 (1972)

[5] J. V. Allaby *et al.*, Tech. Rep. 70-12 (CERN,1970)



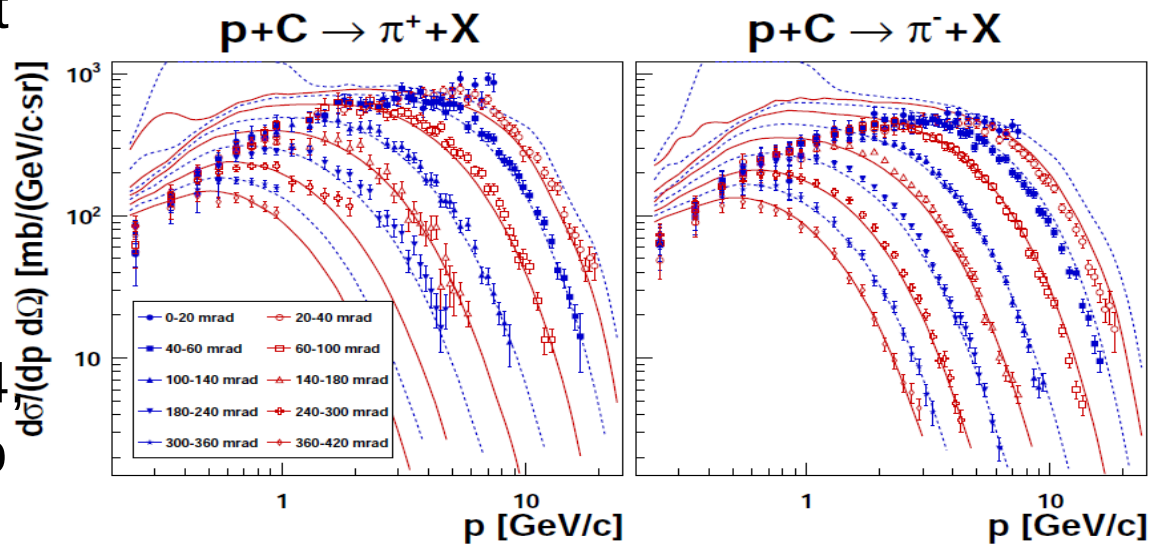
# NA61/SHINE

- Located at the CERN SPS (North Area, H2 beam line)
- Fixed target experiment on primary (ions) and secondary (ions, hadrons) beams



## Inclusive $\pi^+$ spectra in p+C at 31 GeV/c

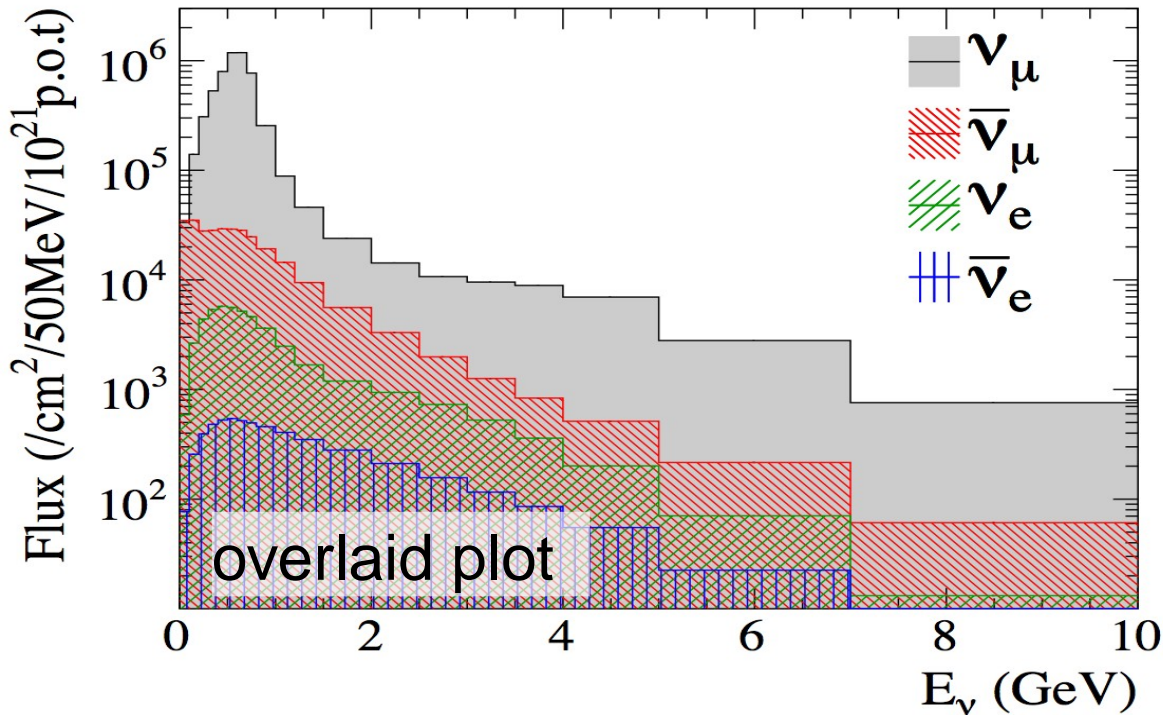
- Pion spectra in p+C interactions at 31 GeV/c are published: Phys. Rev. C84 (2011) 034604
- They are used to improve beam neutrino flux predictions
- Adjust models (UrQMD 1107.0374, Fritiof 1109.6768) used in neutrino and cosmic-ray experiments



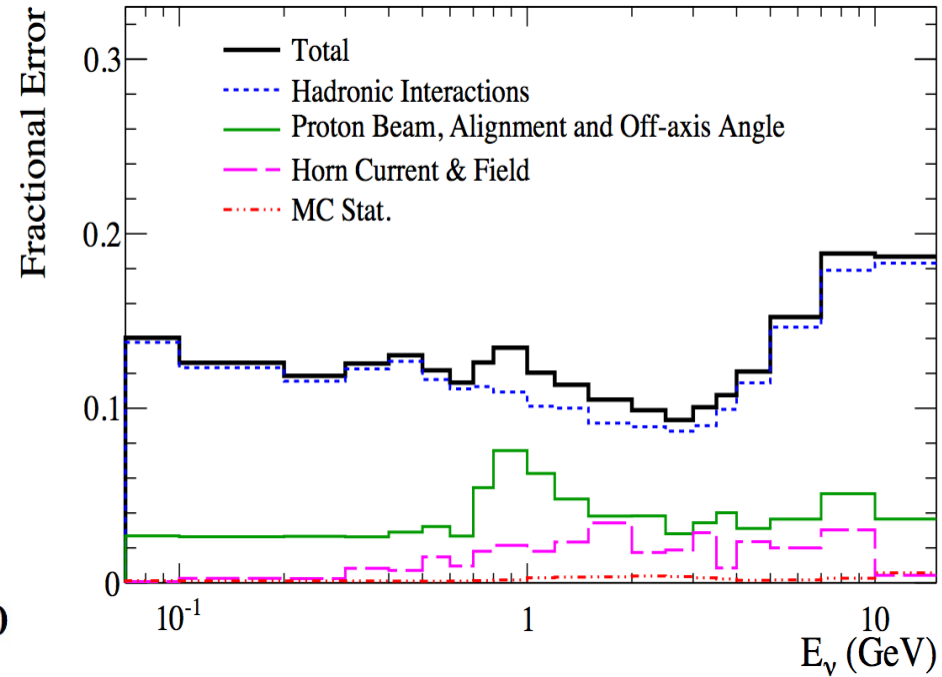
comparison to UrQMD1.3.1

# Flux and Uncertainties

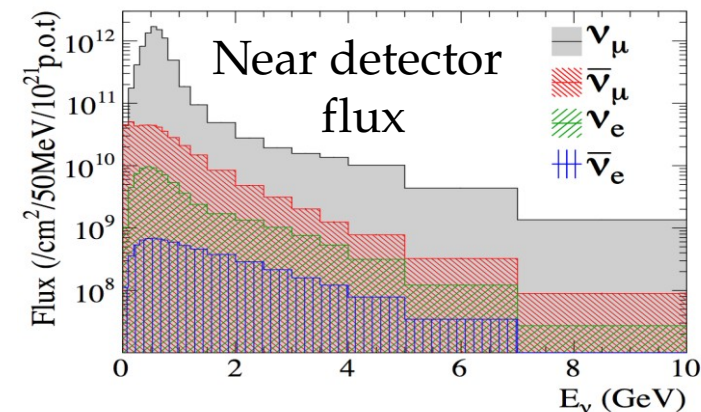
T2K Run1-4 Flux at Super-K



Far detector  $\nu_\mu$  uncertainty



T2K Run1-4 Flux at ND280



- A priori prediction of flux at Super-K has 10-15% uncertainties from 0.1 to 5 GeV
- Off-axis near (ND280) and Far (Super-K) fluxes are not identical, but highly correlated

# fiTQun: Improved SK Reconstruction Algorithm

- Each hit PMT gives charge and time information
- For a given event topology hypothesis, it is possible to produce a charge and time PDF for each PMT
- Based on MiniBooNE likelihood model (NIM A608, 206 (2009))
- Event hypotheses are distinguished by best-fit likelihoods, e.g., electron vs muon or  $\pi^0$

$$\mu^{\text{dir}} = \Phi(p) \int ds g(s, \cos \theta) \Omega(R) T(R) \epsilon(\eta)$$

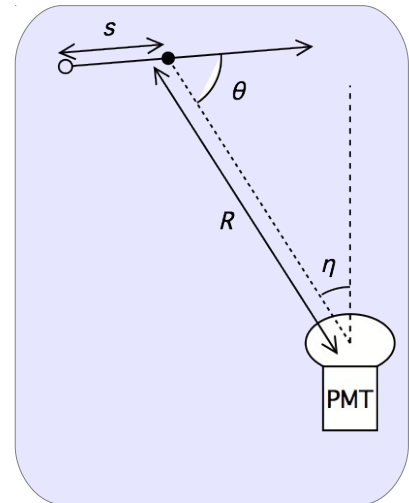
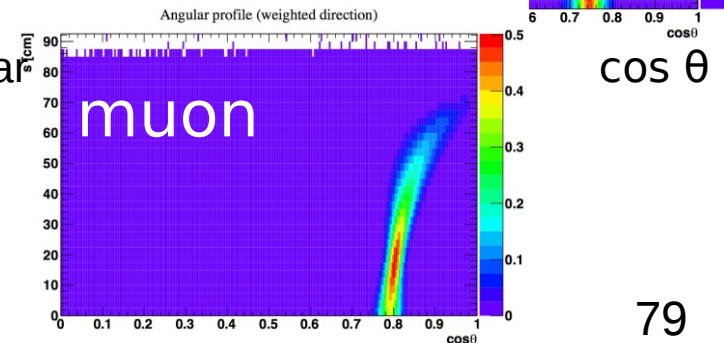
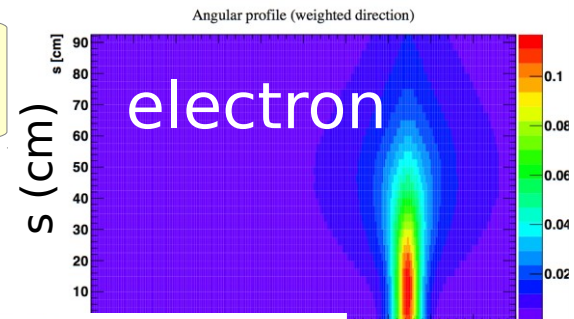
Light Yield

Integral over track length

PMT solid angle

Water attenuation

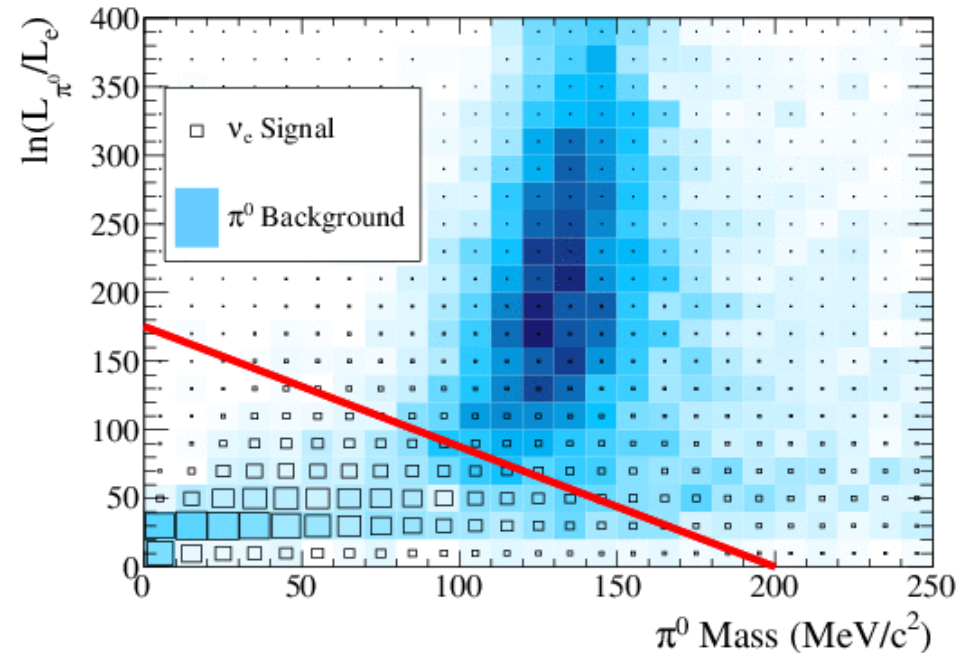
PMT angular response



# Enhanced $\pi^0$ Rejection

- fiTQun can use the mass of the  $\pi^0$  hypothesis and best-fit likelihood ratio of  $e^-$  and  $\pi^0$
- Cut removes 70% more  $\pi^0$  background than previous<sup>§</sup> method for a 2% added loss of signal efficiency.

Likelihood Ratio vs  $\pi^0$  Mass



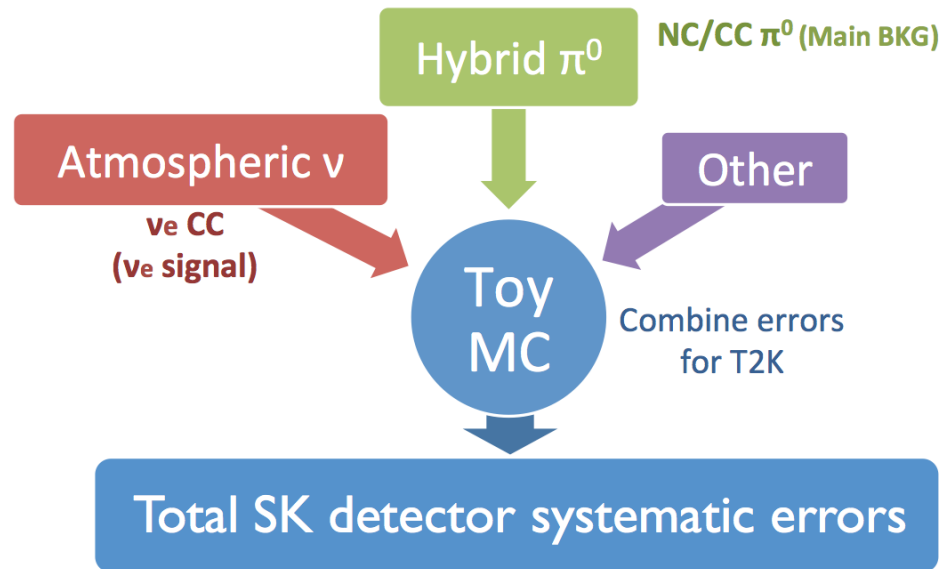
§ Previous approach (POLFit) forced the reconstruction to find two rings and then formed a  $\pi^0$  mass under the two-photon hypothesis



# Selection Results

RUN1-4 6.570x10 <sup>20</sup> POT	MC Expectations w/ $\sin^2 2\theta_{13}=0.1$					Data
	$\nu_{\mu}+\nu_{\mu}$ CC	$\nu_e+\nu_e$ CC	NC	BG total	Signal	
True FV	325.67	15.97	288.11	629.75	27.07	-
FCFV	247.75	15.36	83.02	346.13	26.22	377
One-ring	142.44	9.82	23.46	175.72	22.72	193
e-like	5.63	9.74	16.35	31.72	22.45	60
$E_{\text{vis}} > 100\text{MeV}$	3.66	9.68	13.99	27.32	22.04	57
No decay-e	0.69	7.87	11.84	20.40	19.63	44
$E_{\nu}^{\text{rec}} < 1250\text{MeV}$	0.21	3.73	8.99	12.94	18.82	39
fitQun $\pi^0$	0.07	3.24	0.96	4.27	17.32	28
Efficiency [%]	0.0	20.3	0.3	0.7	64.0	-

# SK Detector Systematic Uncertainties



- Evaluation of Super-K detector systematic uncertainties uses control samples from the data
  - Atmospheric  $\nu_e$
  - Hybrid  $\pi^0$  (electron from  $\nu_e$  CC and MC photon)
  - Cosmic ray muon samples
- Combine errors with Toy MC method

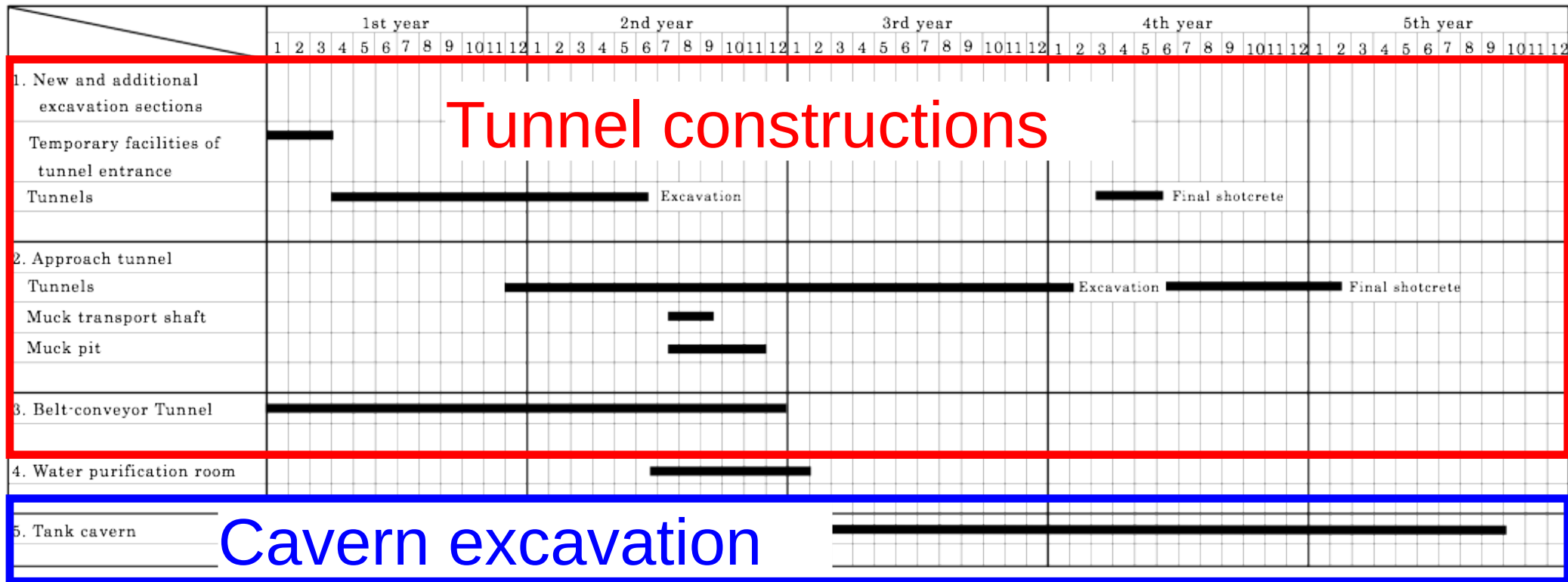
# T2K and J-PARC Run Plans

- T2K's oscillation analyses still statistics limited
  - So far, we have been able to steadily decrease systematics
- T2K will continue to run and benefit from planned J-PARC Main Ring (MR) power improvements:
  - 220 kW operation in CY2013. Integrated  $6.7E20$  POT to date.
  - Linac upgrade to be completed with a year. Expect range of steady MR operation for neutrino between 200-400 kW
  - Planned MR upgrade by 2018 (depends on funding). Up to 750 kW
  - Possible scenario:
    - Double current protons on target by early 2015
    - Next-to-next doubling by early 2017
    - If MR upgrade done in 2018, reach full planned statistics ( $78E20$  POT), roughly 12x the current exposure, roughly end of 2020
- T2K beamline designed to easily switch from neutrino to anti-neutrino beams
  - T2K has made no firm plans for anti-neutrino running

# Summary of J-PARC/Neutrino Facility Upgrade Plan

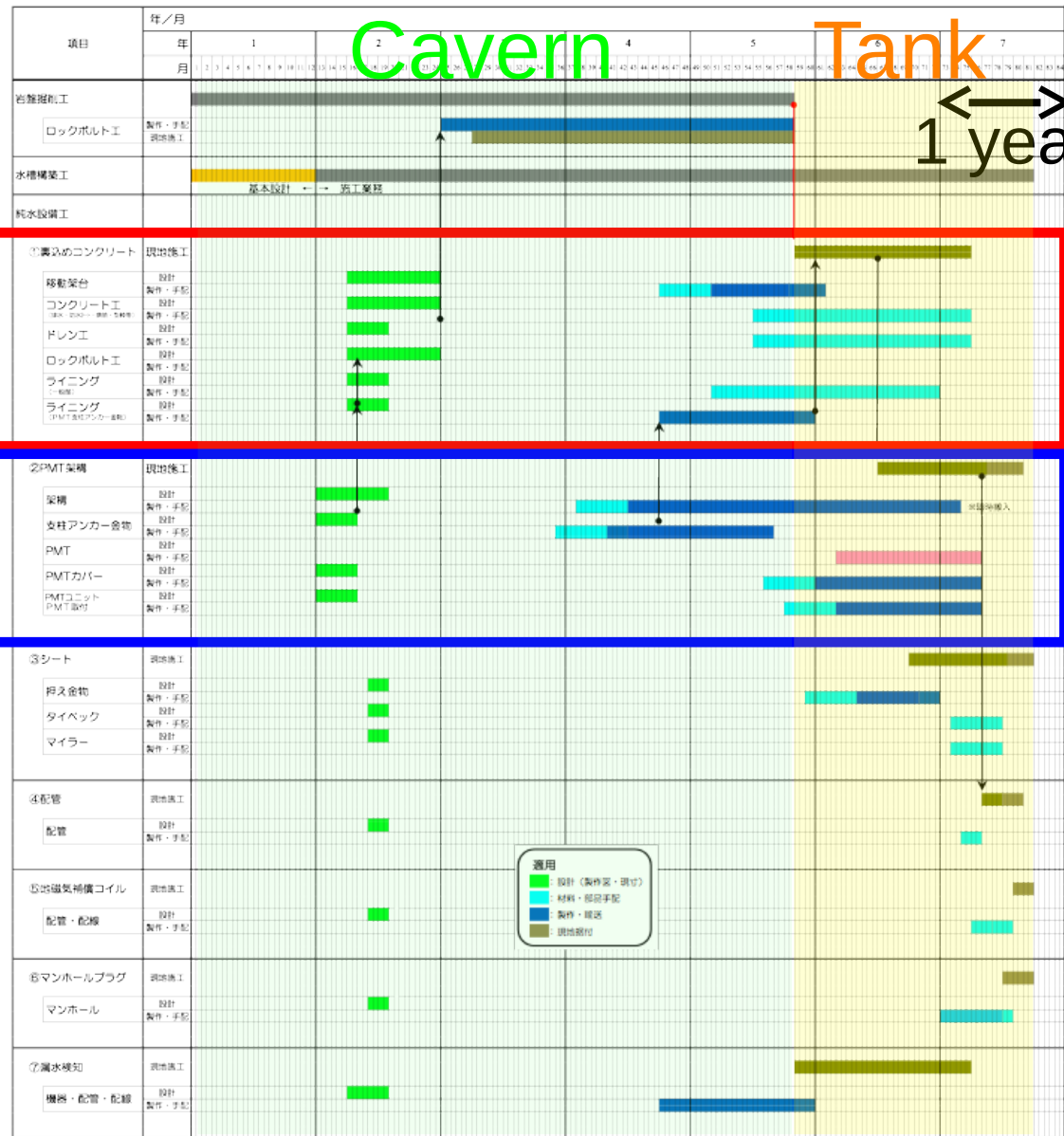
- Achieved beam power for user operation:
  - 350kW max. for T2K
- To increase #p/bunch
  - MR collimator capability 3.5kW
  - LINAC energy upgrade to 400MeV, frontend upgrade in 2014
- MR 750kW operation → to double replication rate operation
  - R&D for MR magnet power supplied well in progress
  - Higher gradient RF core to be ready for installation in 2015
- Neutrino beam-line
  - No problems for critical components so far
  - Works ongoing to replace all 3 horns/target
  - Replacement of Horn-3 completed, radiation well under control.
- Upgrade of neutrino beam-line
  - Double rep.rate: less thermal shock for target/beam window
  - Horn: triple PS operation is necessary for 1Hz \*320kA) operation

# Excavation Schedule



- Cavern construction period: ~5 years
- Transport / approach tunnels: ~3 years
- Excavation of caverns: ~3 years

# Tank construction schedule

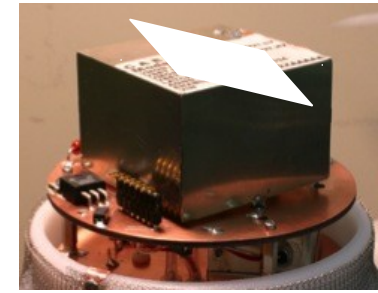


**Lining**

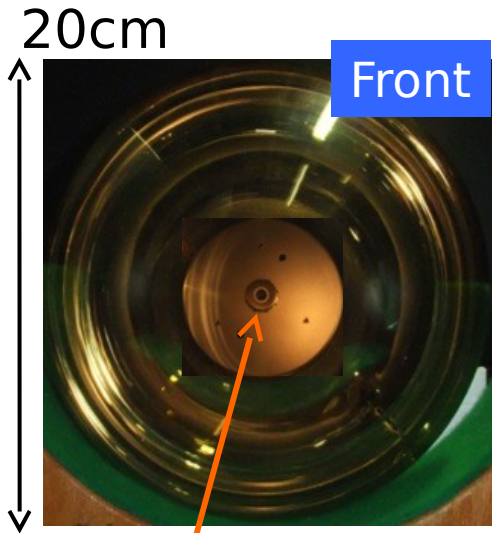
**PMT & support**

- Tank construction: ~2 years
- Lining: 1+ years, PMT installation: ~1 year

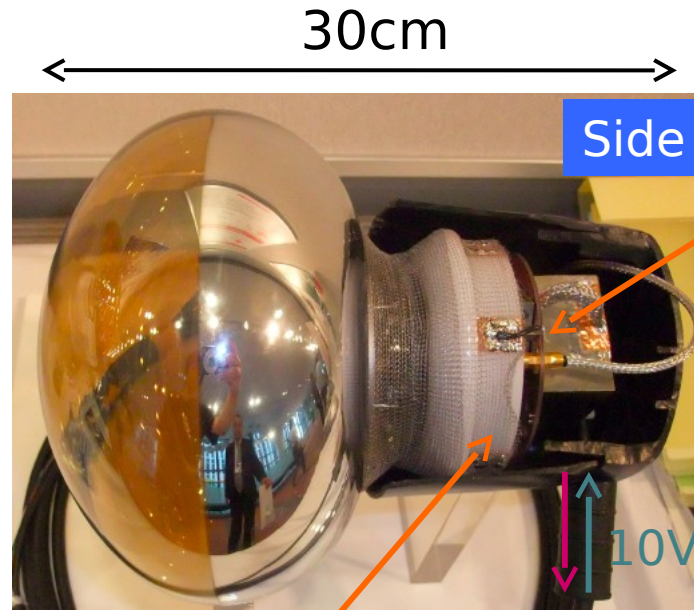
# 8" HPD Prototype



High voltage module  
(2ch 10kV/500V Max.)



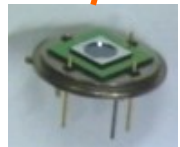
Front



Side

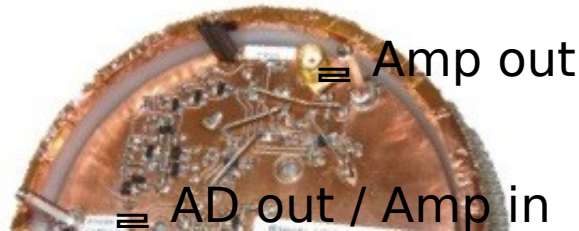
HV module and preamplifier are packed and waterproofed

→ No HV line in water



5mm $\phi$  AD

Preamp board



≡ Amp out

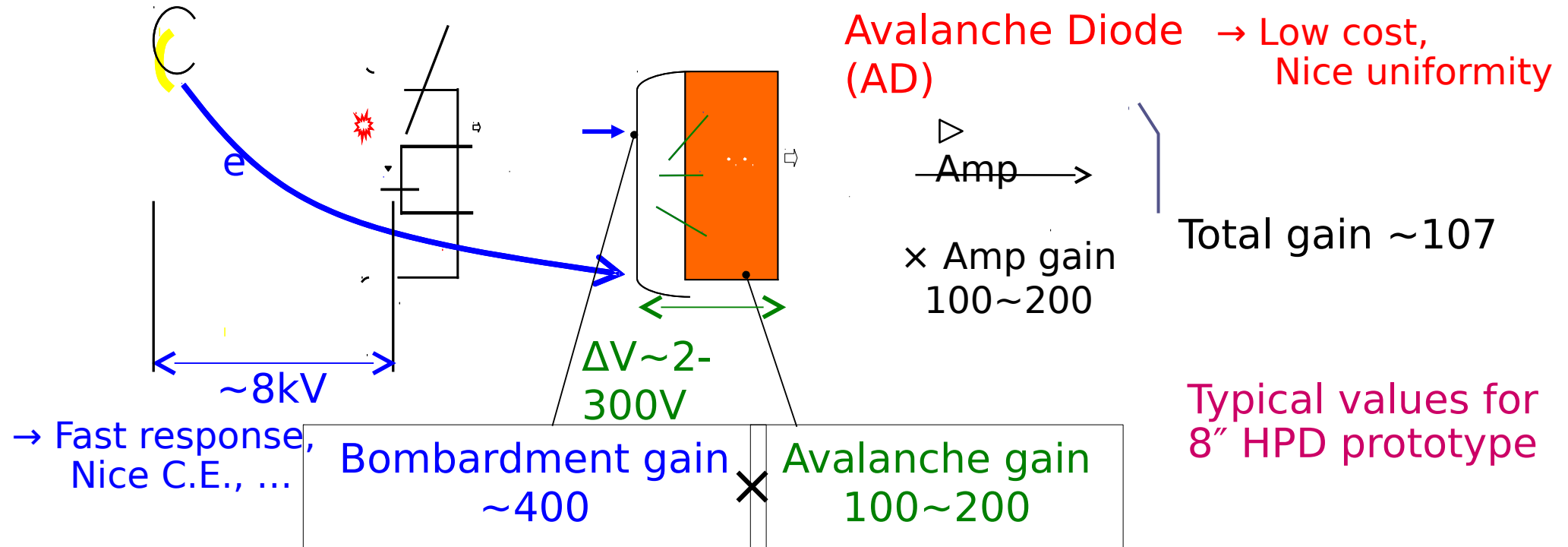
≡ AD out / Amp in

Signal

Spectral response	300 - 650 (420 max.) nm
Photocathode	Bialkali
Window material	Borosilicate glass
Gain	$4 - 9 \times 10^4$
Time Rise	1.7 ns
Fall	2.7 ns
T.T.S.	0.62 ns ( $\sigma$ )
Dynamic range	100 pC ( $1.5 \times 10^4$ p.e.)

Ten 8" HPDs were made for long-term testing

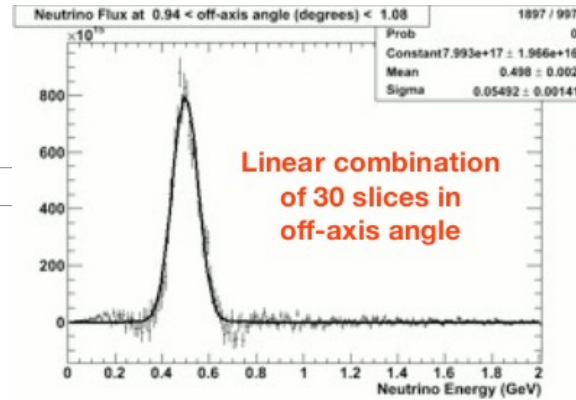
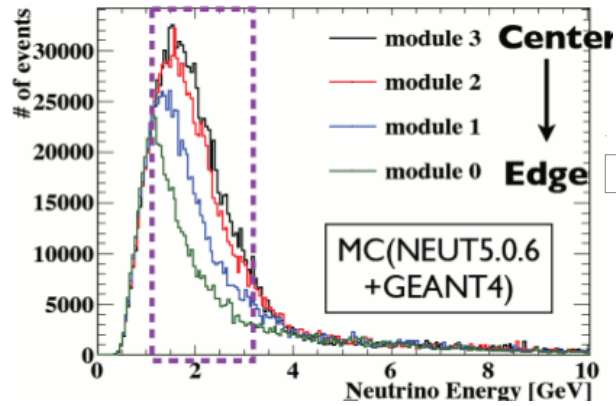
# Hybrid Photodetector (HPD)



- ~8kV supply voltage + Avalanche diode multiplication
  - High voltage to focus photoelectrons into the small AD (5-20mm $\phi$ )
  - Better 1p.e. measurement capability
  - Better timing resolution, faster response
- Simple mechanical structure
  - Axial symmetric response
  - Lower production cost, shorter production period



# Water Column

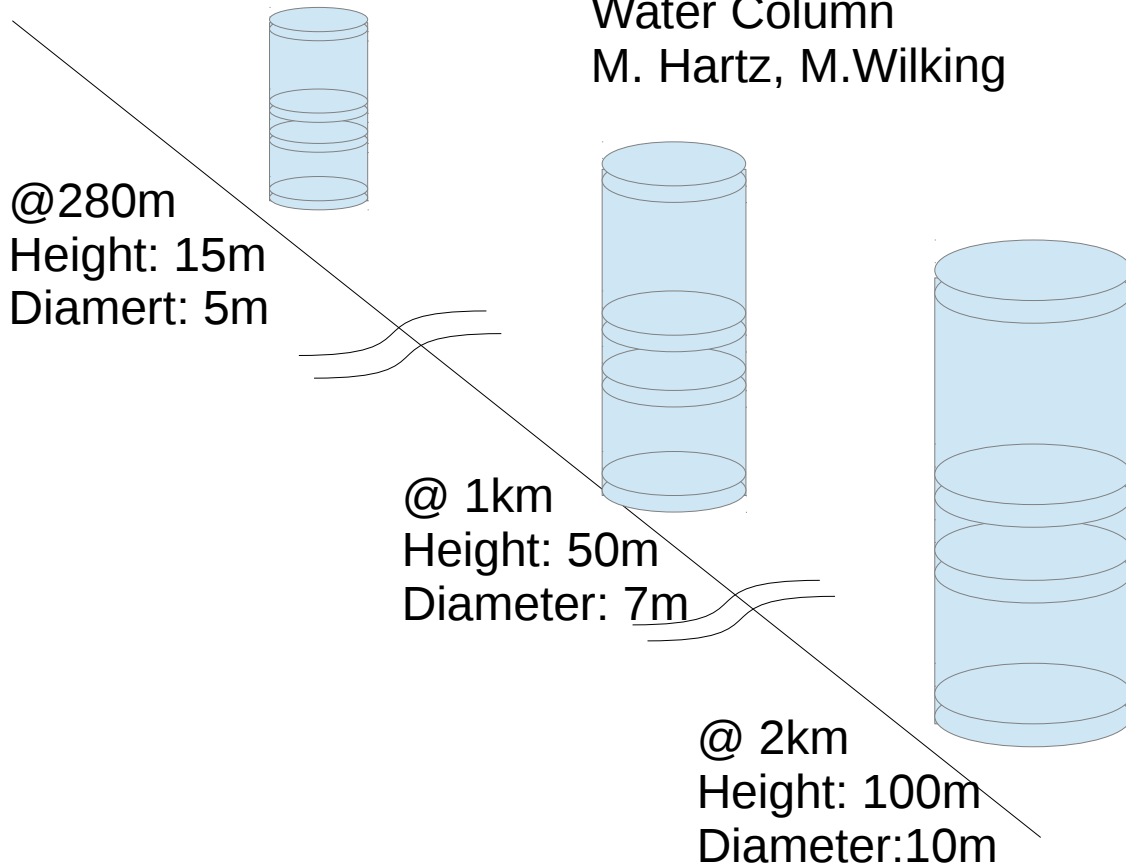


- Minimize dependence on neutrino interaction sampling the beam at several off-axis angles.

- Favoured <1km baseline from engineering point of view.

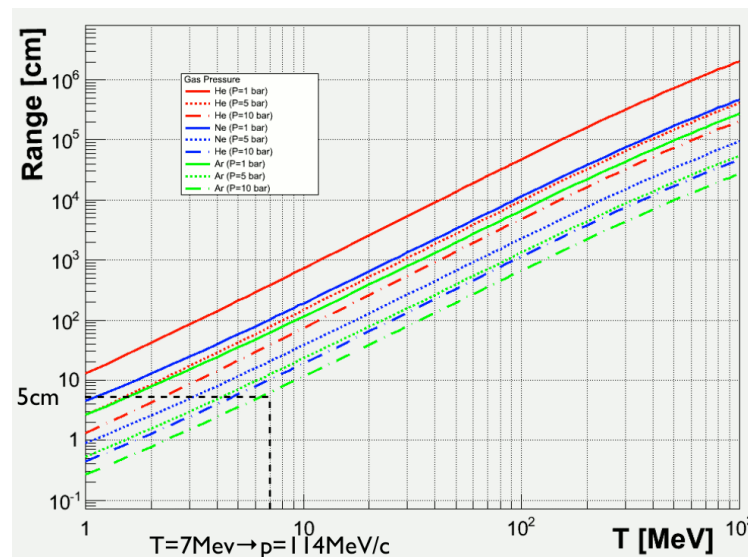
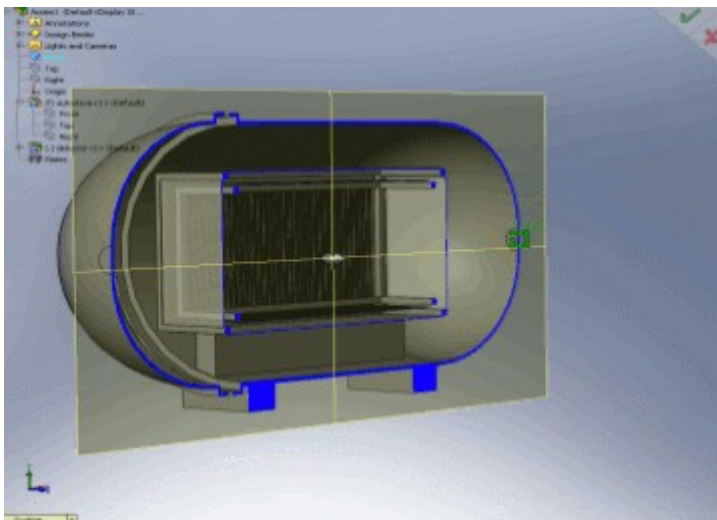
- Possibly add brine (Konaka) around detector to stop muons.

Water Column  
 M. Hartz, M. Wilking



# ND280 Upgrade

- Several studies being performed for a possible upgrade → beneficial for Hyper-K as well. Undergoing study.
- Improve ND280 to optimize cross-section measurements.
- Proposed high pressure TPC to access the low energy nuclear debris and help in the study for neutrino-nucleus interactions. Investigated 3 basic gases (He, Ne, Ar and CF<sub>4</sub>) and 2 pressures.



High Pressure TPC  
(Barcelona)