

Particle Identification, Trigger & Modern Detectors

- Particle Identification
 - Neutral particles
 - Methods for charged particles
- Trigger
 - Basics
 - Modern Example
- not presented: Modern Detectors
 - “slide show”

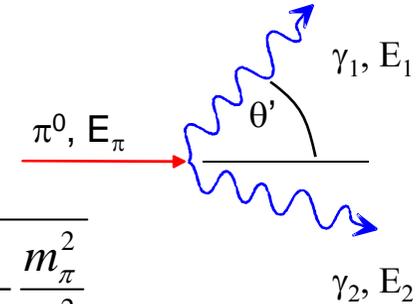


π⁰

□ lightest neutral hadron:

- life time: τ=0.084fs (π[±]: τ=26ns)
- decay: π⁰→γγ; BR = 98.798%
- γ energy and angle:

$$E_\gamma = \frac{1}{2} E_\pi (1 + \beta \cos \theta_{CMS}) \quad \beta = \sqrt{1 - \frac{m_\pi^2}{E_\pi^2}}$$



- J_π = 0 → isotropic θ distribution in CMS
- deterministic kinematics
- reconstruction from el.mag. showers induced by the 2γ:
 - needs el.mag. calorimeter: E_{γ1}, E_{γ2}
 - disparity: D > 3 for 50%, D > 7 for 25%
 - good angular resolution: distance d → angle θ
 - for π⁰ hypothesis: use mass restriction m_{π0}=135MeV

$$D = \frac{E_1}{E_2} \approx \frac{1 + \cos \theta}{1 - \cos \theta} \quad \text{for } \beta \rightarrow 1$$

□ heavier neutral hadrons:

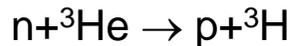
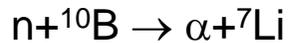
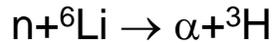
- long living: neutron
- short living: η, ρ⁰, ω, φ, K_S⁰(90fs), [K_L⁰(51ns)], D⁰(0.41ps), B⁰(1.5ps), Δ, Λ, Σ, Ξ, ...
 - reconstructed kinematically from secondary particles : π⁰, γ & charged particles

Neutron Counters

- no direct detection possible → 4 main methods:
 - thermal n: neutron activation reactions, e.g. $^{63}\text{Cu}(n,\gamma)^{64}\text{Cu}$ → measure delayed γ
 - $E_n < 20\text{MeV}$: prompt nuclear reactions with charged secondaries
 - $E_n < 1\text{GeV}$: elastic scattering of n or p → measure recoil partner
 - $E_n > 5\text{GeV}$: cascade of inelastic scatterings → calorimeter

□ neutron monitor:

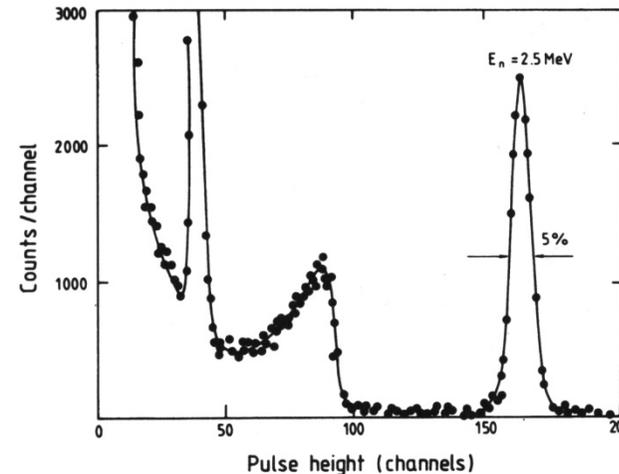
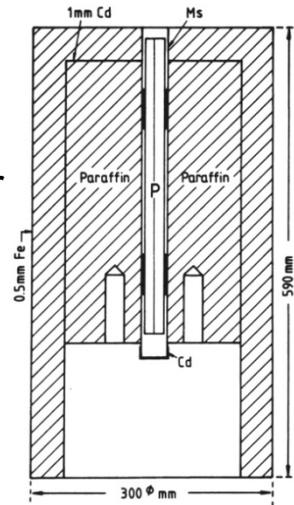
- up to 20MeV
- uses paraffin (H2 rich) as moderator to thermalise neutrons
→ large cross section for prompt nuclear reactions (1-1000 barn)



- use e.g. BF_3 gas counter



BF_3
10keV-10MeV



${}^3\text{He}$ + Krypton
proportional counter:
neutrons (2.5MeV)

Time of Flight Measurement

- particle identification through flight time:

$$\Delta t = \frac{L}{\beta_1 c} - \frac{L}{\beta_2 c} = \frac{L}{c} \left[\sqrt{1 + \frac{m_1^2 c^2}{P^2}} - \sqrt{1 + \frac{m_2^2 c^2}{P^2}} \right] \approx \frac{Lc(m_1^2 - m_2^2)}{2P^2} \text{ for } P^2 \gg m^2 c^2$$

- time resolution:

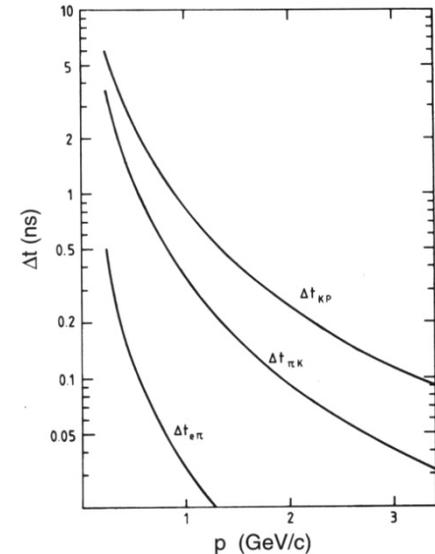
- $\sigma_t = 300\text{ps}$ for organic scintillation counter
- $\sigma_t = 50\text{ps}$ for parallel-plate counter

- $4\sigma_t$ separation : $\sigma_t = 300\text{ps}$ $\sigma_t = 50\text{ps}$

π -K @ 1GeV	3.4m	0.6m
π -K @ 2GeV	13m	2.2m
e- π @ 200MeV	1.0m	0.16m
e- π @ 400MeV	6.0m	1.0m

- method limited to low momenta (<2GeV/c)

time-of-flight
differences
for 1m



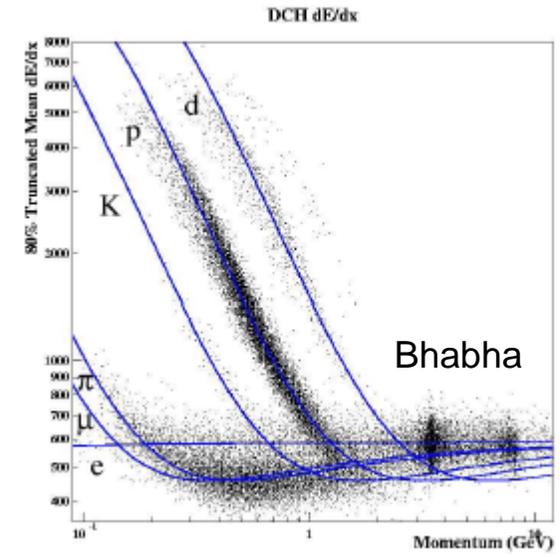
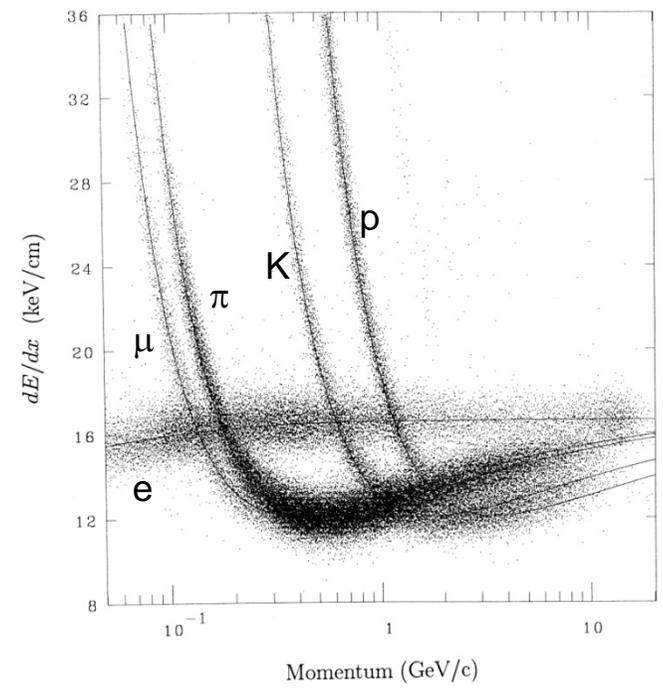
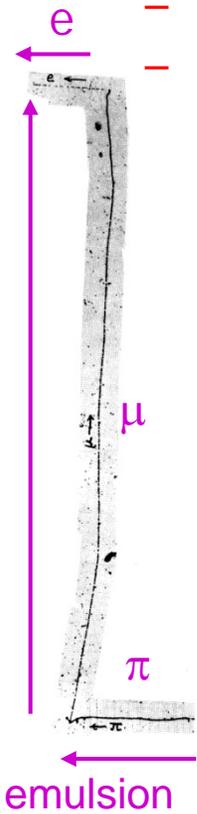
COSY TOF
barrel & end cap



dE/dx

Low energy range:

- heavier particles polarise medium stronger → larger energy loss via ionisation
- Bethe-Bloch: dE/dx rises strongly for $p < m$ [GeV]
- used in emulsions, cloud & bubble chambers and tracking chambers
- several samples taken per track: increased efficiency
- “multiple ionisation measurement” in region of relativistic rise



BaBar drift chamber

Two Gamma TPC

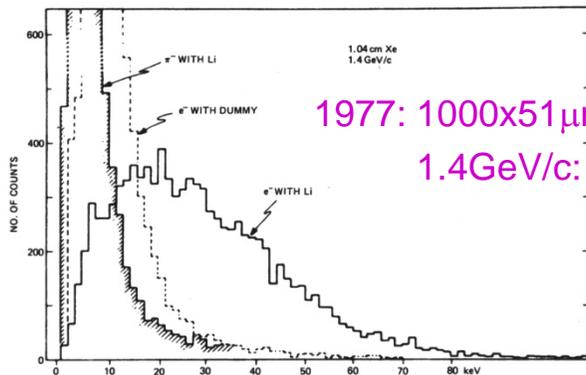
Transition Radiation

□ X-ray regime:

- $\varepsilon=n^2$; $\varepsilon=\varepsilon_1+i\varepsilon_2$; $\varepsilon_1<1$; $\varepsilon_2\ll 1$: $v>v_C$, i.e. below Cherenkov radiation threshold
- but still photons emitted at n-boundaries, i.e. change in dielectric constant
- charge in vacuum and mirror charge in medium form moving dipole \rightarrow el.mag. "TR"

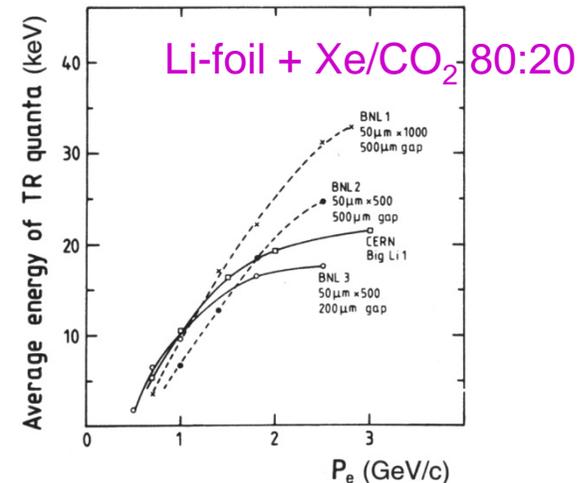
□ periodic radiator + detector, e.g. proportional chamber:

- intensity $\sim \gamma$, concentrated in half opening angle $\phi \sim 1/\gamma$ $\gamma=E/mc^2$
- periodic arrangement: foils or air gaps \rightarrow interference \rightarrow threshold in γ
- X-ray absorption $\sim Z^{3.5} \rightarrow$ low Z radiator needed \rightarrow best candidate: Li



1977: 1000x51 μ m Li-foils, Xe detector
1.4GeV/c: π : $\gamma=10$, e^- : $\gamma=2740$

- usable range: $\gamma>1000$, i.e. $E_e>0.5\text{GeV}/c$, $E_\pi>140\text{GeV}/c$
- for $\gamma<1000$: 1-5keV X-ray radiation detection is needed
- alternative: 12x(n 35 μ m Li-foils + Xe chamber): $\varepsilon_{e^-}=90\%$, π -rejection: 8×10^{-4}
- $\varepsilon_\pi=90\%$, K-rejection: 10^{-2}



Li-foil + Xe/CO₂ 80:20

Recap: Threshold Cherenkov Counter

- particle identification through velocity measurement:

- for charged particles: $\beta c = v > \frac{c}{n}$

- 'shock-wave': coherent wave front with angle

$$\cos \theta_c = \frac{1}{\beta n(\omega)}$$

- + finite radiator effects

- directed photon signal

- use: $p=mv$

- concepts:

- threshold counter
 - differential counter
 - ring imaging

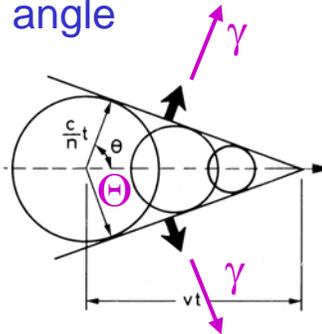


Fig. 5.11. Velocity distribution of charged hyperons with 15 GeV/c momentum in a short secondary beam [LI 73].

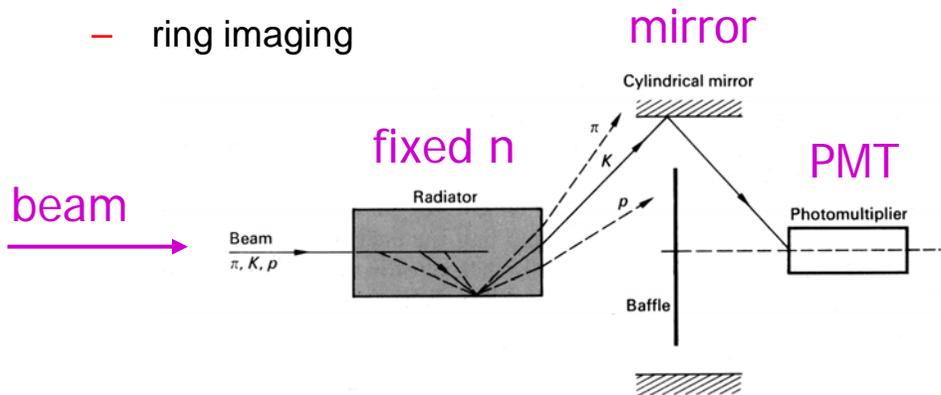
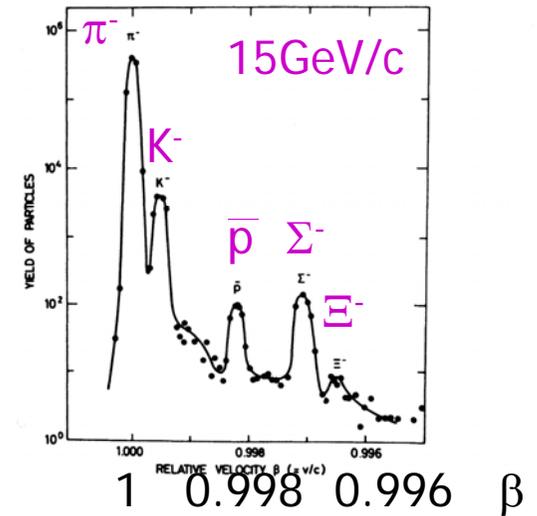
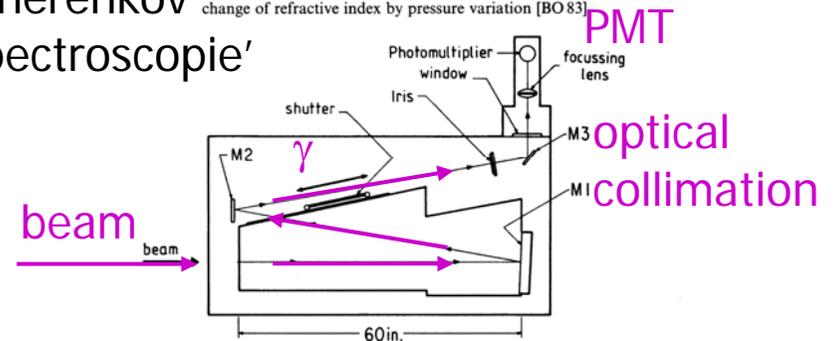


Figure 2.18 Early design of differential Čerenkov counter. The arrangement is intended to select light from one of three components of the beam (K-mesons, in the case shown).

'Cherenkov spectroscopie'

Fig. 5.12. Differential Cherenkov counter with fixed diaphragm and change of refractive index by pressure variation [BO 83]



variable pressure: $n(P)$

Cherenkov Counter - Details

angular distribution of radiation:

- produced photons: N
- maxima due to diffraction
- for a long radiator: $L \gg \lambda$

$$\frac{d^2 N}{d\lambda d \cos \theta} = \frac{2\pi\alpha}{\lambda} \left(\frac{L}{\lambda}\right)^2 \left(\frac{\sin x}{x}\right)^2 \sin^2 \theta \quad x(\theta) = \frac{\pi\lambda}{L} \left[\frac{1}{n\beta} - \cos \theta \right]$$

$$\frac{dN}{d\lambda} \approx \frac{2\pi\alpha}{\lambda^2} L \sin^2 \theta_C$$

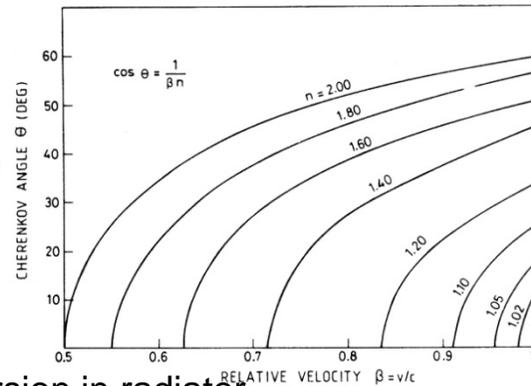
- number of produced photons: flat in energy!!

$$N = 2\pi\alpha L \int_{\lambda_2}^{\lambda_1} \frac{\sin^2 \theta_C}{\lambda^2} d\lambda = \frac{2\pi\alpha L}{\hbar c} \int_{E_1}^{E_2} \sin^2 \theta_C dE \quad ; \quad c = \lambda \nu \quad ; \quad E = \hbar \nu$$

- figure of merit : $N_{\text{eff}} = \varepsilon N$, with detector efficiency ε

Counter types:

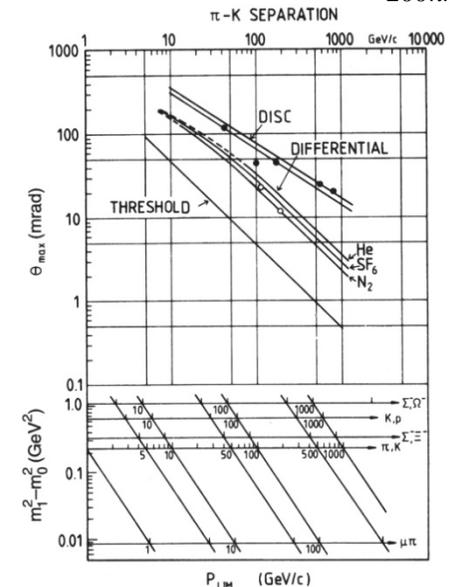
- threshold:
 - limited by choice of material & momentum
- differential:
 - better selection \rightarrow better separation
- DISC:
 - differential, correcting for chromatic dispersion in radiator
 - achieved: $\Delta\beta/\beta \sim 10^{-7} \rightarrow \pi$ -K separation up to 500 GeV/c



$$\begin{aligned} \frac{N}{1 \text{ cm}} &= 490 \sin^2 \theta_C \left|_{400 \text{ nm}}^{700 \text{ nm}} \right. \\ &= 870 \sin^2 \theta_C \left|_{300 \text{ nm}}^{700 \text{ nm}} \right. \\ &= 1630 \sin^2 \theta_C \left|_{200 \text{ nm}}^{700 \text{ nm}} \right. \end{aligned}$$

Problem: only for particles parallel to optical axis of detector, i.e.

Particle Physics Detectors, 2010 along beam line Stephan Eisenhardt



Cherenkov Media

- chose material n to match separation of $m_1 < m_2$:

- heavier particle m_2 does not yet radiate
- or is just below threshold: $\beta_2 < 1/n$

- selection of radiator materials:

- gases at normal conditions
- Aerogel fills gap between gases and solids/fluids

- example light yield:

- e.g. π -K separation
- radiator: $L=1\text{m}$ of C_4F_{10}
- thresholds in C_4F_{10} : $E_\pi = 2.6\text{GeV}$; $E_K = 9.3\text{GeV}$
- $\langle \text{QE} \rangle$ of photodetector: $\varepsilon_{\text{QE}} = 0.2$
- detector efficiency: ε_D
- light yield:

- π @ 9GeV : $\rightarrow \beta_\pi = 0.999879 \rightarrow \theta_\pi = 50\text{mrad}$

$$N = \varepsilon_{\text{QE}} \varepsilon_D L \cdot 870 \sin^2 \theta_C \Big|_{300\text{nm}}^{700\text{nm}} = \varepsilon_{\text{QE}} \varepsilon_D \cdot 220 = \varepsilon_D \cdot 44$$

- π @ 10GeV : $\rightarrow \beta_\pi = 0.999902 \rightarrow \theta_\pi = 51\text{mrad}$

- K @ 10GeV : $\rightarrow \beta_K = 0.998780 \rightarrow \theta_K = 19\text{mrad}$

medium	n	γ_s (threshold)	β_s (threshold)
Diamond	2.42	1.10	0.41
ZnS(Ag)	2.37	1.10	0.42
lead fluoride	1.80	1.20	0.56
Glass	1.46-1.75	1.22-1.37	0.57-0.68
Scintillator (toluene)	1.58	1.29	0.63
Plexiglas (acrylic)	1.48	1.36	0.68
Water	1.33	1.52	0.75
Aerogel	1.025-1.075	2.7-4.5	0.93-0.976
Pentane	1.0017	17.2	0.9983
C_4F_{10}	1.0014	18.9	0.9986
CF_4	1.00050	31.6	0.9995
CO_2	1.00043	34.1	0.9996
He	1.000033	123	0.99997

$$\beta^2 = 1 - \frac{m^2 c^4}{E^2} \quad \gamma^2 = \frac{1}{1 - \beta^2}$$

$$N = \varepsilon_{\text{QE}} \varepsilon_D \cdot 226 = \varepsilon_D \cdot 45.2$$

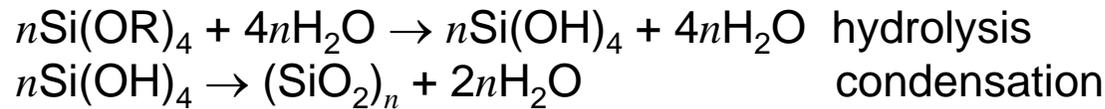
$$N = \varepsilon_{\text{QE}} \varepsilon_D \cdot 31 = \varepsilon_D \cdot 6.2$$

Aerogel I

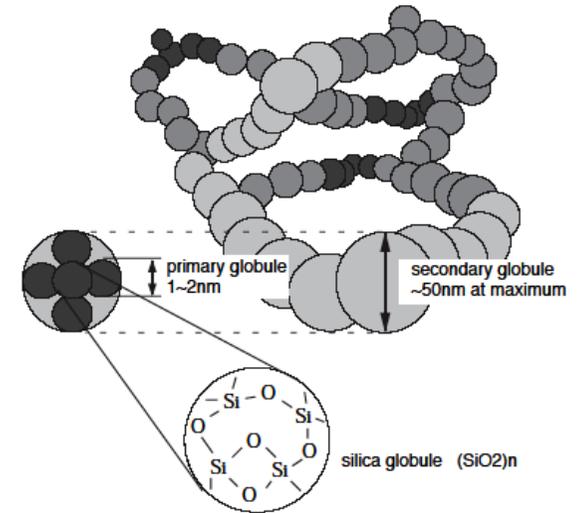
□ structure: $n(\text{SiO}_2)+2n(\text{H}_2\text{O})$

- “foamed silicon” → light: 22 litres = 3 kg
- baked out in tiles of up to $\sim 15 \times 15 \times 6 \text{cm}^3$

□ production: sol-gel process



- chemical treatment to make hydrophobic
- supercritical drying: CO_2 extraction method (31°C , 7.5 MPa)

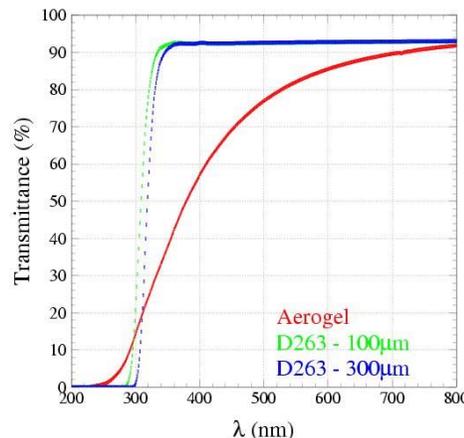


□ transmission T:

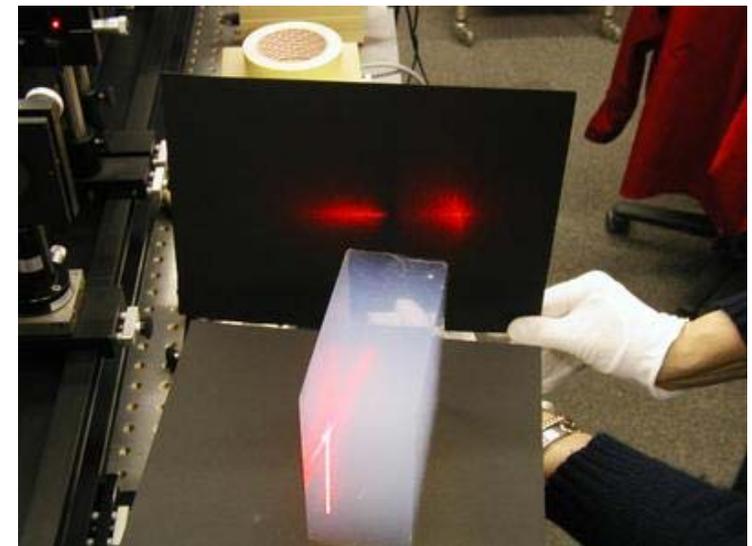
- exponential λ^4 dependence:

$$T = Ae^{\left(-\frac{Cx}{\lambda^4}\right)}$$

- limited by Rayleigh scattering

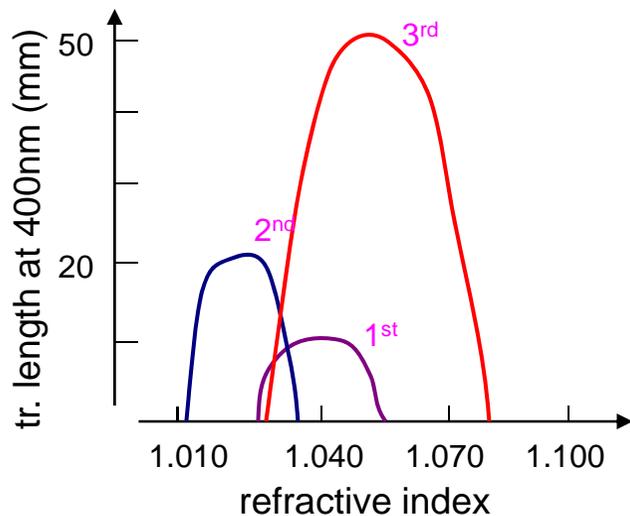
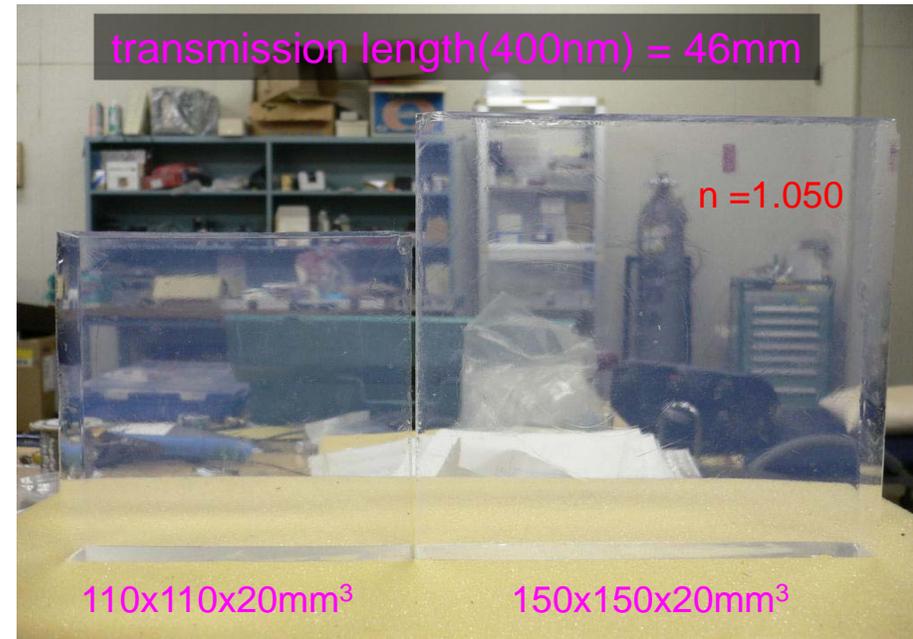


- red emission dominant

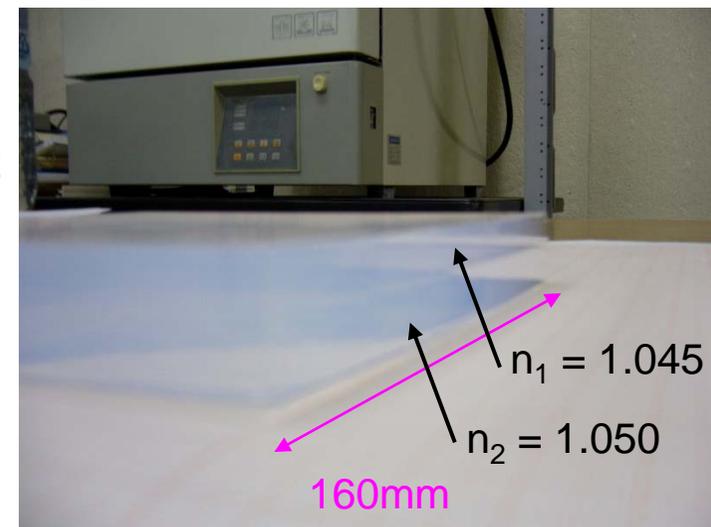


Aerogel II

- clarity C: (with $n=1.030\pm 0.001$ @ 400nm)
 - Matsushita (hydrophobic): $C \sim 0.009 \mu\text{m}^4 \text{cm}^{-1}$
 - Novosibirsk (hydroscopic): $C \sim 0.005 \mu\text{m}^4 \text{cm}^{-1}$
 - larger tiles
 - higher yield of unscattered photons
 - but more difficult to handle
- new developments:
 - higher index aerogels
 - stacking of 2...3 different indices for better proximity focussing

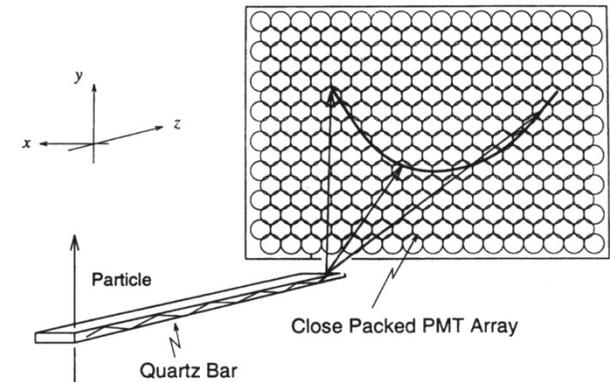
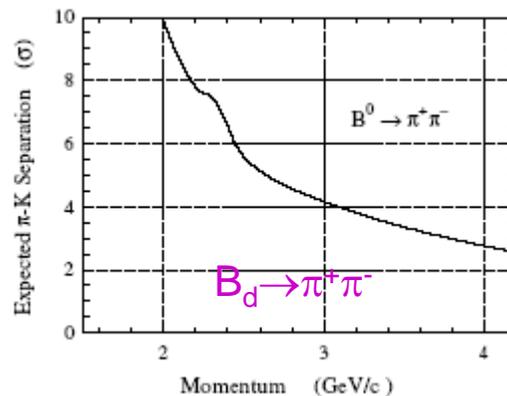
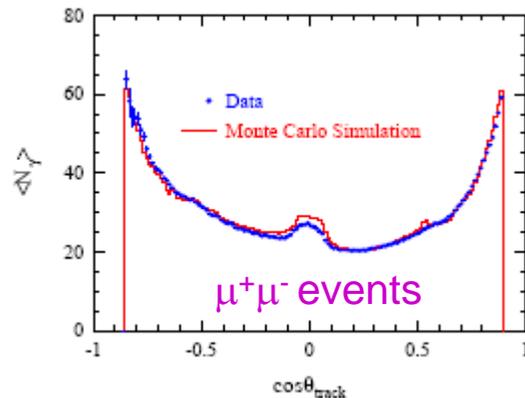
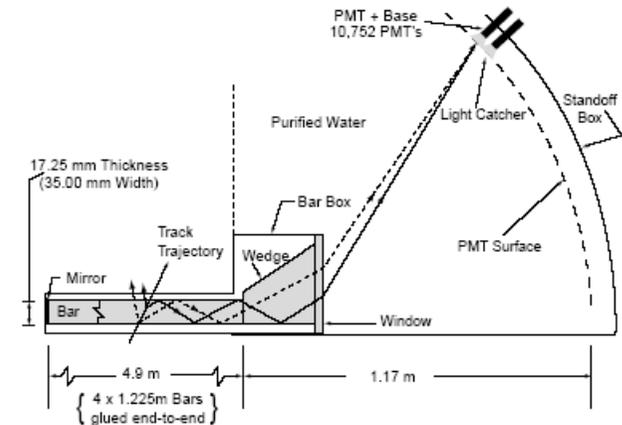
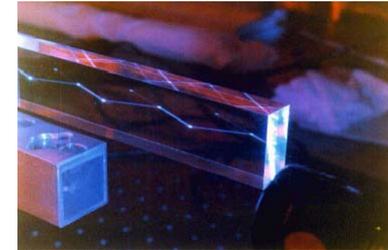


- 1st generation: 70's-80's
 - $n = 1.025-1.055$
- 2nd generation: 1992-2002
 - $n = 1.010...1.030$
 - new: hydrophobic
- 3rd generation: 2002-
 - $n = 1.030...1.080$
 - new: solvent



DIRC

- BaBar: Detector of Internally Reflected Cherenkov light:
 - 144 quartz rods, $1.7 \times 3.5 \times 490 \text{ cm}^3$, highest grade optical polish
 - angle of Cherenkov light wrt. track conserved
 - glued to quartz wedges to fold image
 - 6000l pure H_2O expansion tank ($n_{\text{H}_2\text{O}} \sim n_{\text{quartz}}$)
 - readout: 10752 PMT
 - single photon resolution: $\sigma_{C,\gamma} = 10.2 \text{ mrad}$
 - track resolution: $\sigma_{t,\gamma} = 1.7 \text{ ns}$
 - (for no systematic errors) $\sigma_{\text{track},\gamma} = \frac{\sigma_{C,\gamma}}{\sqrt{N_{pe}}}$
 - photon yield & K- π separation:

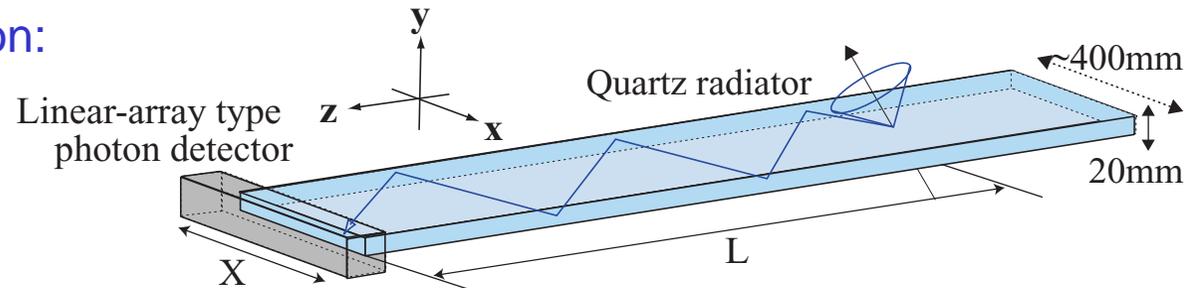
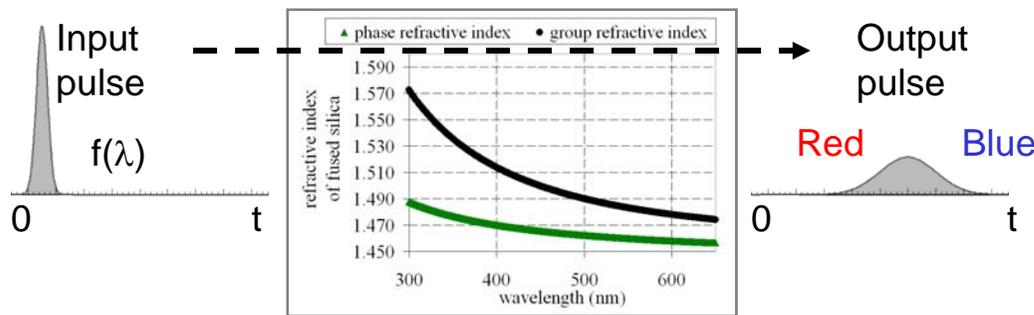


Time of Propagation

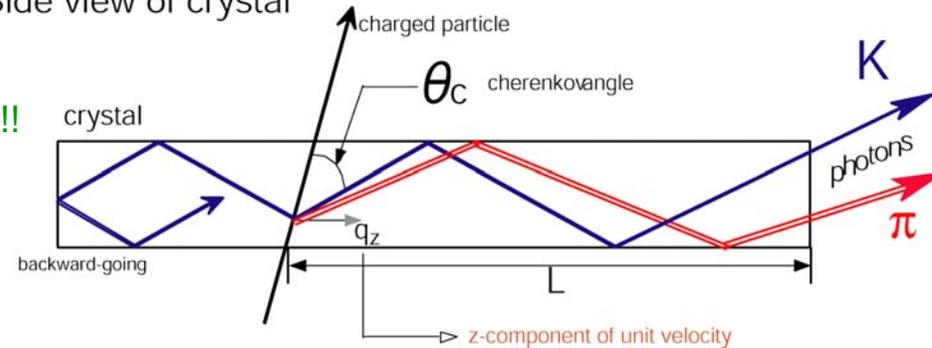
DIRC + propagation time information:

– 3 processes add in time:

- TOF from IP
- TOP in crystal due to path length:
 $L_\pi < L_K$, caused by $\theta_{c,\pi} > \theta_{c,K}$
 Belle upgrade: Barrel TOP: $\Delta t(\pi-K) \sim 200\text{ps}$
- TOP due to dispersion in optical medium:
 red photons travel faster than blue photons
 fused silica: $\Delta t/L \sim 40\text{ps/m}$



Side view of crystal



□ quest: $O(1\text{ps})$ time resolution!

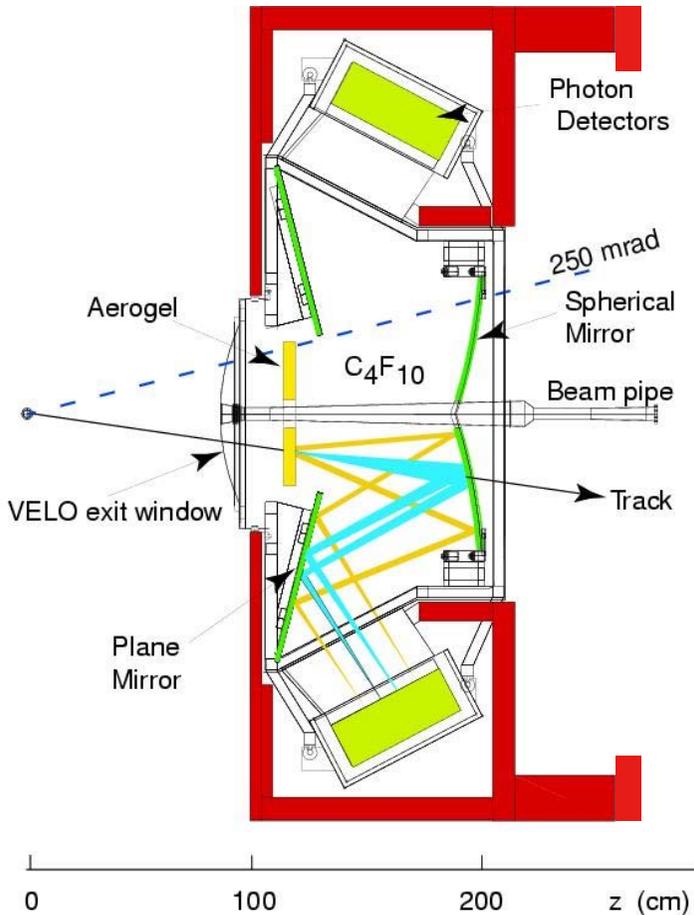
- recent testbeam: 27ps ($\sim 1 N_{pe}$),
 7.2ps ($\sim 50 N_{pe}$)

today's time resolution (of the shelf):

- Hamamatsu H-9500 Flat Panel MaPMT (256 pixels, 3x12mm pad, $\sigma_{TTS} \sim 220\text{ps}$)
- Hamamatsu H-8500 MaPMT (64 pixels, 6x6mm pad, $\sigma_{TTS} \sim 140\text{ps}$)
- Burle 85011-501 MCP-PMT (64 pixels, 6x6mm pad, $\sigma_{TTS} \sim 50-70\text{ps}$)

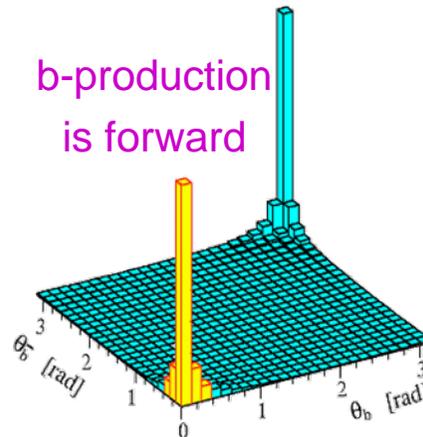
Ring Imaging Cherenkov Counter

RICH1
side view

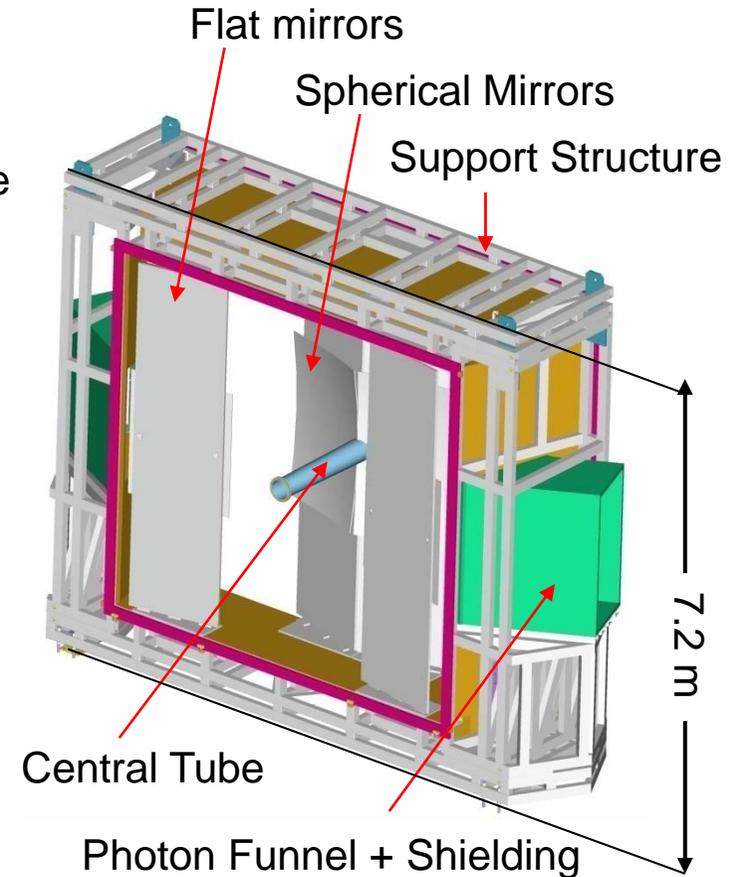


- LHCb RICH:
 - 3 radiator media: aerogel, C_4F_{10} , CF_4
- spherical mirrors:
 - focusing → ring image
 - tilt → outside acceptance
- secondary flat mirrors
 - magnetic shielding
- acceptance
 - RICH1: 25-300mrad
 - RICH2 :15-120mrad

b-production
is forward



RICH2



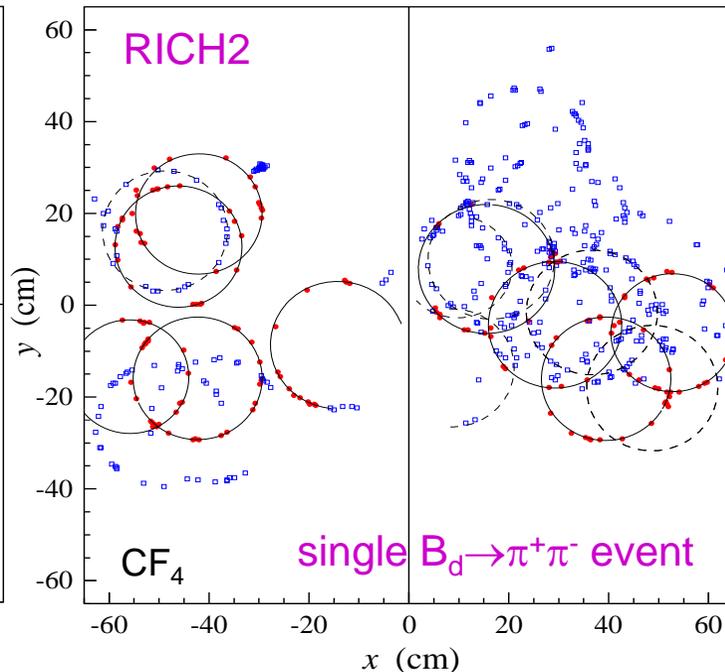
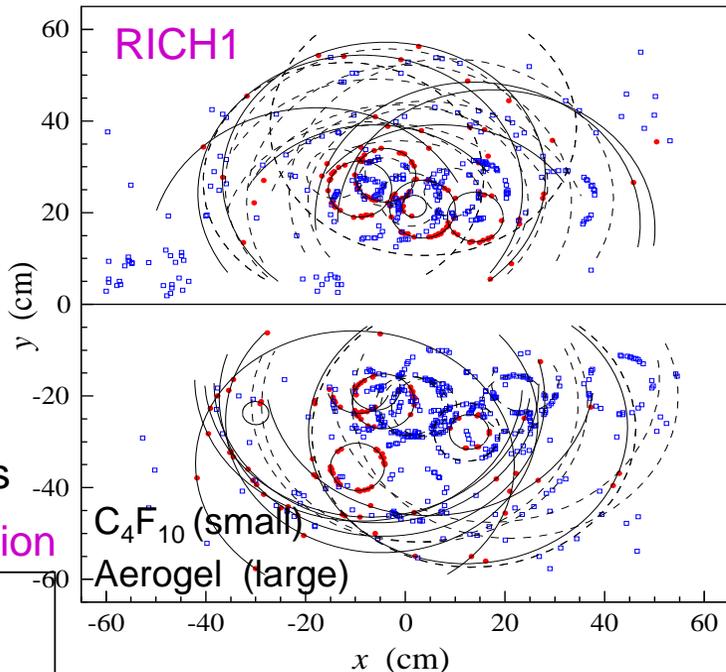
Pattern Recognition

□ LHCb RICH detector:

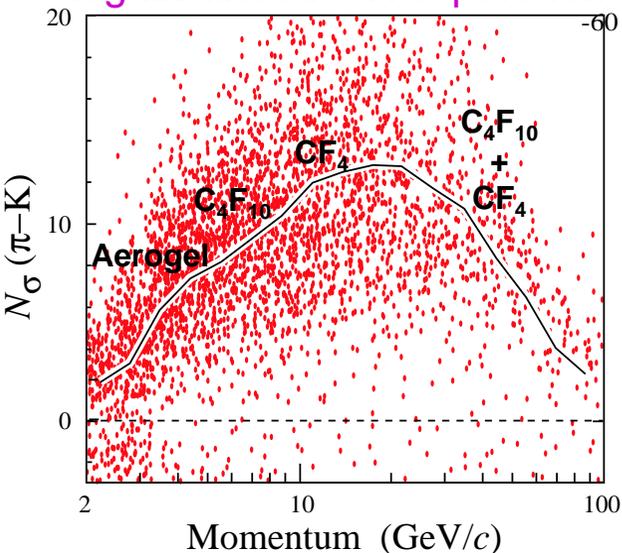
- find ellipses
- match to tracks
- ghosts: no track
- angle \rightarrow velocity

$$\cos(\theta_c) = \frac{1}{n \cdot v/c}$$

- track mom. \rightarrow mass



log-likelihood π -K separation



	Cherenkov angle resolution	detected photons
Aerogel	1.82 mrad	6.8
C_4F_{10}	1.27 mrad	30
CF_4	0.59 mrad	23

active area fraction
 $\sim 44 \times 0.7 \times 0.99 \times 0.99$
 mirror reflectivities

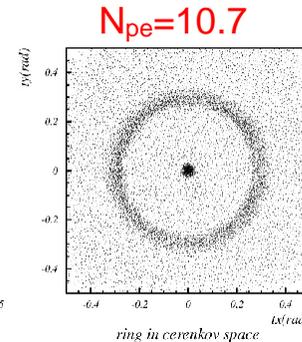
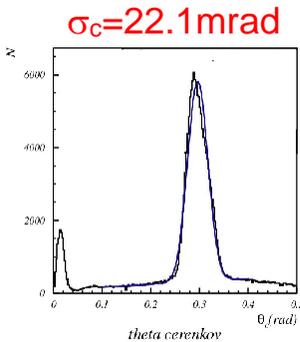
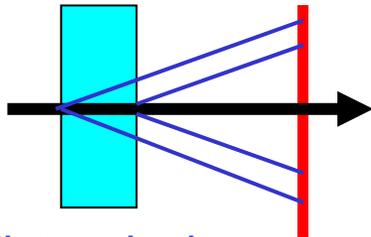
Proximity Focusing RICH

□ Belle upgrade: end-cap RICH

- proximity focusing = compact design → usable in storage ring detector!
- aerogel radiator, gap: O(20cm), high spatial resolution γ detection

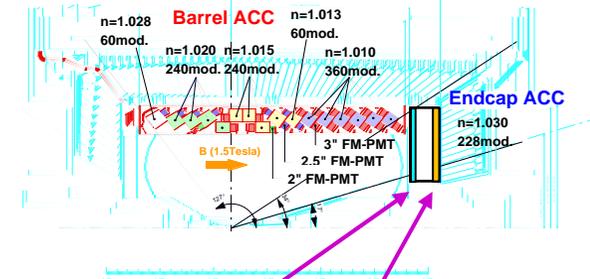
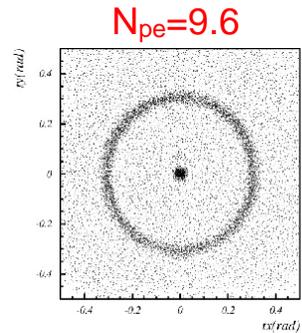
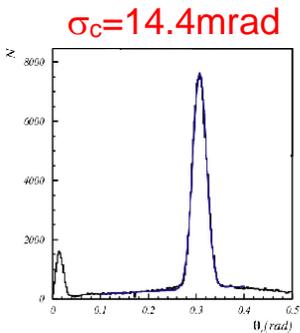
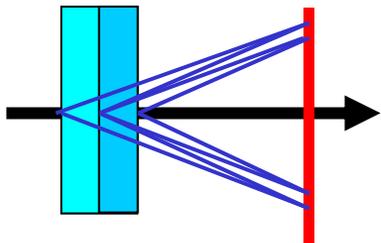
□ conventional design:

- 4cm thick aerogel
- $n=1.047$

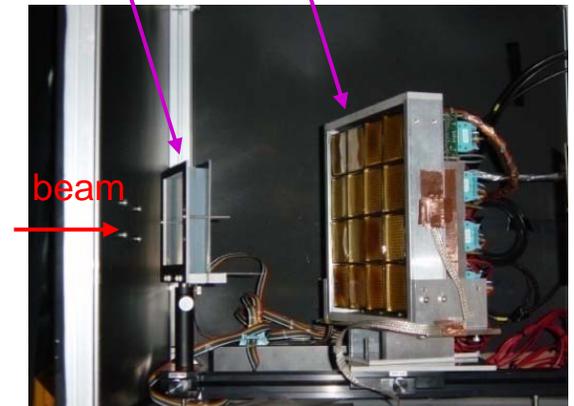


□ multiple radiator design:

- 2 layers, each 2cm thick
- $n_1=1.047, n_2=1.057$



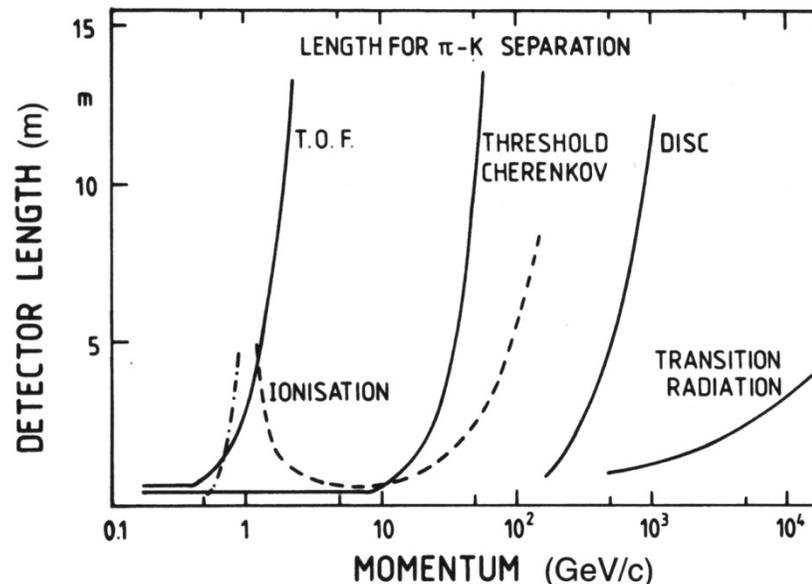
array of
2 layer
aerogel
flat-panel MaPMT



- π/K separation with focusing configuration: $\sim 4.8\sigma$ @4GeV/c

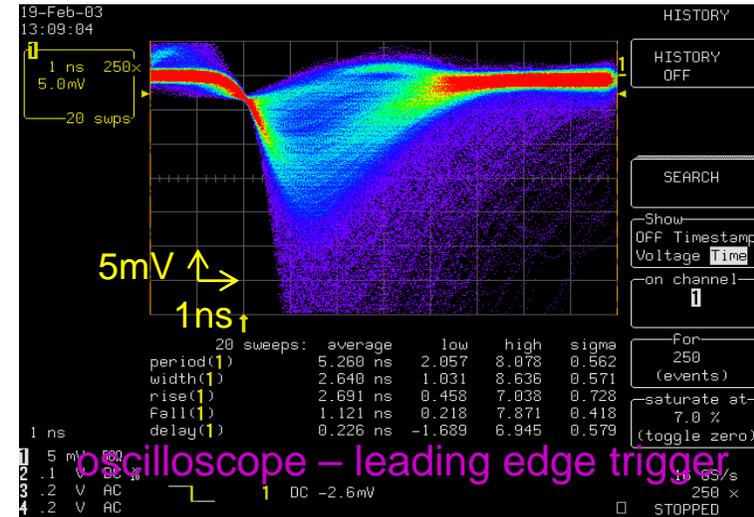
Particle Identification - Summary

method:	momentum range		requirements
	fixed target L=30m	storage ring L=3m	
dE/dx	0.2...2GeV/c	0.2...2GeV/c	$\sigma_r=2\%$ (3%) for 30m (3m)
time of flight	<4GeV/c	<1GeV/c	$\sigma_t=300\text{ps}$
DIRC, TOP	n.a.	2...4GeV/c	highest optical quality surface
Cherenkov threshold	<80GeV/c	<25GeV/c	10 photoelectrons
RICH, prox. focus. RICH	1...150GeV/c	0.7...4.5GeV/c	single γ , O(mrad) resolution
DISC	<2000GeV/c	n.a.	achromatic gas counter
dE/dx multiple ionisation meas.	1.2...100GeV/c	1.5...45GeV/c	$\sigma_r=2\%$ (3%) for 30m (3m)
transition radiation	$\gamma>1000$	$\gamma>1000$	X-ray detection with $E>10\text{keV}$

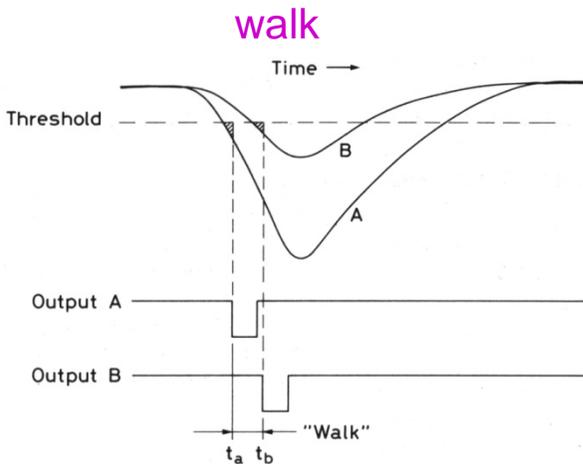


Trigger - Basics

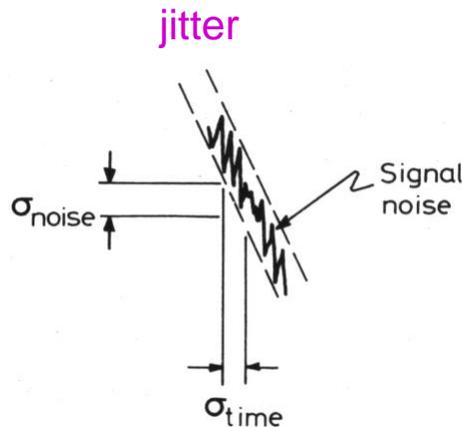
- **Trigger:**
 - time-stamp for occurrence of defined event
- **Walk:**
 - variations in amplitude or rise time
 - different rel. timing of leading edge wrt. signal
 - finite amount of charge to trigger discriminator
 - slope dependent excess over threshold needed
- **Jitter:**
 - noise & statistical fluctuations in signal
 - identical signal don't trigger at the same point



oscilloscope – leading edge trigger

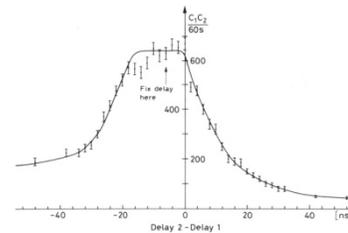


particle Physics Detectors, 2010

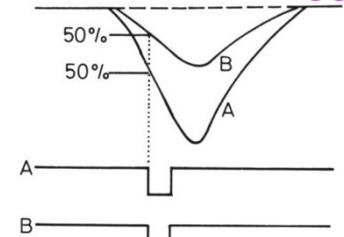


Stephan Eisenhardt

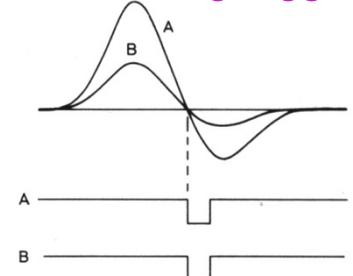
coincidence curve with finite resolution



constant fraction trigger

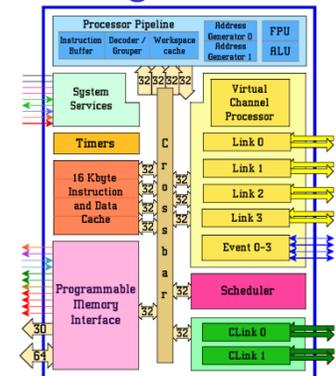


zero crossing trigger



Trigger Concepts

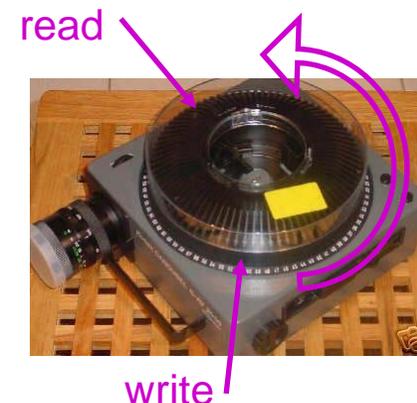
- **Electron collider:**
 - large cross section for studied processes
 - often: data taking in resonance → “take all” approach
 - i.e. buffer event bursts and write to disc between event bursts → just provide the bandwidth...
- **Hadron collider:**
 - dominated by background
 - seek “needle in haystack”
 - sophisticated, highly efficient online event selection needed
- **Tiered trigger for online event selection: cut background, leave maximum of signal**
 - level 0: hardware coded
 - fast: $O(\mu\text{s})$, deterministic in time
 - e.g. hit multiplicities, (transversal) energy sums for single detector sub systems
 - level 1: hardware or (preferably) software coded
 - factor 10-100 more time, still deterministic in time
 - merging level 0 data for fast detector sub-systems
 - level 2,3: parallel computing
 - event building → physical parameters as offline
 - parallel processing of events, variable time, single event might be time consuming
 - application of “physics filters”: cuts on high level parameters as close as possible to offline analysis



1990: transputer
today: PC farm

Dead Time, Latency & Bandwidth

- Accelerator clock:
 - gives time structure for events
 - defines data rate and requirements for data buffering and time to decide
- Dead time:
 - level 0 decision needs longer than clock cycle → may miss valid data for BG event
→ signal buffering: circular pipeline, continuously filled, event readout on L0 trigger
- Latency:
 - a) individual event: delay between event and trigger signal
 - b) trigger system: maximum allowed time for trigger decision
 - dead time free:
if trigger decision guaranteed to be faster than one pipeline revolution
- Bandwidth:
 - regard at each stage: bandwidth = event size x output rate
 - usually limited by: available technology and cost for computing and networking
 - total bandwidth get split into fixed or tuneable “physics channels”
 - pre-scaling: reduce large contributions by known fraction to enhance rare events



Modern Example: LHCb

Level 0:

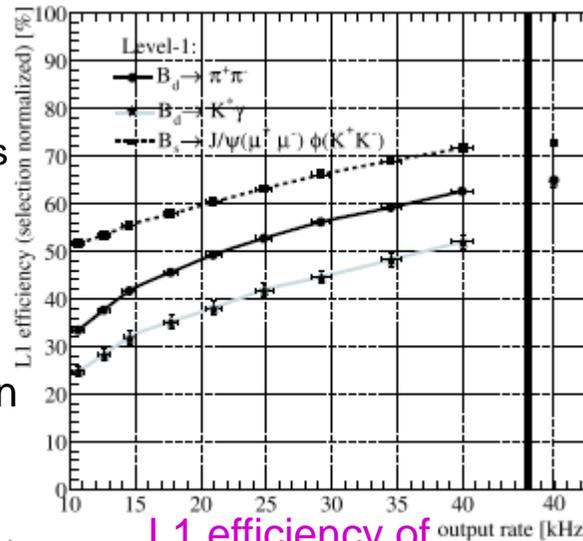
- 40MHz input
- 1MHz output: 1/40 reduction
- 4 μ s latency:
 - TOF+cables: <1000ns
 - processing: <1200ns
 - decision unit: <500ns
 - readout supervisor: <800ns
 - contingency: 500ns

Level 1:

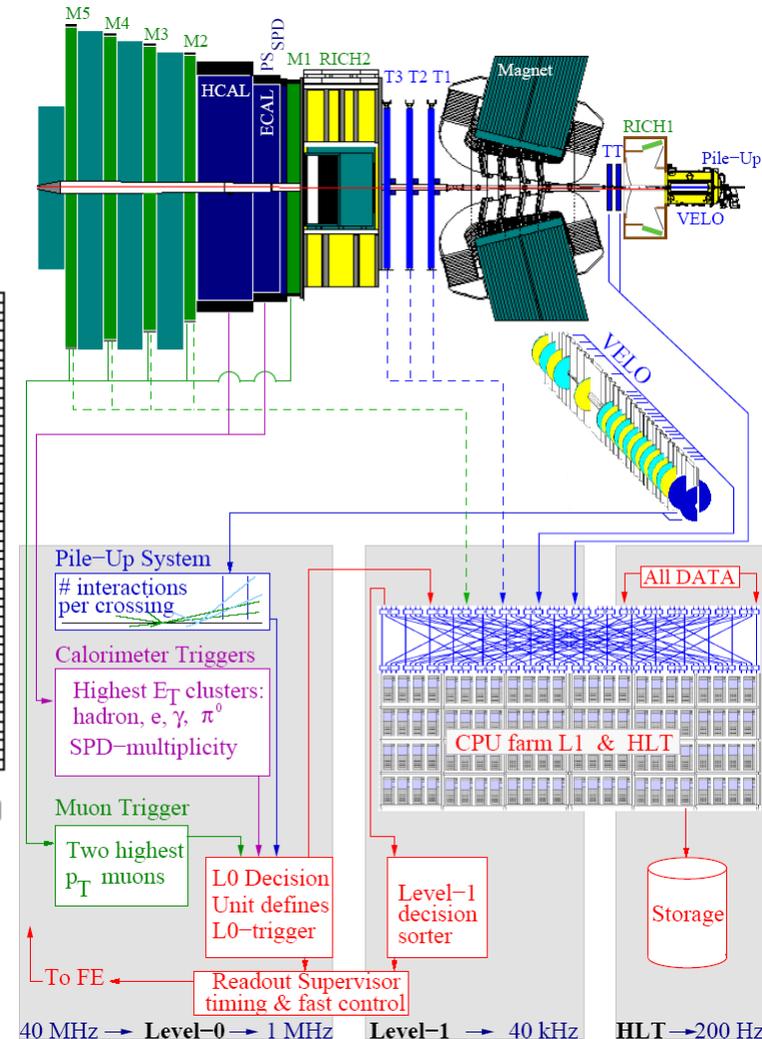
- 1MHz input
- 40kHz output: 1/25 reduction
- variable latency up to 58ms

HLT:

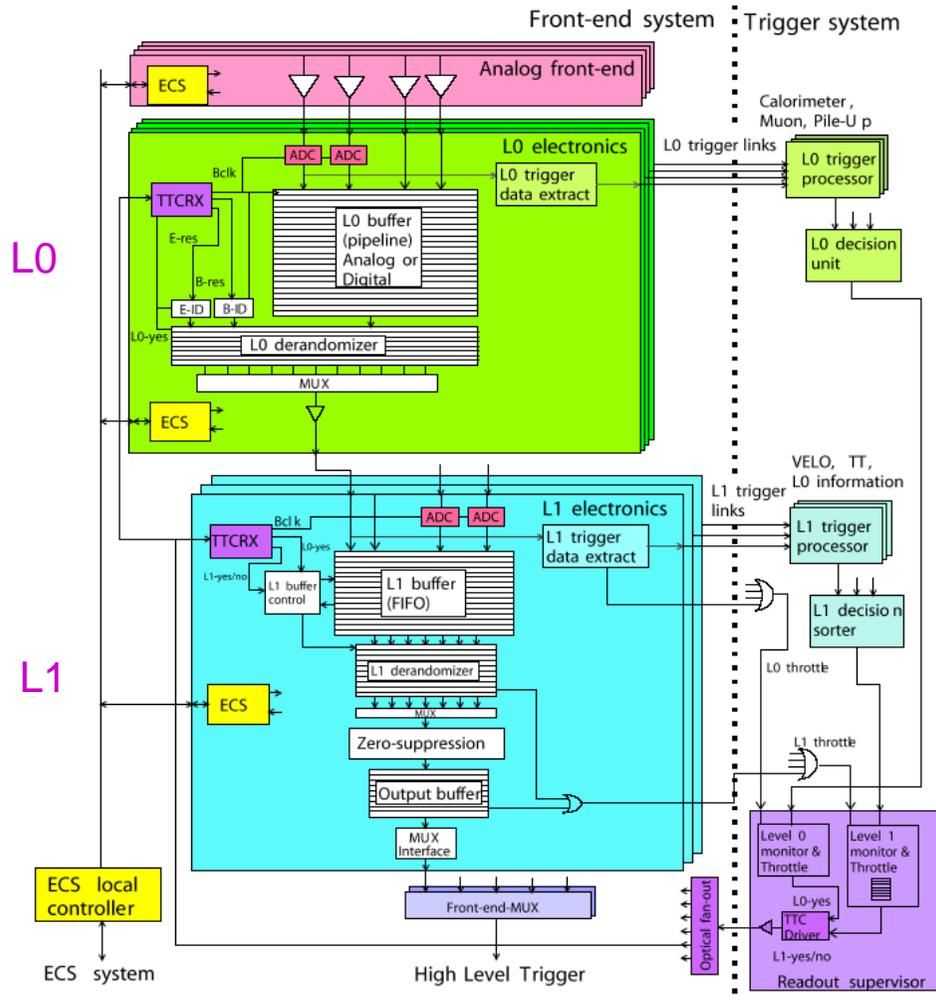
- L2+L3
- 40kHz input
- 200Hz output to disc/tape: 1/200 reduction



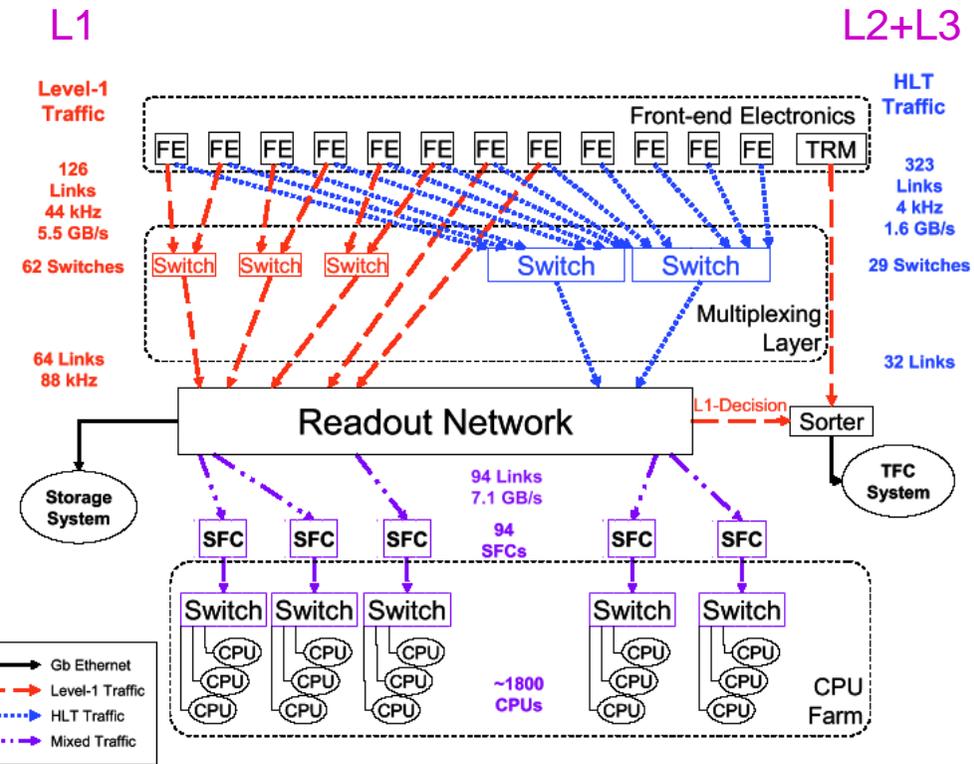
L1 efficiency of
L0 selected and
offline accepted events



Modern Example: LHCb



- derandomiser (L0 & L1):
 - synchronise data from sub-systems
- zero-suppression: @ L1
- data package sorting and packaging: @ L1

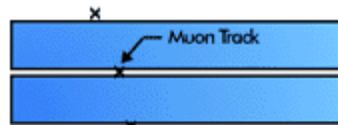


Trailer – not presented

- Trailer to course:
 - beyond the time scale of the lectures...
 - summary of integrated detector concepts
 - as ideas for self-study to dig further

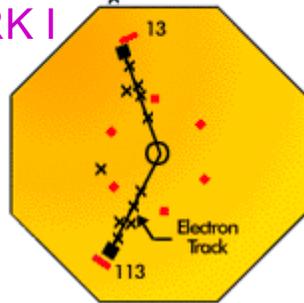
70's – integrated detectors

1973: MARK I

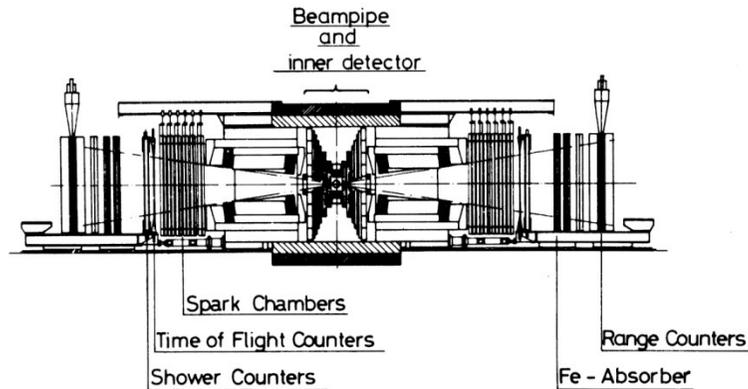


MARK I

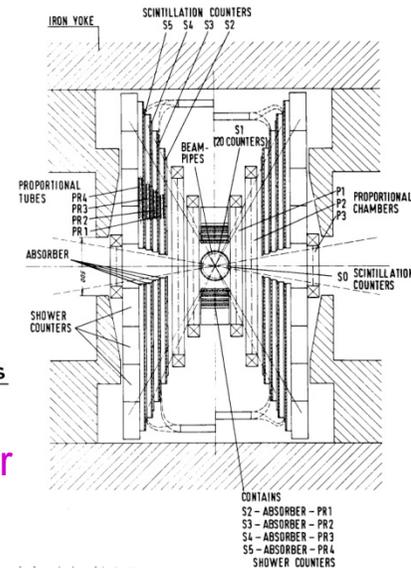
1974: DASP



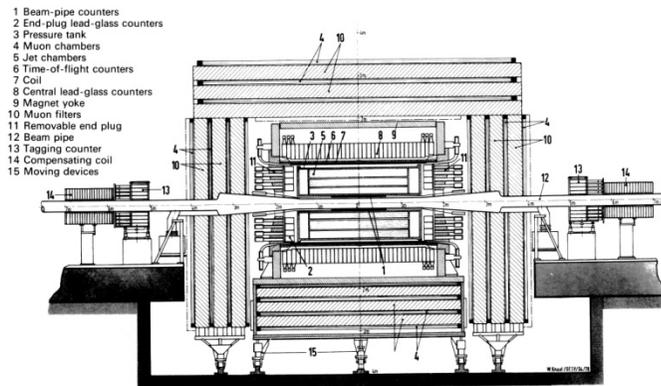
1977: MARK II



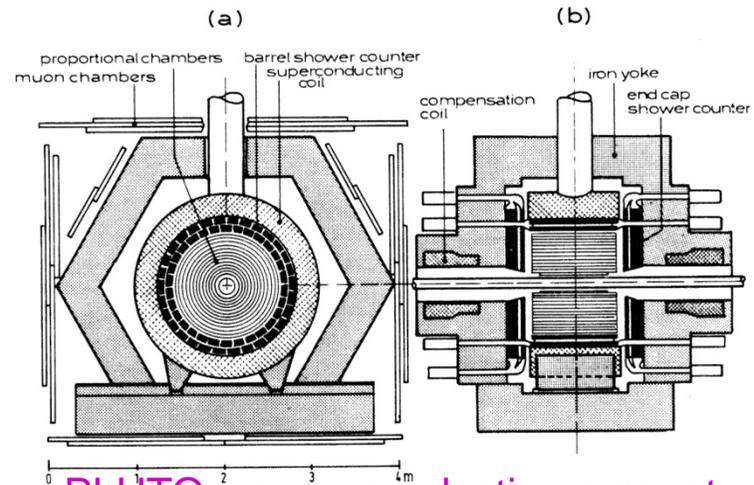
DASP – double arm spectrometer



1979: JADE, TASSO, PLUTO, MARK J



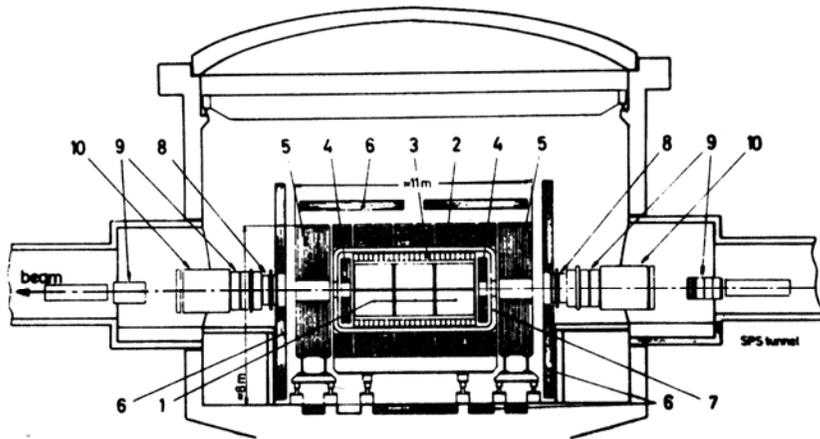
JADE – jet drift chamber



PLUTO – superconducting magnet

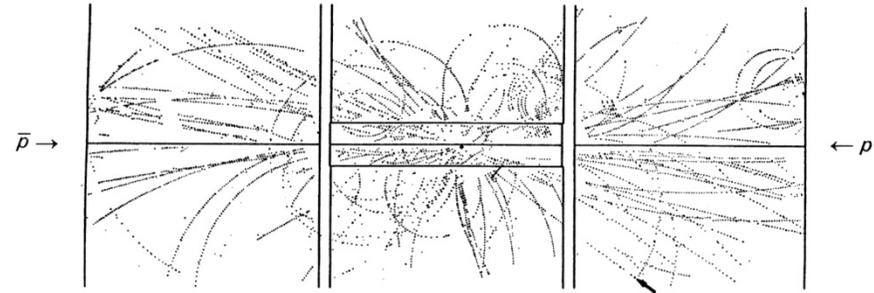
Vertex Detectors

- 1983: UA1 & UA2: gaseous tracking chambers



UA1 – central tracking

Fig.8.16: Seitenansicht des UA1-Detektors zum Nachweis von Proton-Antiproton-Wechselwirkungen bei 540 GeV Schwerpunktsenergie: 1. Zentraldetektor 2. und 5. Hadron-Kalorimeter, 3. und 4. Elektron-Photon-Schauerzähler 6. Myon-Detektor, 7. Spule für Dipolfeld, 8. und 9. Kleinwinkeldetektor mit Kammern und Kalorimetern, 10. Kompensator-Magnete [UA1].



first W → ev

- 1982: SLC: silicon detectors

SLC – first CCD vertex

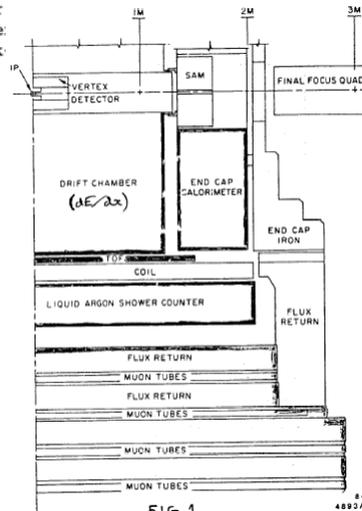
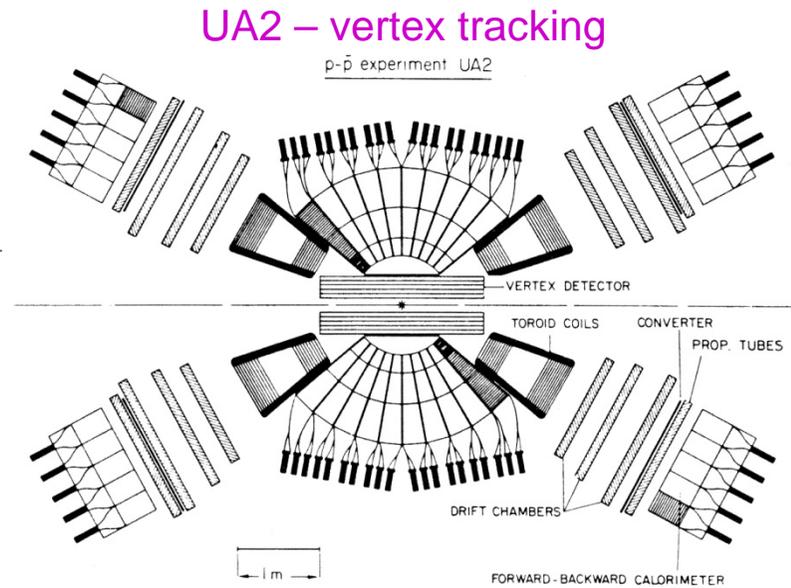


FIG. 1

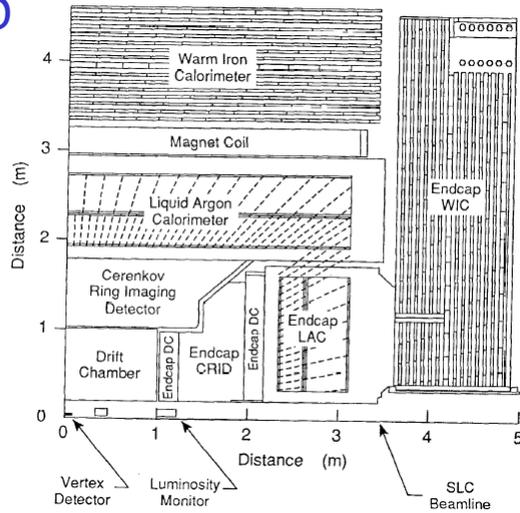


UA2 – vertex tracking

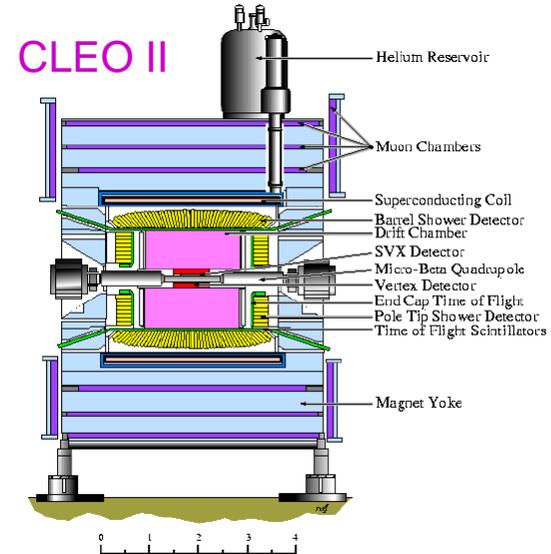
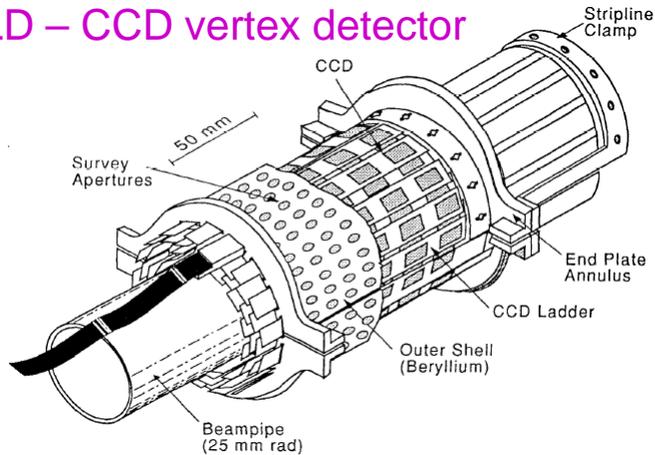
p-p̄ experiment UA2

SLD & CLEO

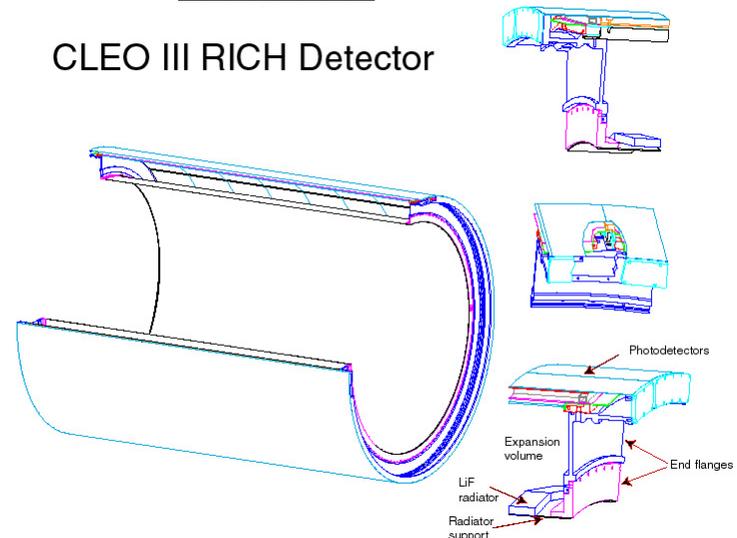
- 1982: CLEO I, 1993: CLEO II, 2001: CLEO III
- 1990: SLD



SLD – CCD vertex detector

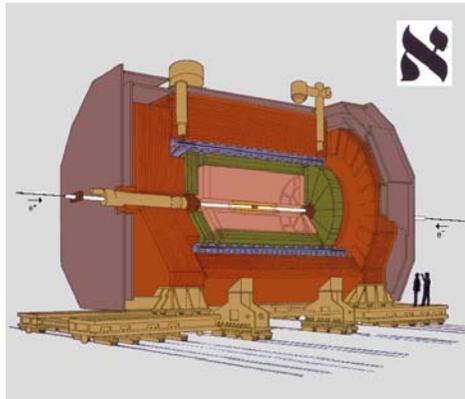


CLEO III RICH Detector



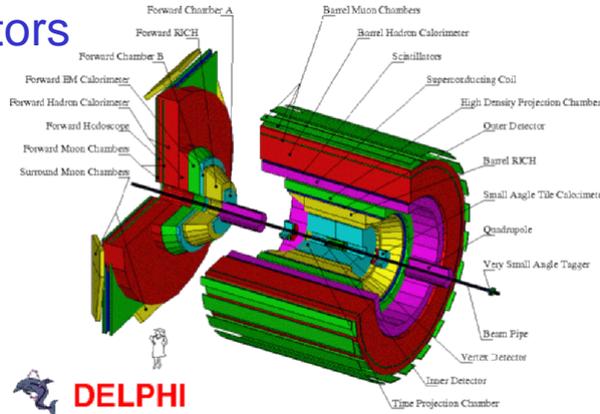
LEP

- 1989: $\sqrt{s} = 80-205\text{GeV}$ e^+e^- accelerator
- hermeticity & silicon vertex detectors

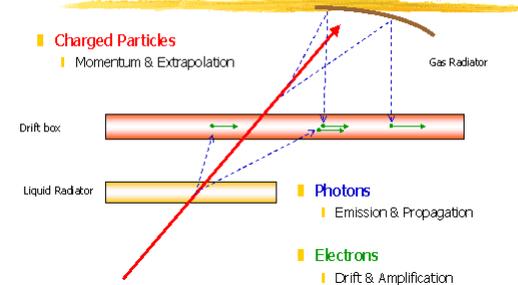


The ALEPH Detector

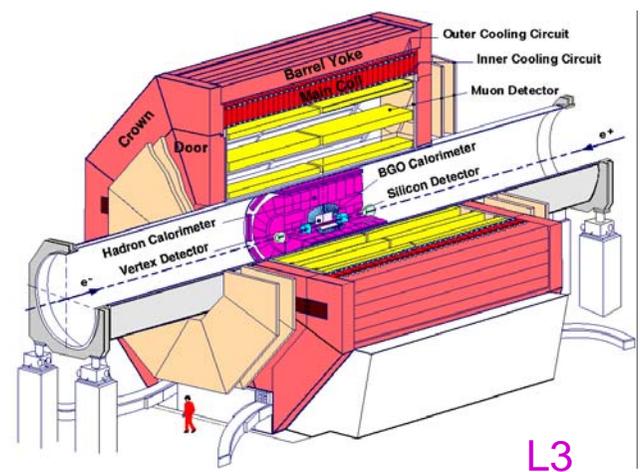
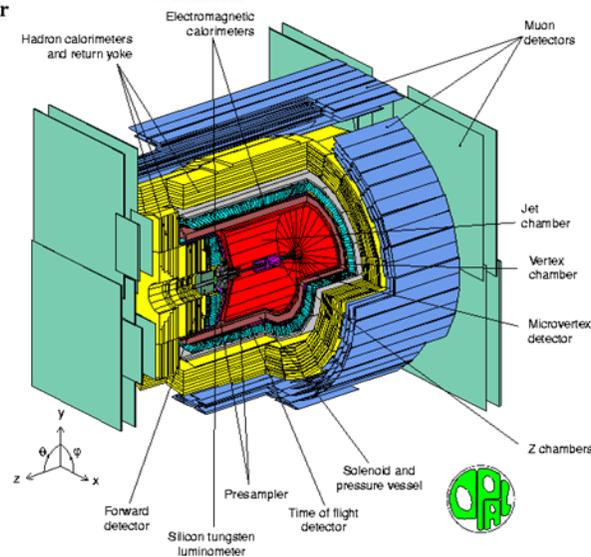
- Vertex Detector
- Inner Tracking Chamber
- Time Projection Chamber
- Electromagnetic Calorimeter
- Superconducting Magnet Coil
- Hadron Calorimeter
- Muon Chambers
- Luminosity Monitors



Detector Effects due to ...

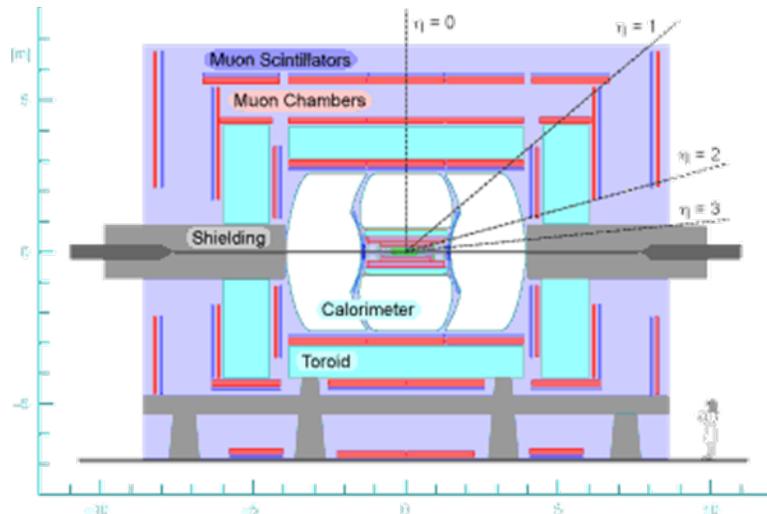


DELPHI - RICH

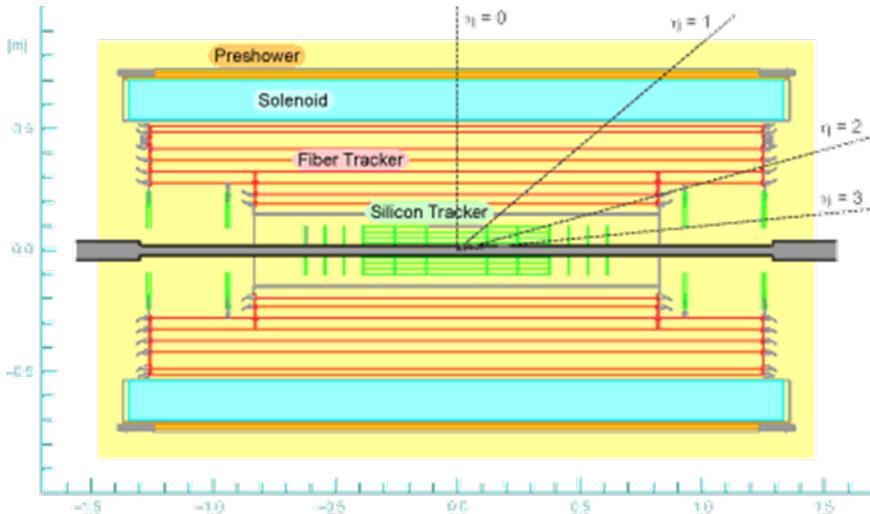


Tevatron

1992: $\sqrt{s} = 1.8\text{TeV}$ D0



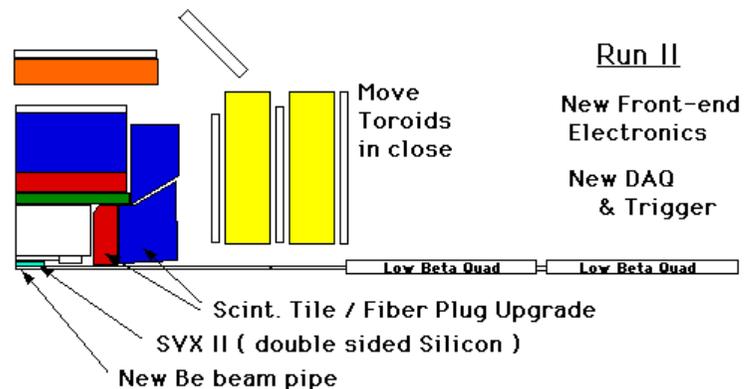
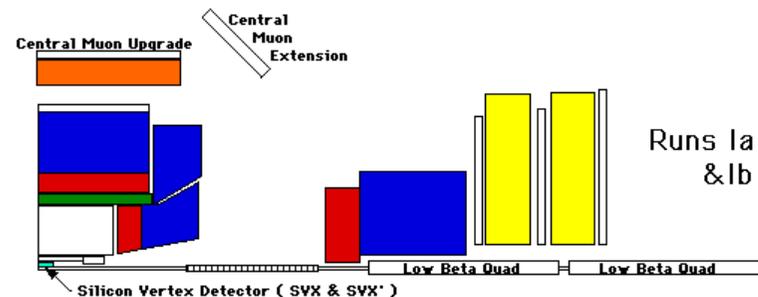
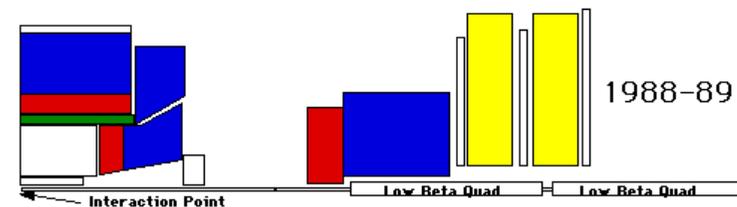
silicon vertex detector



CDF CDF Detector Evolution

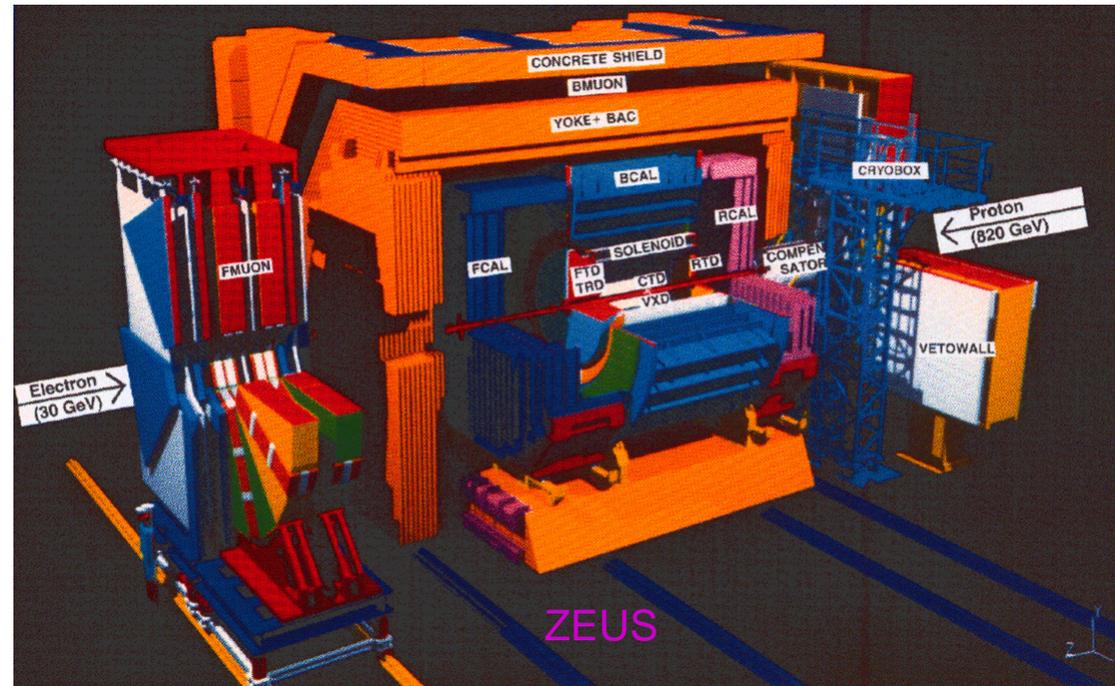
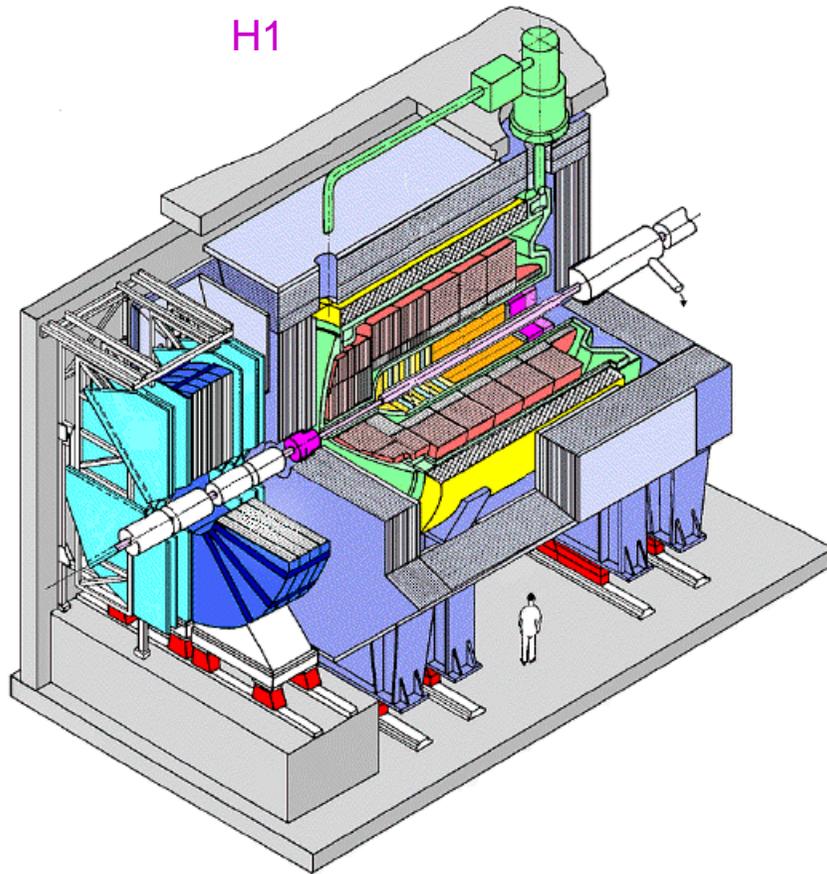
Key:

- Solenoid Coil
- Toroid
- Steel Shielding
- Track Chamber
- Electromagnetic Calorimeter
- Hadronic Calorimeter

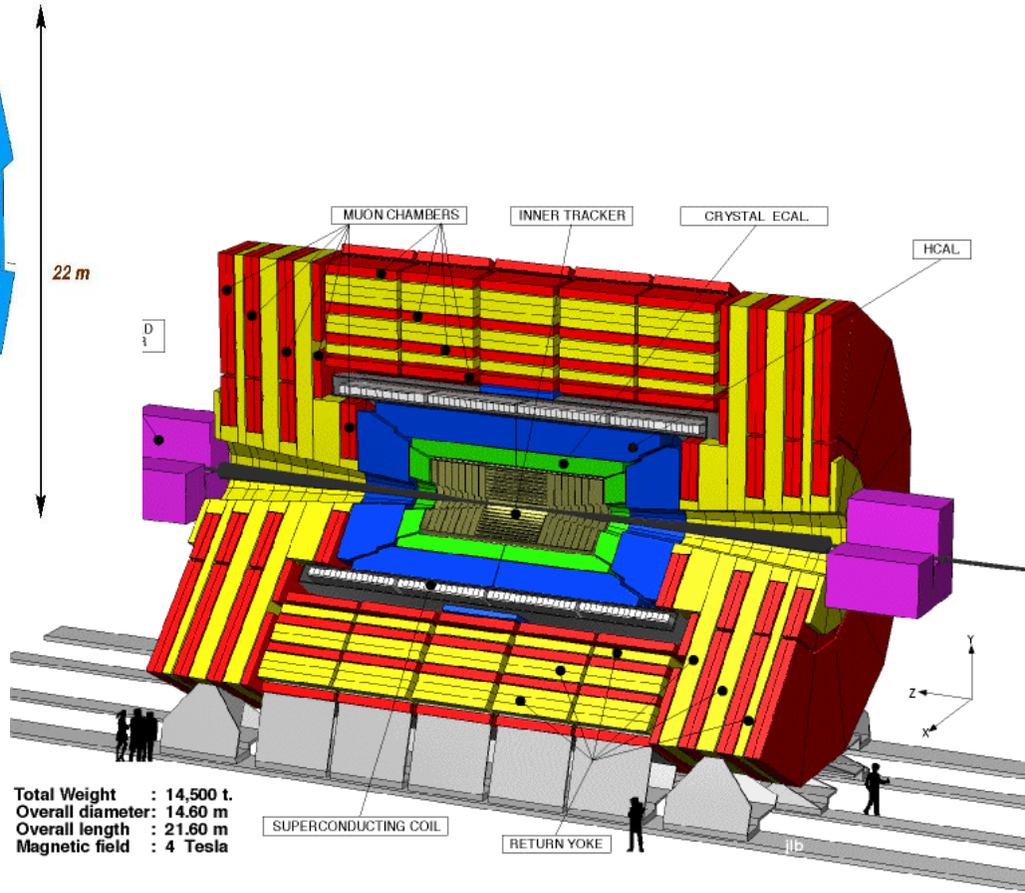
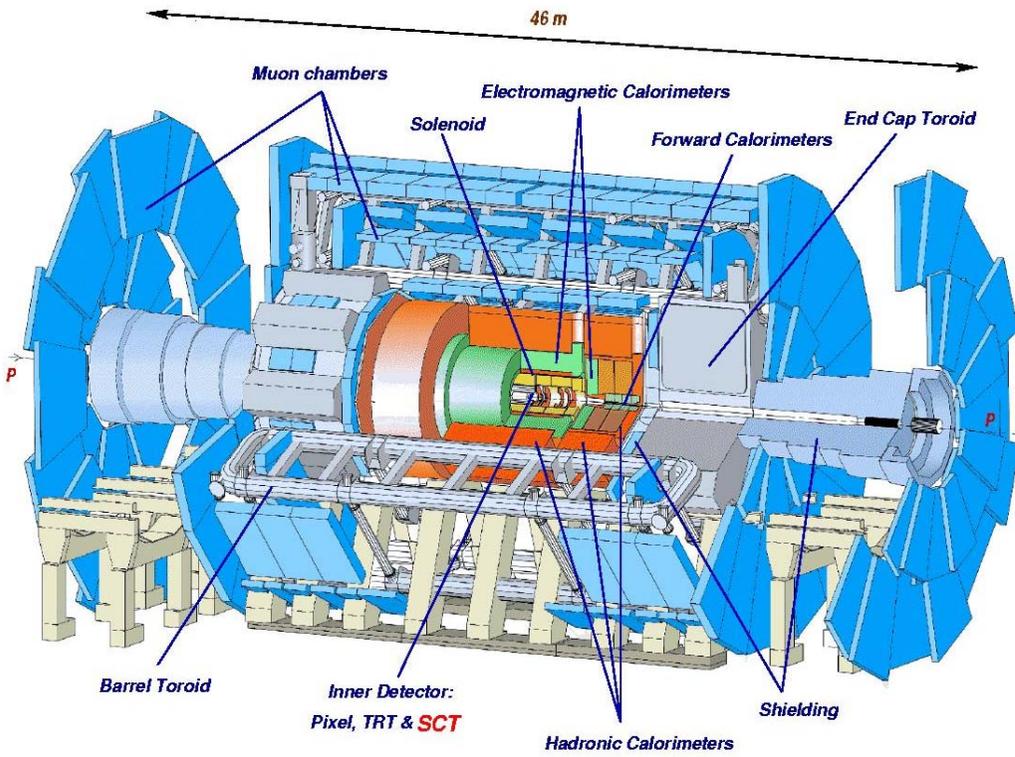


HERA

- 1992: asymmetric ep-accelerator
- hadron calorimetry



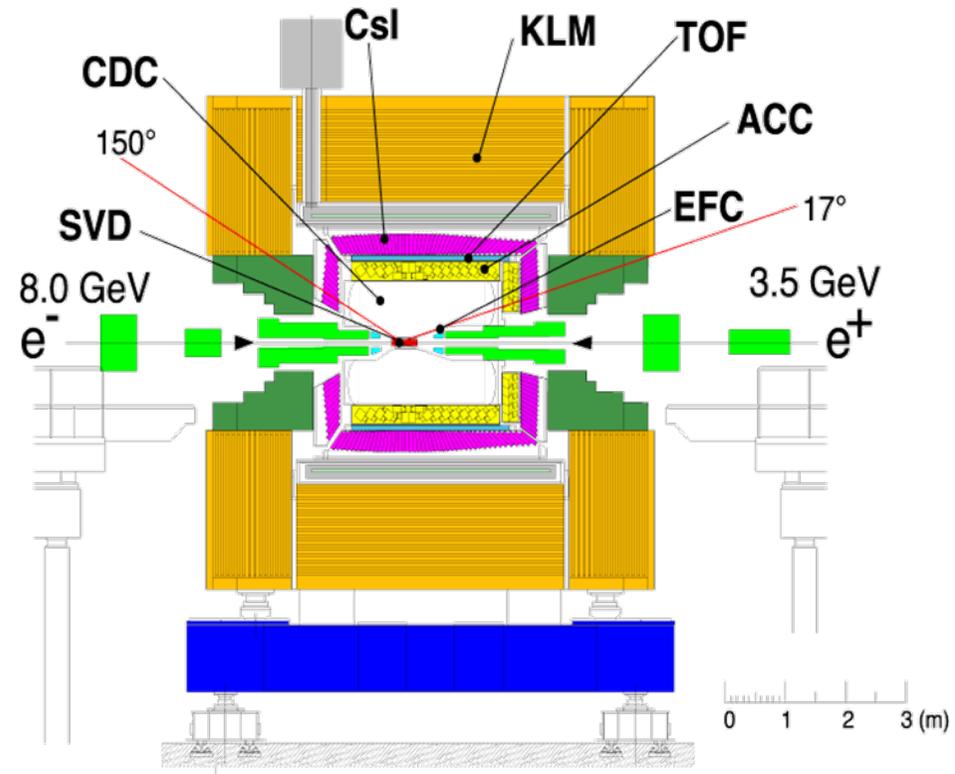
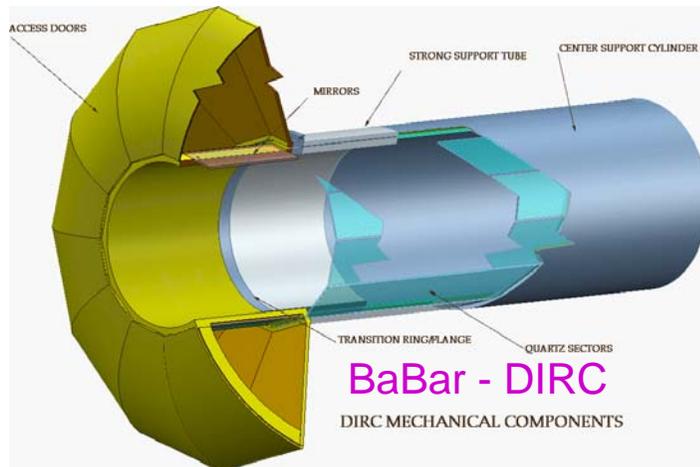
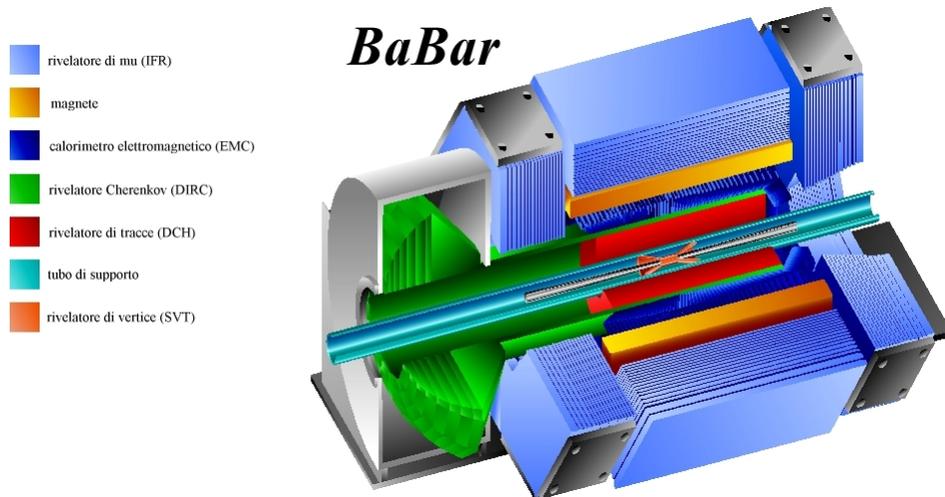
Atlas & CMS



Total Weight : 14,500 t.
 Overall diameter: 14.60 m
 Overall length : 21.60 m
 Magnetic field : 4 Tesla

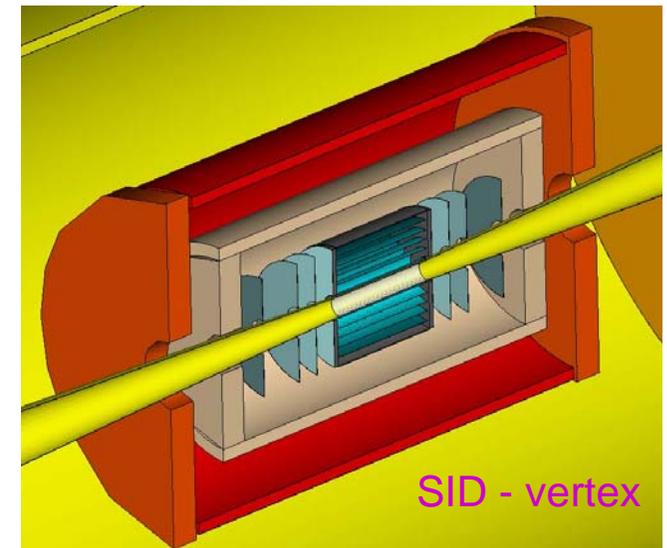
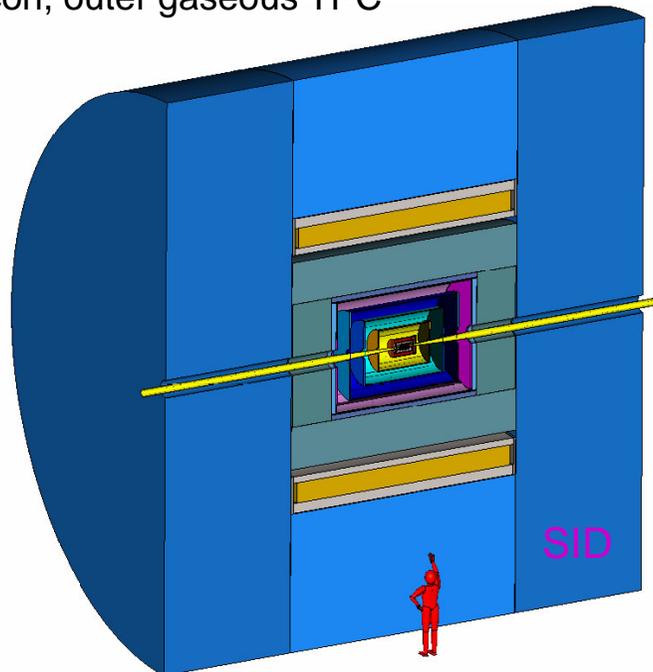
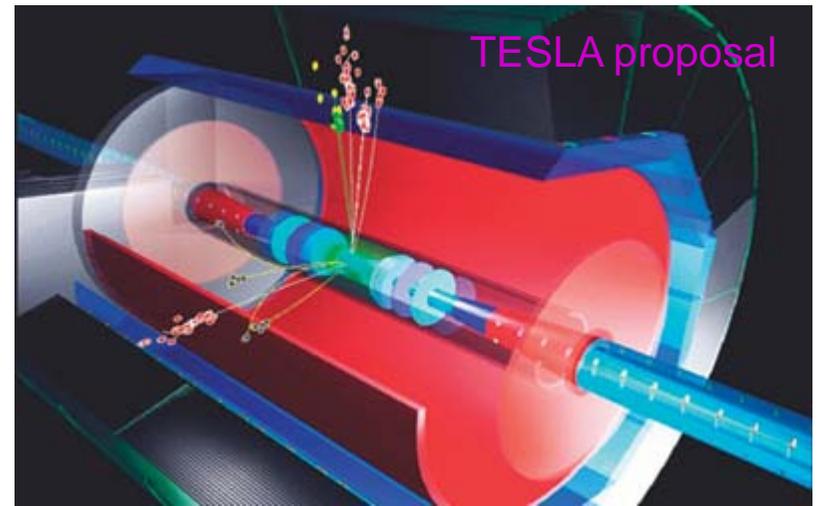
BaBar & Belle

- 2000: B – physics at asymmetric e^+e^- accelerators



ILC Detectors

- ~ 2015: three design collaborations:
 - each is “global”
 - SiD: Silicon Detector
 - silicon only
 - LDC: Large Detector Concept
 - large gaseous TPC
 - GLC: Global Large Detector
 - inner silicon, outer gaseous TPC



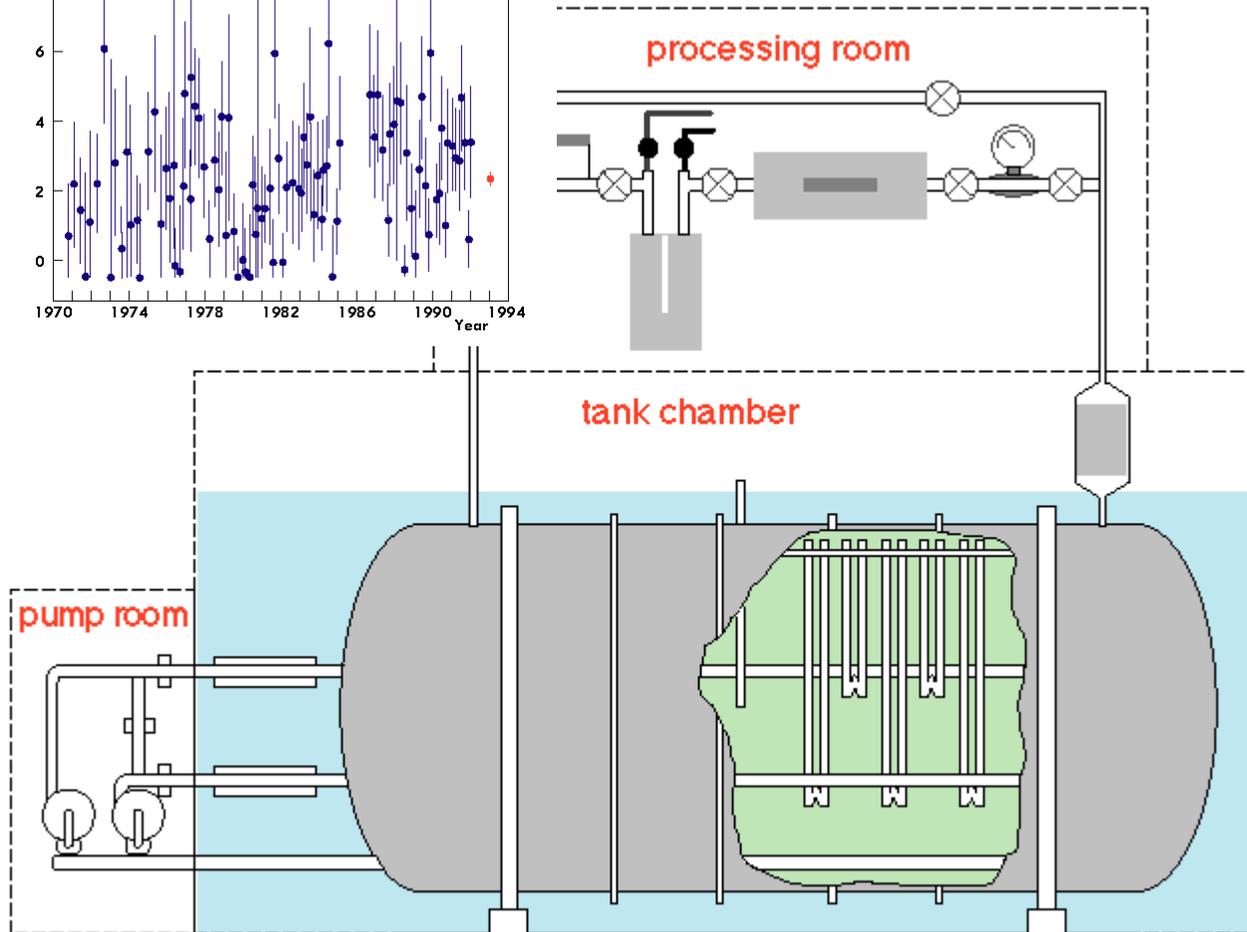
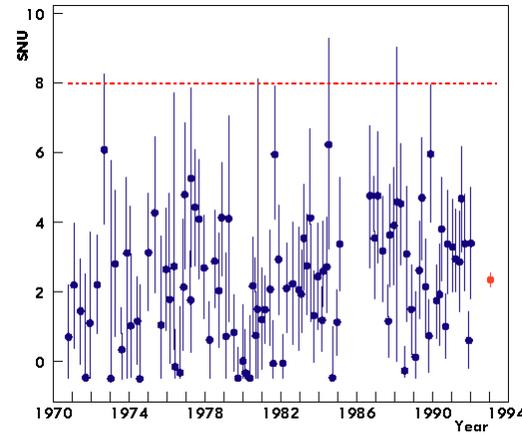
Homestake

□ 1969-1993:

- 615 tons tetrachloroethylene
- $\nu + {}^{37}\text{Cl} \rightarrow {}^{37}\text{Ar}(\tau=35\text{days})$
- wash O(atom) a day!

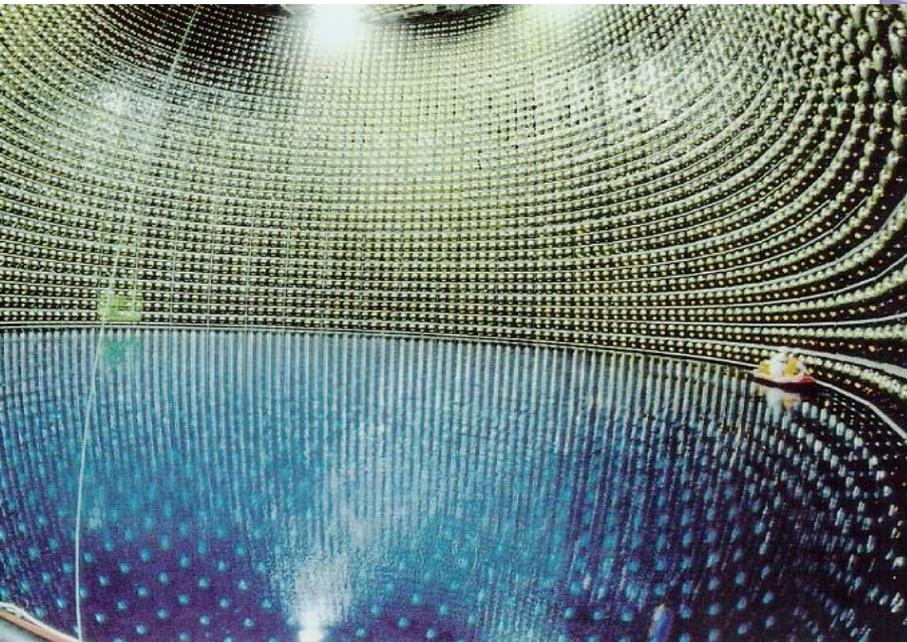


- new: SAGE, GALLEX

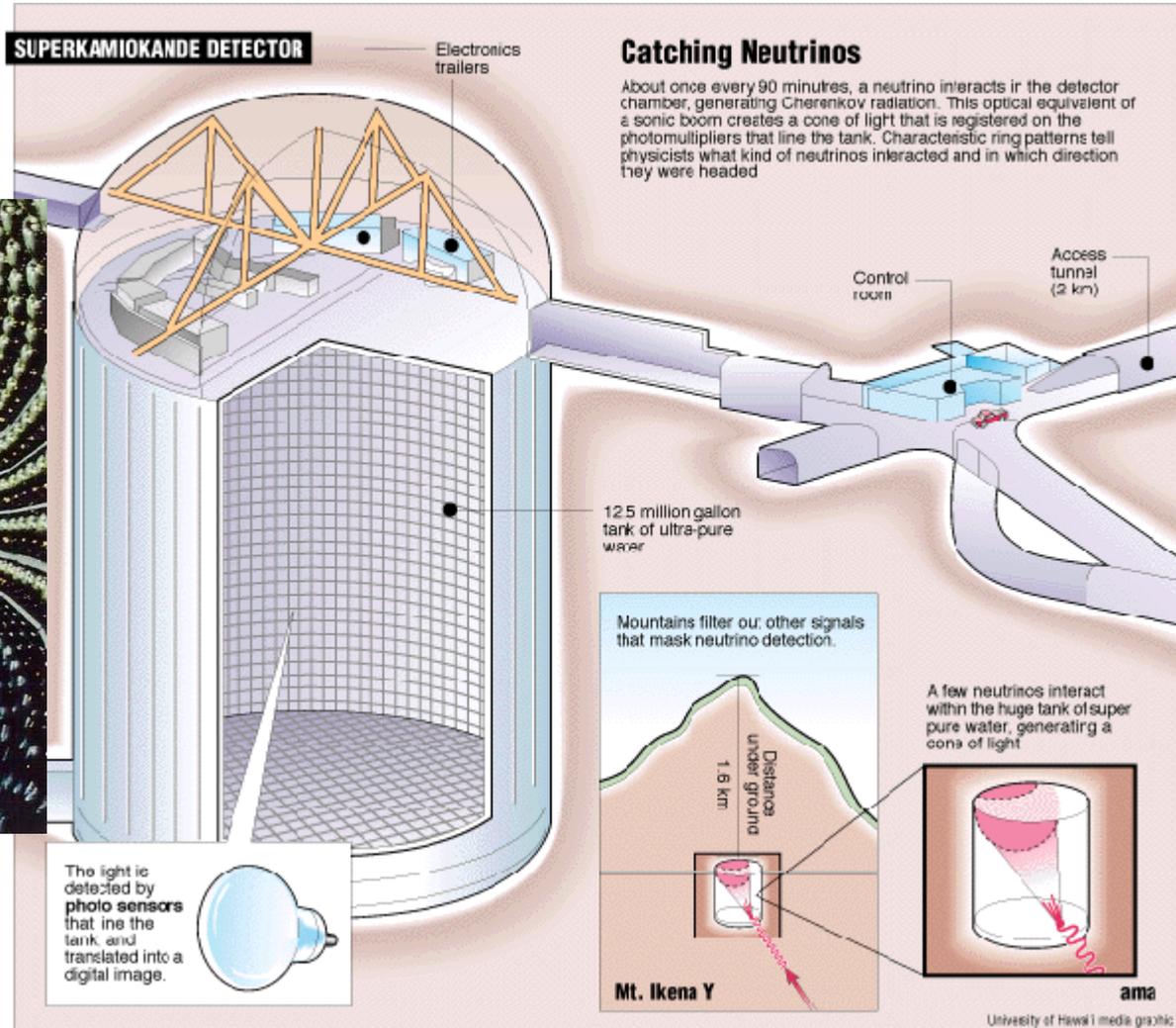


Super-Kamiokande

- ❑ 1996: 50000 tons ultra-pure water
 - ❑ 11146 PMT of 50cm diameter
- (1983: Kamiokande, 3000 tons, 1000 PMT)

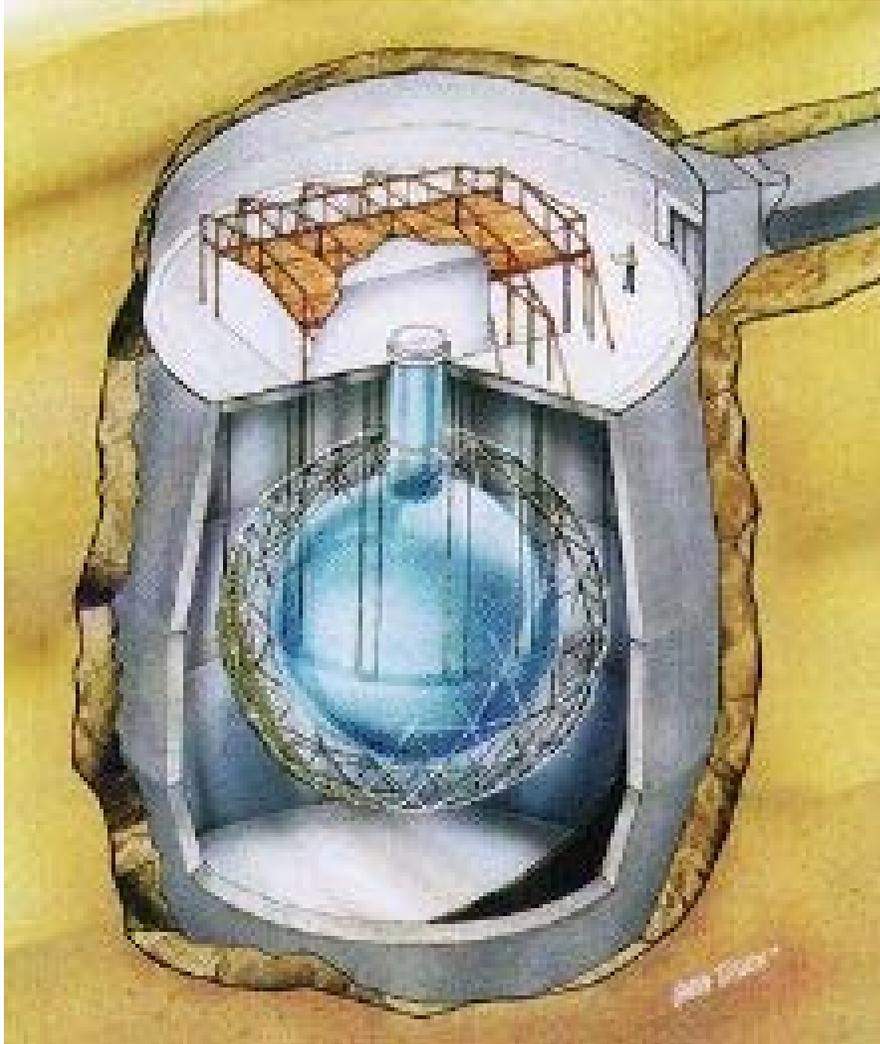


2001: implosion of 6777 PMT in chain-reaction



SNO

- 1999: Sudbury Neutrino Observatory: 1000 tons of heavy water (D_2O)



Particle Physics Detectors, 2010



Stephan Eisenhardt

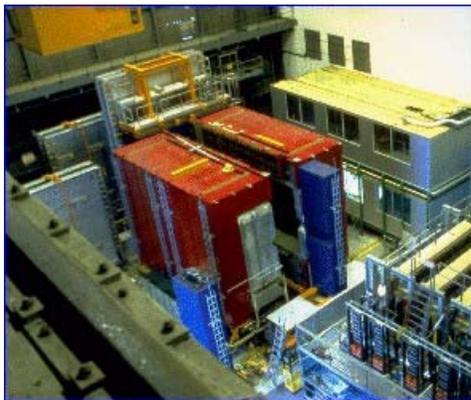
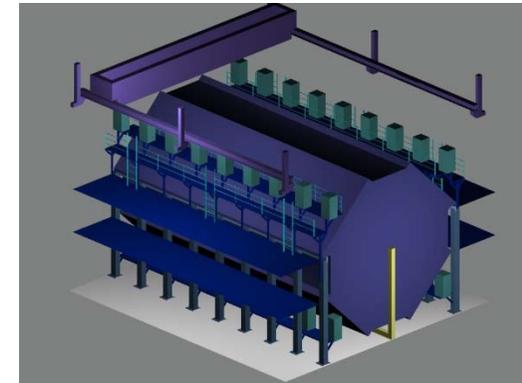
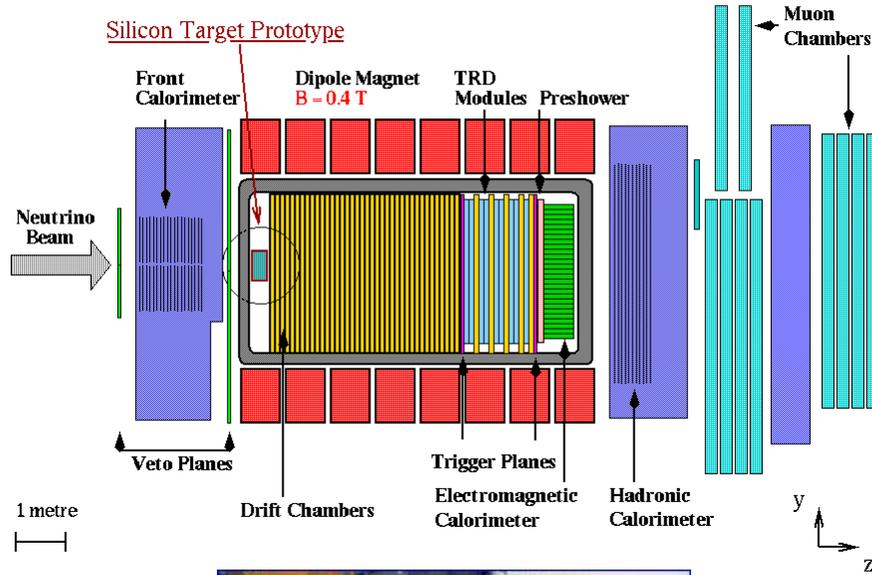
NOMAD & MINOS

1993: Neutrino Oscillation Magnetic Detector

2003: Main Injector Neutrino Oscillation Search

short "long base line": 835m

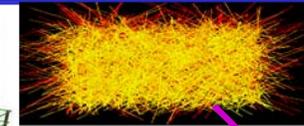
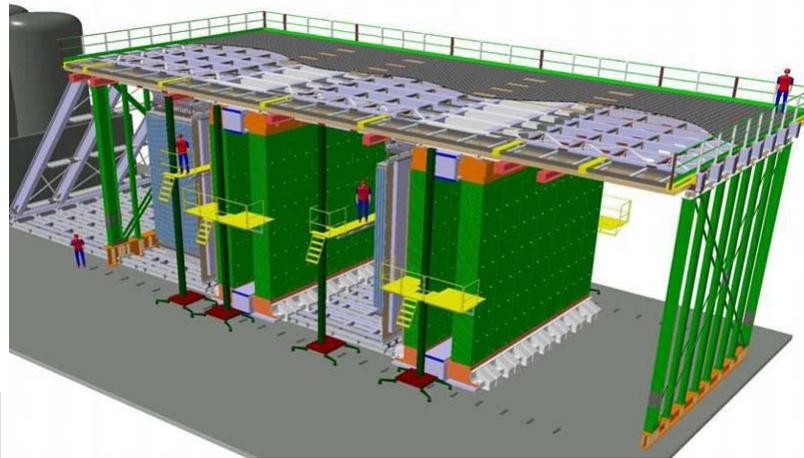
long base line: 730 km



Opera

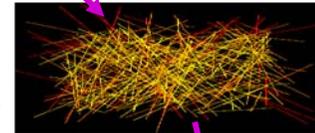
□ Goal: ν_τ appearance

- 0.15MW source
- high energy ν_μ beam
- long baseline: 732km
- handfuls of events/yr

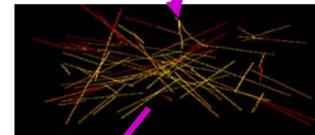


hits in brick

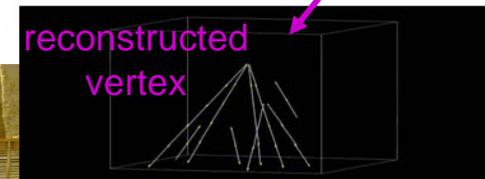
tracks in brick



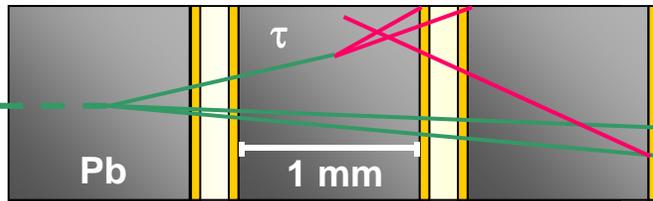
appearing tracks



reconstructed vertex

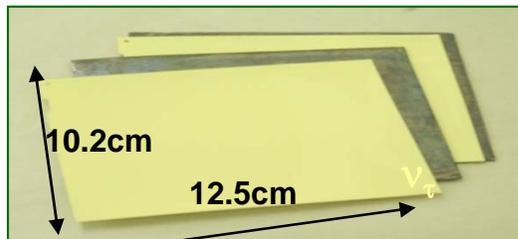


□ return of the nuclear emulsion

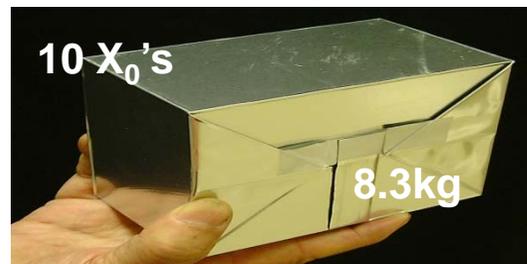


emulsion layers (44 μ m thick)
+ 200 μ m plastic spacer

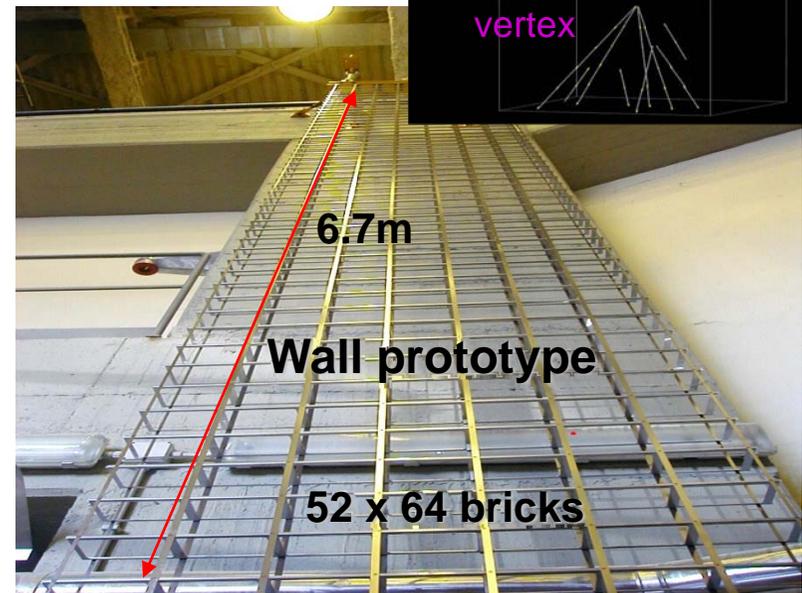
...swap brick if
Scintillator trigger
has noticed some
tracks...



layers



BRICK: 57 emulsion foils &
56 interleaved Pb plates



6.7m

Wall prototype

52 x 64 bricks

ICARUS

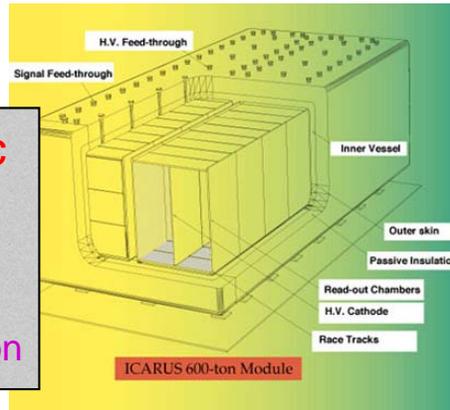
- Liquid Argon TPC: to study ν_τ appearance
 - Electronic Bubble chamber
 - Planes of wires (3mm pitch) widely separated (1.5m)
 - 55K readout channels!
 - Very Pure Liquid Argon
 - Density: 1.4, $X_0=14\text{cm}$ $\lambda_{\text{had}}=83\text{cm}$
 - $3.6 \times 3.9 \times 19.1\text{m}^3$: 600 ton module (480fid)

ν_τ appearance: signature event

e^- , 9.5 GeV, $p_T=0.47$ GeV/c



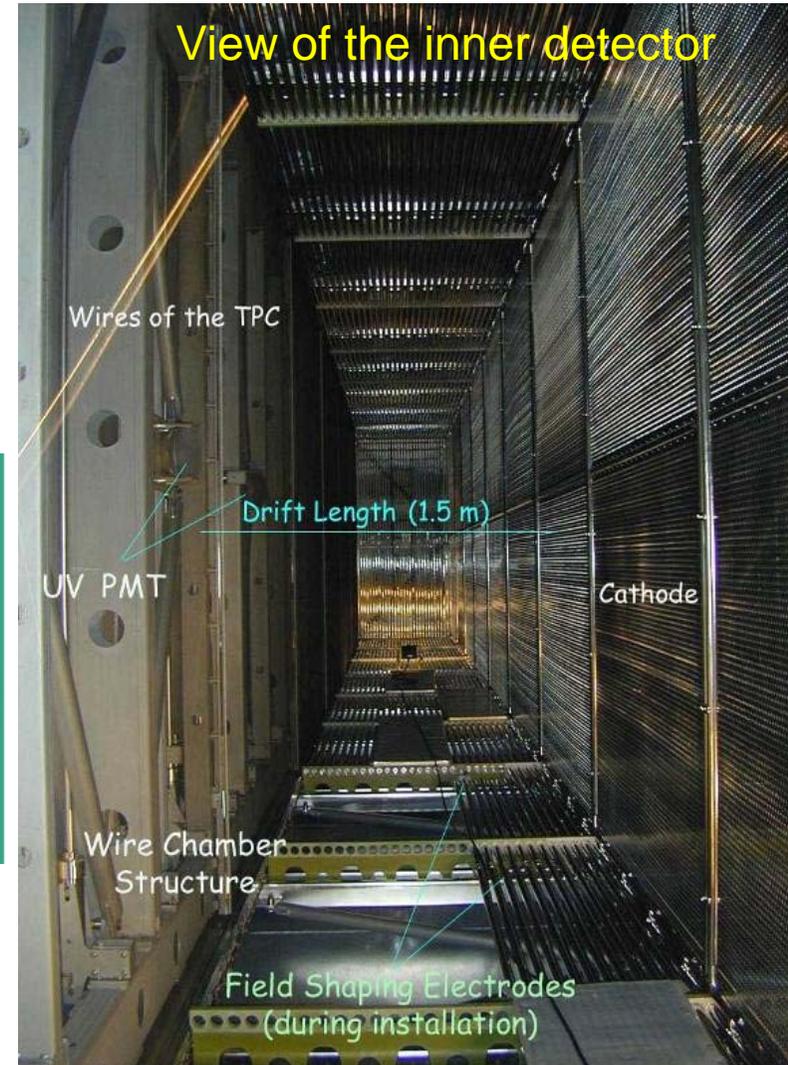
MC simulation



elmag. showers from 2γ

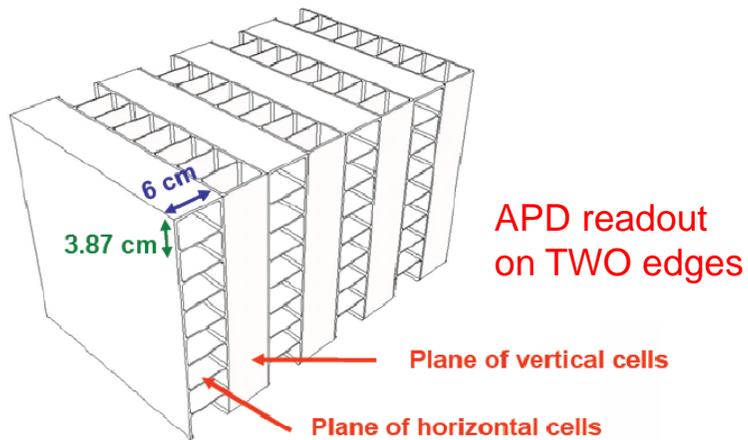


MC simulation



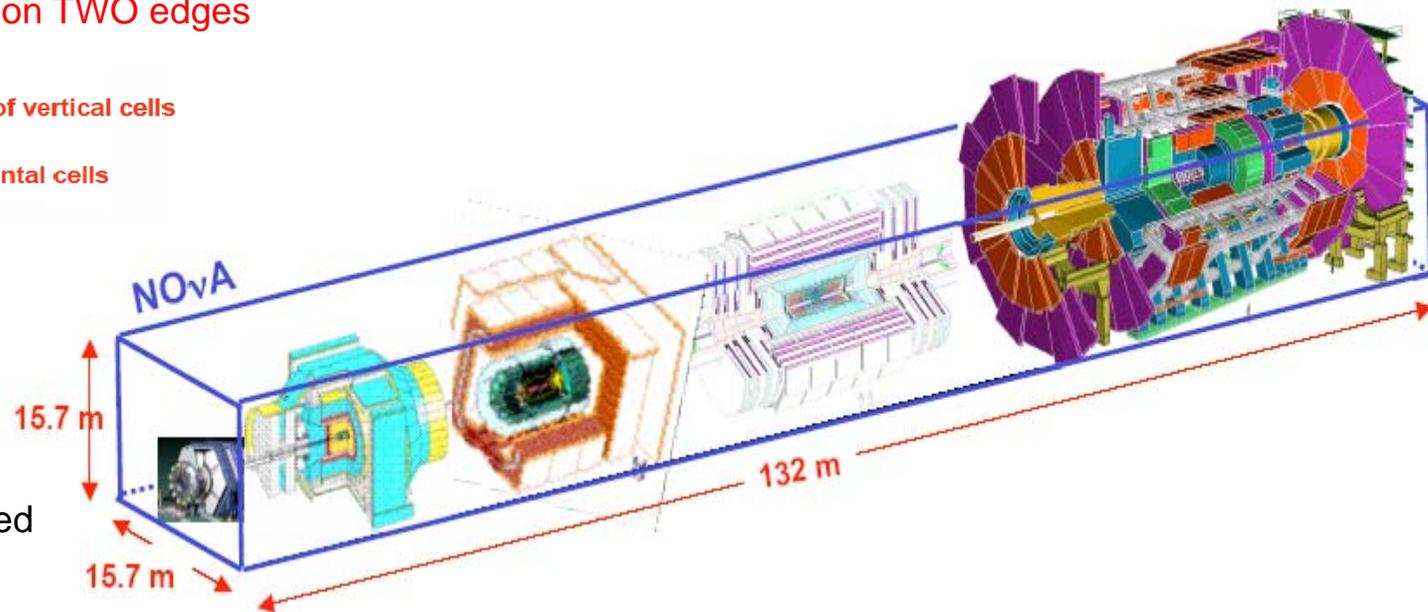
NO_vA

- most massive detector for $\nu_e \rightarrow \nu_\tau$ oscillation search
 - 'all' liquid scintillator (85% Sci, 15% PVC)



- construction & cost challenge:
 - scaling detector volume is not so trivial
 - want monolithic, manufacturable structures
 - seek scaling as surface rather than volume if possible

- size matters:
 - 30 ktons
 - big and massive as:
BaBar, CDF, D0,
CMS and ATLAS combined



HEGRA & Auger

air shower detectors:

- 1998: HEGRA – atmospheric Cherenkov telescope
- 2004: Auger – fluorescence & Cherenkov observatory

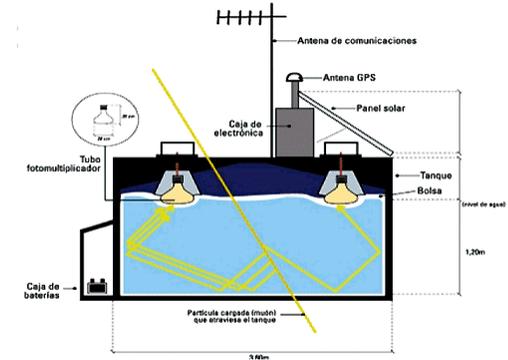
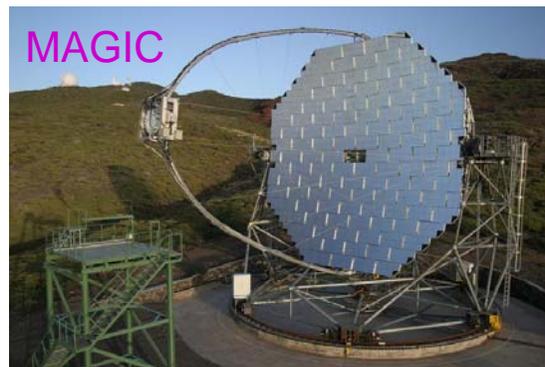
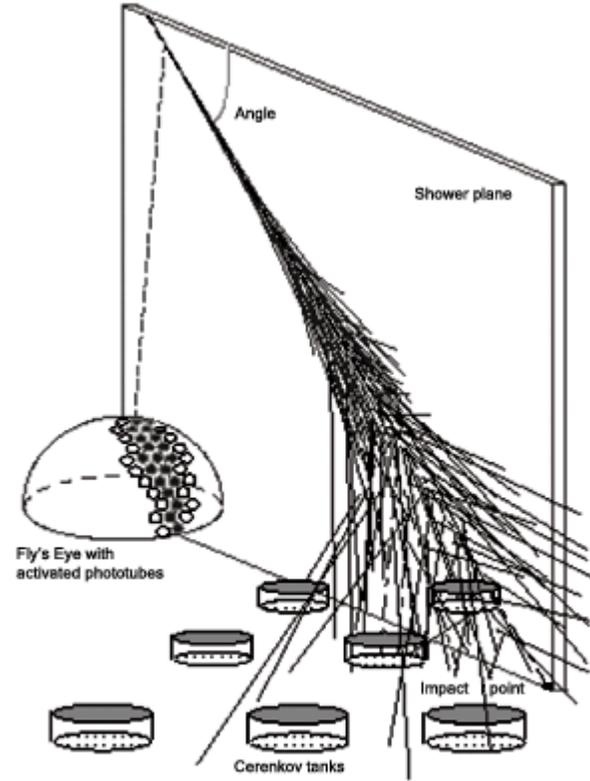
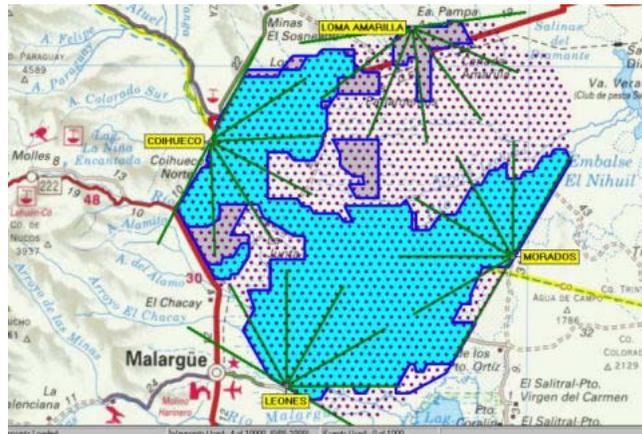
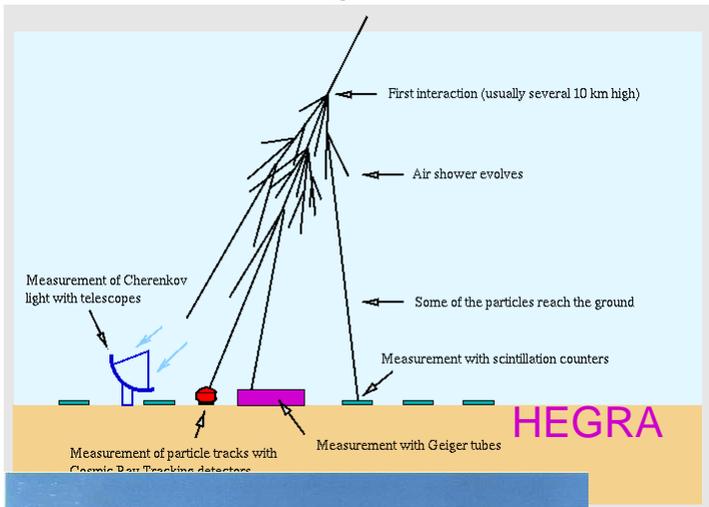
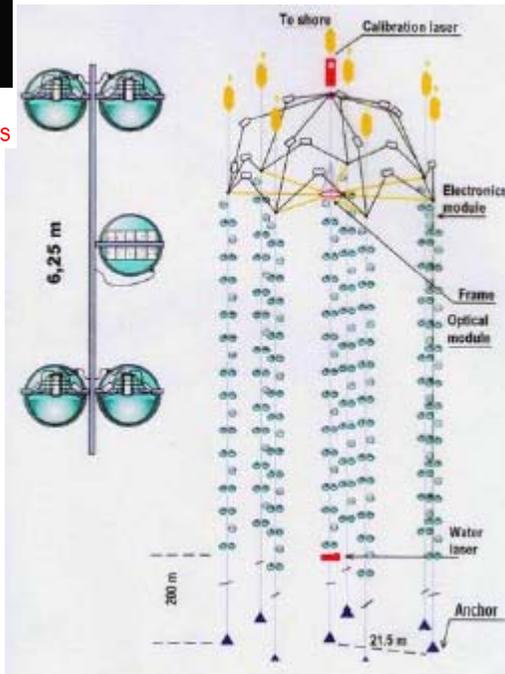
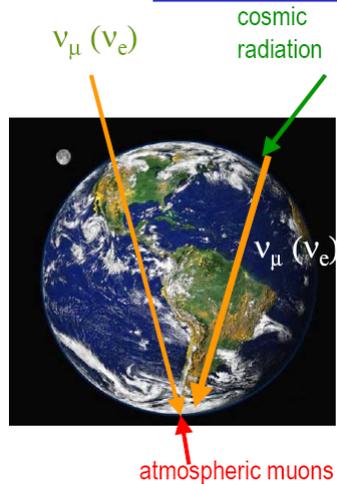


Figure 2. The water Cherenkov tank.

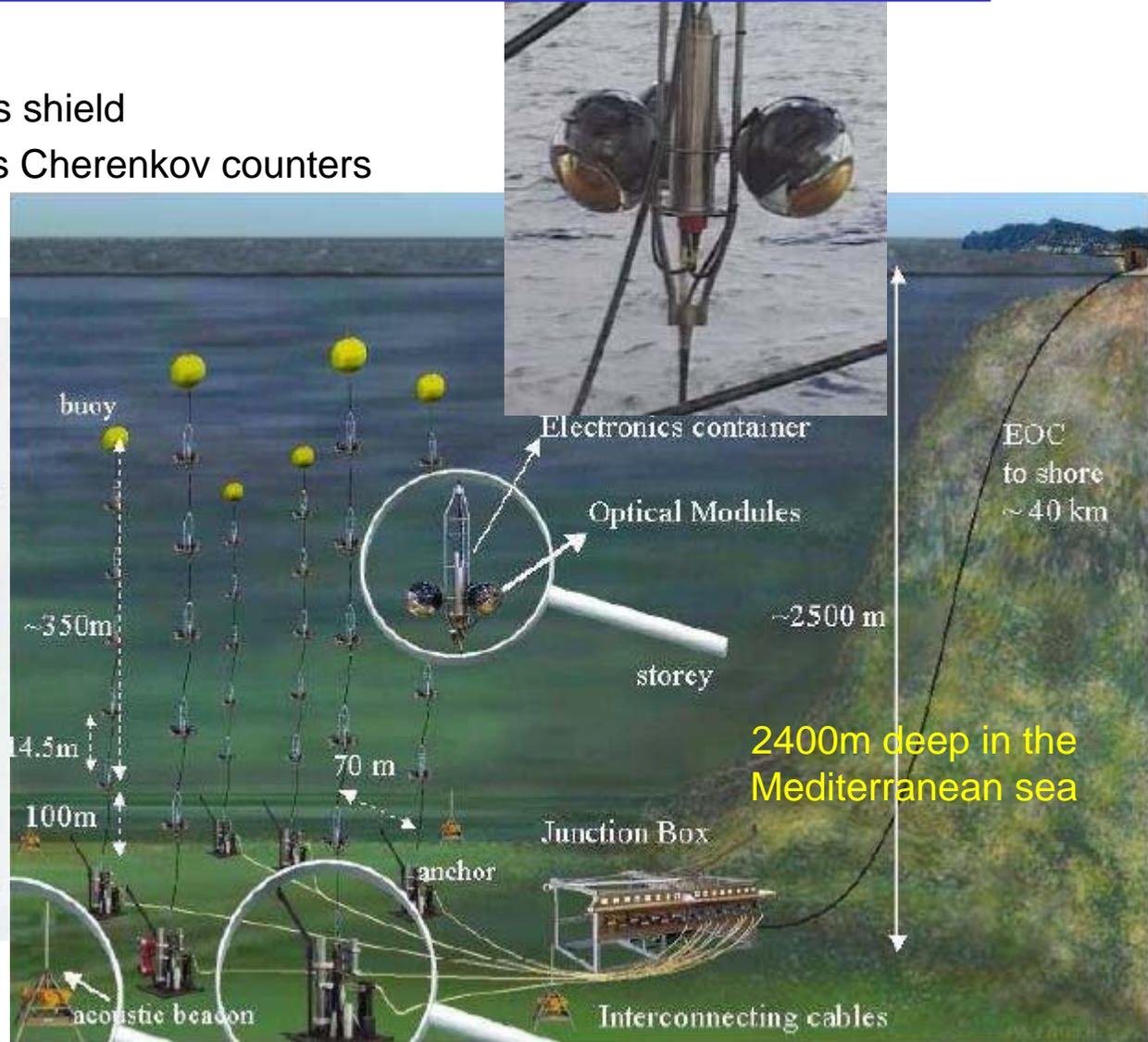
Baikal NT200, ANTARES & KM3Net

□ ν detectors:

- using earth as shield
- large volumes Cherenkov counters

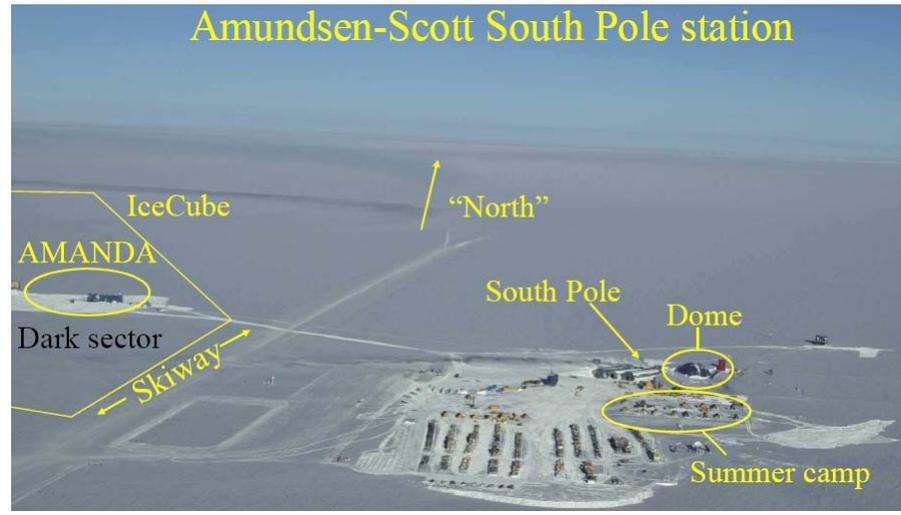
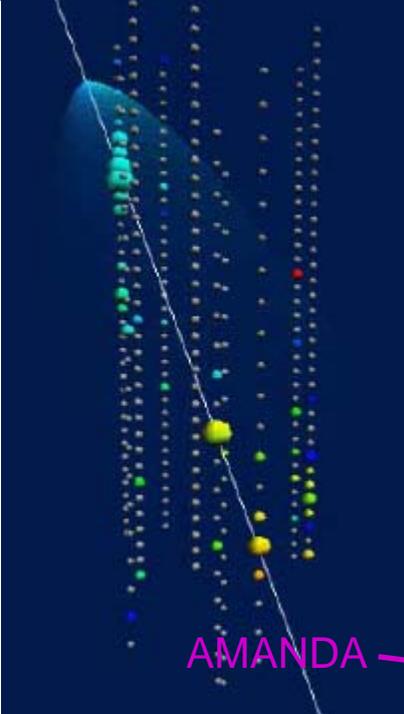
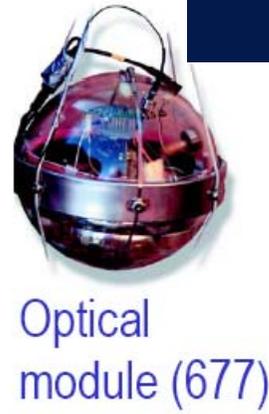
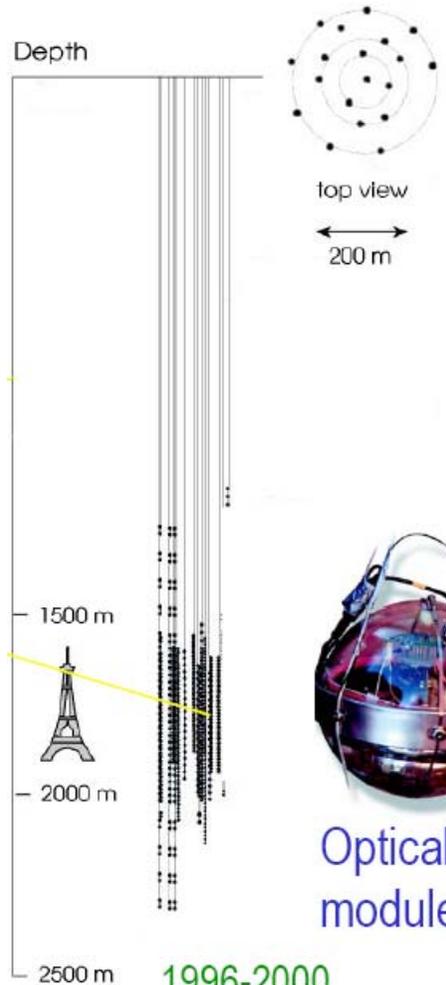


1100m deep in the Baikal lake

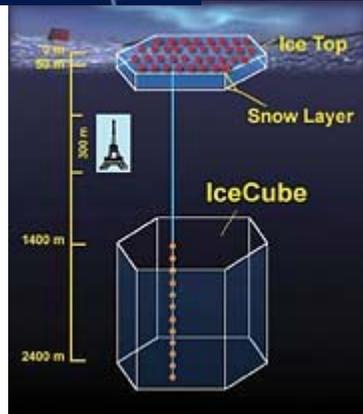
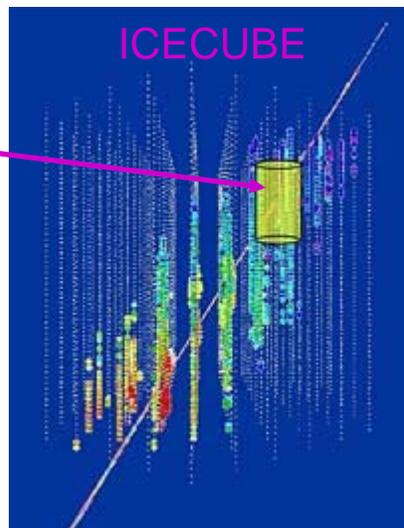


AMANDA & ICECUBE

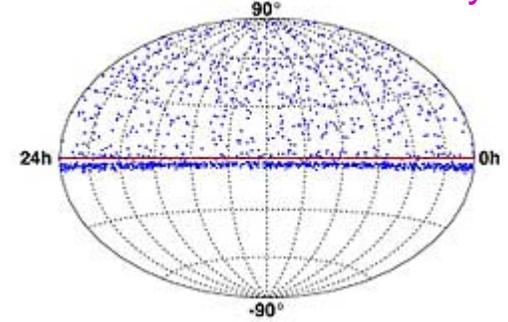
□ at south pole:



AMANDA



AMANDA neutrino sky



1996-2000
Particle Physics Detectors, 2010

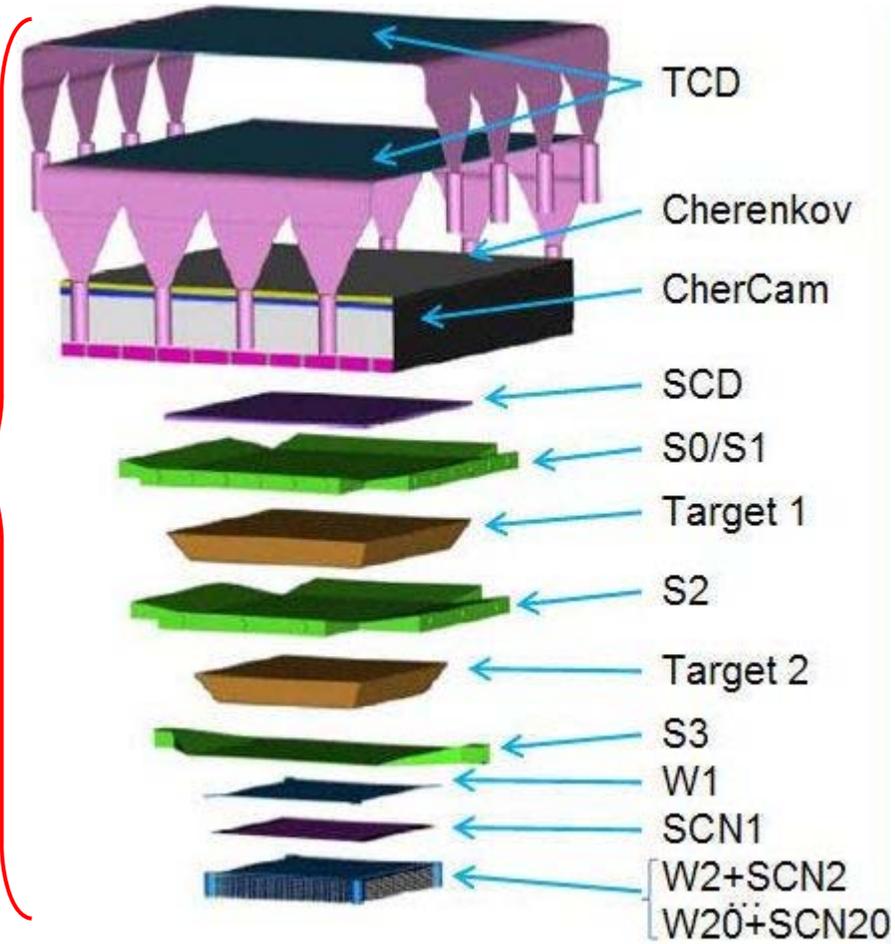
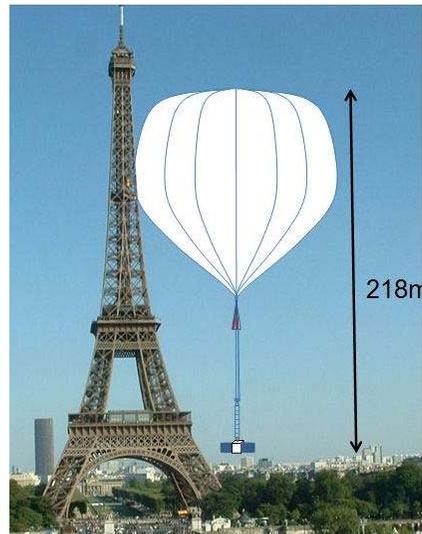
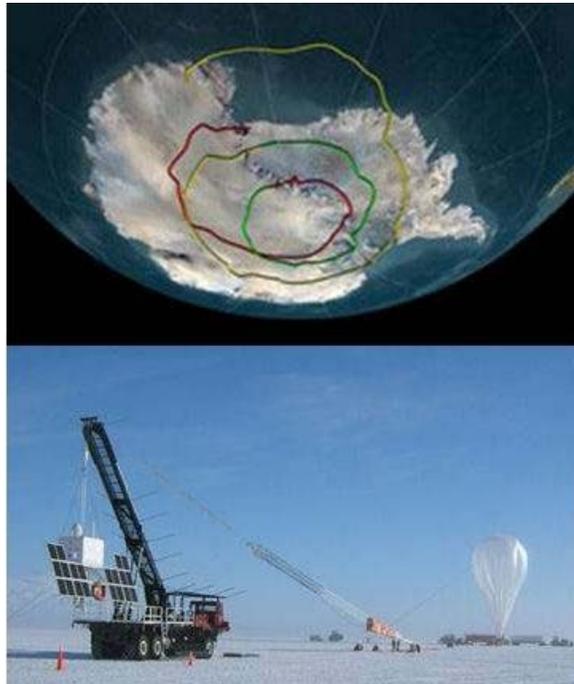
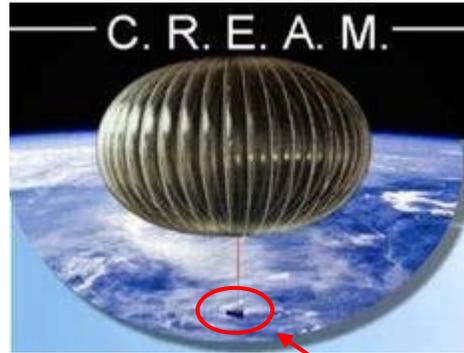
Stephan Eisenhardt

VI/43

Atmosphere-bound - CREAM

□ Cosmic Ray Energetics and Mass

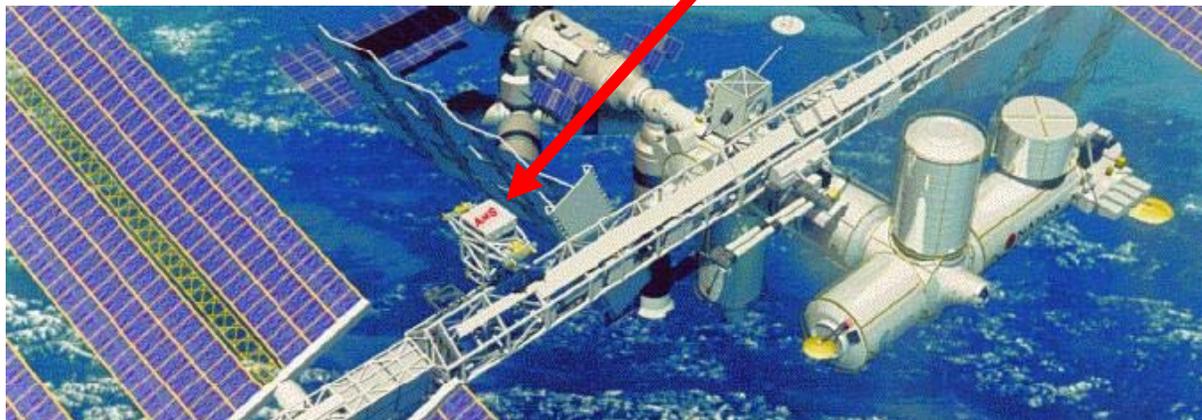
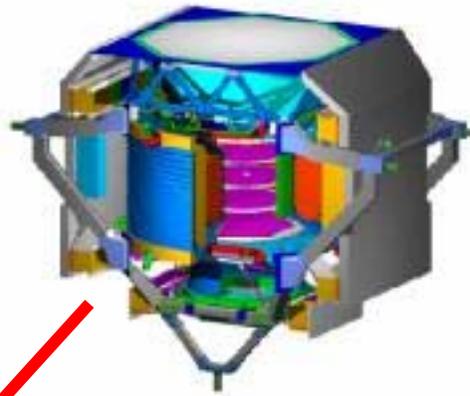
1st CREAM flight:
16.12.05-26.01.05



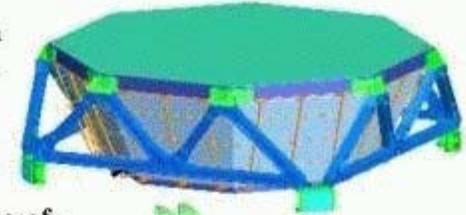
Space-bound - AMS

□ Alpha Magnetic Spectrometer:

- to be mounted on ISS
- launch: "2008"
- particle physics detector in space:
 - antimatter
 - gamma rays
 - cold dark matter
 - earth's particle environment
 - ...



Transition
Radiation
Detector



Upper Time-of-
Flight



Star tracker

Silicon
tracker

Super-
conducting
Magnet

Anti-coincidence
Counter

Lower Time-of-
Of-Flight

Ring-imaging
Cerenkov detector

Electromagnetic
calorimeter

