

Search for Light Dark Matter with Spherical Proportional Counters



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This project has received funding from the European Union's Horizon 2020 research and innovation programme under grant agreement no 841261 (DarkSphere)

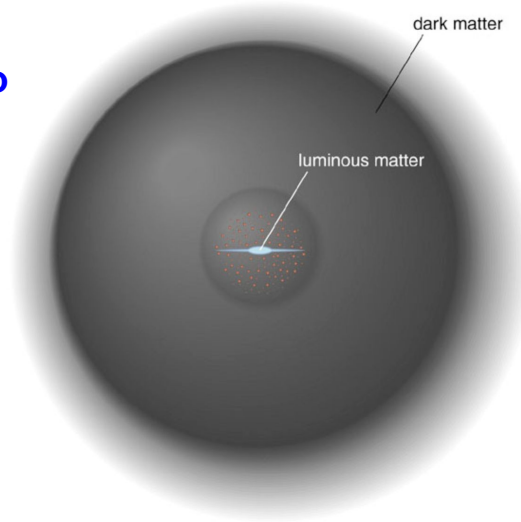


The dark matter conundrum

What should it be from astrophysical constraints:

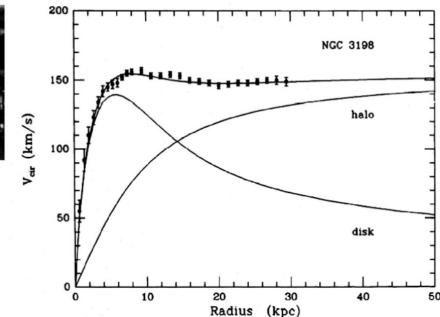
- Mostly “Cold”
- Non-Baryonic
- “Weakly” interacting
- $\Omega_{\text{DM}} = 0.265$
- Stable or $\tau_X \gg \tau_U$

No Standard Model particle matches the criteria



“In a spiral galaxy, the ratio of dark-to-light matter is about a factor of ten. That's probably a good number for the ratio of our ignorance-to-knowledge. We're out of kindergarten, but only in about third grade.”

Vera Rubin



Observations of Zwicky 87 years ago

Die Rotverschiebung von extragalaktischen Nebeln
von F. Zwicky.
(16. II. 33.)

"The Redshift of Extragalactic Nebulae",
published in German in Helvetica Physica
Acta in 1933



Dark matter detection

PHYSICAL REVIEW D

VOLUME 31, NUMBER 12

15 JUNE 1985

Detectability of certain dark-matter candidates

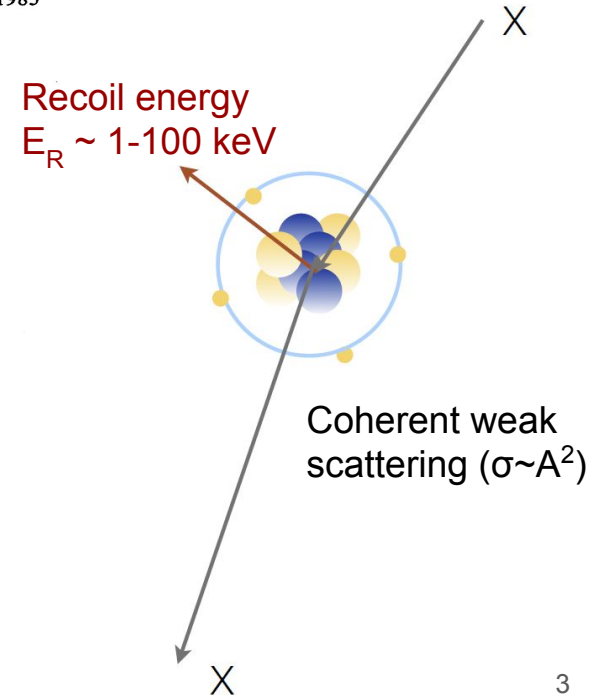
Mark W. Goodman and Edward Witten

Joseph Henry Laboratories, Princeton University, Princeton, New Jersey 08544

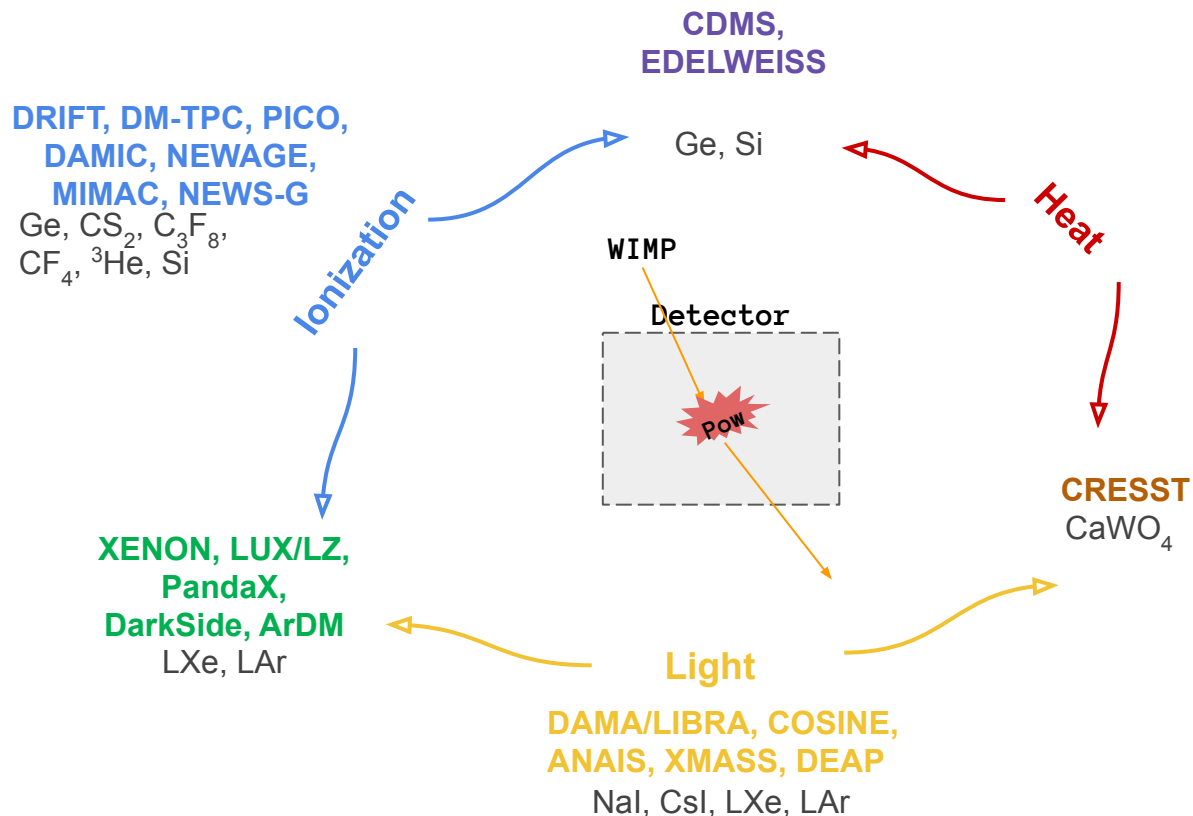
(Received 7 January 1985)

We consider the possibility that the neutral-current neutrino detector recently proposed by Drukier and Stodolsky could be used to detect some possible candidates for the dark matter in galactic halos. This may be feasible if the galactic halos are made of particles with coherent weak interactions and masses $1-10^6$ GeV; particles with spin-dependent interactions of typical weak strength and masses $1-10^2$ GeV; or strongly interacting particles of masses $1-10^{13}$ GeV.

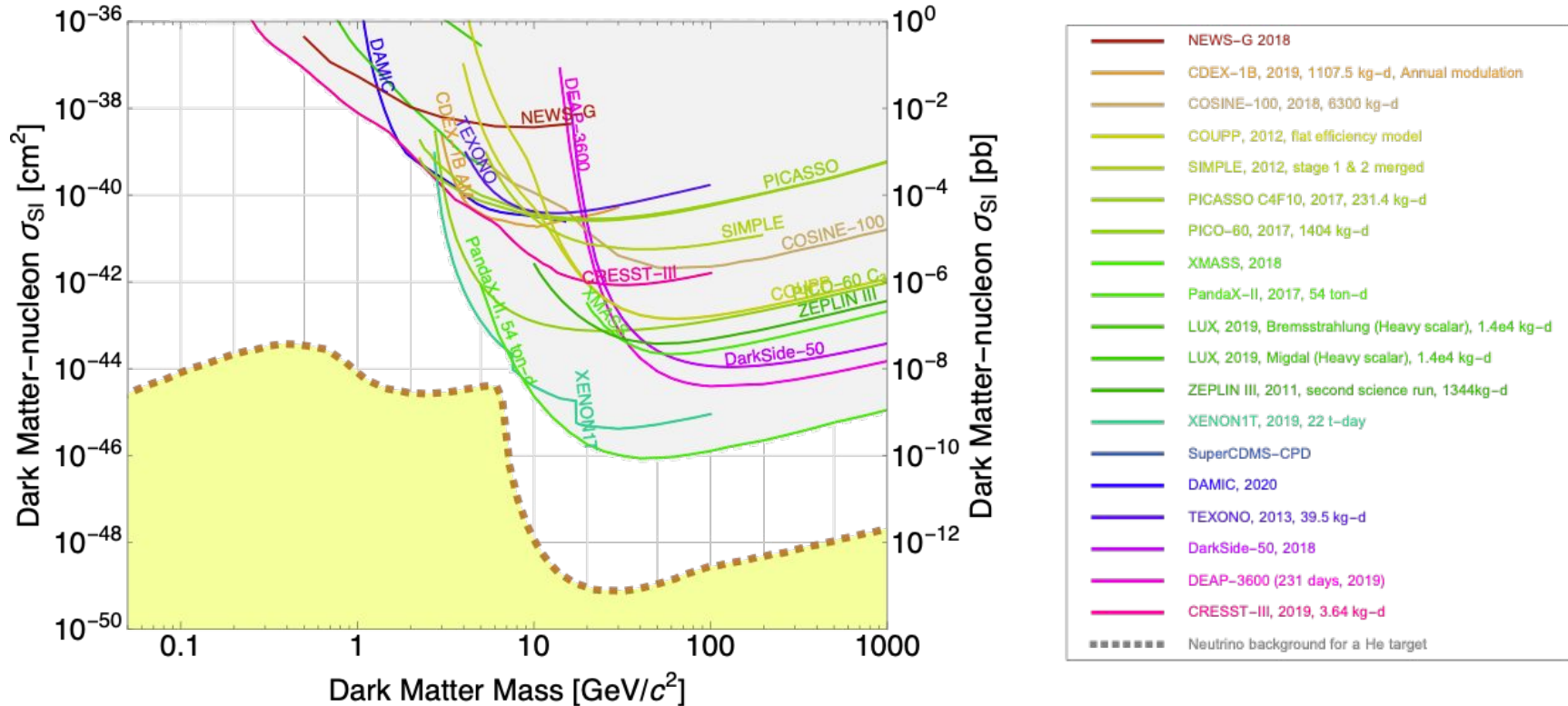
“WIMP miracle” \Rightarrow Relic abundance explained by a massive particle ($5 \text{ GeV}/c^2$ - few TeV/c^2) interacting through weak scale interaction with baryonic matter



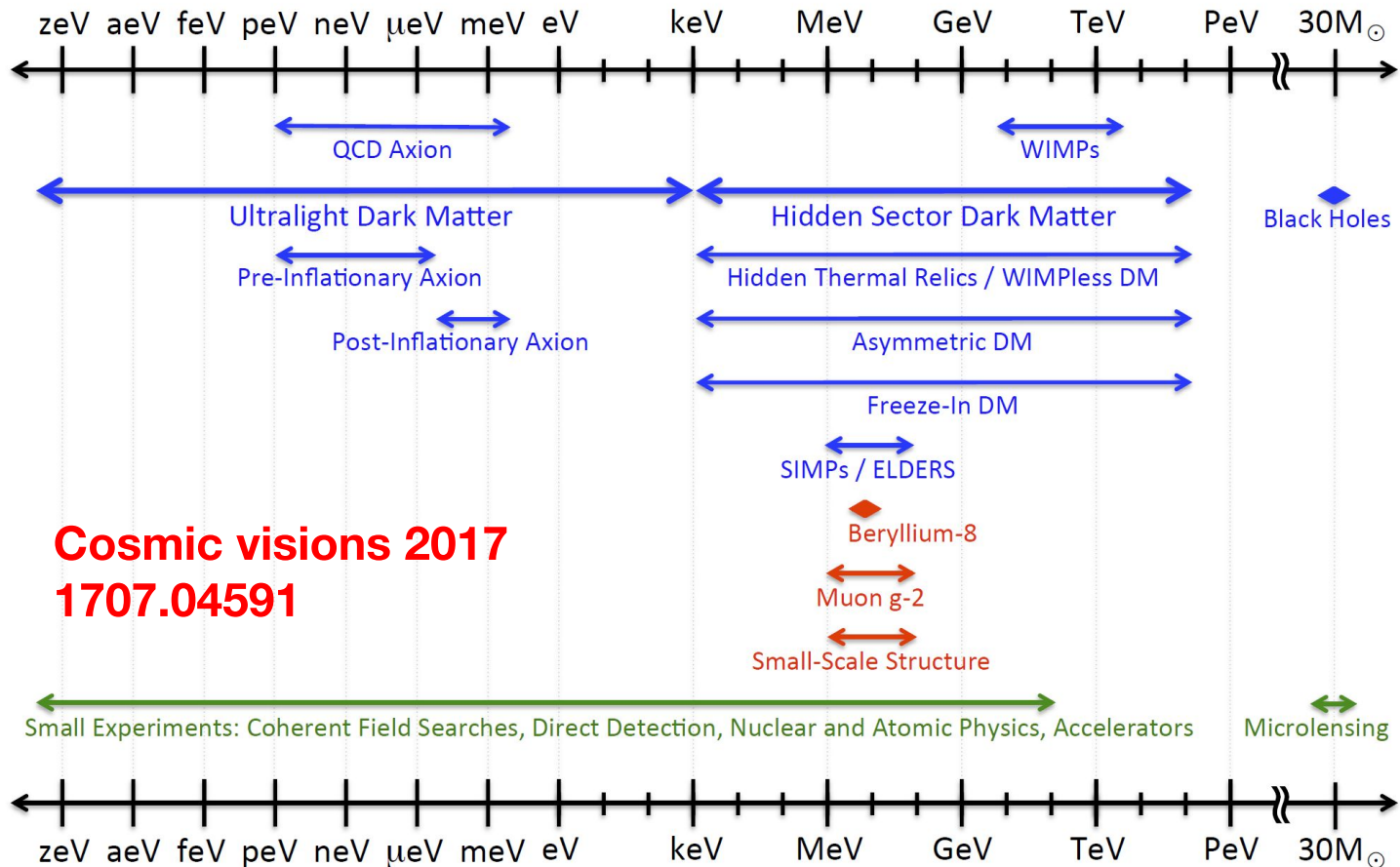
State of the art for dark matter detectors



Direct detection landscape

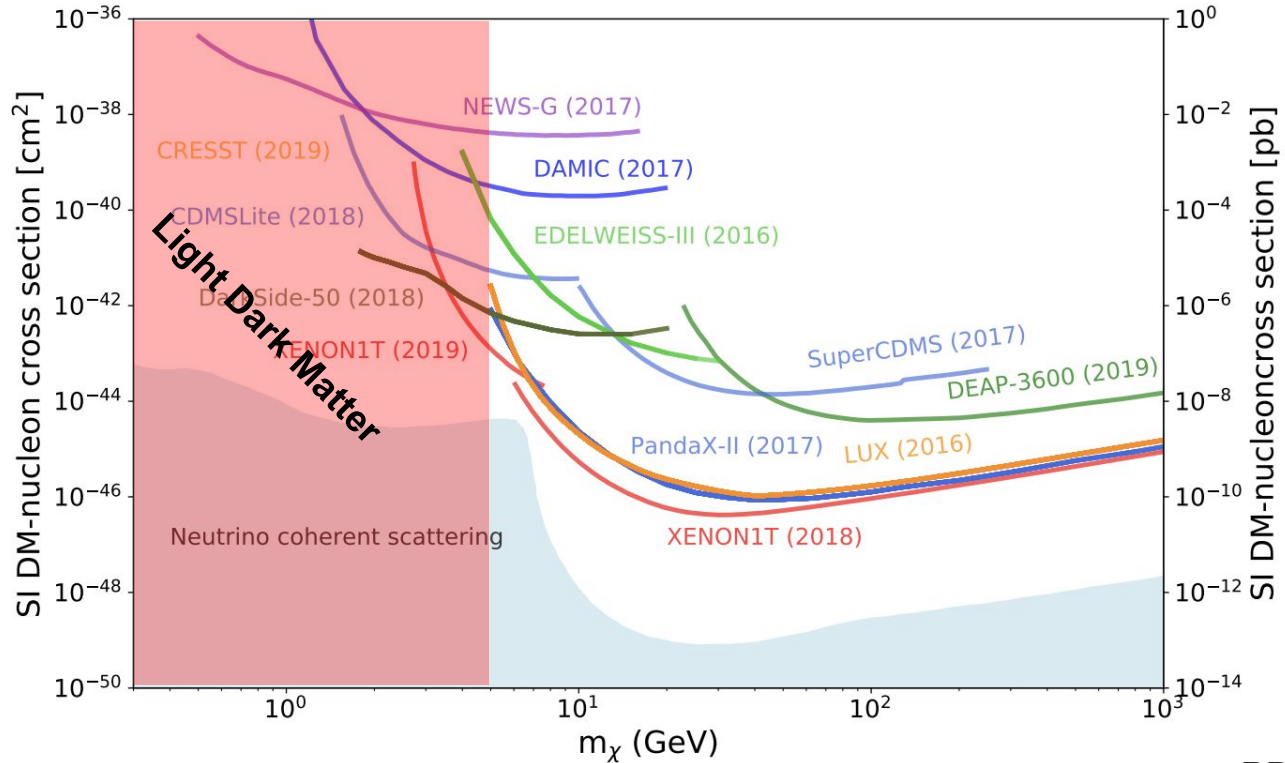


Dark Sector Candidates, Anomalies, and Search Techniques

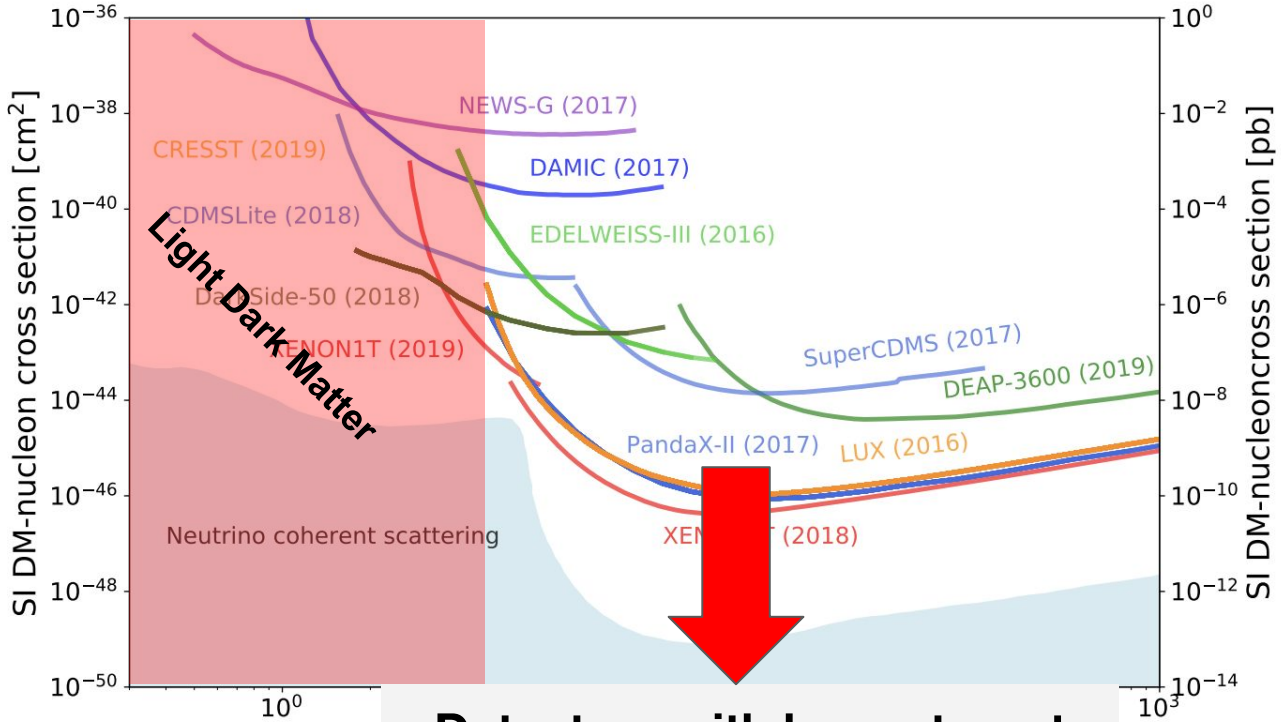


Cosmic visions 2017
1707.04591

Light Dark Matter (LDM) mass region



How does one search the LDM region?

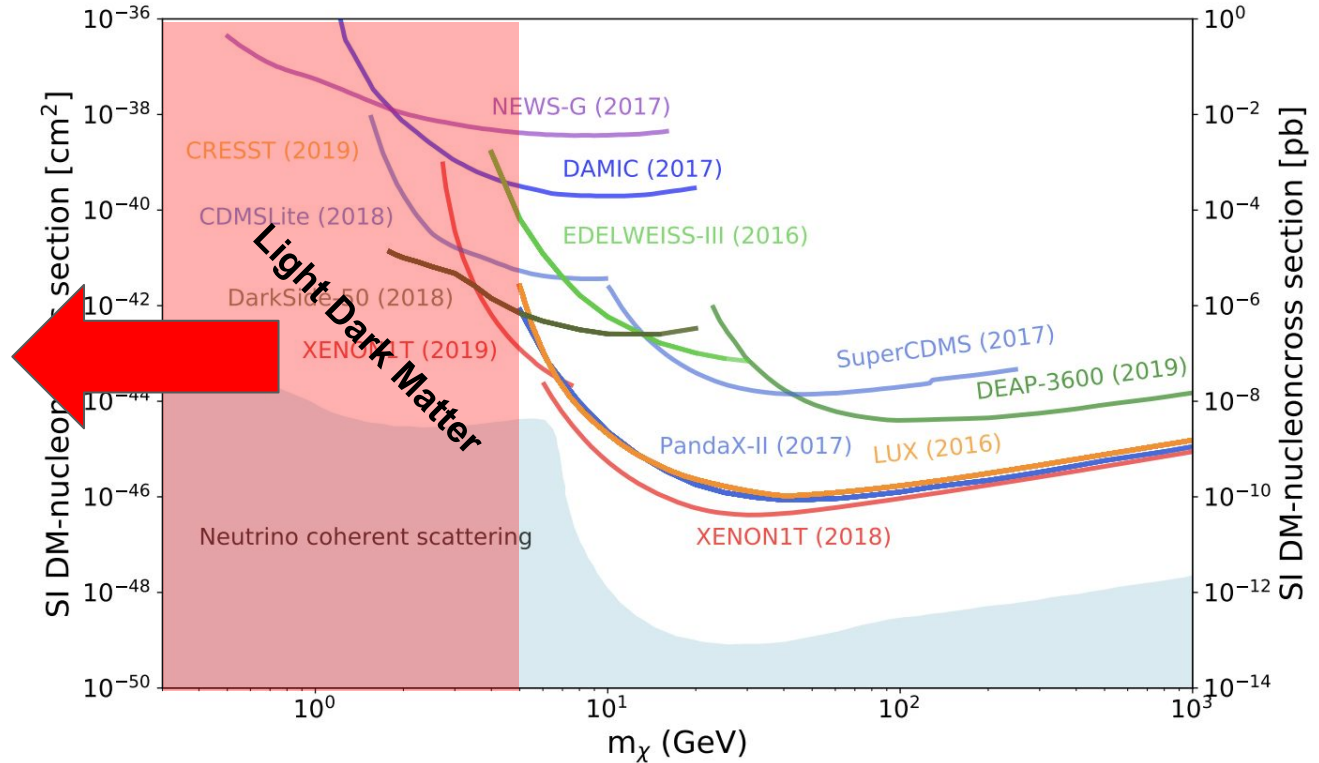


Detectors with larger target mass - exposure kg · days

Direct detection landscape

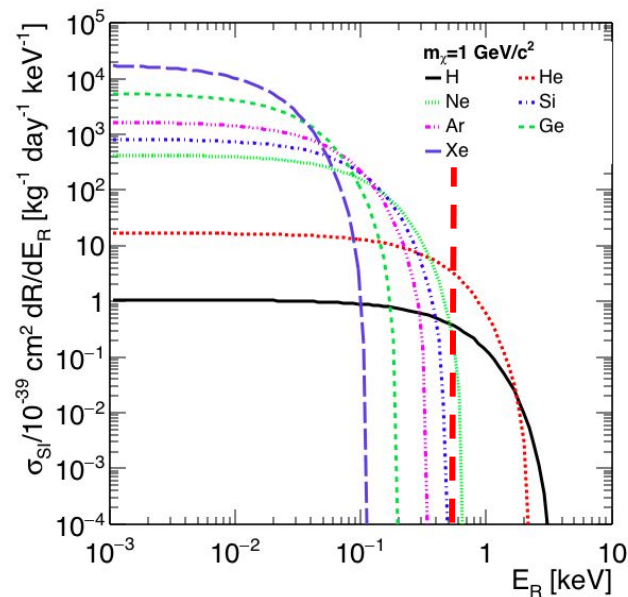
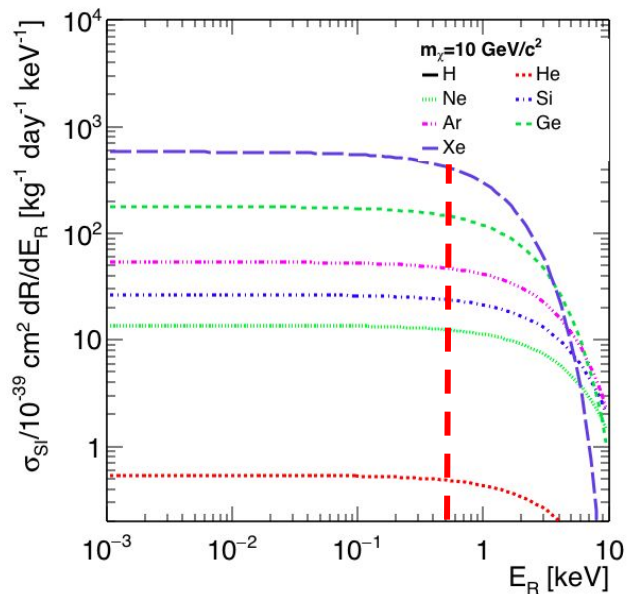
New detector technologies:

- Lower threshold
- Lighter targets



Light-DM detection particularities - A

Target kinematics



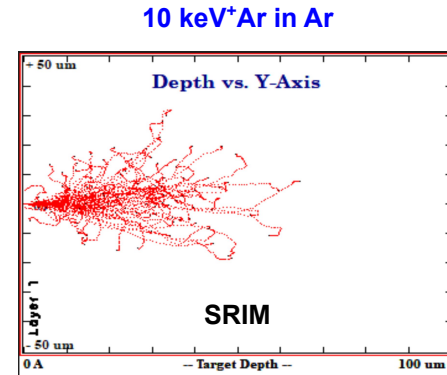
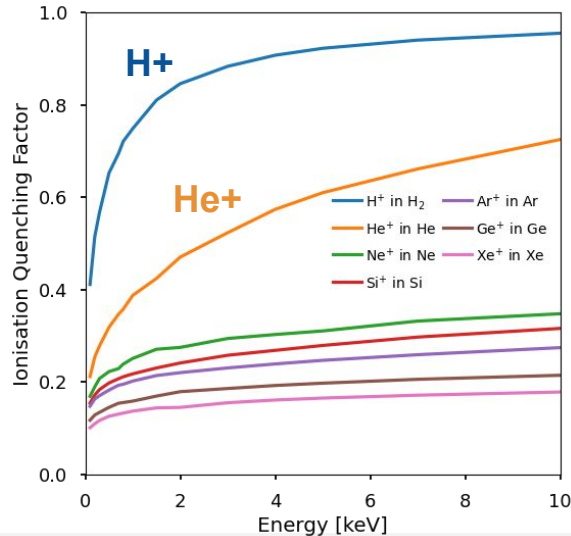
--- 500 eV

Light Projectile + Light target \Rightarrow Better kinematic match

Light-DM detection particularities - B

Ionization quenching

Quenching factor: fraction of ion kinetic energy dissipated in a medium in the form of ionization electrons and excitation of the atomic and quasi-molecular states.



Light Projectile + Light target \Rightarrow Less demanding detector threshold



Light Dark Matter searches with an innovative gaseous detector the Spherical Proportional Counter

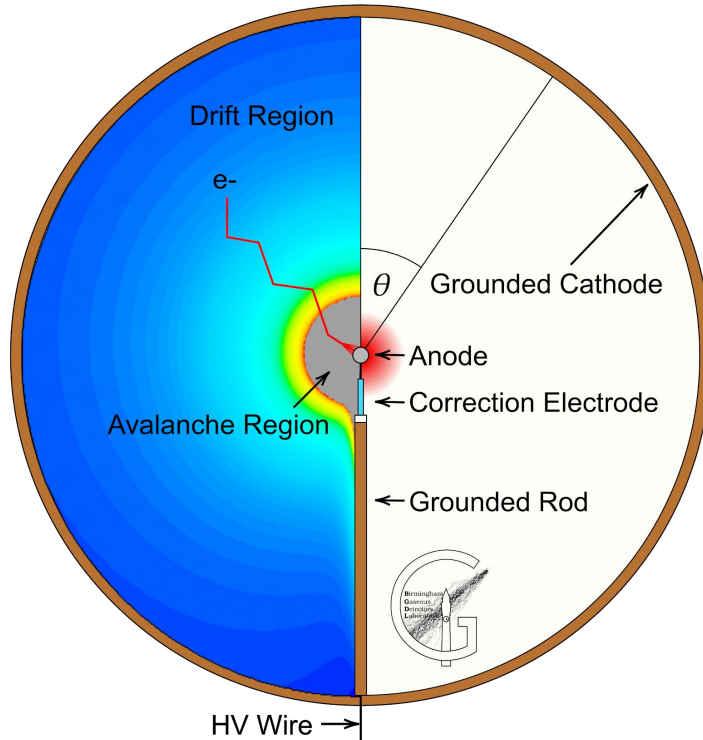


NEWS-G collaboration UK, France, Greece, Canada, US



The Spherical Proportional Counter (SPC)

[I.Giomataris et al., JINST, 2008, P09007](#)



Electric field

Strong radial dependence

$$E(r) = \frac{V_0}{r^2} \frac{r_A r_C}{r_C - r_A} \approx \frac{V_0}{r^2} r_A$$

$r_A = \text{anode radius}$

$r_C = \text{cathode radius}$

Detector volume naturally divided in:

- Drift region
- Amplification region
- Simple design
- Single readout

Spherical Proportional Counter (SPC)

The “birth” of a detector

Old LEP RF cavities



Spherical gaseous detectors



*In the picture:
I.Giomataris, G.Charpak*

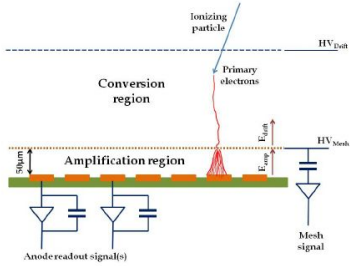


Advantage of spherical geometry

Large detectors - Low threshold

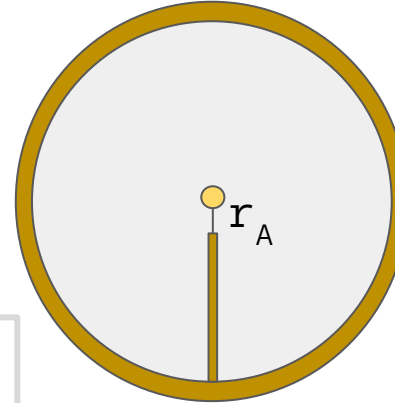
Capacities for a 1 m³ detector in different geometries

Parallel Plate



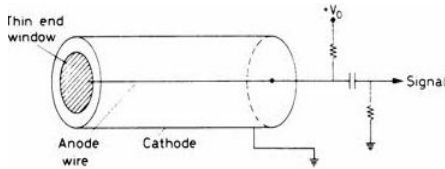
$$C = \epsilon_0 \frac{S}{d} \approx 3500 \text{ pF}$$

Spherical



$$C \approx 4\pi\epsilon r_A \approx 1.5 \text{ pF}$$

Cylindrical



$$C = 2\pi\epsilon_0 \frac{L}{\ln(b/a)} \approx 115 \text{ pF}$$

Lower Capacitance → Lower Electronic Noise → Lower Threshold

Advantage of spherical geometry

Construction with radiopure materials



Advantages of the spherical geometry

- Lowest surface to volume ratio
- Sustains higher pressure
- Robustness (anode \varnothing 1 mm - 6.3 mm)

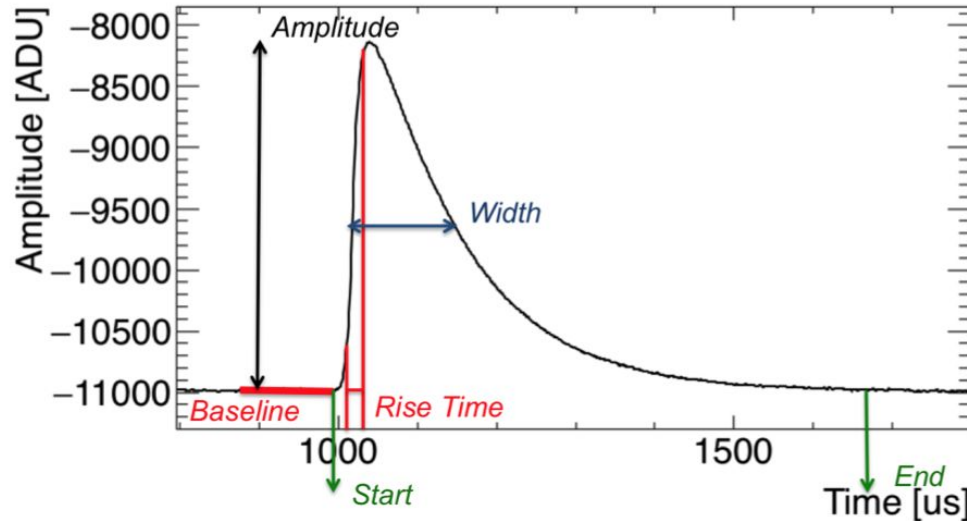
Built solely by radiopure materials

- Vessel made of Cu (~tens of kg)
- Rod made of Cu (~hundreds of gr)
- All the rest less **< 1 g**

Induced Pulses

Pulse Shape Analysis (PSA) parameters

Long Tail Pulse



Rise time & Width \propto Drift time dispersion

Basic Parameters

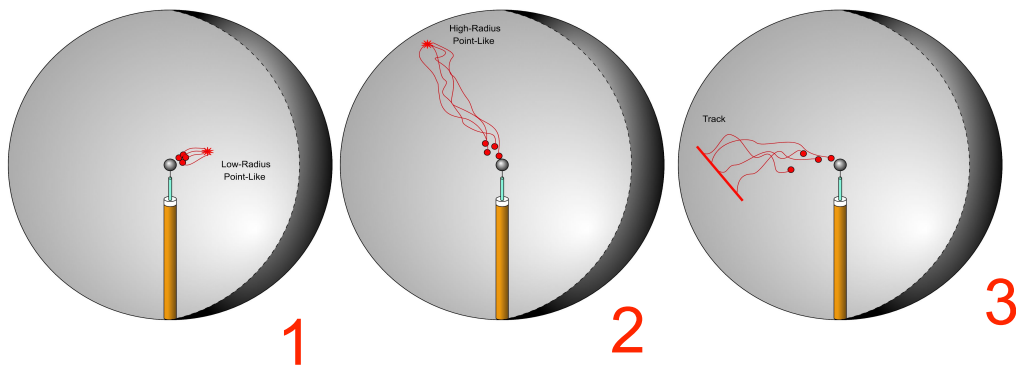
- Baseline
- Noise
- Amplitude (Pulse Height)
- Rise time
- Width
- Integral
- Number of peaks

A lot of
information in
the pulse shape

Pulse-Shape Discrimination

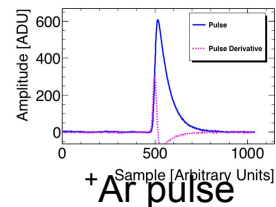
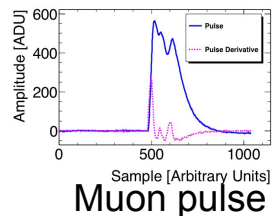
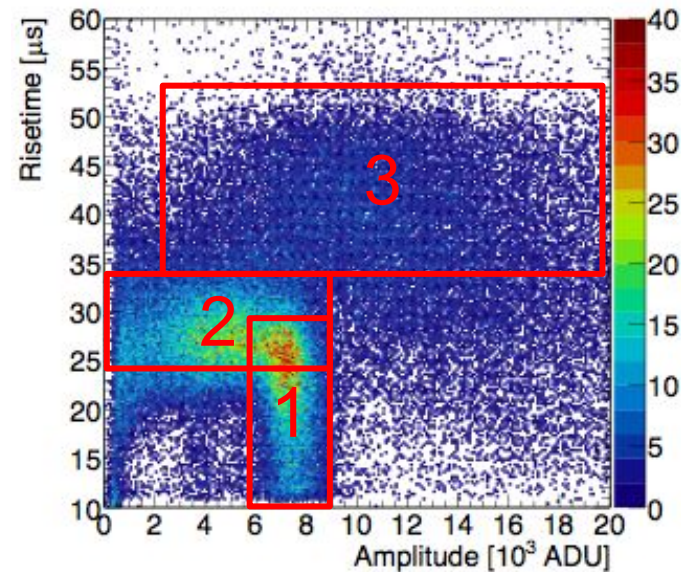
Rise time (~charge collection time) selections to:

- Distinguish point-like versus extended ionisations
- Fiducialise detector
 - Majority of background from cathode material
 - Can select against near-cathode events
- Reliant on homogeneous electric field and high electric field at large radii (for charge collection)



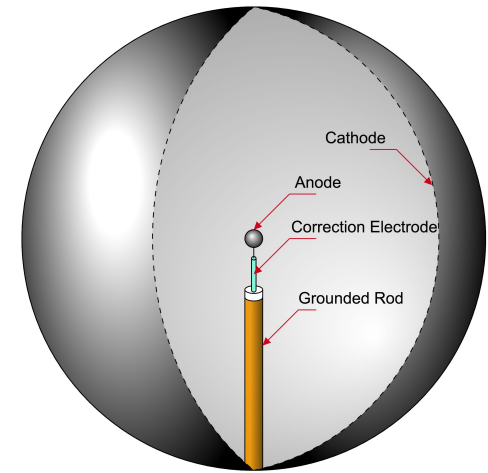
- (1) X-rays in volume,
- (2) X-rays near cathode
- (3) Cosmic Muons

ø30 cm detector
1.3 bar
He:Ar:CH₄
(51.7:46%:2.3%)
Fe-55 Source
Inside



Detector features

- **Large volume read out with a small number of channels**
- **Single electron threshold due to:**
 - **Low capacitance**
 - **High gain**
- **Radio-pure construction**
- **Background rejection handles**
- **Flexible operation**
 - **Swappable gases-targets**
 - **Variable pressure choice**

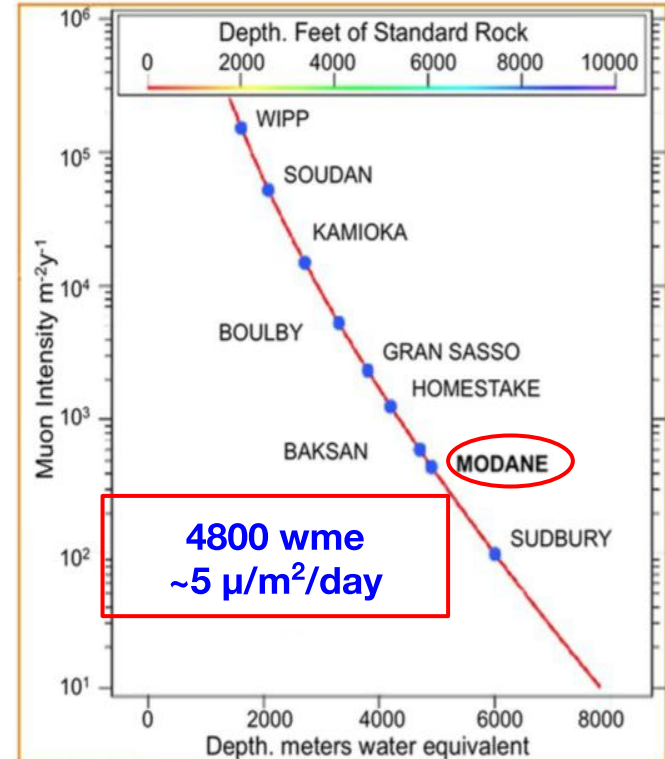
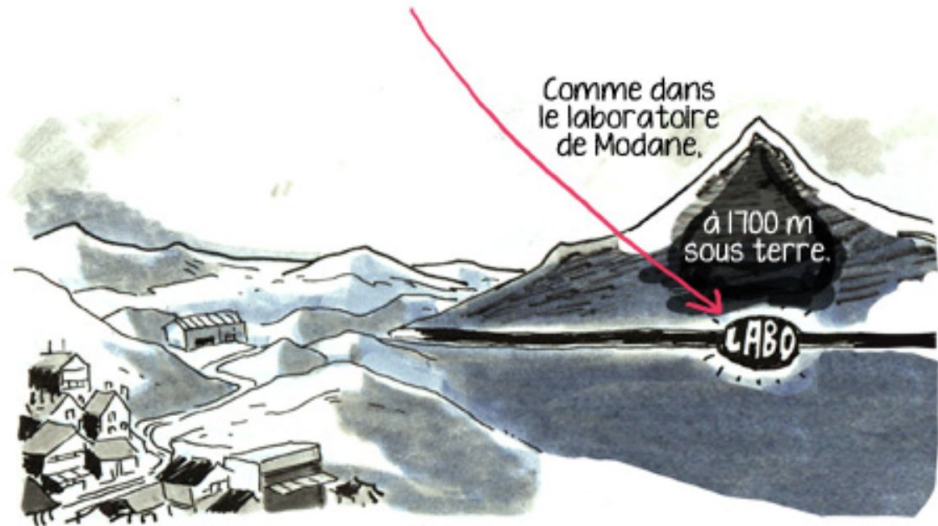


Applications include:

- Dark matter searches
- $2\beta 0\nu$ decay searches
- CEvNS physics
- Neutron spectroscopy (potentially useful for proton therapy)

NEWS-G at Modane

IL faudrait que
ce matériau,
(le germanium)
soit protégé au maximum
de toutes les sortes
de radiations naturelles ...



NEWS-G at Modane

SEDINE detector

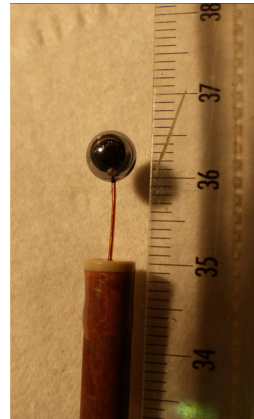
Vessel

Ø 60cm copper



Sensor

Ø 6.3mm Si

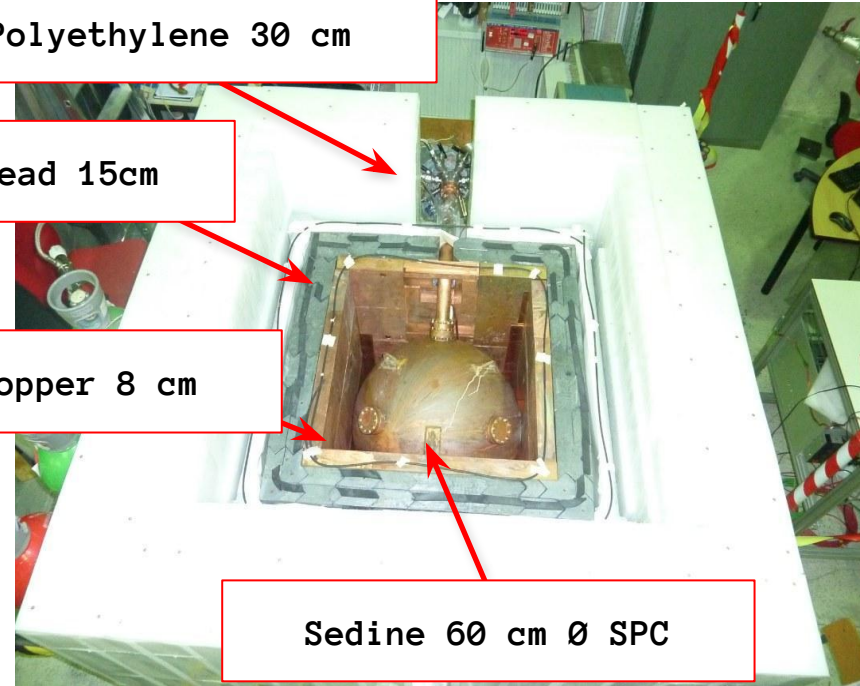


Polyethylene 30 cm

Lead 15cm

Copper 8 cm

Sedine 60 cm Ø SPC

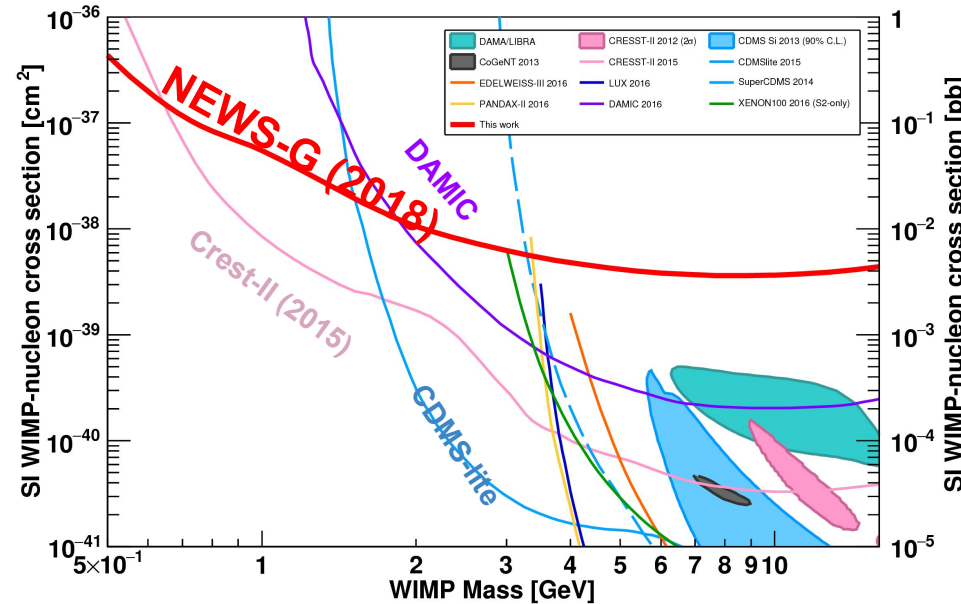


Gas Mixture: Ne+0.7%CH₄ at 3.1 bar (280 g)

Exposure: 9.6 kg*days (34.1 live-days x 0.28 kg)

First results of NEWS-G with SEDINE (2018)

[NEWS-G collaboration, Astropart. Phys. 97, 54 \(2018\), doi: 10.1016/j.astropartphys.2017.10.009](https://doi.org/10.1016/j.astropartphys.2017.10.009)



Astroparticle Physics 97 (2018) 54-62

Contents lists available at ScienceDirect

Astroparticle Physics

journal homepage: www.elsevier.com/locate/astropartphys




First results from the NEWS-G direct dark matter search experiment at the LSM



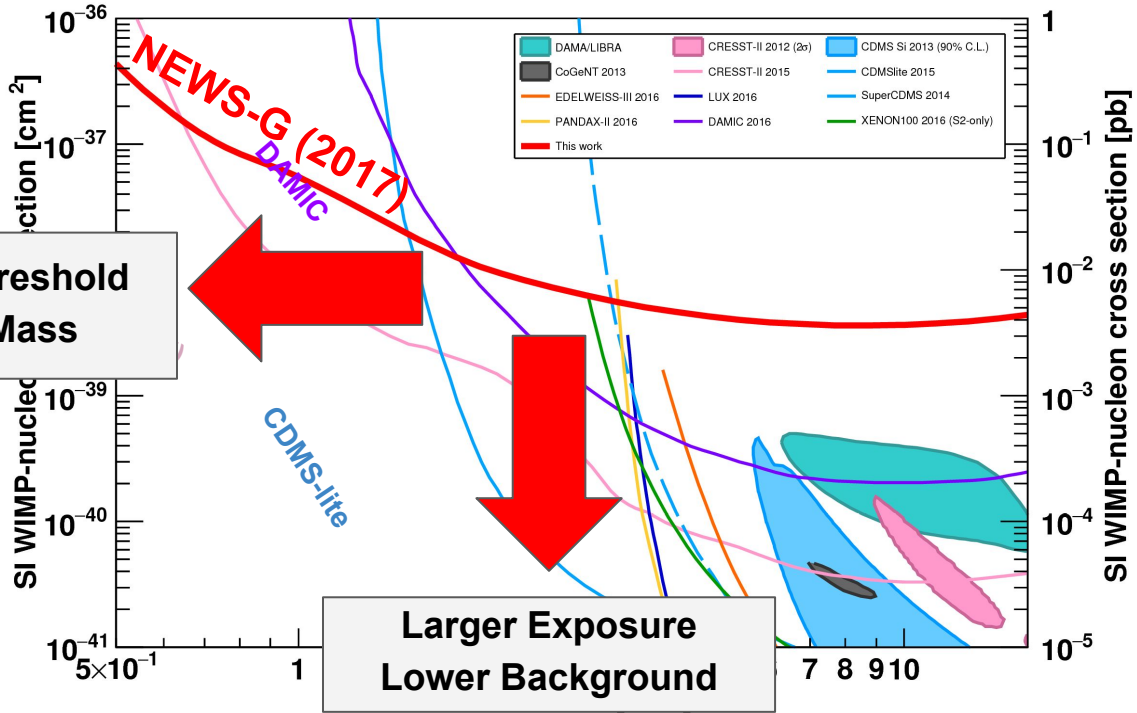
Q. Arnaud^{a,*}, D. Asner^b, J.-P. Bard^c, A. Brossard^{a,c}, B. Cai^d, M. Chapellier^a, M. Clark^a, E.C. Corcoran^d, T. Dandl^e, A. Dasgheibi-Fard^f, K. Dering^g, P. Di Stefano^h, D. Durnford^g, G. Gerbier^g, I. Giomataris^g, P. Gorel^g, M. Gros^g, O. Guillaudin^h, E.W. Hoppe^h, A. Kamaha^g, I. Katsioulas^g, D.G. Kelly^g, R.D. Martin^g, J. McDonald^g, J.-F. Muraz^h, J.-P. Mols^g, X.-F. Navick^g, T. Papaevangelou^g, F. Piquemal^g, S. Roth^{a,d}, D. Santos^h, I. Savvidis^g, A. Ulrich^g, F. Vazquez de Sola Fernandez^g, M. Zampaolo^g

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^fLSM, CNRS/IN2P3, Université Grenoble-Alpes, Modane, France
^gSMOLAB, Lively, Ontario, P3Y 1N2, Canada
^hLPC, Université Grenoble-Alpes, CNRS/IN2P3, Grenoble, France
ⁱDepartment of Physics, Aristotle University of Thessaloniki, GR-55124 Thessaloniki, Greece

Exclusion at 90% confidence level (C.L.) of cross-sections above $4.4 \cdot 10^{-37} \text{ cm}^2$ for a **0.5 GeV/c² WIMP**

Limit set on spin independent WIMP coupling with standard assumptions on WIMP velocities, escape velocity and with quenching factor of Neon nuclear recoils in Neon calculated from SRIM

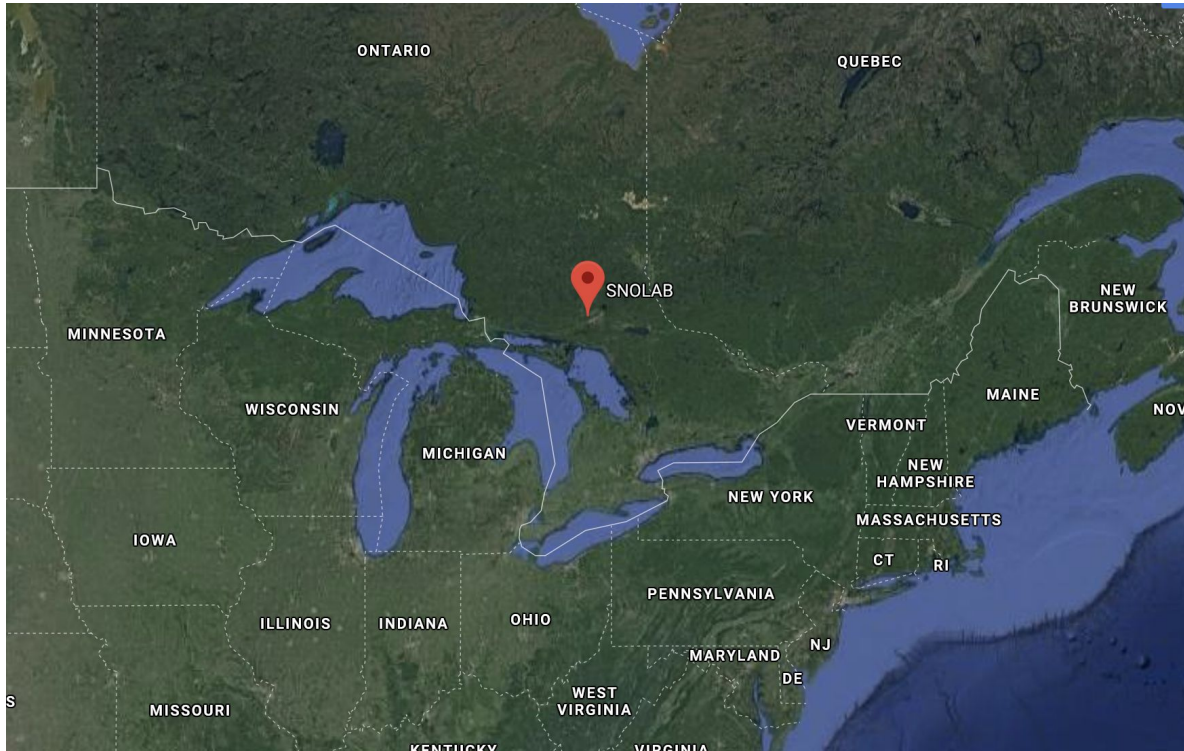
NEWS-G moving forward



Lower Energy Threshold
Lower Target Mass

Larger Exposure
Lower Background

NEWS-G at SNOLAB

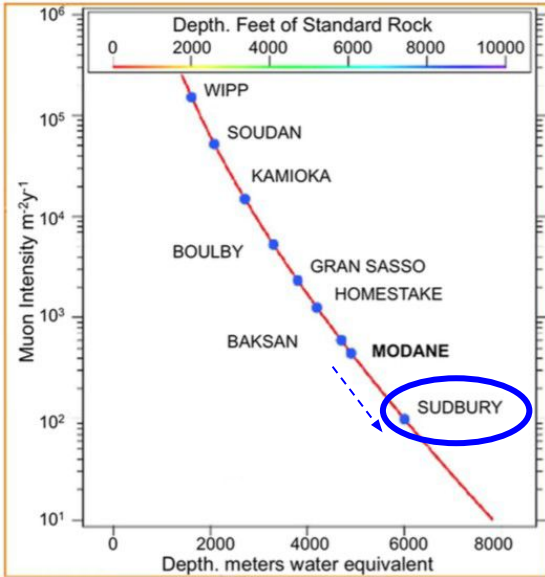
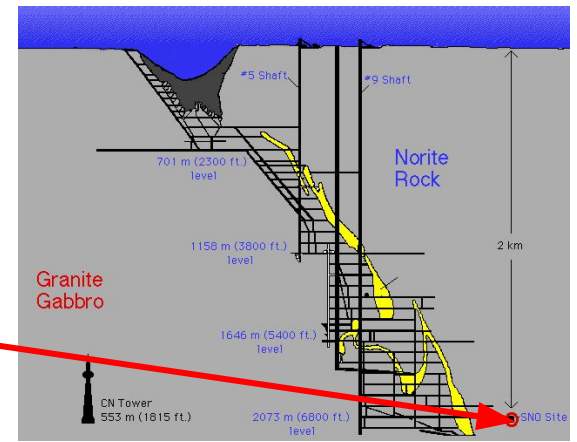


NEWS-G at SNOLAB

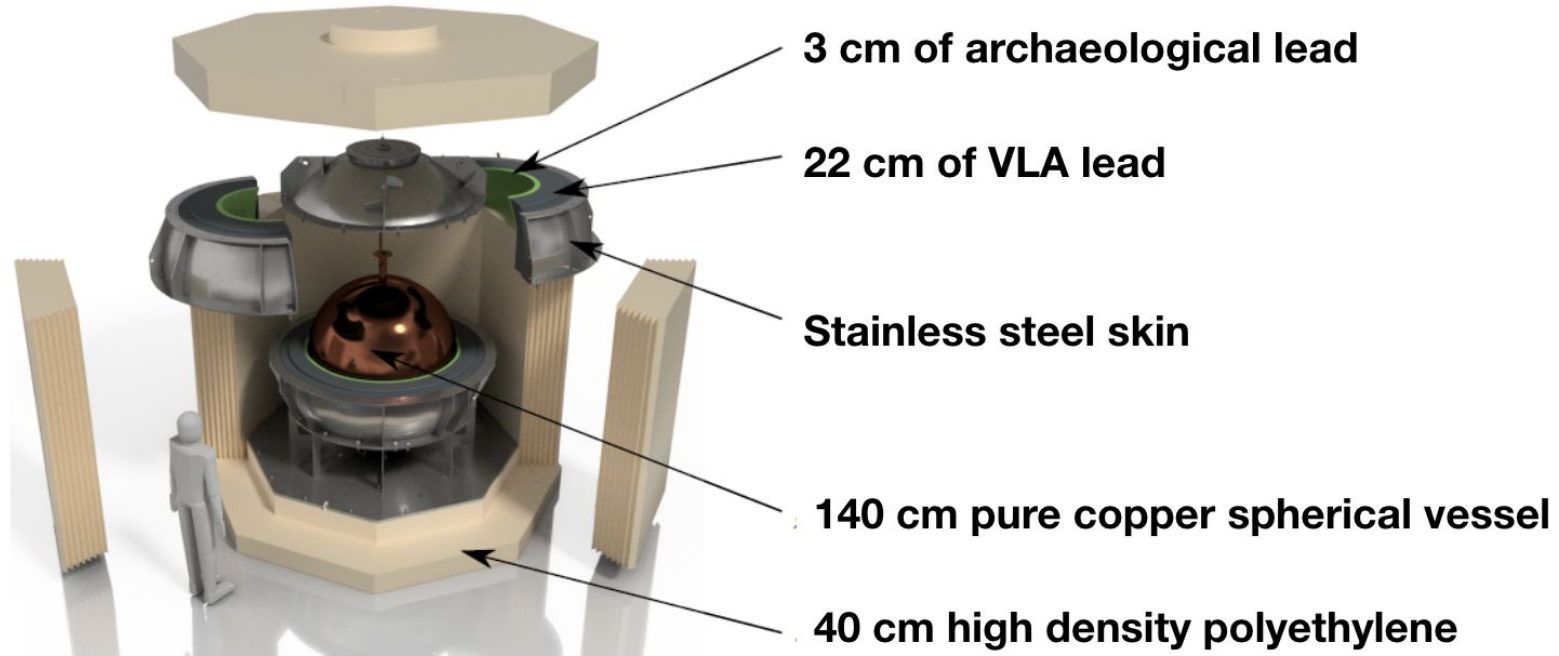
The underground laboratory in the Sudbury, Canada

Deeper underground
0.25 $\mu\text{m}^2/\text{day}$
~8x lower μ flux than LSM

Practically, at 2 km is the deepest clean room in the world

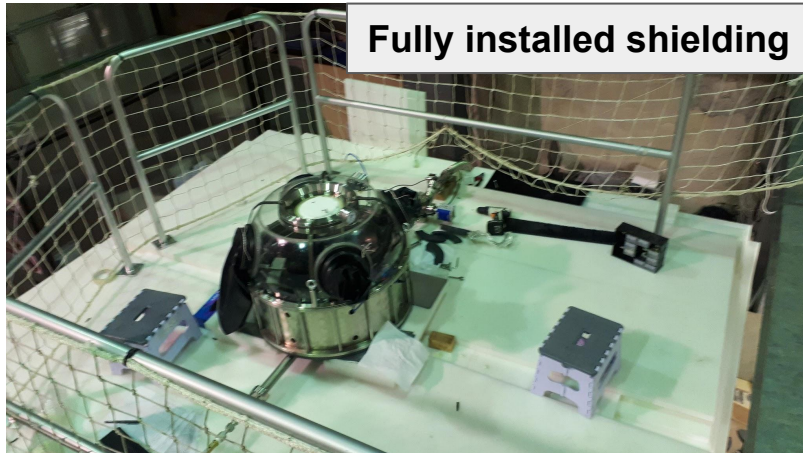


The NEWS-G detector - SNOGLOBE



SNOGlobe at LSM

- Detector already commissioned at LSM
 - Assembled and operated for a month
 - Gases used Ne/CH₄ (1 bar), CH₄ (135 mbar)
 - Sensor and electronics performance tested
 - Backgrounds under study



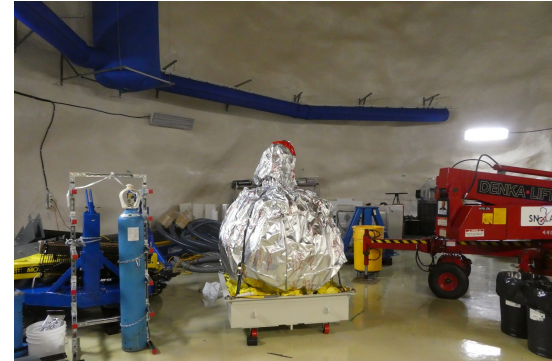
Status of installation at SNOLAB



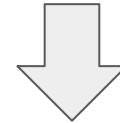
Shipment December 2019



Moving underground



Placement



Flushed with N₂



Status of installation at SNOLAB

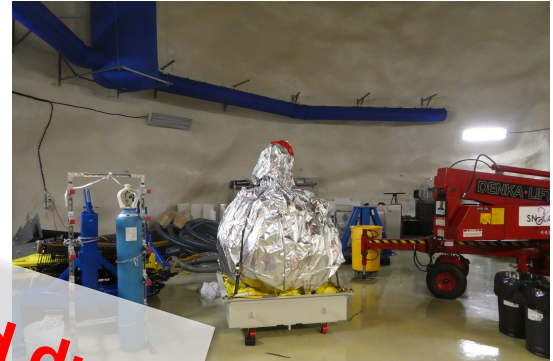


Shipment December 2019

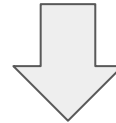
**Installation work halted due to COVID-19 lockdown
Restarted in Sep 2020**



Moving underground

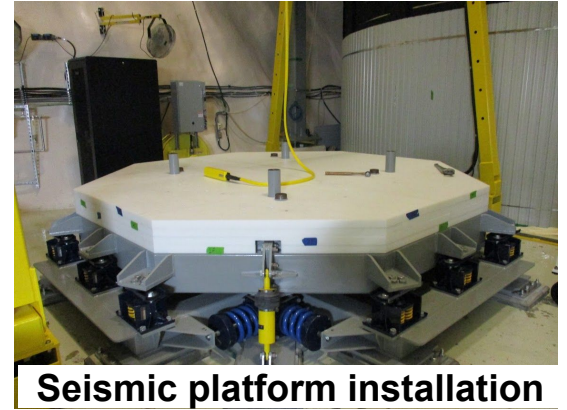


Placement



Flushing with N₂

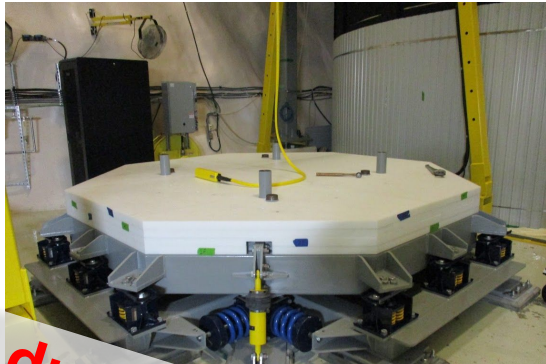
Status of installation at SNOLAB



Status of installation at SNOLAB



Unwrapped and baked Sep 2020



mic platform installation

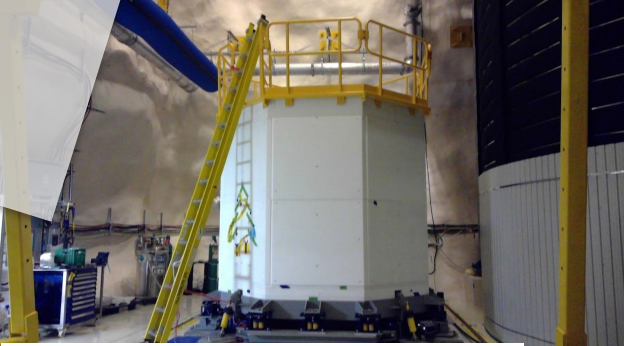
**Installation work halted due to COVID-19 lockdown
Expected to restart in February 2021**



SPC inserted in Lead shield

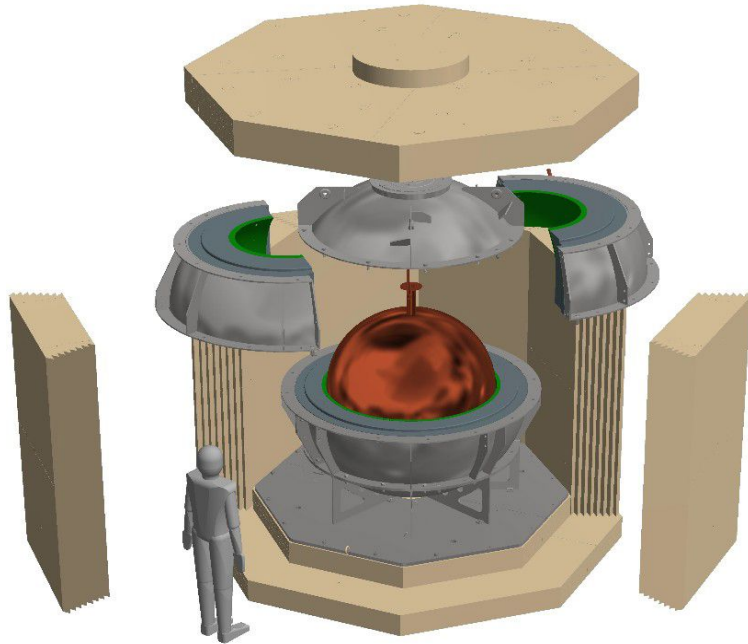


PE shielding installation



SNOGLOBE built Dec 2020

Challenges with the NEWS-G detector



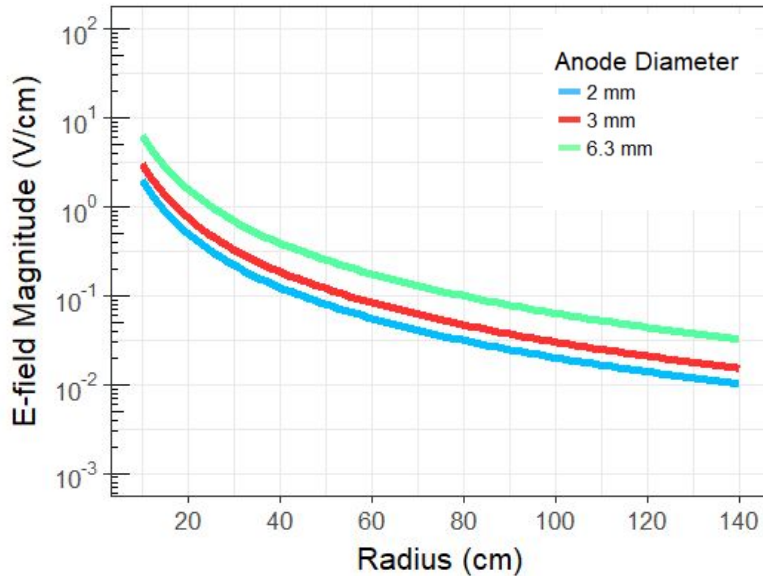
- **Charge collection at large radii and high pressure operation**
 - **Electric field strength**
 - **Contaminants**
- **Detector response uniformity**
- **Background**
 - **Material purity**
- **Monitoring and calibration**
- **Detector simulation/response**

Electric field strength in large volume SPCs

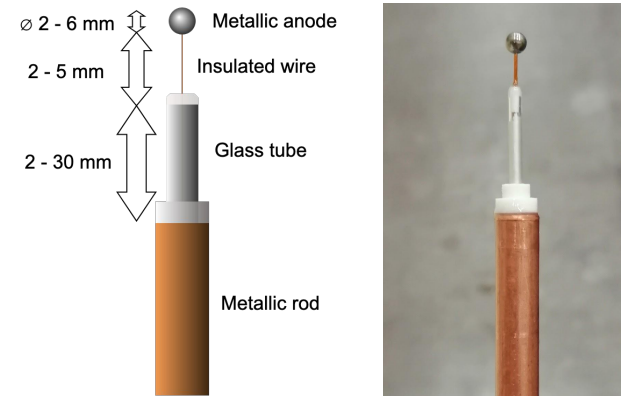
Scaling-up

$$v(r), E(r) \sim r_A / r^2$$

Comparison of E-field magnitude for different anode diameters



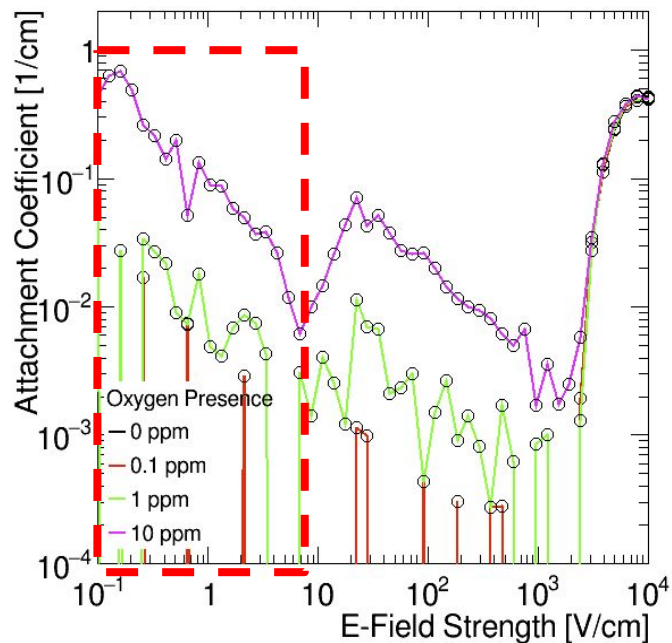
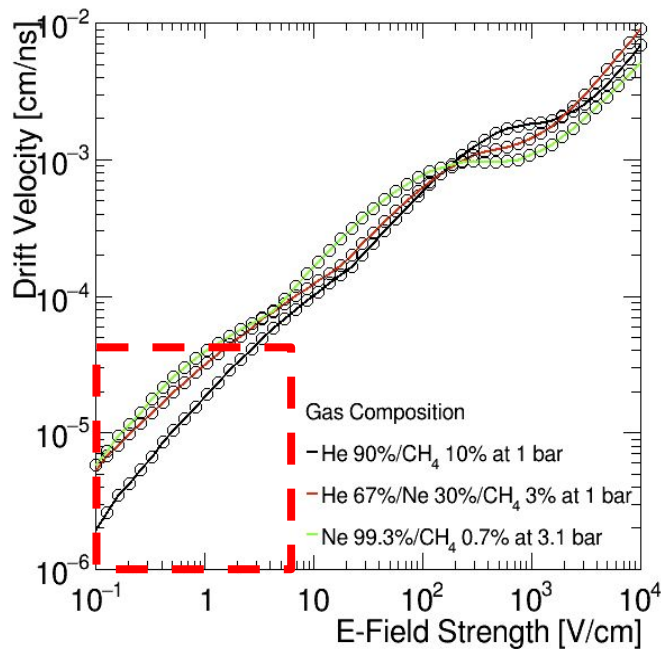
Single anode glass sensors



[I. Katsioulas et al, JINST, 13, 11, P11006, 2018](#)
[10.1088/1748-0221/13/11/P11006](#)

Charge collection in low electric field

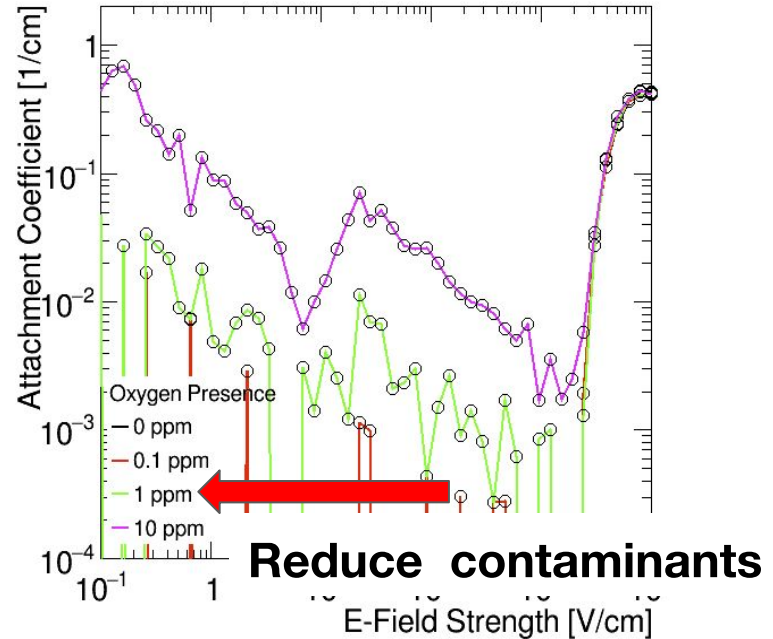
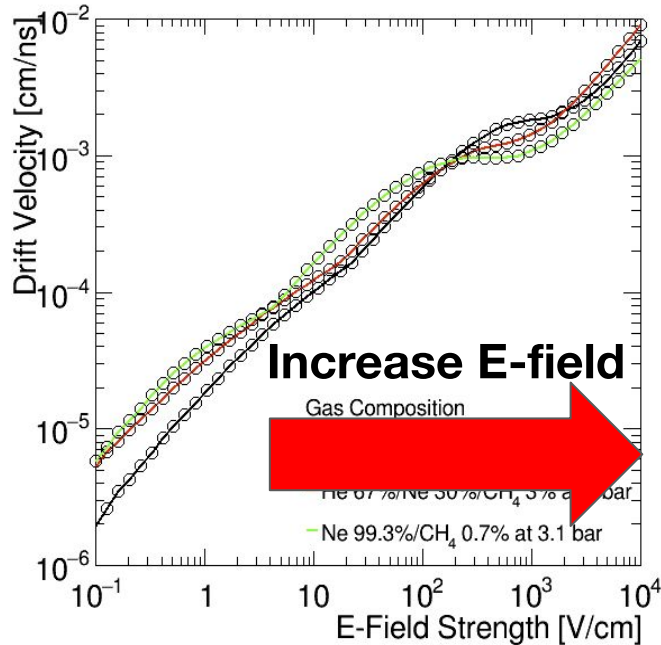
Magboltz study on the sensitivity to contaminants



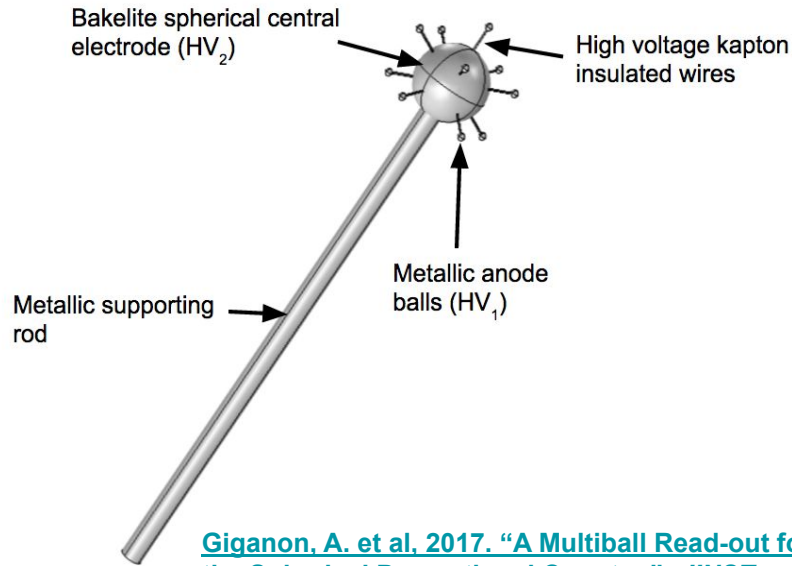
Low E-field region - - -

Charge collection in low electric field

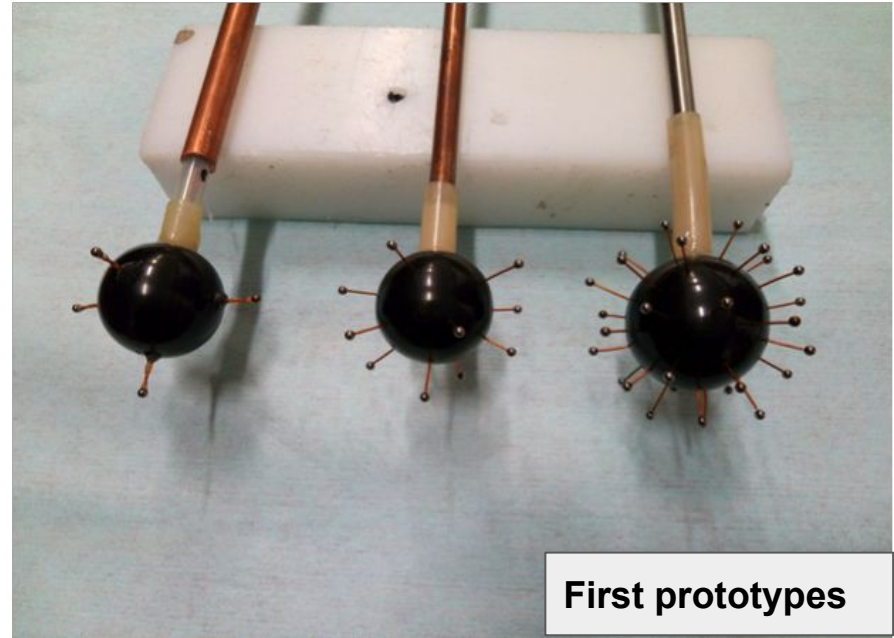
Magboltz study on the sensitivity to contaminants



The multi-anode sensor - ACHINOS

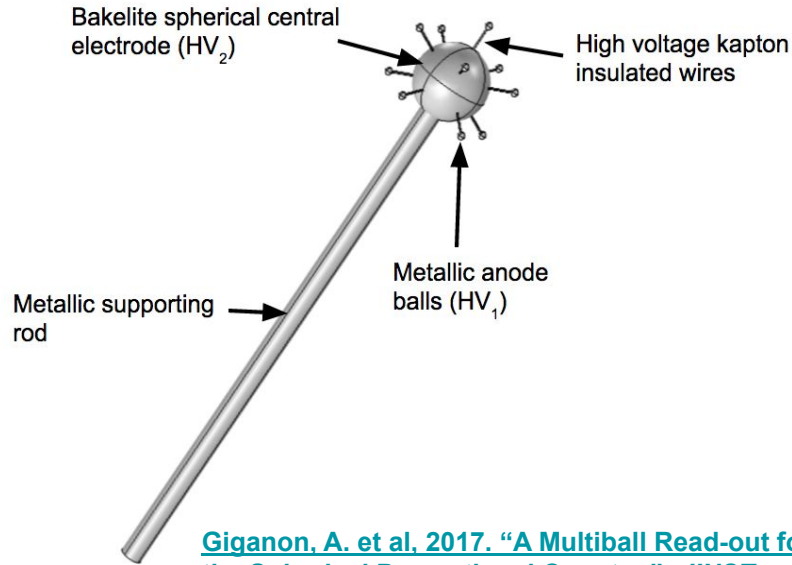


[Giganon, A. et al, 2017. "A Multiball Read-out for the Spherical Proportional Counter.", JINST](#)



Instead of one => Use multiple anodes set to the same potential !!!

The multi-anode sensor - ACHINOS



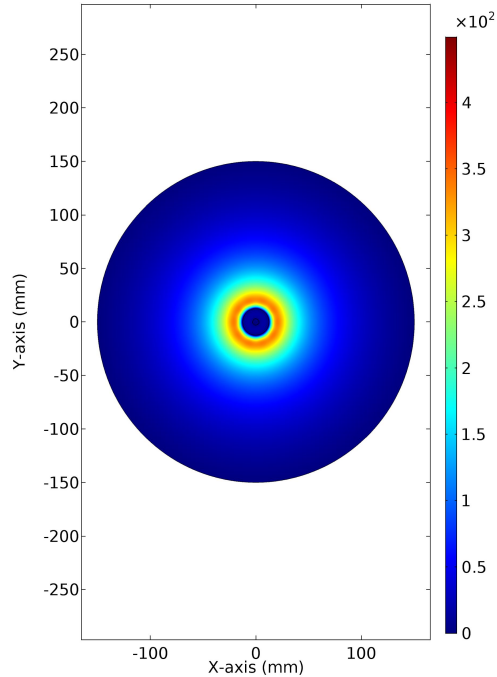
[Giganon, A. et al, 2017. "A Multiball Read-out for the Spherical Proportional Counter.", JINST](#)



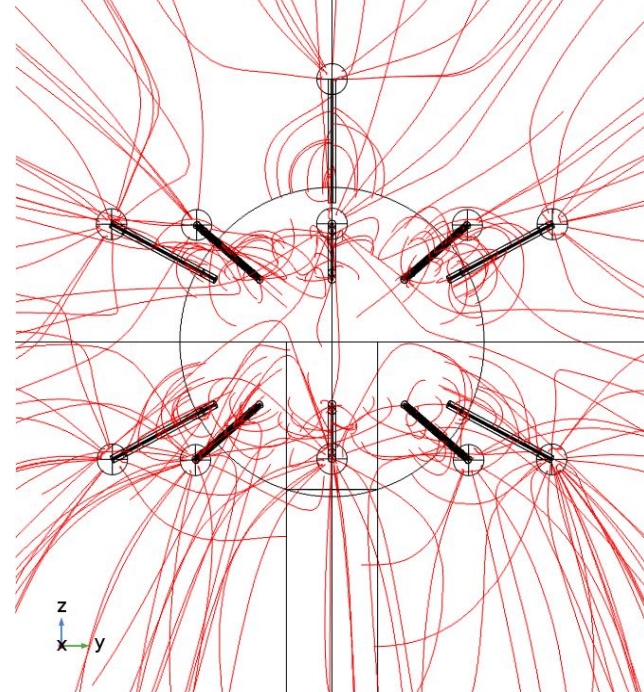
Instead of one => Use multiple anodes set to the same potential !!!

Electric field configuration with an ACHINOS

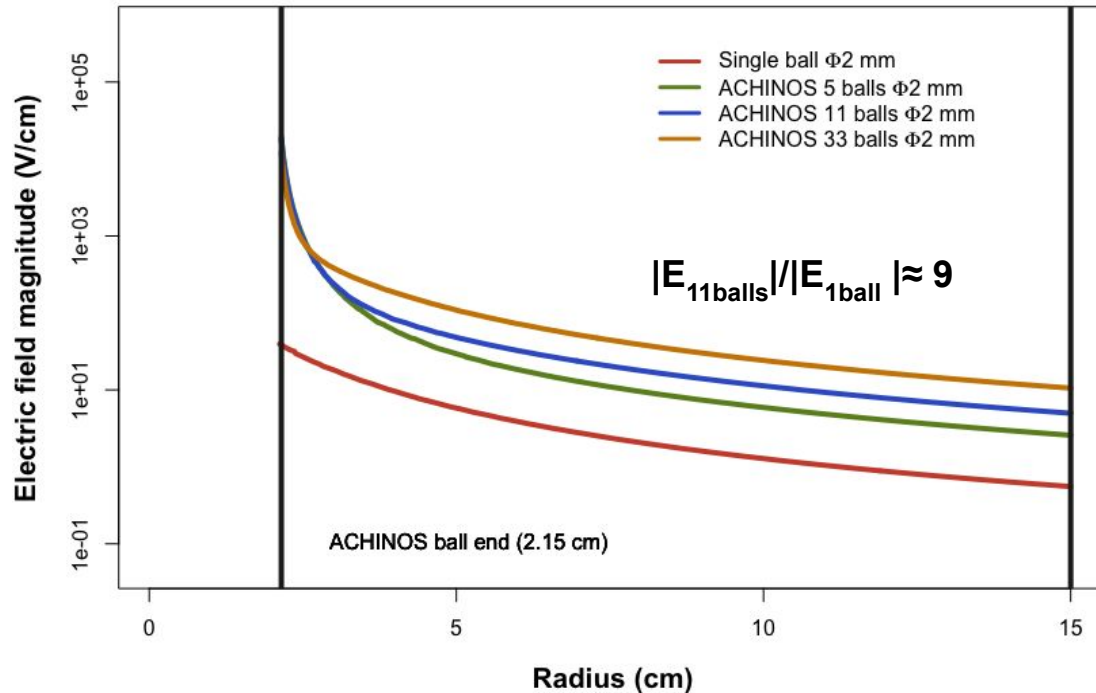
Collective iso-potentials



Electric field lines near the anodes



Advantages of the ACHINOS sensor

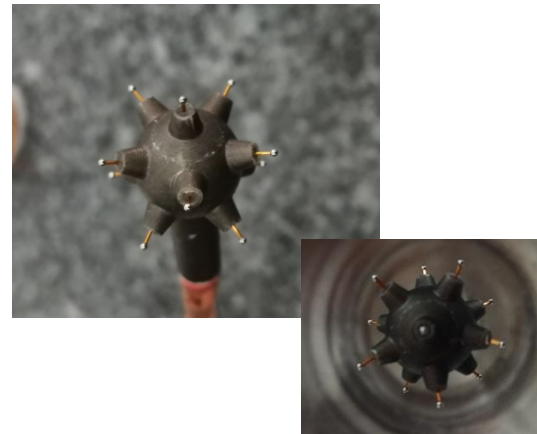
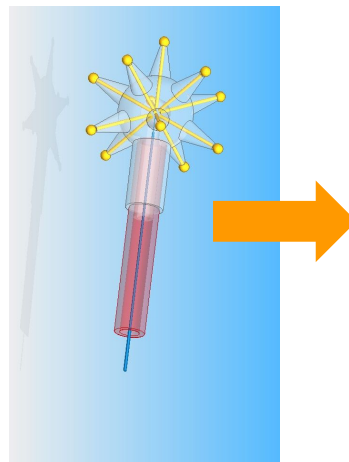


- **Decouples electron drift and multiplication**
 - Small anode size → high gain
 - More anodes → Efficient charge collection
- **Allows for increased target mass**
 - Larger volume
 - Higher pressure
- **Individual readout TPC-like capabilities to the SPC**

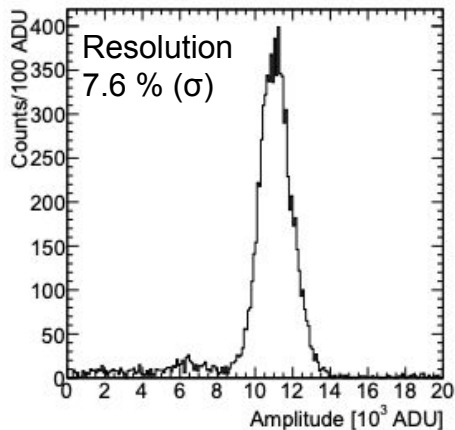
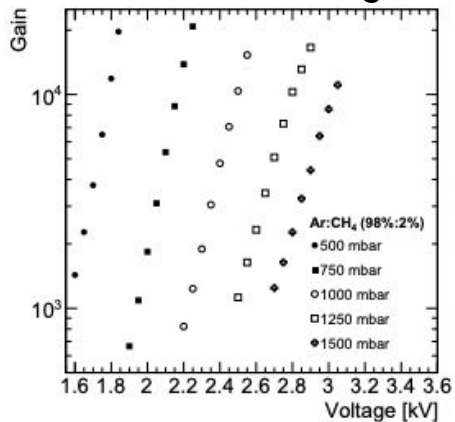
Performance of ACHINOS with DLC coating

3D design

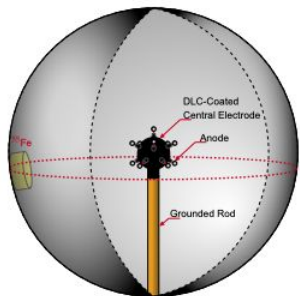
Modules using 3D printing



Gain vs Voltage



Measurement of the 5.9 keV ⁵⁵Fe X-ray line



- Good energy resolution
- High pressure operation (> 2 bar)
- High gain
- Stability
- 2 channel readout

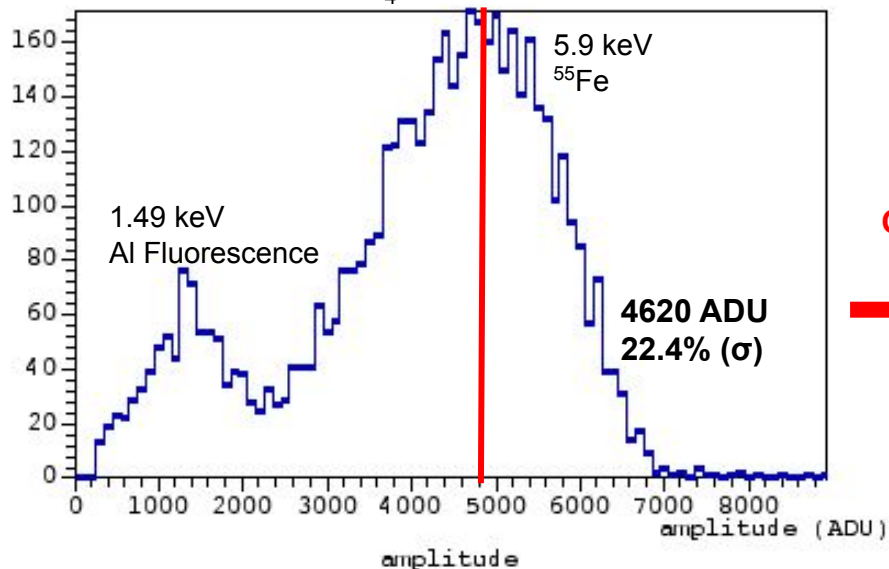
I. Giomataris *et al* 2020 JINST 15 P11023

Gas purification



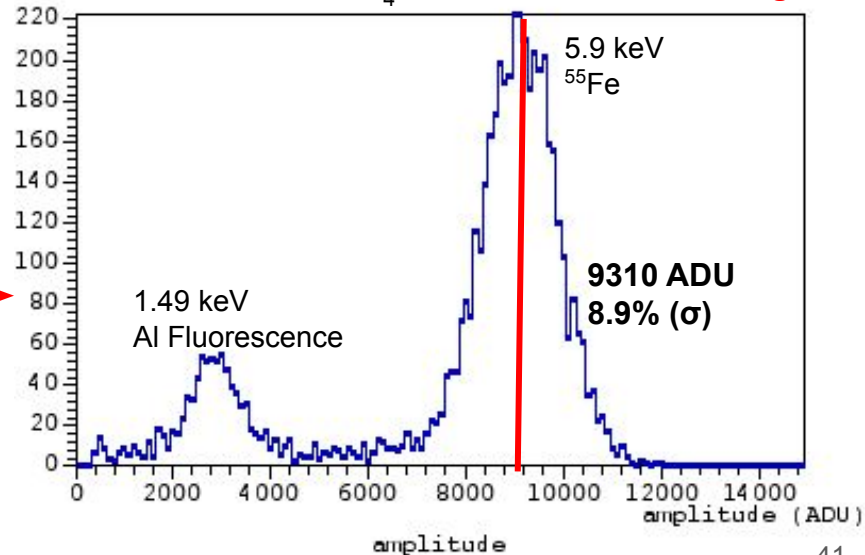
- Contaminants: O₂, H₂O, electronegative gases
- Filtering with: Getter, Oxysorb, Custom filter
- Filtering in a gas recirculation system
 - SAES MicroTorr Purifier (MC700 902-F)
 - Incorporated with Residual Gas Analyser
- Improved filtering efficiency in large sphere

600 mbar He+10% CH₄ **without contaminant filtering**



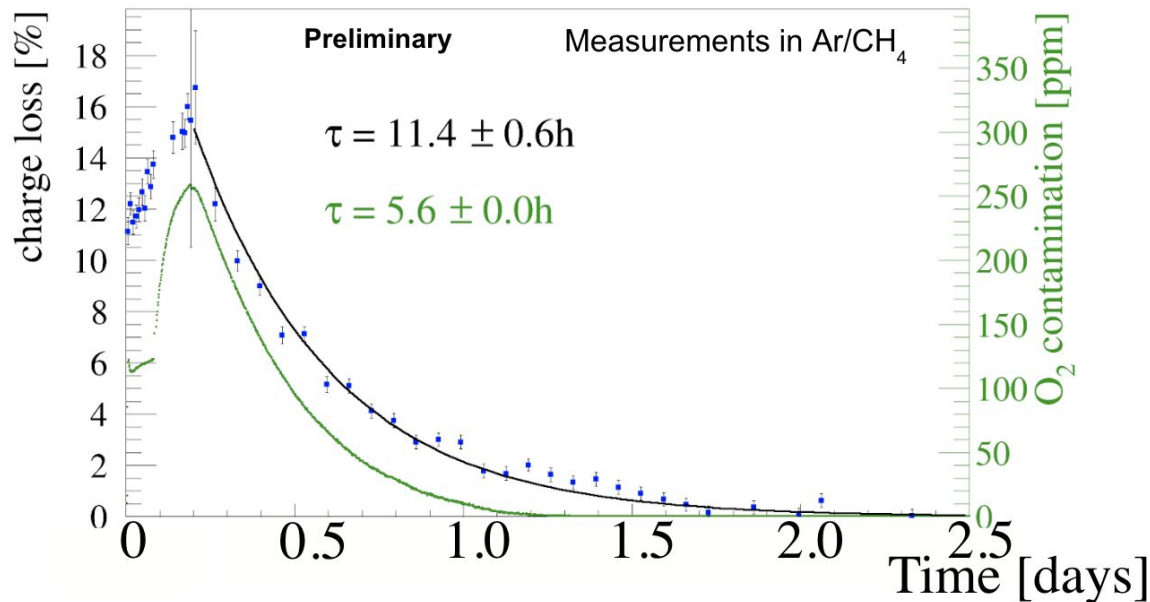
Using
Oxisorb
→

600 mbar He+10% CH₄ **with contaminant filtering**



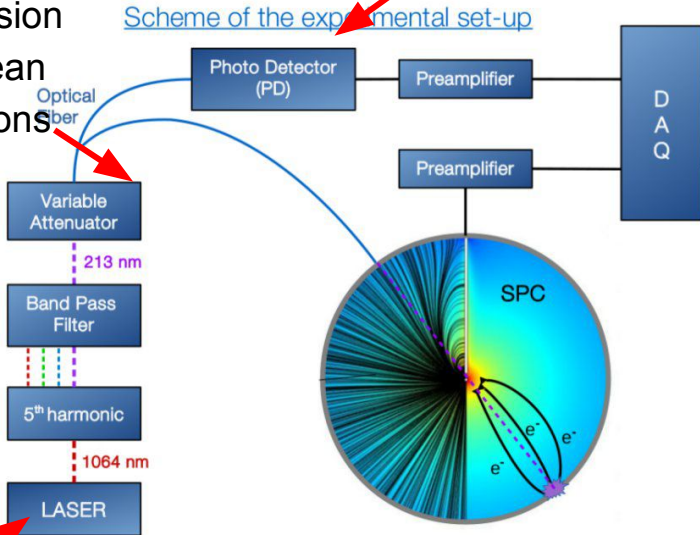
Filtering within a recirculation system

Charge Loss and
Oxygen Concentration
over Time while gas
passes circulated
through MicroTorr
Purifier



Laser calibrations

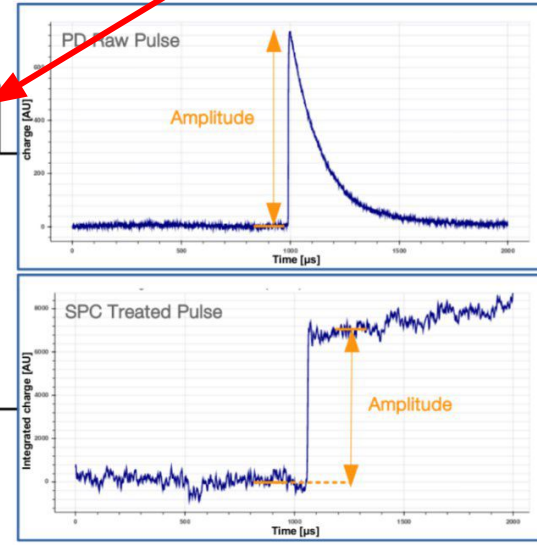
Tunable transmission
to control the mean
number of electrons



A powerful UV laser
capable of extracting 100s
of electrons

Parallel photodetector to
tag laser events

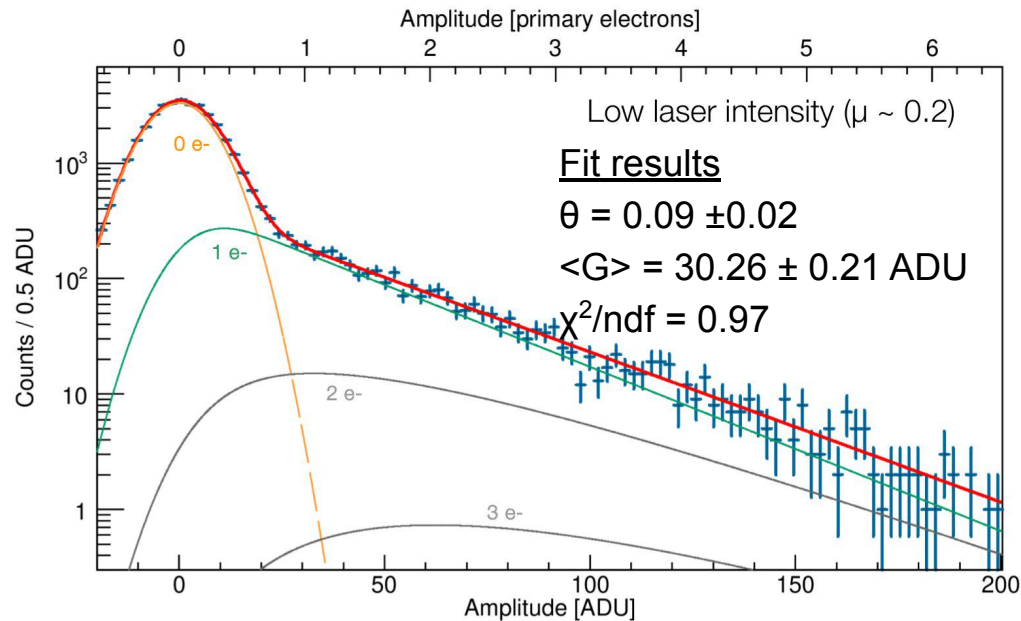
Common DAQ for timing
analysis between two
channels



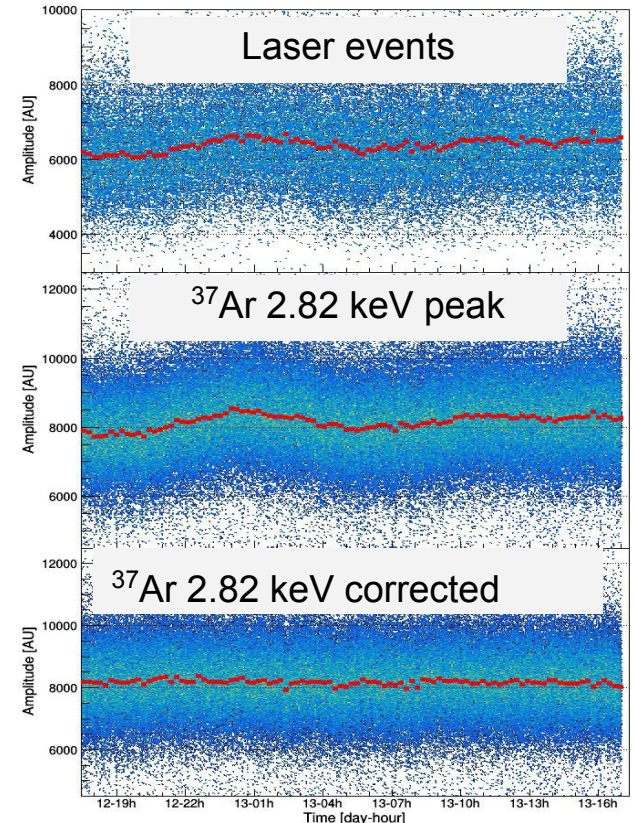
- 213 nm laser used to extract primary electrons from wall of SPC
- Photo detector in parallel tags events and monitors laser power
- Laser intensity can be tuned to extract 1 to 100 photo electrons

Laser calibrations

Single electron response modeling

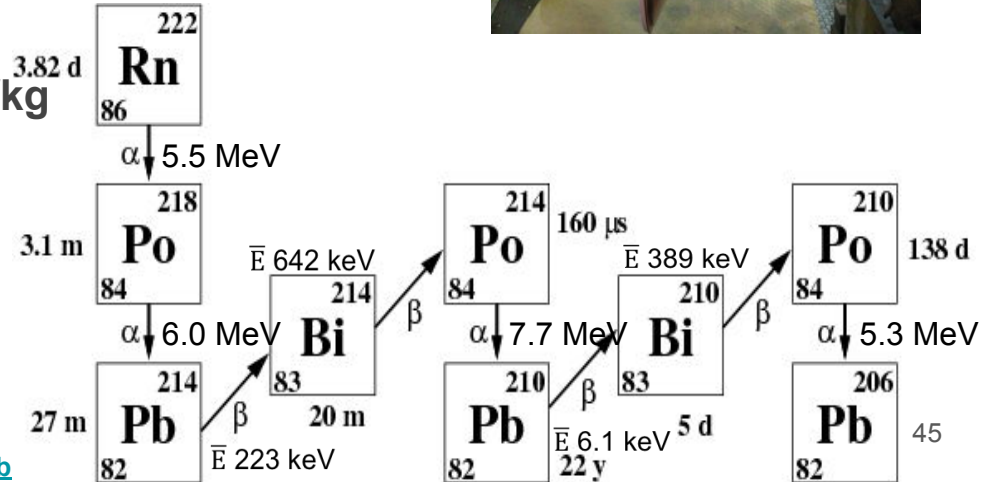
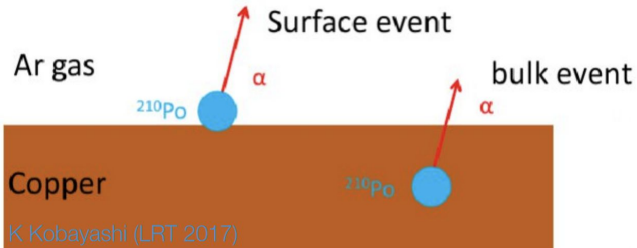


Detector monitoring



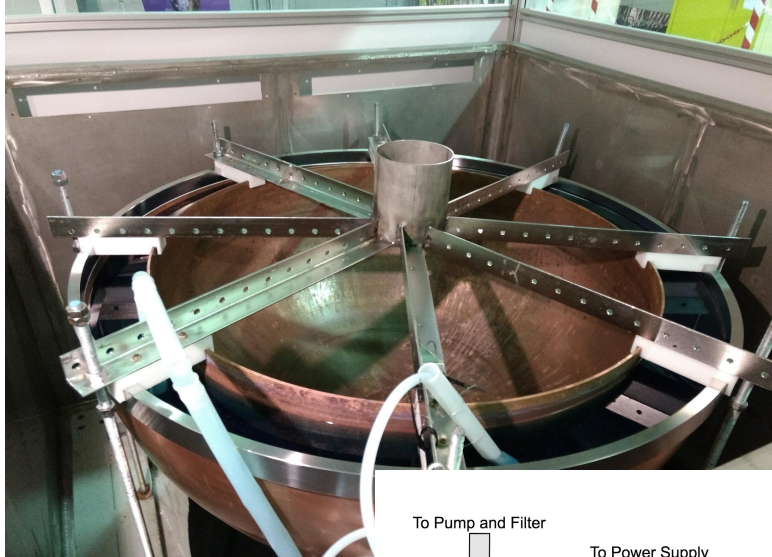
Background in NEWS-G copper

- 4N Aurubis Oxygen Free Copper C10100 (99.99% pure)
 - Spun into two hemispheres
- Copper has no long-lived isotopes
- $^{63}\text{Cu}(n,\alpha)^{60}\text{Co}$ from fast neutrons – mostly cosmic muon spallation
- Contaminants : U and Th decay chain traces
 - Measured for NEWS-G $\sim 10 \mu\text{Bq/kg}$
 - ^{210}Pb out of equilibrium - 28.5 mBq/kg

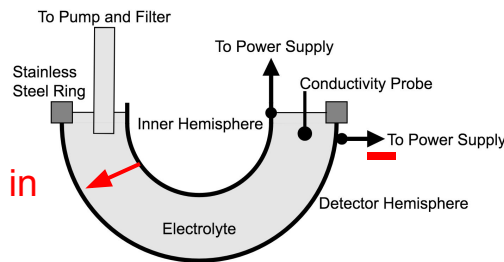


Electroplating Copper

The setup during electroplating at LSM



Cu Movement in Electroplating

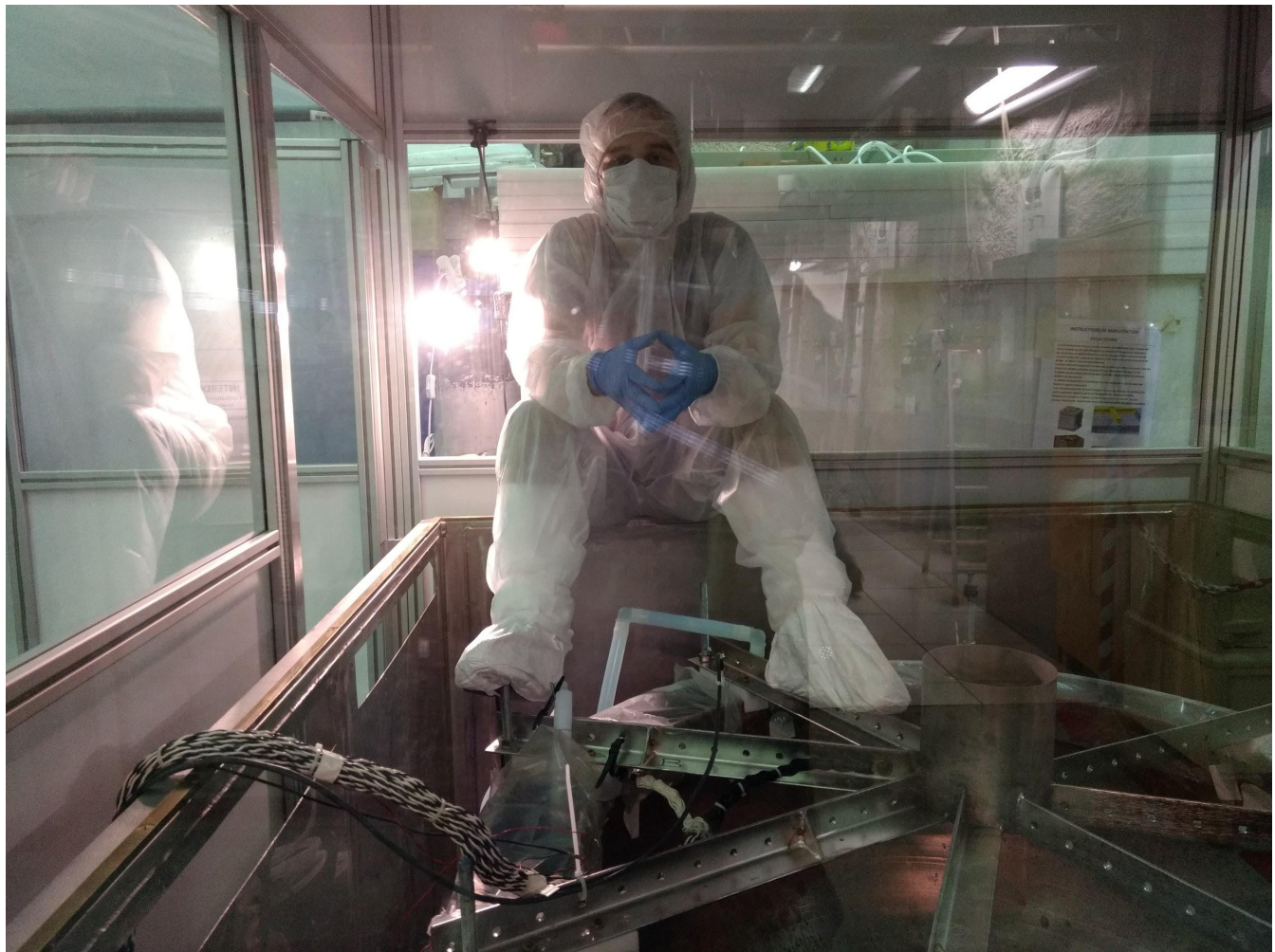


Copper electroplating for background suppression in the NEWS-G experiment

L. Balogh^a, C. Beaufort^b, A. Brossard^a, R. Bunker^c, J.-F. Caron^a, M. Chapellier^a, J.-M. Coquillat^a, E.C. Corcoran^d, S. Crawford^e, A. Dastgheibi Fard^b, Y. Deng^a, K. Dering^a, D. Durnford^a, G. Gerbier^a, I. Giomataris^f, G. Giroux^a, P. Gorel^g, M. Gros^a, P. Gros^a, O. Guillaudin^b, E.W. Hoppe^a, I. Katsioulas^h, F. Kelly^d, P. Knights^{i,j,k,l}, L. Kwon^a, S. Langrock^b, P. Lautridou^a, R.D. Martin^a, J.-P. Mols^f, J.-F. Muraz^a, X.-F. Navick^f, T. Neep^l, K. Nikolopoulos^l, P. O'Brien^a, R. Owen^l, M.-C. Piro^a, D. Santos^g, G. Savvidis^l, I. Savvidis^l, F. Vazquez de Sola Fernandez^a, M. Vidal^a, R. Ward^l, M. Zampaolo^b, NEWS-G Collaboration, S. Alcantar Anguiano^a, I.J. Aronquist^a, M.L. di Vacri^a, K. Harouaka^a, K. Kobayashi^{m,n,o}, K.S. Thommasson^a

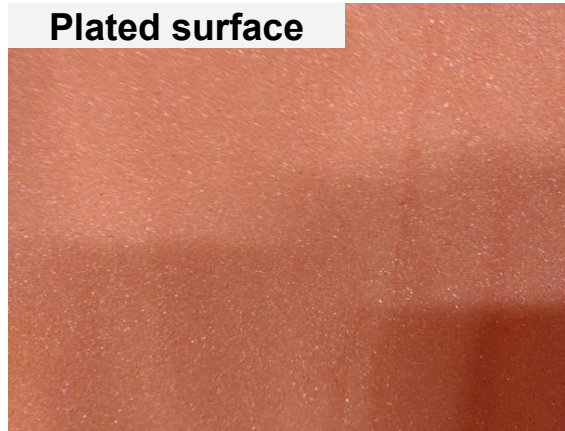
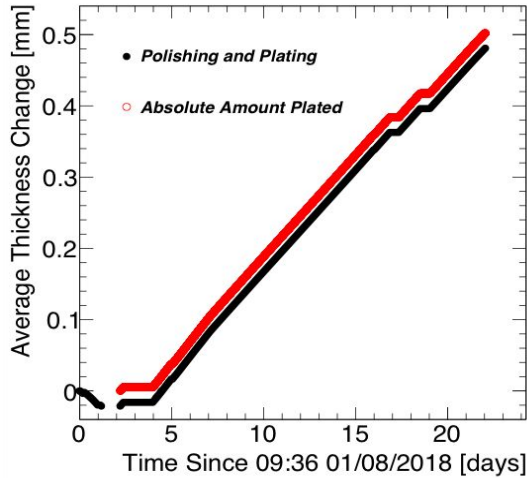
- Using PNNL expertise - Strong participation from UoB in electroforming copper
- The inner surface of the detector was electroplated to stop Bremsstrahlung X-rays from ^{210}Pb and ^{210}Bi β -decays in copper
- 0.5 mm pure copper plated on inner surface at LSM: expected background from ^{210}Pb and ^{210}Bi under 1 keV reduced by a factor 2.6
- Total background in the experiment expected to be 1.96 dru

**P.Knights
UoB - Paris-Saclay**



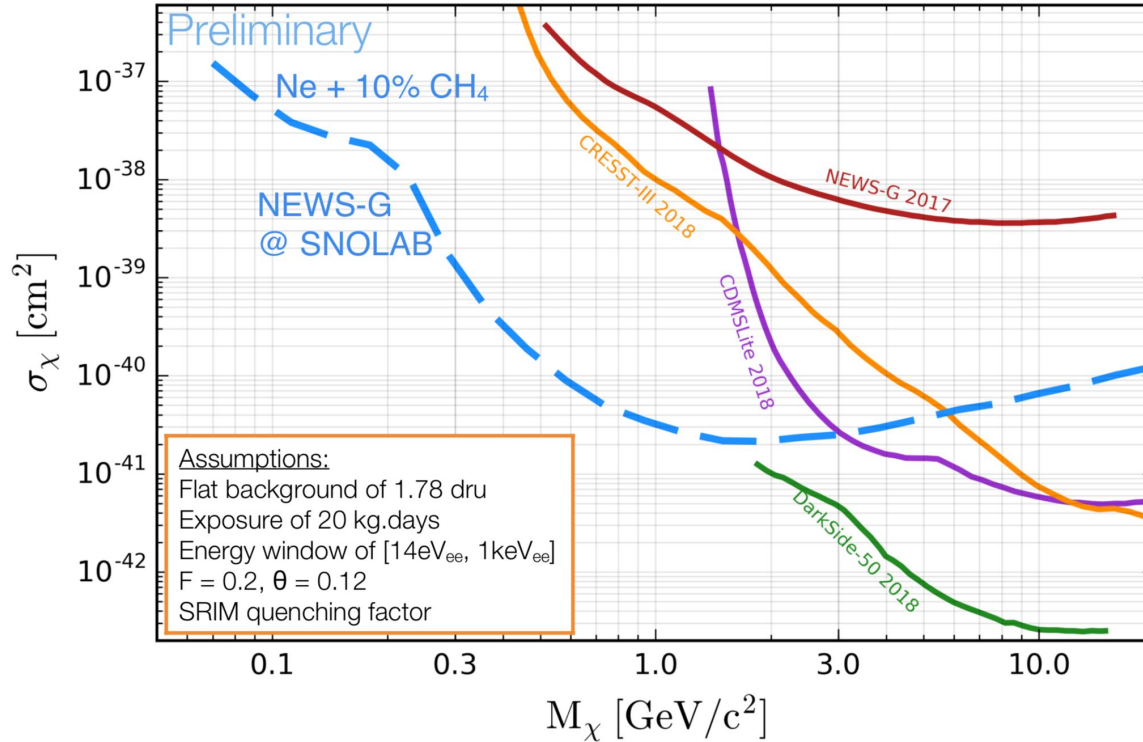


Result after plating



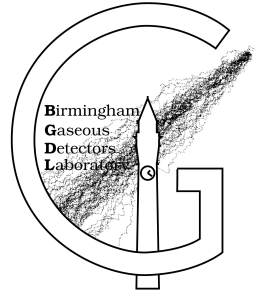
- **Good surface quality achieved**
- **Hemispheres electron-beam welded together**
- **Detector already operating at LSM**
- **Copper was deposited at a rate of $\sim 36 \mu\text{m}/\text{day}$**
 - **Result is promising for possibly a whole detector electroformed underground**

Projected sensitivity for SNOGLOBE



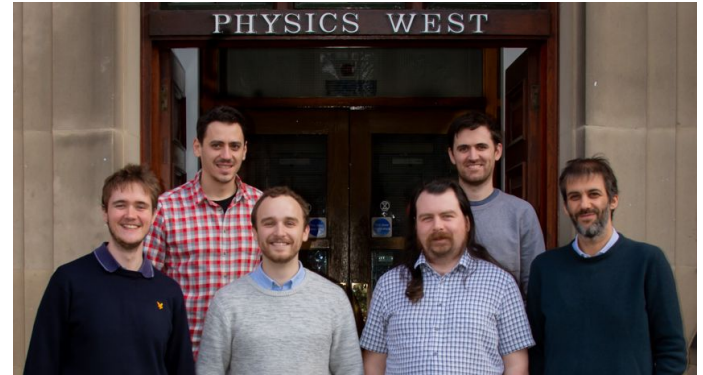
- Low threshold
- Strong background rejection handles
- Hydrogen-rich mixtures
- Lower backgrounds

NEWS-G @ Birmingham

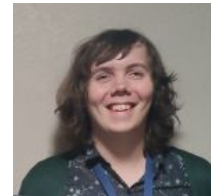


I. Katsioulas, P. Knights, J. Mathews I. Manthos, T. Neep, K. Nikolopoulos, R. Ward

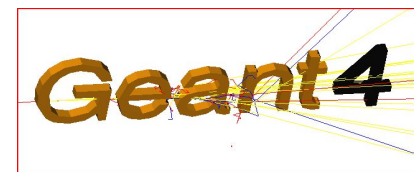
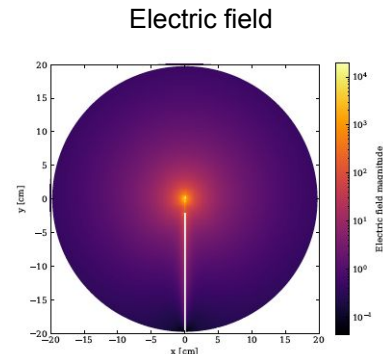
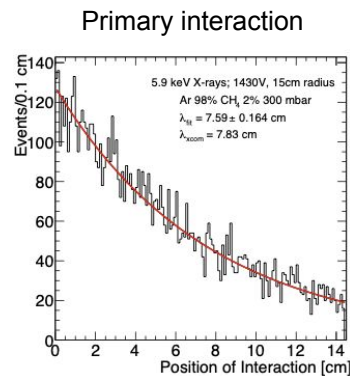
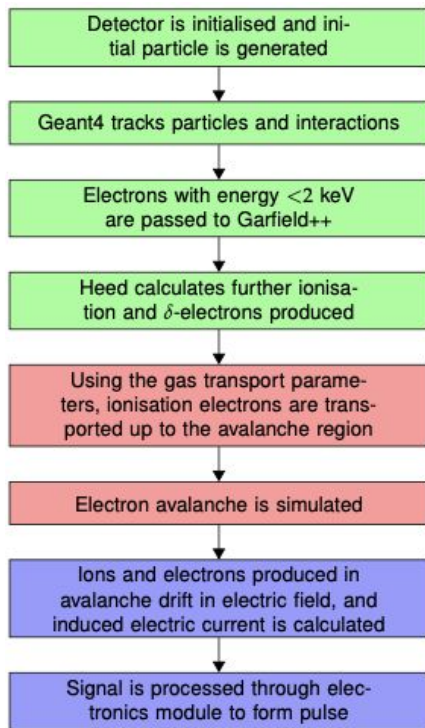
- Co-spokesperson (April 2021)
- Physics Run coordination
- Radiopurity-Background suppression techniques
- Detector physics simulations
- Multivariate analysis techniques
- Ionisation quenching studies
- Neutron background studies



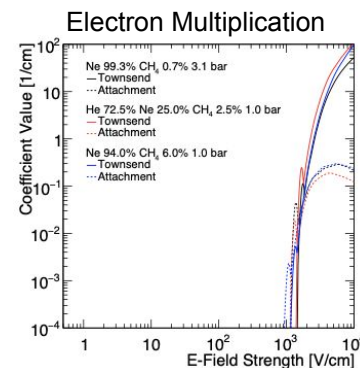
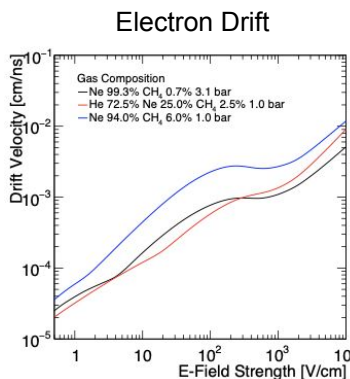
Newest additions!



Simulating the detector response

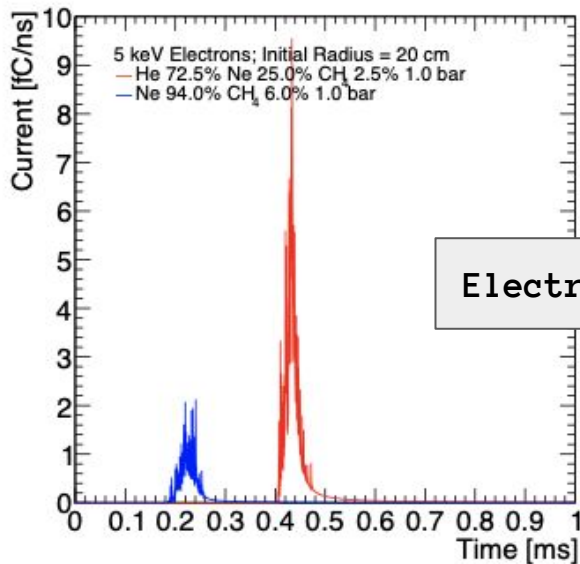


Garfield++

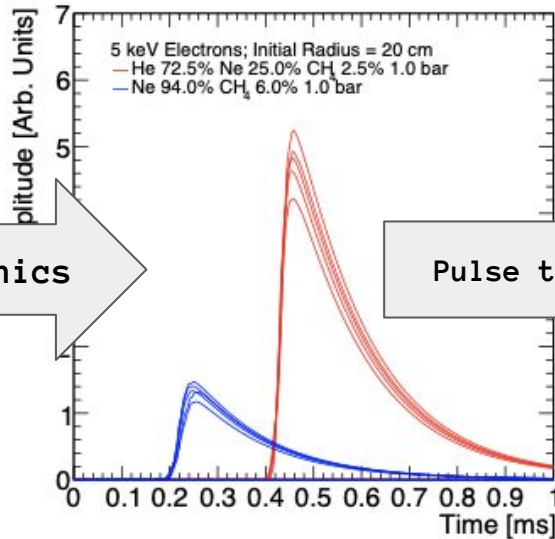


[Katsioulas, I. et al, 2017. "Development of a Simulation Framework for Spherical Proportional Counters", arXiv:2002.02718](#)

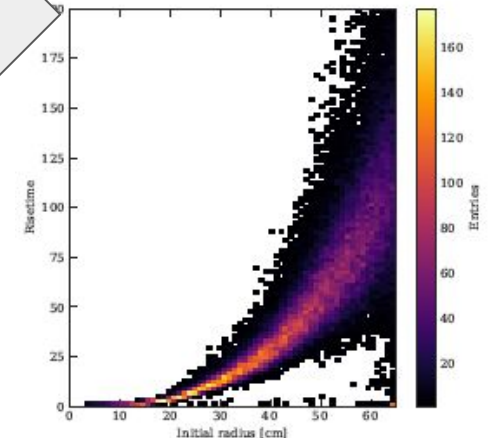
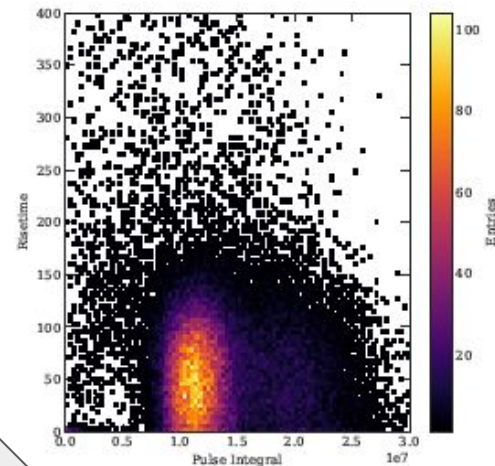
Simulating the detector response



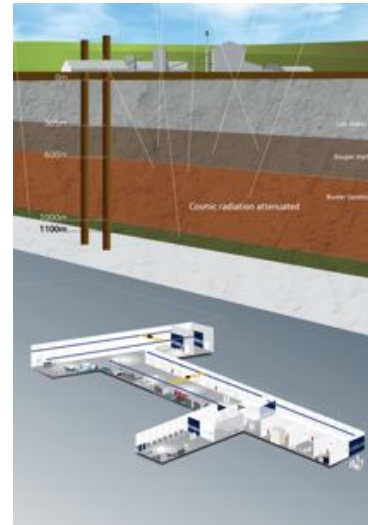
Electronics



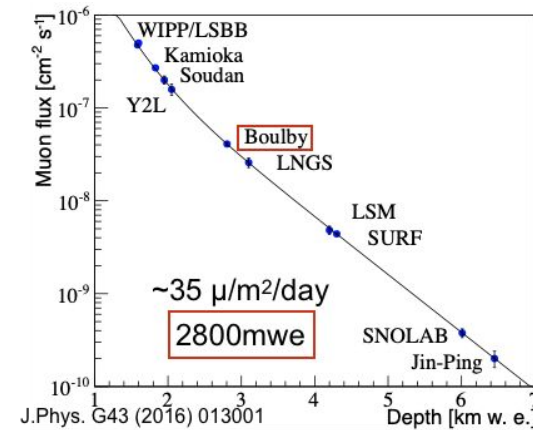
Pulse treatment



Collaboration with Boulby underground laboratory



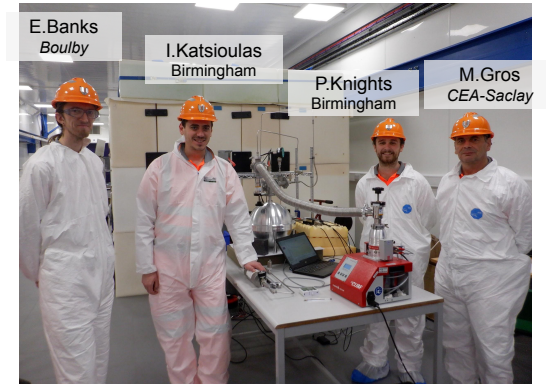
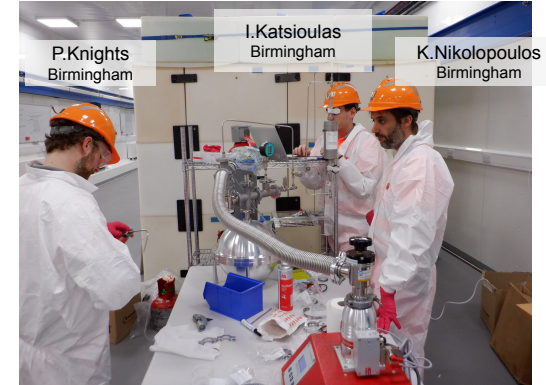
Lowest Radon Levels 3 Bq/m³



Background measurements at Boulby

- Instrumentation R&D at controlled environment
- Neutron flux measurement
 - Thermal neutron
 - Fast neutron
- Including energy information
- Method applicable to all other underground laboratories

Aluminium S30



Neutron detection with SPCs

❑ Neutron spectroscopy with Spherical Proportional Counter

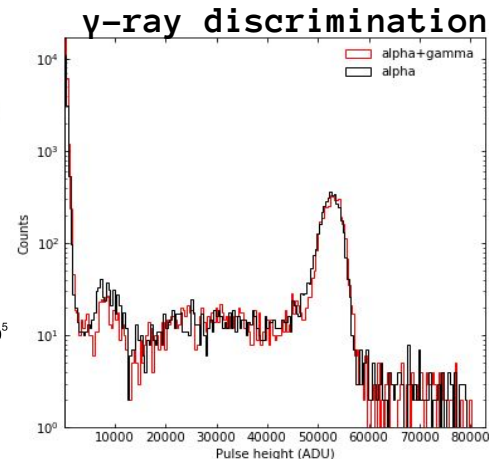
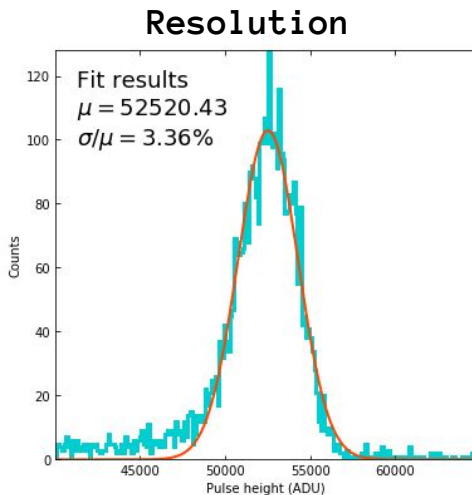
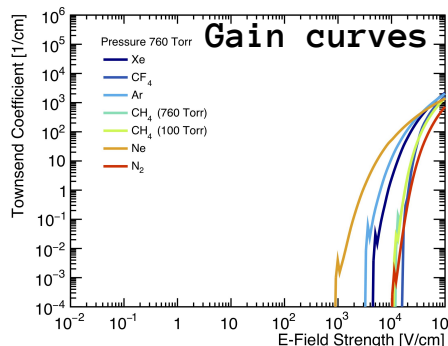
- ❑ Using Nitrogen as gas
- ❑ $^{14}\text{N}+n \rightarrow ^{14}\text{C}+p + 625 \text{ keV}$
- ❑ $^{14}\text{N}+n \rightarrow ^{11}\text{B}+\alpha - 159 \text{ keV}$

❑ Initially demonstrated: *Bougamont, E et al (2017). NIM A, 847, 10–1*

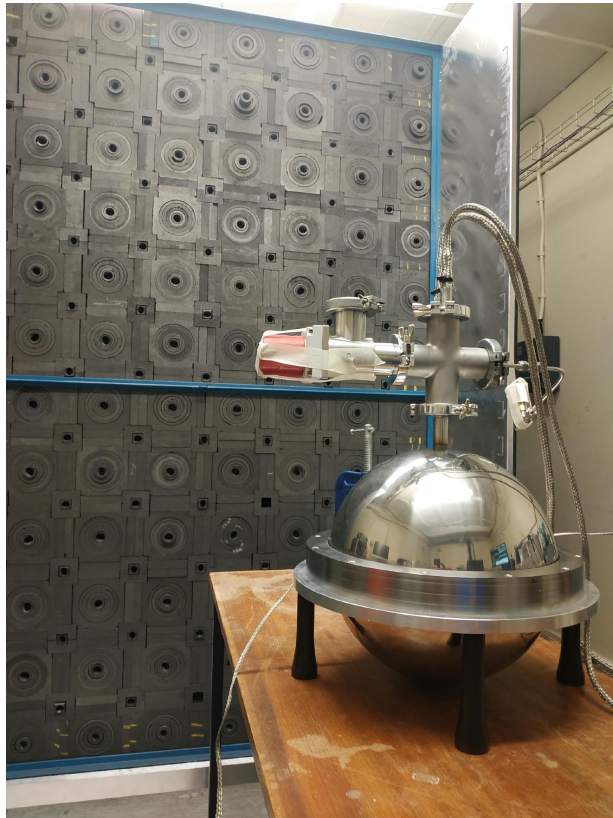
- ❑ ^{252}Cf , ^{241}Am , ^9Be , ambient fast neutrons
- ❑ Thermal neutrons
- ❑ Operation at 0.2-0.5 bar
- ❑ HV reached 6 kV

Goals:

- Operation in high pressure (~5 bar)
 - Minimisation of wall effect
 - Larger target mass - Sensitivity
- Calibration with mono-energetic neutrons
- Measurements with thermal and fast neutrons



Neutrons measurements at UoB



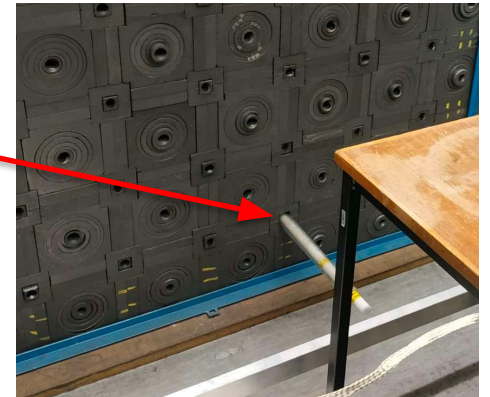
$^{241}\text{Am}^9\text{Be}$ neutron source
 $A = \sim 10^{10}$ Bq

SPC

- 30 cm \varnothing
- N_2 gas filling

Multi-anode sensor

- 11 anodes
- 1mm \varnothing
- Reading in 2 channels

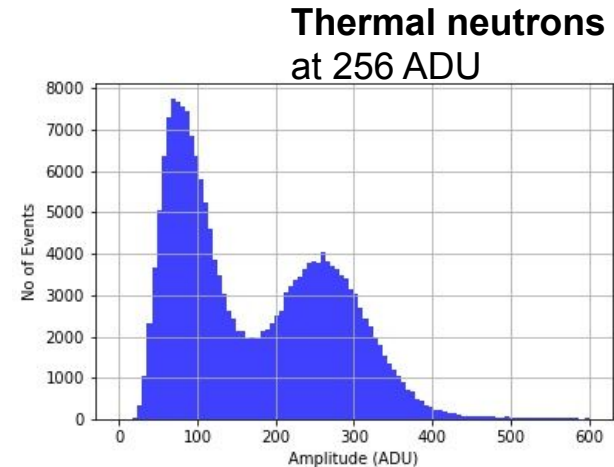
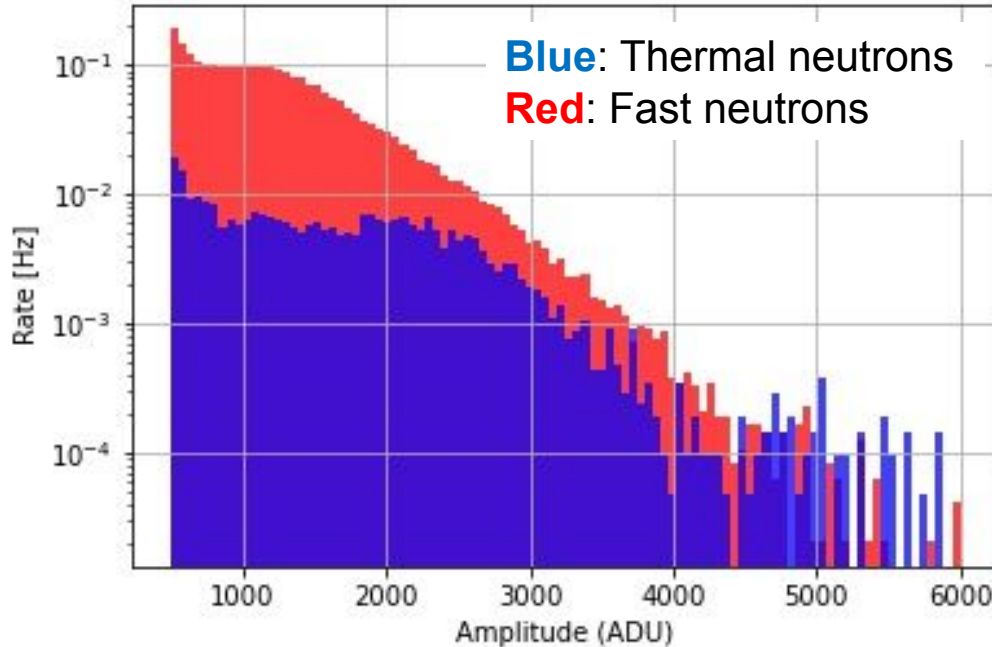


Investigate the capability of the SPC to detect fast neutrons and neutrons thermalized by the graphite.

Neutron measurements with the SPC

$^{241}\text{Am}^9\text{Be}$ neutron source
1 bar N_2 , 3.8 kV

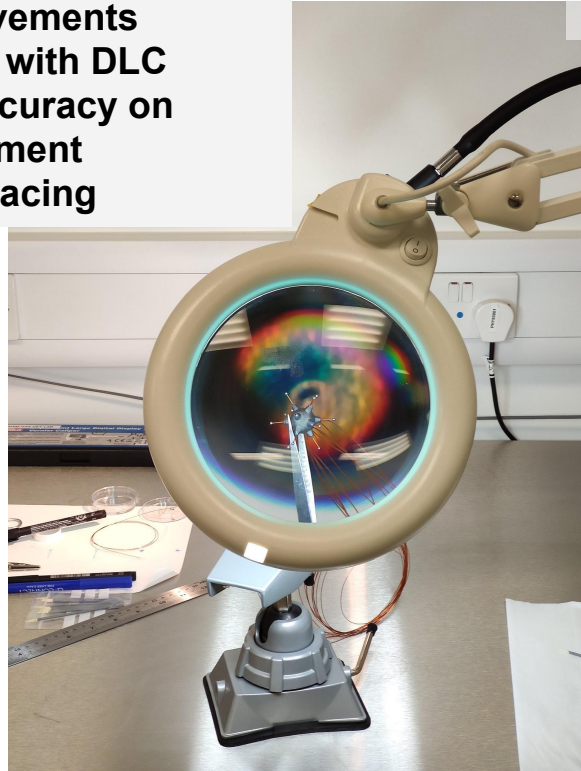
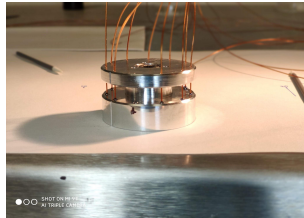
Preliminary results (ongoing analysis)



Instrumentation developments at UoB

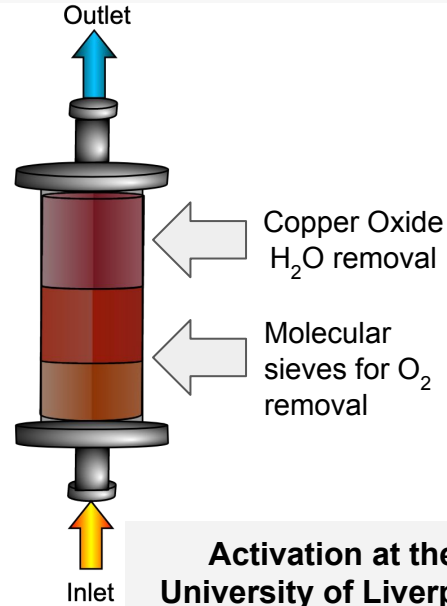
ACHINOS - Achievements

- Fully coated with DLC
- Improved accuracy on anode placement
- Improved spacing



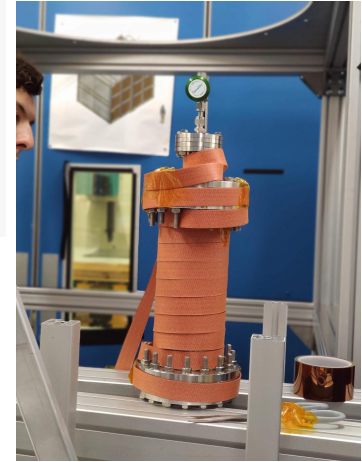
Gas Filter - Aims

- Replace commercial filters
- Reduce Rn emanation
- Maintain efficiency
- Known ingredients
- Cost effective



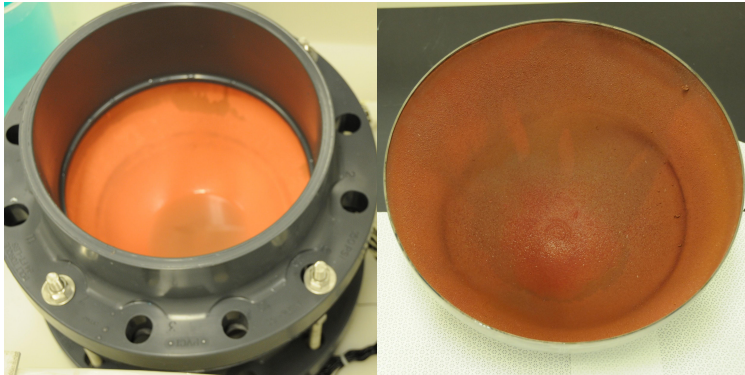
Activation at the
University of Liverpool

Collaboration with
Dr. K.Mavrokoridis Team



ECUME - Electroformed CUprum Manufacturing Experiment

- Underground electroformed $\varnothing 140$ cm sphere
 - Minimised cosmogenic activation - Electroformed in SNOLAB
 - No machining or welding - grow sphere directly
- Based on what was achieved for current NEWS-G sphere
 - $36 \mu\text{m/day} \rightarrow \sim 1 \text{ mm/month}$



Scale hemi-spherical model (PNNL) used for previous electroplating of detector

Current Status:

- R&D bath for prototype at PNNL - underway
- $\varnothing 30$ cm prototype will then be produced
- Full-scale scheduled for late 2021

ECUME - Electroformed CUprum Manufacturing Experiment

	Source	Contamination / flux	Unit	Events rate <1 keV [dru]	Events rate in [1;5] keV [dru]	Total rate [mHz]
Gas mixture	³ H	13	μBq/kg	0.05	0.06	0.005
	²²² Rn	111	μBq/kg	0.05	0.04	0.2
Copper sphere 500 μm electrolyte	²¹⁰ Pb	98.5	mBq/kg	1.04	1.01	0.86
	²³⁸ U	3	μBq/kg	0.0117	0.115	0.028
	²³² Th	13	μBq/kg	0.0754	0.0602	0.163
	⁴⁰ K	0.1	mBq/kg	0.0157	0.0186	0.0022
Roman lead	²¹⁰ Pb	<25	mBq/kg	<0.14	<0.12	0.057
	²³⁸ U	44.5	μBq/kg	0.142	0.094	0.277
	²³² Th	9.1	μBq/kg	0.0256	0.0161	0.0577
	⁴⁰ K	<1.3	mBq/kg	<0.28	0.23	0.65
Low activity lead	²¹⁰ Pb	4.6	Bq/kg	0.053	0.055	0.17
	²³⁸ U	79	μBq/kg	0.17	0.132	0.5
	²³² Th	9	μBq/kg	0.0251	0.0201	0.075
	⁴⁰ K	<1.46	mBq/kg	<0.35	0.26	0.67
Cavern	Gamma	4.87×10^{-8}	γ/cm ² /s	0.0084	0.0095	0.00464
	Neutron	4000	neutron/m ² /day	0.0044	0.0004	3.54×10^{-11}
	Muon	0.27	muon/m ² /day	0.00062	0.00044	5.04×10^{-8}
Total				1.67	1.54	2.4
Total + cosmogenic activation of the copper sphere				5.20	5.20	5.1
Total + cosmogenic activation of the copper sphere and 6 months of cooling				2.8	2.9	3.4
Total + cosmogenic activation of the copper sphere and 1 years of cooling				2.1	1.0	2.0
Total + cosmogenic activation of the copper sphere and 9 years of cooling				1.0	1.7	2.0

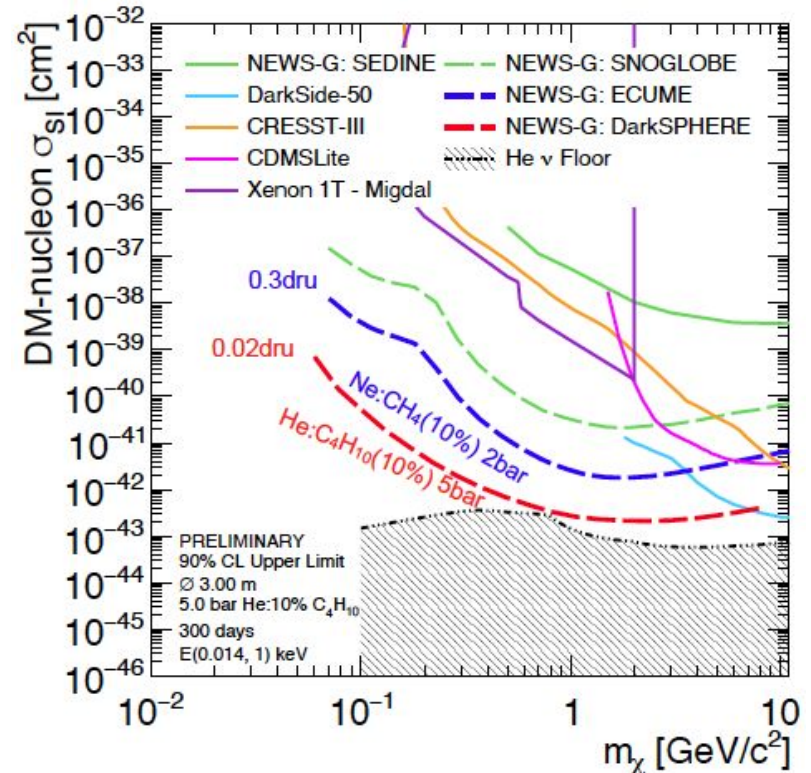
Removing contributions from copper, lead shielding becomes dominant background

DarkSPHERE: Exploring light Dark Matter with Spherical Proportional Counters electroformed underground

Conceptual parameters:

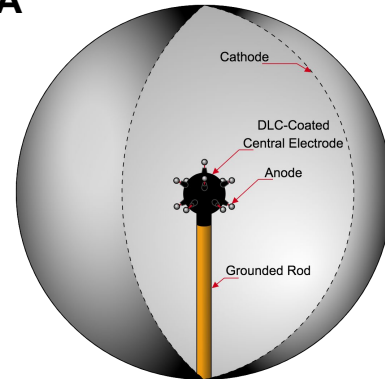
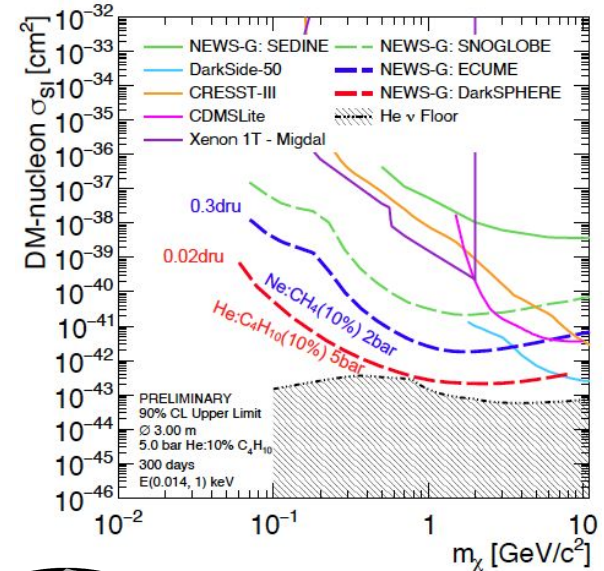
- Installation at Boubly
- 3 m \varnothing SPC
- Fully electroformed underground
- Operation with He/iC₄H₁₂ and possibly Xe
- Pressure up to 5 bar
- Large target mass $O(100\text{kg})$
- Sensitivity down to the ν -floor
- Multiphysics platform

Dimensions in mm



Summary

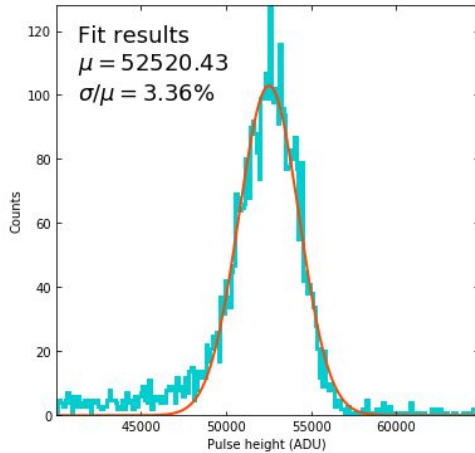
- **NEWS-G searches for light DM candidates**
 - Lighter targets
 - Improved shielding/materials/procedure
 - Lower energy threshold
- **Sensor Development**
 - Improved electric field uniformity
 - **ACHINOS: electric field in large detectors**
- **Improved gas quality: Filtering, Recirculation,**
- **Improved calibration/monitoring: ^{37}Ar , Laser, RGA**
- **Simulation framework development**
- **On going work at Boulby!**
- **Paving the way for future NEWS-G!**
- **Many physics opportunities!**



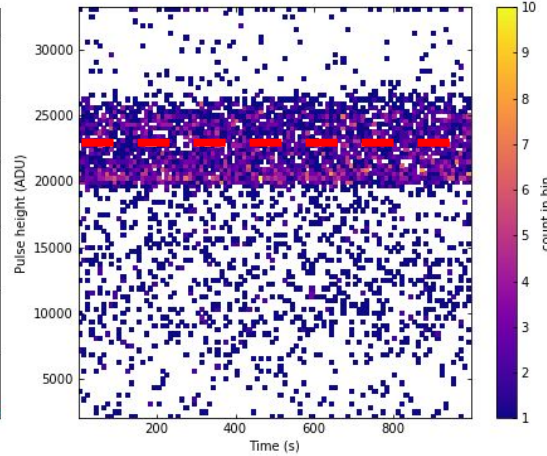
BACKUP SLIDES

Detector characterisation - In progress

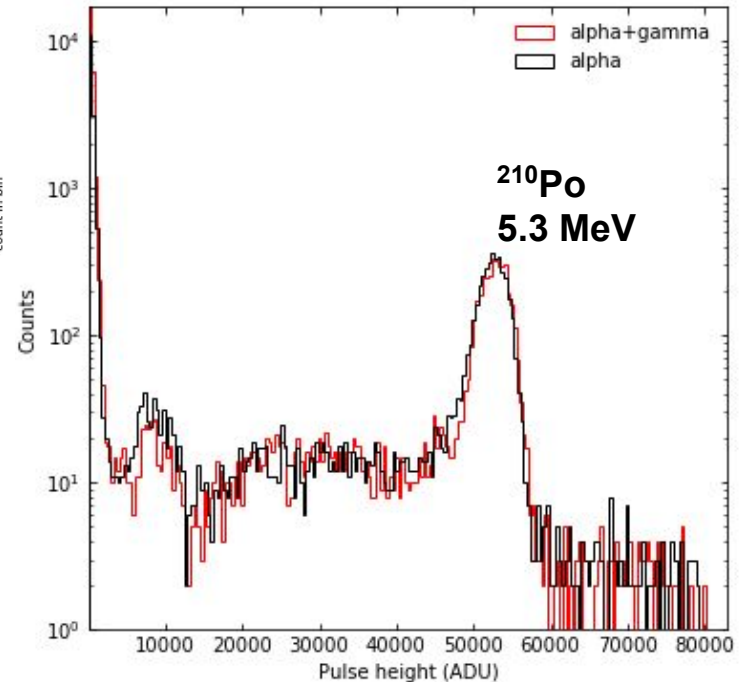
Resolution



Stability



γ -ray discrimination



Planning:

- Operation at up to 2 bar
 - Minimisation of wall effect
 - Larger target mass - Sensitivity
- Calibration with mono-energetic neutrons
- Measurements with thermal and fast neutrons

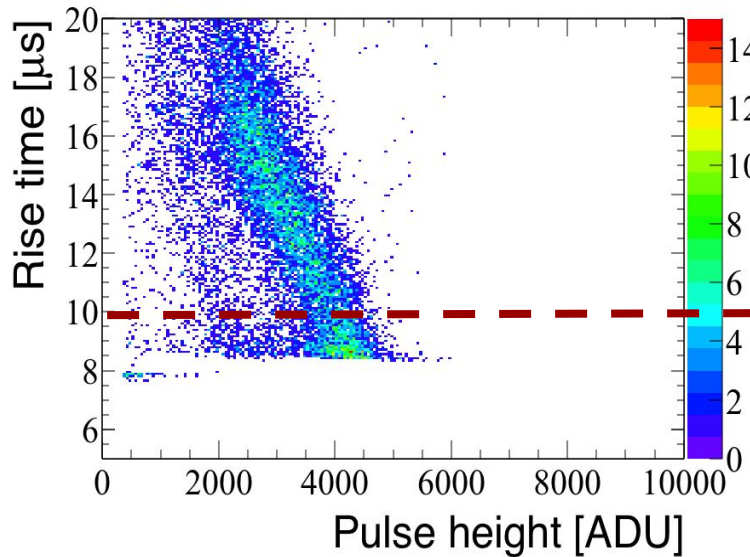
BG budget

	Source	Contamination / flux	Unit	Events rate <1 keV [dru]	Events rate in [1,5] keV [dru]	Total rate [mHz]
Gas mixture	³ H	13	μBq/kg	0.05	0.06	0.005
	²²² Rn	111	μBq/kg	0.05	0.04	0.2
Copper sphere 500 μm electrolyte	²¹⁰ Pb	28.5	mBq/kg	1.04	1.01	0.86
	²³⁸ U	3	μBq/kg	0.0117	0.115	0.028
	²³² Th	13	μBq/kg	0.0754	0.0692	0.163
	⁴⁰ K	0.1	mBq/kg	0.0157	0.0186	0.0622
Roman lead	²¹⁰ Pb	<25	mBq/kg	<0.14	<0.12	0.057
	²³⁸ U	44.5	μBq/kg	0.142	0.094	0.277
	²³² Th	9.1	μBq/kg	0.0256	0.0161	0.0577
	⁴⁰ K	<1.3	mBq/kg	<0.28	0.23	0.65
Low activity lead	²¹⁰ Pb	4.6	Bq/kg	0.053	0.055	0.17
	²³⁸ U	79	μBq/kg	0.17	0.132	0.5
	²³² Th	9	μBq/kg	0.0251	0.0201	0.075
	⁴⁰ K	<1.46	mBq/kg	<0.35	0.26	0.67
Cavern	Gamma	4.87×10^{-8}	γ/cm ² /s	0.0084	0.0095	0.00464
	Neutron	4000	neutron/m ² /day	0.0044	0.0004	3.54×10^{-11}
	Muon	0.27	muon/m ² /day	0.00062	0.00044	5.04×10^{-8}
Total				1.67	1.54	2.4
Total + cosmogenic activation of the copper sphere				5.20	5.20	5.4
Total + cosmogenic activation of the copper sphere and 6 months of cooling				2.8	2.5	3.4
Total + cosmogenic activation of the copper sphere and 1 years of cooling				2.1	1.9	3.0
Total + cosmogenic activation of the copper sphere and 2 years of cooling				1.9	1.7	2.9

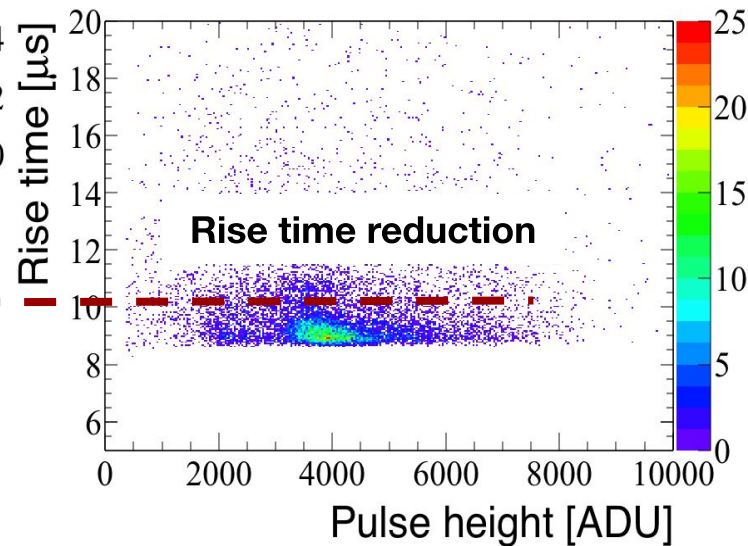
Results with the prototypes

Rise time reduction

Single anode



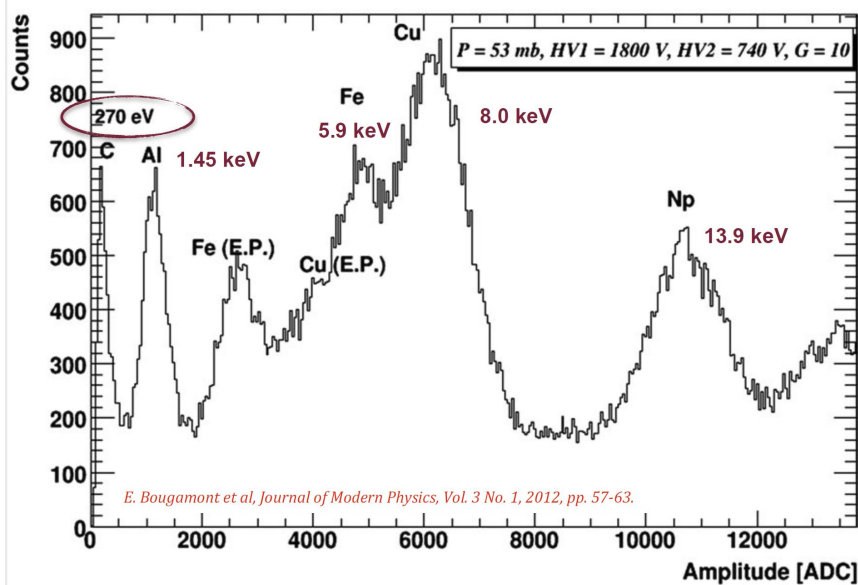
11-anode ACHINOS



--- Maximum rise time

He:Ar:CH₄ (80:11:9)
640 mbar
HV₁ = 2015 V
HV₂ = -200 V
2 mm \emptyset anodes

Low energy detection capabilities of a large volume SPC



SPC \varnothing 130 cm

Gas: Ar+2%CH₄

Detection of fluorescence X-rays

$^{241}\text{Am} \rightarrow ^{237}\text{Np} + ^4\text{He} + 5.6 \text{ MeV}$

Lines

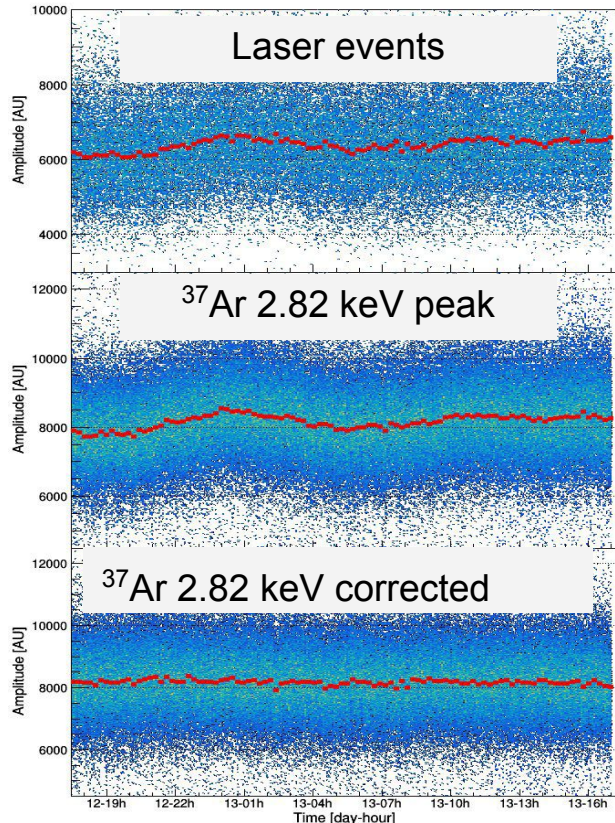
Al \rightarrow 1.45 keV

Cu \rightarrow 13.93 keV

$^{237}\text{Np} \rightarrow 13.93 \text{ keV}(L_{\alpha}) 17.60 \text{ keV}(L_{\beta})$

-Energy threshold at the single electron level

Detector monitoring



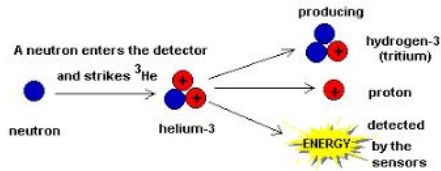
The laser can be used to monitor the detector response during physics runs

Long-term fluctuations in gain can be caused by temperature changes, O₂ contamination, sensor damage...

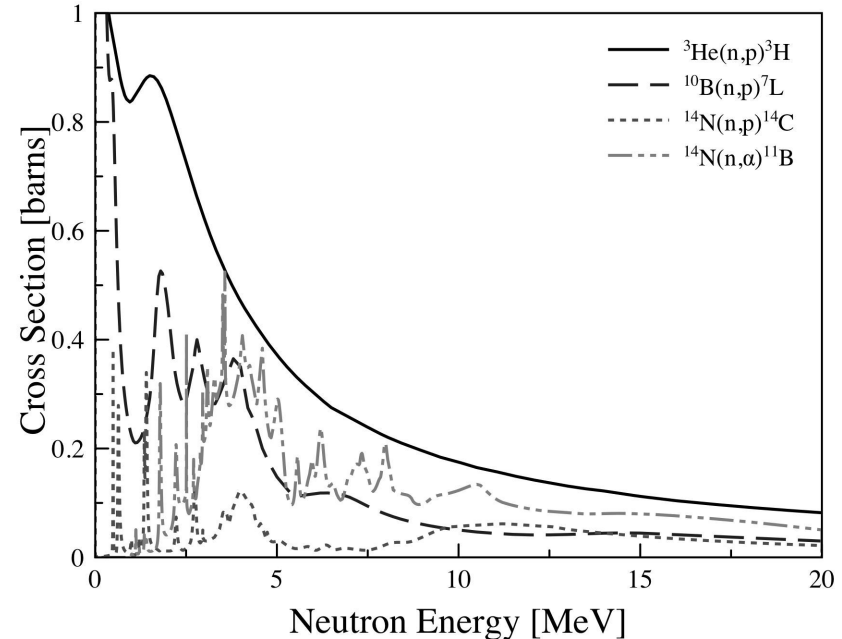
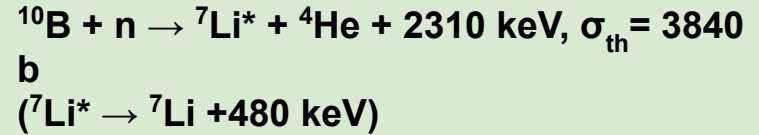
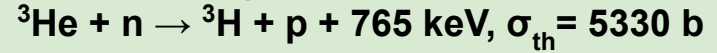
Laser monitoring data could even be used to correct for long-term fluctuations

Neutron spectroscopy

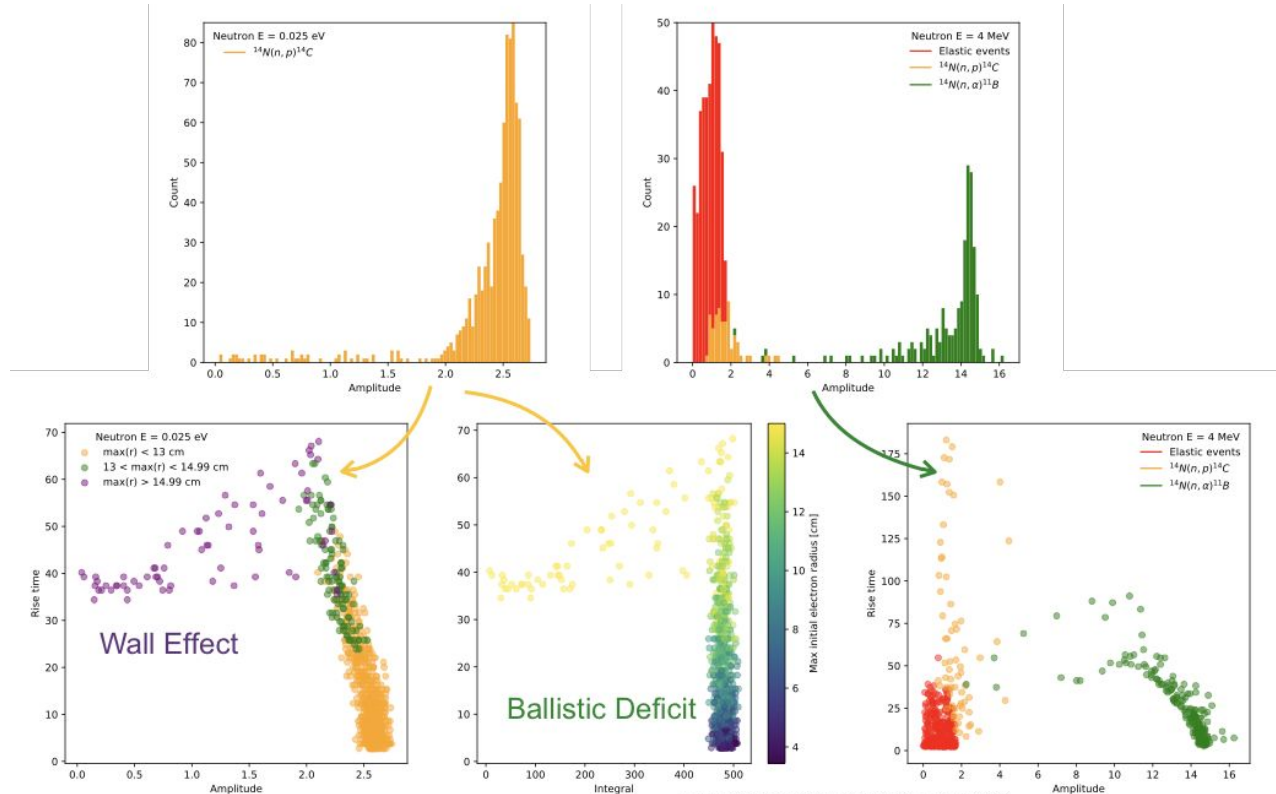
- **Neutrons: important background in DM searches**
 - Identical signature to signal events
 - Stored material activation
- **Few measurements at underground laboratories**
 - ^3He -based detectors extremely expensive



Common targets:



Simulation of neutron transport



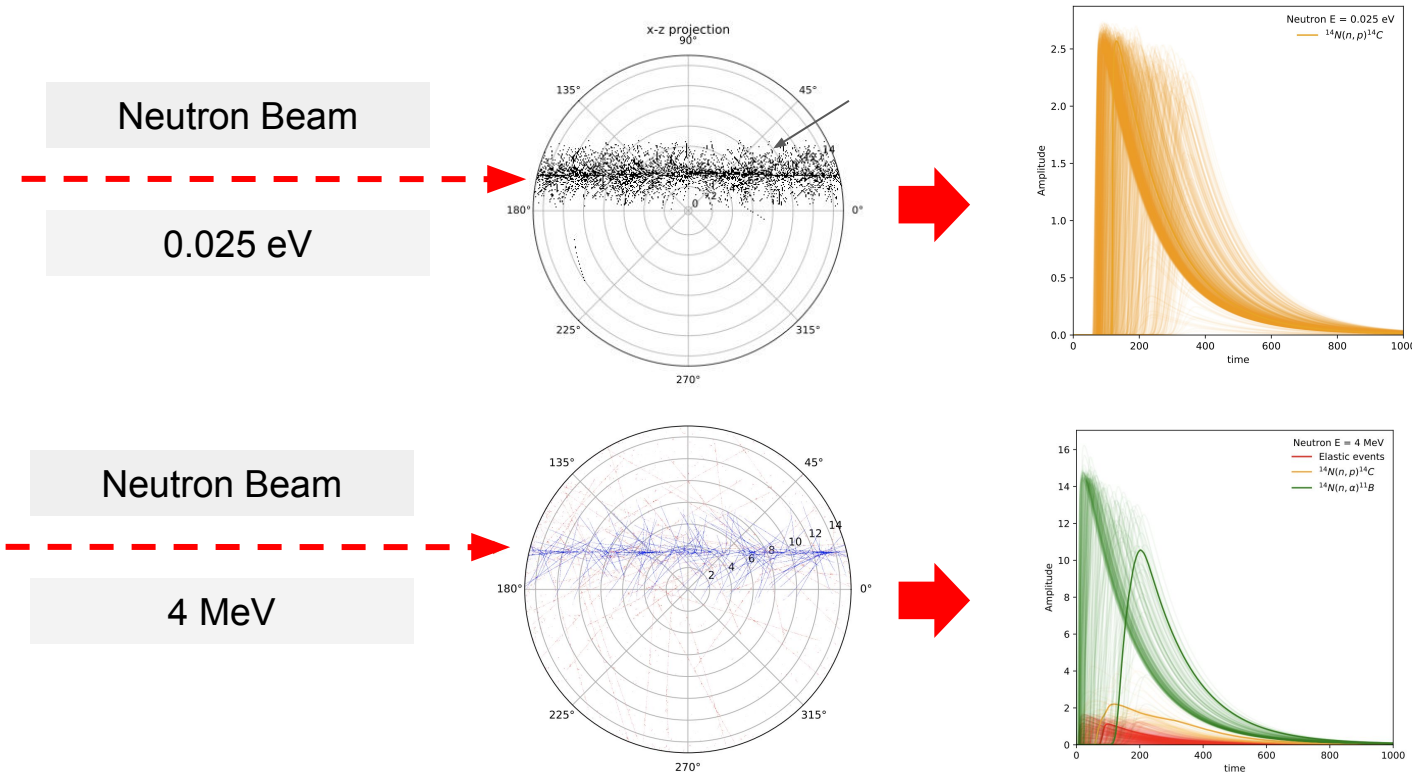
Simulation of neutron transport

Simulation Parameters:

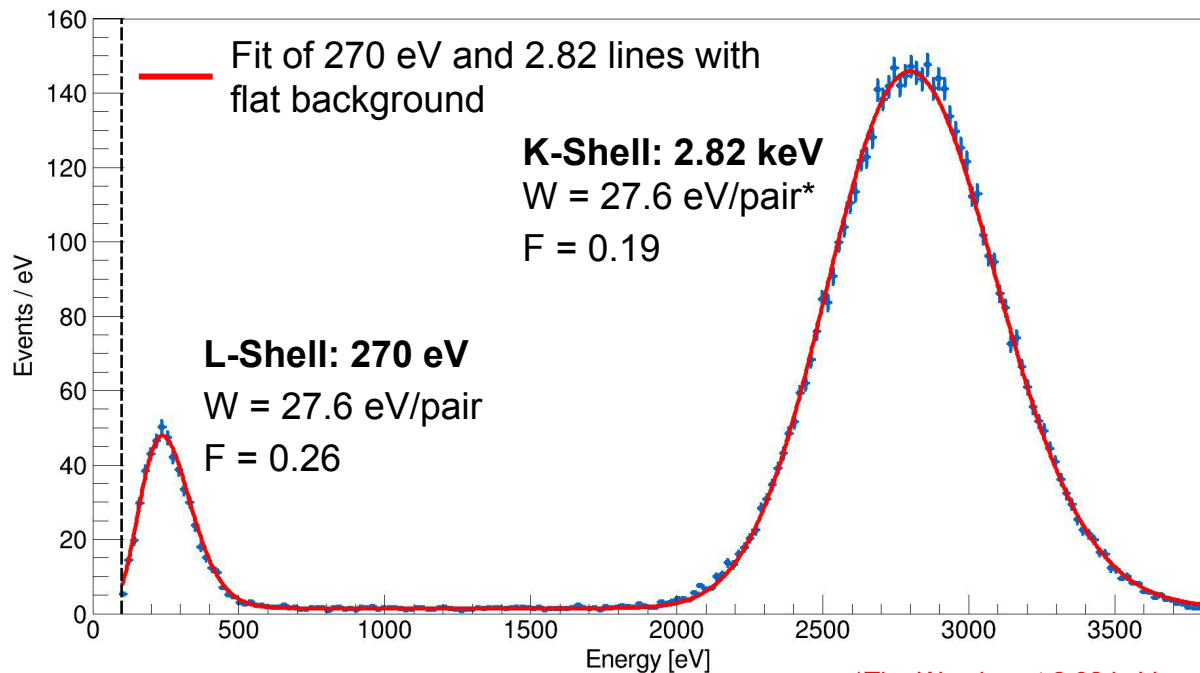
∅ vessel 30 cm

Nitrogen at 300 mbar

Anode ∅ 2 mm



Ar37 calibrations and gas fundamentals properties measurement



- Ar37 produced by irradiating Ca power with a high flux of fast neutrons
- Together with laser calibrations, can find W (mean ionization energy) with 1% precision for target gas, and set upper limits on F (Fano factor)

Detector response modeled:

- Primary ionisation (COM-Poisson)

D. Durnford et al, Phys. Rev. D 98, 103013 (2018)

- Avalanche (Polya)

$$f(E') = \sum_{N=1}^{N_{\max}} P_{\text{CMP}}(N|\mu, F) \times P_{\text{Polya}}^{(N)}(E'|\langle G \rangle, \theta)$$

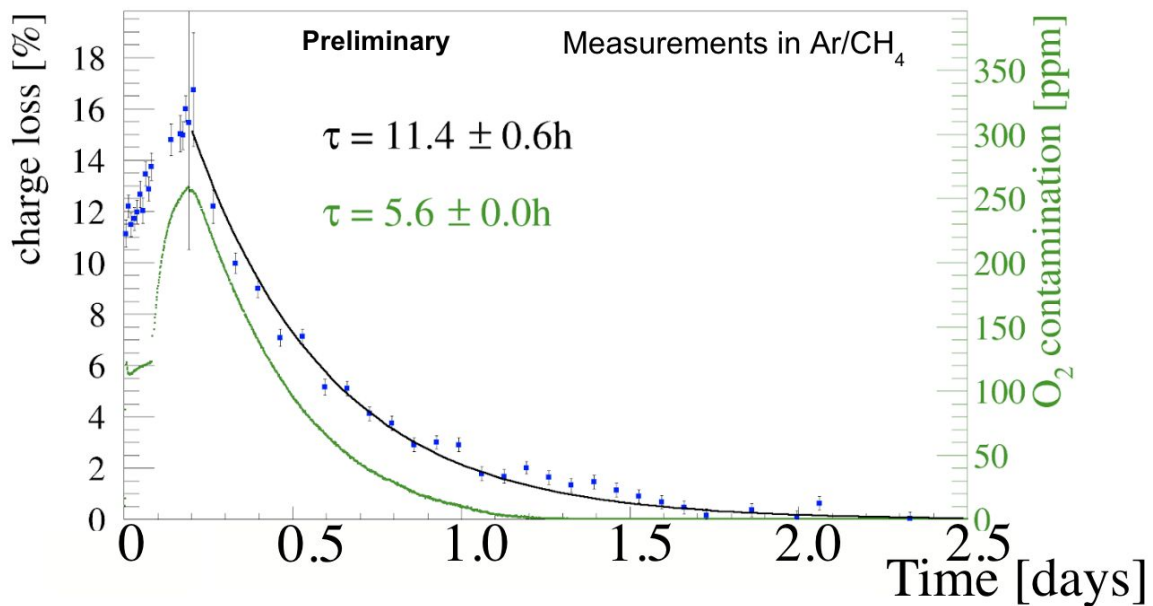
*The W -value at 2.82 keV was calculated directly from $\langle G \rangle$ and

fixed for this fit

Filtering within a recirculation system

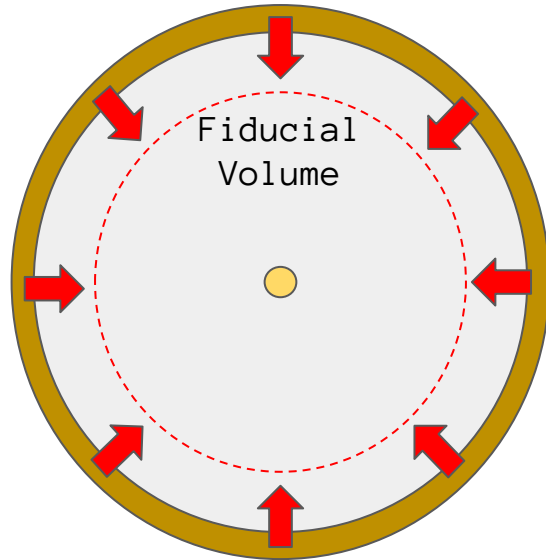
- SAES MicroTorr Purifier (MC700 902-F) then used
- Improved filtering efficiency in large sphere – attachment problem ‘solved’
- Incorporated into recirculation system with RGA

Charge Loss and Oxygen Concentration over Time while gas passes circulated through MicroTorr Purifier

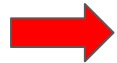


Background rejection capabilities-A

Fiducialisation



Background comes from the materials of the vessel

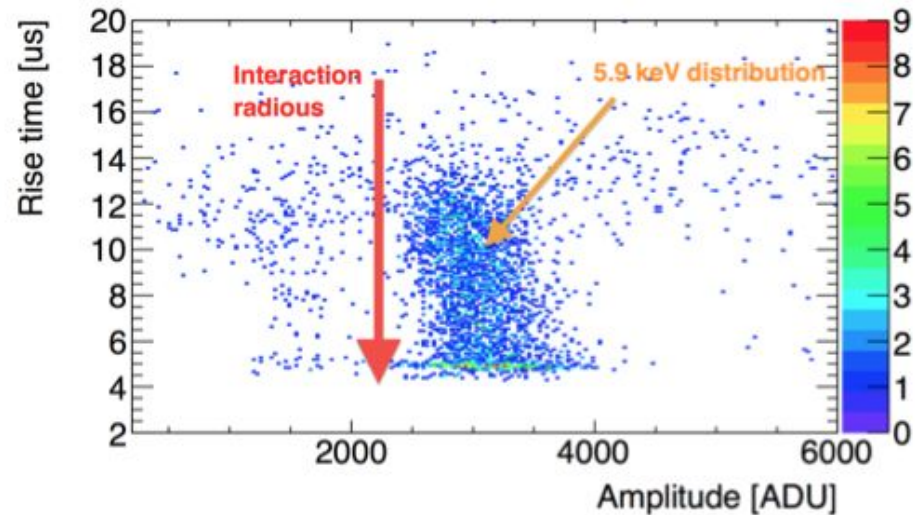


Surface

Primary e- drift time dispersion

$$\sigma(r) \propto (r/r_{\text{sphere}})^3$$

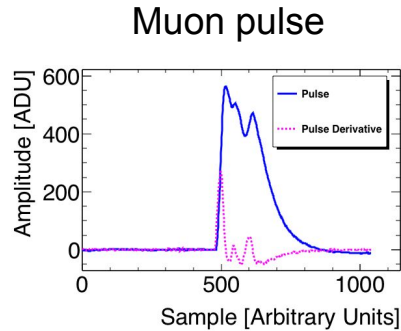
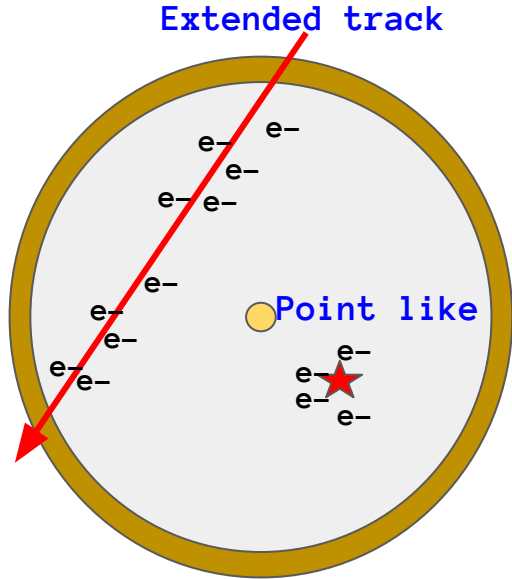
5.9 keV X-rays line



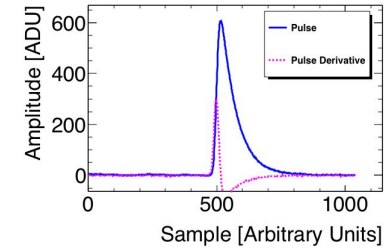
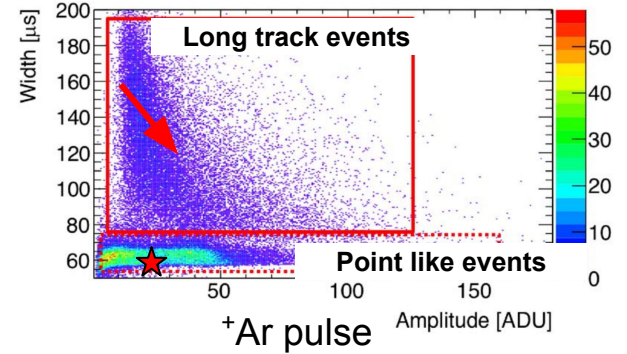
Rise time $\rightarrow \Delta t$ between 90% - 10% of pulse height

Background rejection capabilities-B

Event discrimination

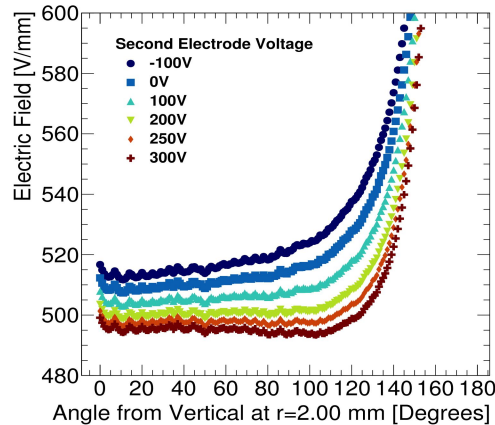
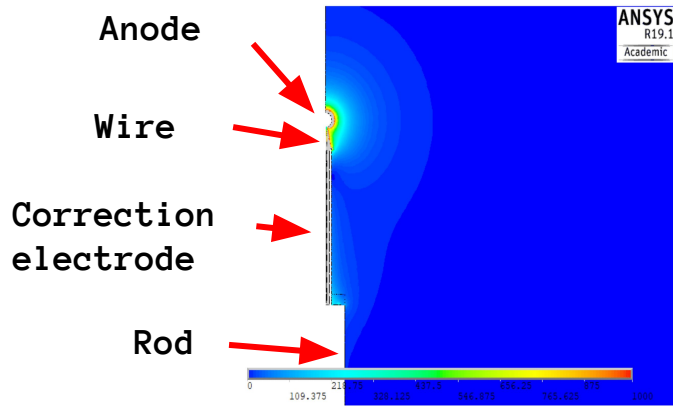
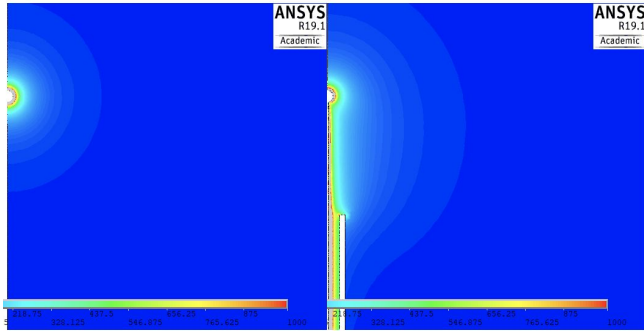


Amplitude = 575 ADU
Width (FWHM) = 155.5 μ s
Rise time = 18.2 μ s



Amplitude = 606 ADU
Width (FWHM) = 63.4 μ s
Rise time = 16.3 μ s

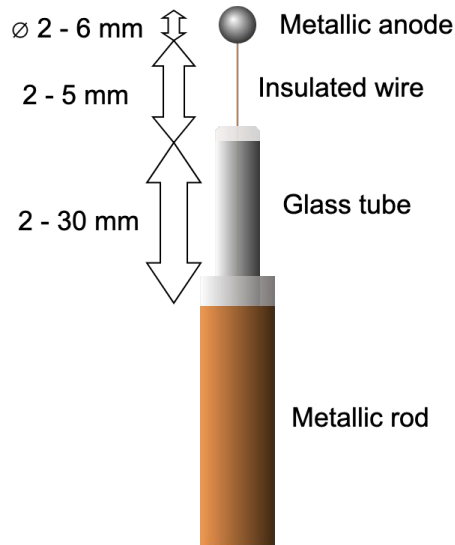
Electric field homogeneity



- Ideally, electric field:
 - Purely radial
 - Strength $1/r^2$
- Reality more complex, as support structure needed for sensor
 - $E=E(r,\theta)$
 - Non-uniform detector
 - Response
- Improved field uniformity by adding correction electrode

The resistive glass electrode

Design



Implementation



Provides

- Spark quenching
- Charge evacuation

Advantages

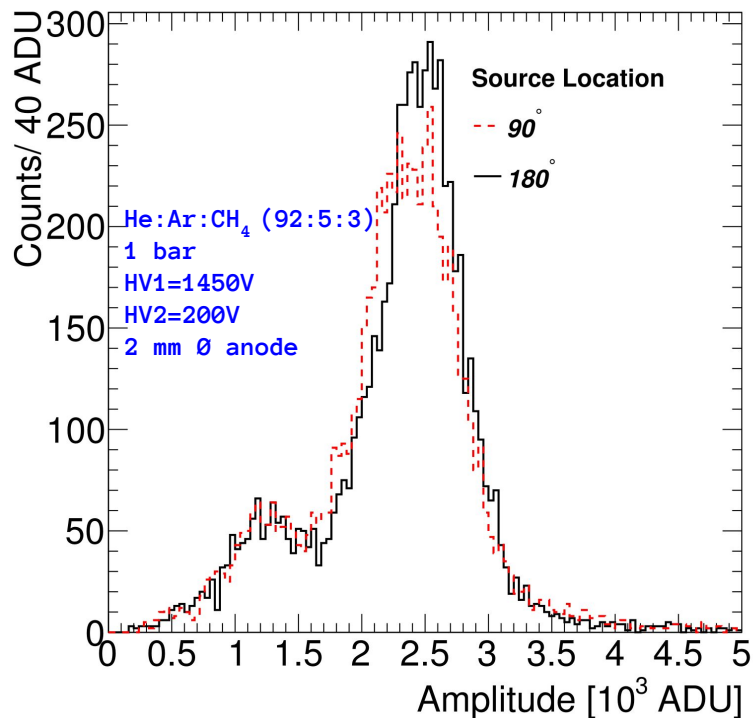
- Simple
- Symmetric
- Robust
- Low material budget

Material properties

- Soda-lime glass
- $\rho = 5.05 \times 10^{10} \Omega \cdot \text{cm}$
- $d = 2.1 - 2.25 \text{ g/cm}^3$
- $A = 14.48 \text{ mBq/g}$

Performance

Measurements with the SPC in the two positions



Detector stability

