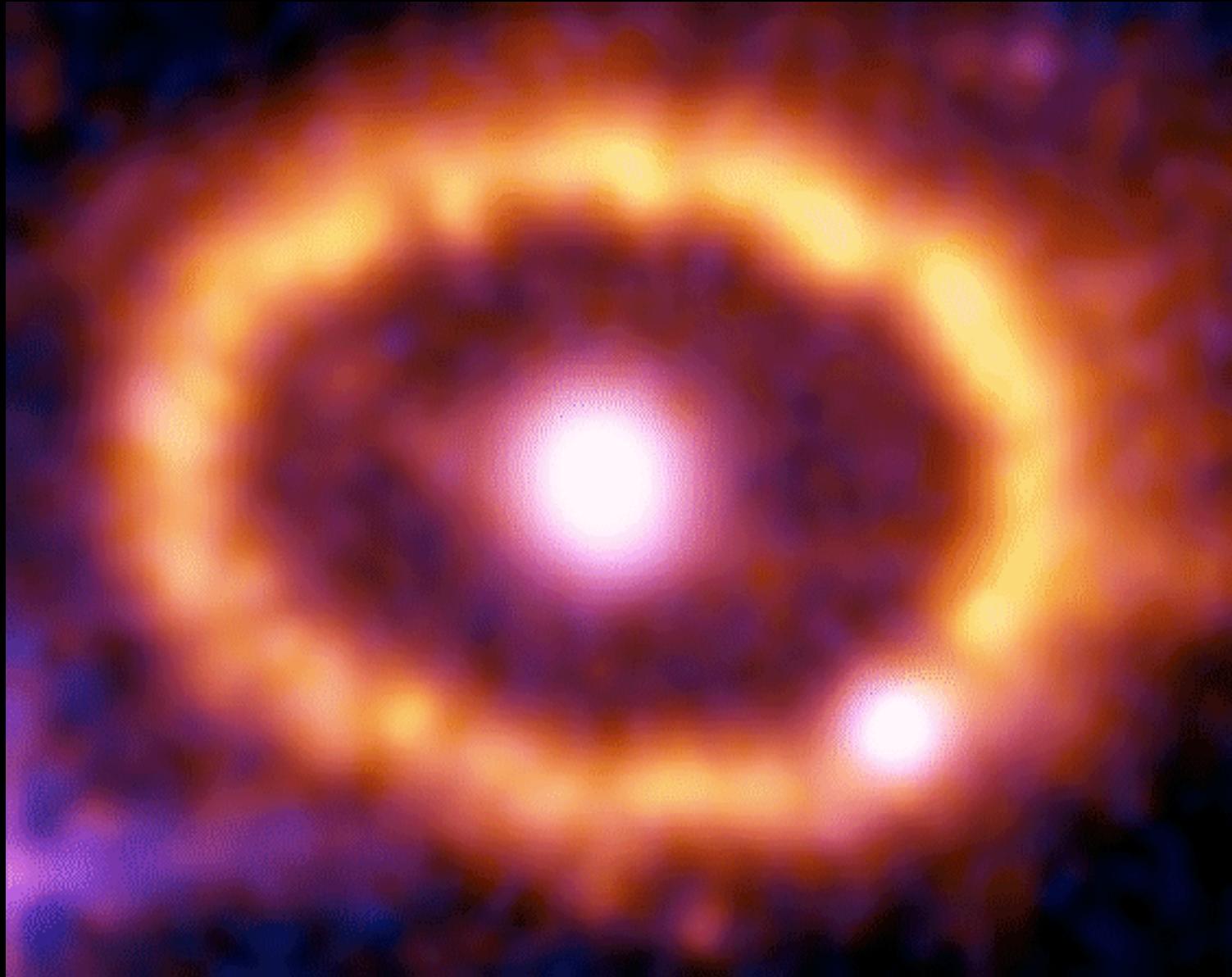




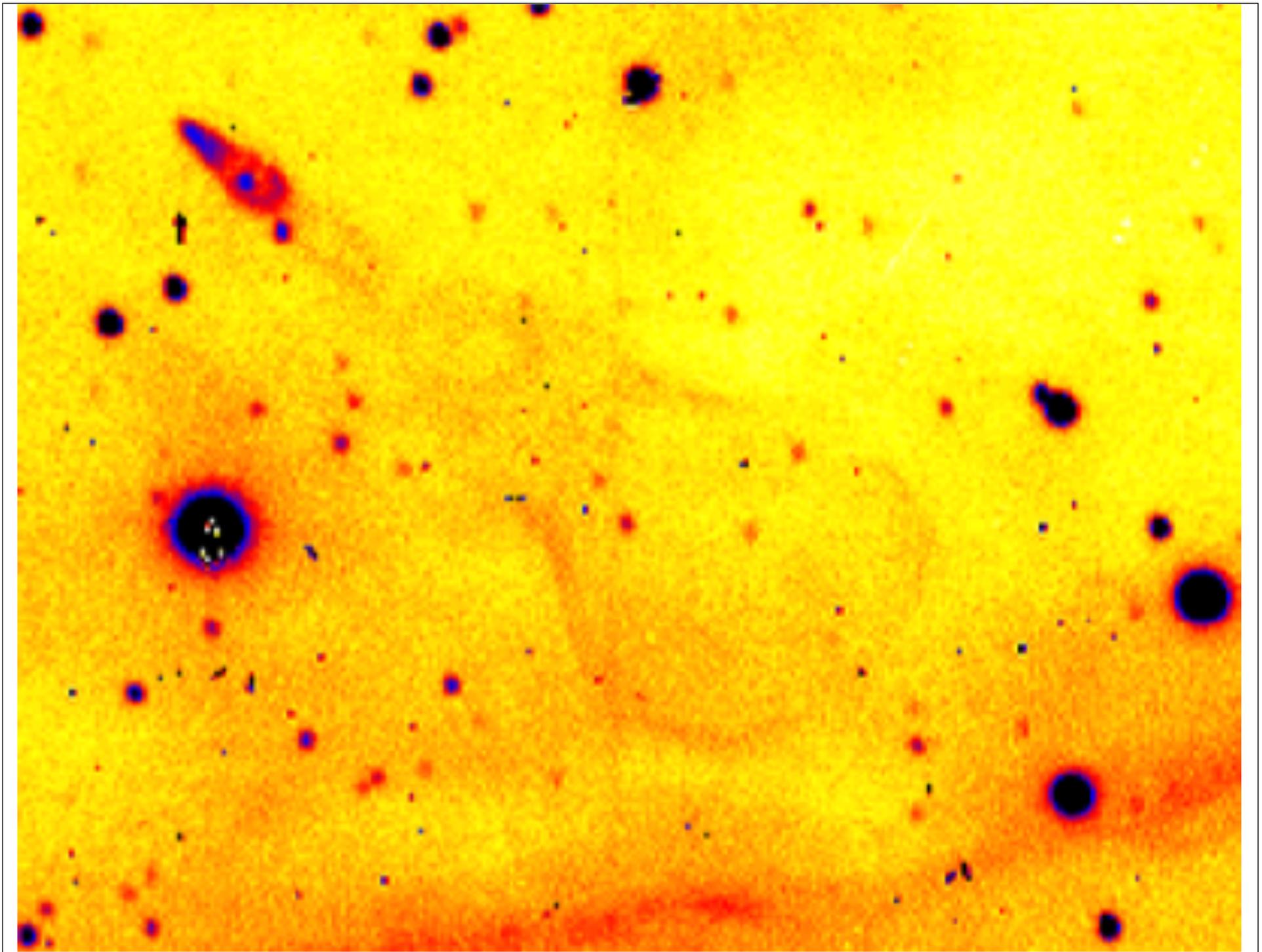
^{44}Ti and Core Collapse Supernovae

Alex Murphy

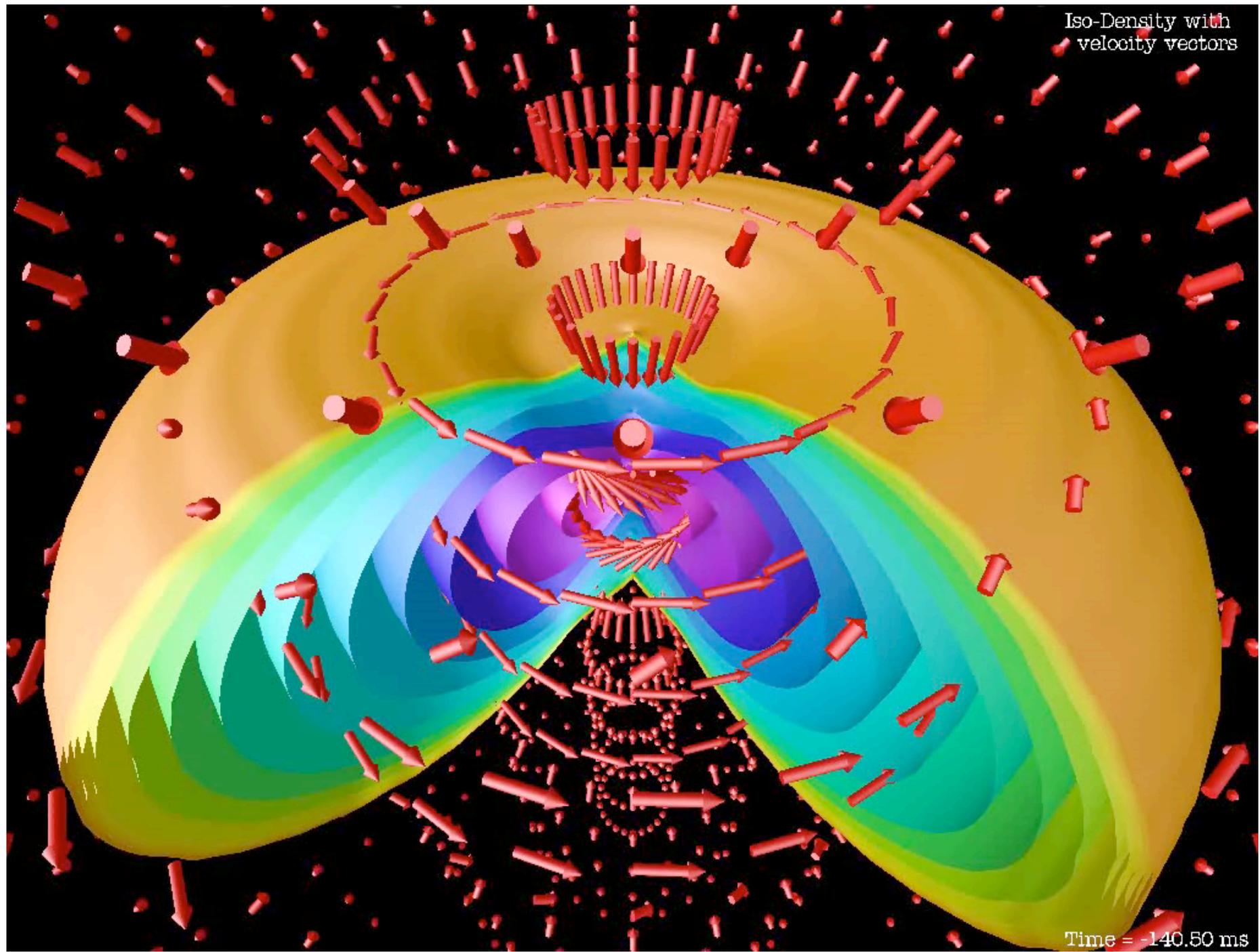








Iso-Density with
velocity vectors



Time = -140.50 ms

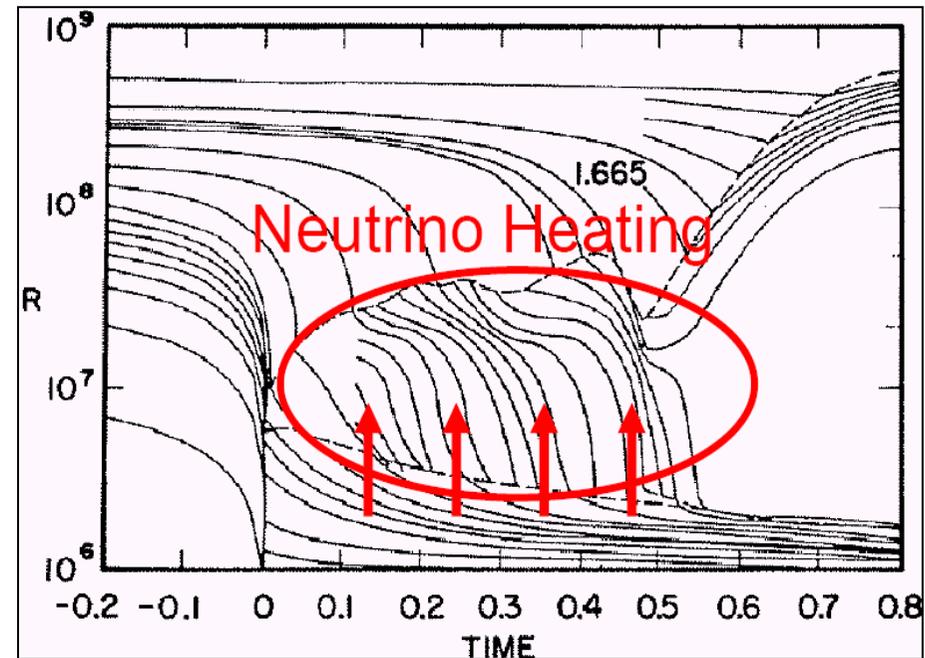
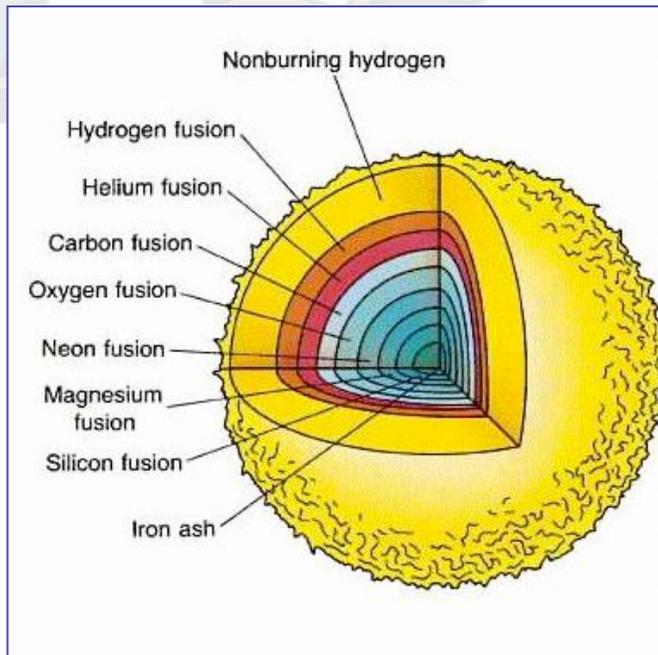
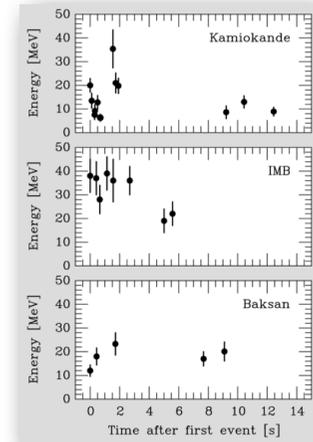
Core Collapse Supernovae are...

- Some of the most powerful explosions in the Cosmos
- Some of the most scrutinised objects in the Universe
- Responsible for production of many heavy elements
- Responsible for shaping our local distribution of stars
- Until recently there was consensus on the basic mechanism (perhaps there still is...)
- The best simulations still don't (*) reliably explode
- Extremely complex

We need a good diagnostic

The Neutrino Mechanism

- Massive star ($>8-10 M_{\odot}$)
- Stellar evolution \rightarrow onion-skin-like structure
- At maximum of BE/A , thermal support lost \rightarrow Collapse
- Huge flux of neutrinos “re-energises” explosion
- Neutrino driven wind – an excellent candidate site for the r-process



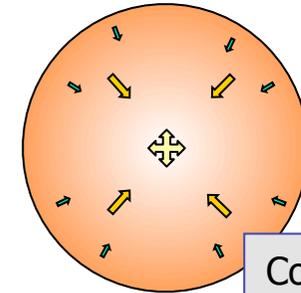
$>10 M_{\odot}$ evolves $\sim 10^7$ yr

Rapid collapse of core

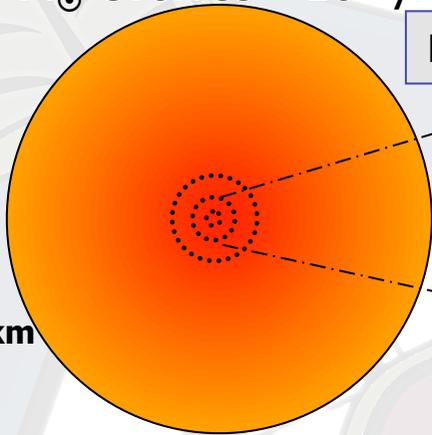
3000 km

Extreme temp: photo-dissociates nuclei back to protons, neutrons and alphas.

Core bounces

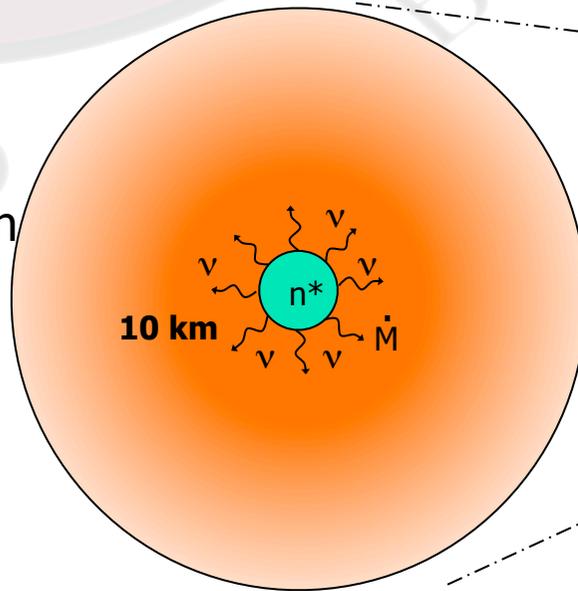


3×10^7 km



Alpha-rich freezeout, r-process in neutrino-driven wind

10 km



Huge thermal emission of neutrinos
 $\sim 5-10$ seconds

$\rho \sim \text{few} \times \rho_{\text{nuclear}}$

CCSN Simulation status

(Adapted from A. Burrows)

- Rapidly advancing field
- Consensus that the 'true' mechanism should not be 'marginal'
- Prompt shock ALWAYS fizzles
- 1D neutrino-driven model ALWAYS fizzles
- GR... makes no difference
- Improved neutrino physics... makes no difference
- EOS probably has no surprises
- All groups get the same results for all progenitors

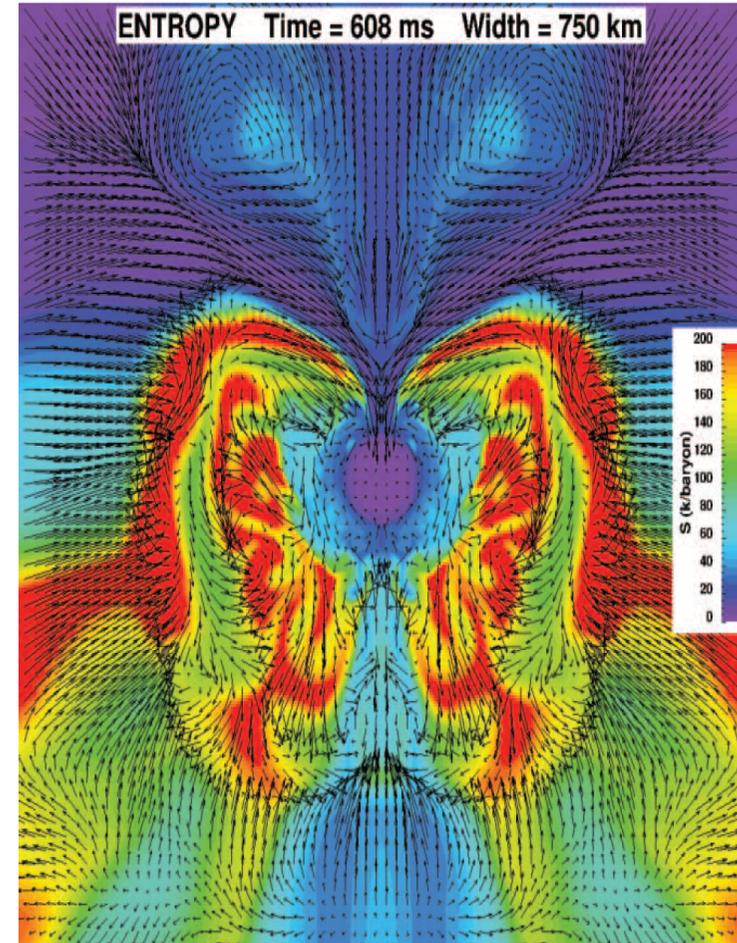
CCSN Simulation status

(Adapted from A. Burrows)

- 2D/3D model are showing more promise
- Asymmetry looks like it is the key
 - Convection
 - 'Fingers'
- The asymmetry probably comes from rotation
 - 'Naturally' produces bi-polar explosions
 - Collapse amplifies rotation
 - Increases gain region
 - Reduces effective gravity
 - Likely explanation of pulsar kicks
 - Likely source of gravitational radiation
 - Possible role of g-wave acoustics

The Acoustic Mechanism

- 'New' mechanism: A. Burrows *et al.* (ApJ 640 (April 2006) 878-890)
 - In-fall of matter on to the core induces strong gravity waves
 - These set up acoustic oscillations a few hundred milliseconds after core collapse.
 - Oscillations couple efficiently with the outer core/overlying material,
 - Intense sound waves radiated.
 - Appears to lead to robust explosion



Parametric instability?

International Journal of High-Energy Physics

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Physics Letters B 705 (2011) 148–151

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Parametric instability induced scalar gravitational waves from a model pulsating neutron star

Charles H.-T. Wang^{a,b}, Paolo M. Bonifacio^{a,*}, Robert Bingham^{b,c}, J. Tito Mendonça^{d,b}

^a SUPA Department of Physics, University of Aberdeen, King's College, Aberdeen AB24 3UE, UK

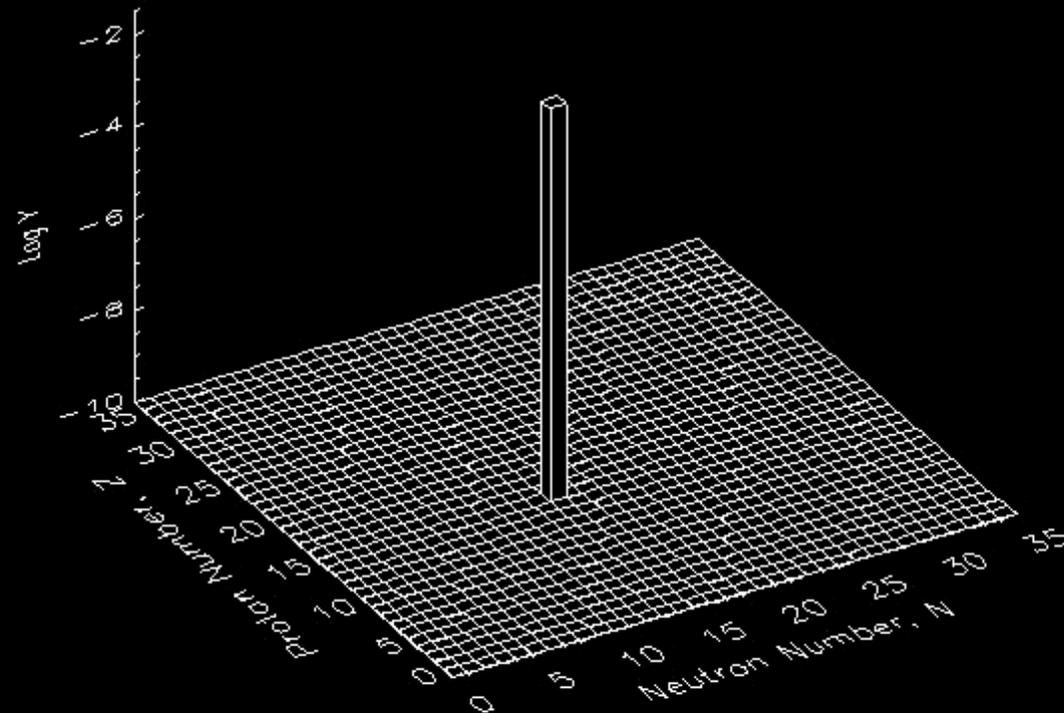
^b STFC Rutherford Appleton Laboratory, Chilton, Didcot, Oxfordshire OX11 0QX, UK

^c SUPA Department of Physics, University of Strathclyde, Glasgow G4 0NG, UK

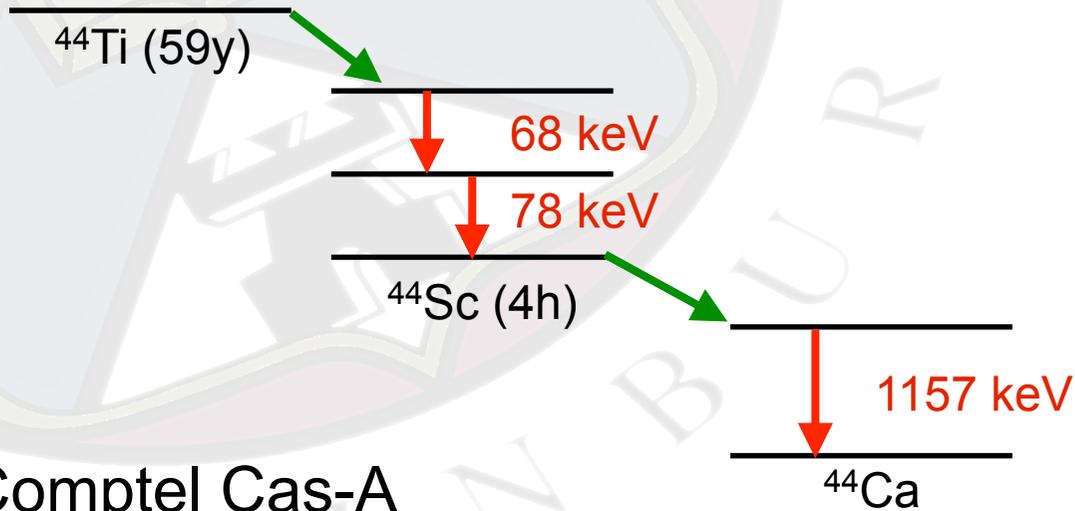
^d GOLP/Centro de Física de Plasmas, Instituto Superior Técnico, 1049-001 Lisboa, Portugal

Nucleosynthesis

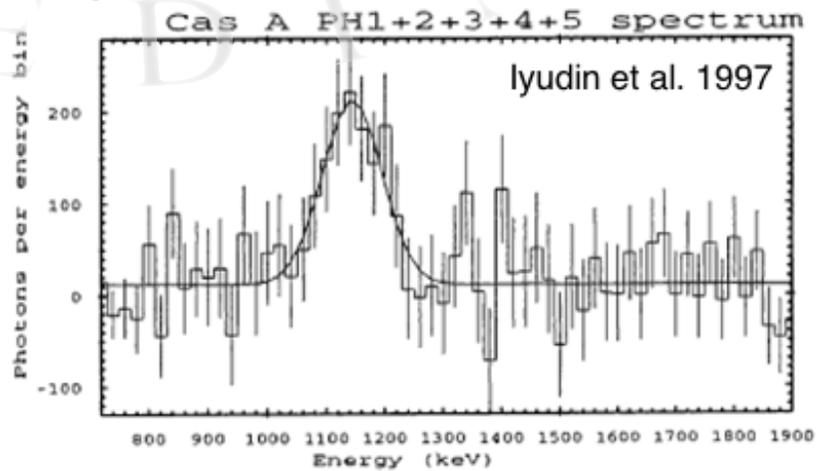
$t \text{ (s)} = 6.74200\text{e-}20$ $T_9 = 5.50$ $\rho \text{ (g/cc)} = 1.00000\text{e+}07$



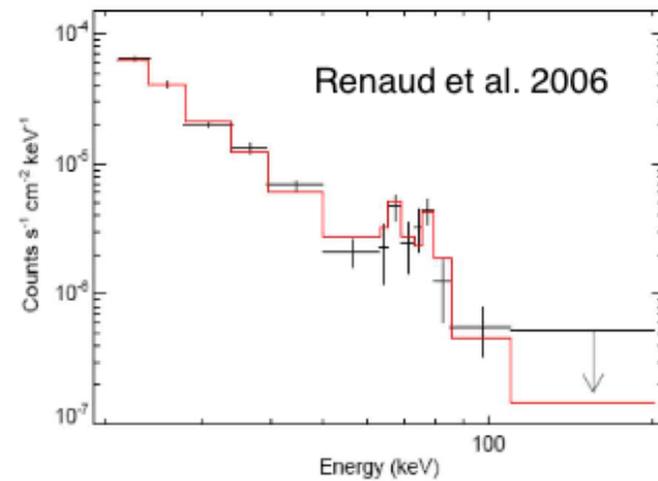
Gamma-ray emission from ^{44}Ti



Comptel Cas-A

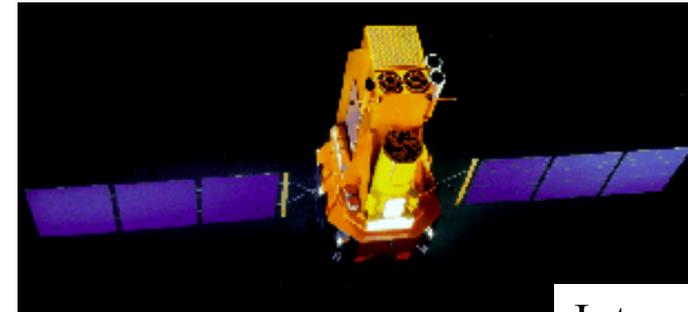


IBIS Cas-A

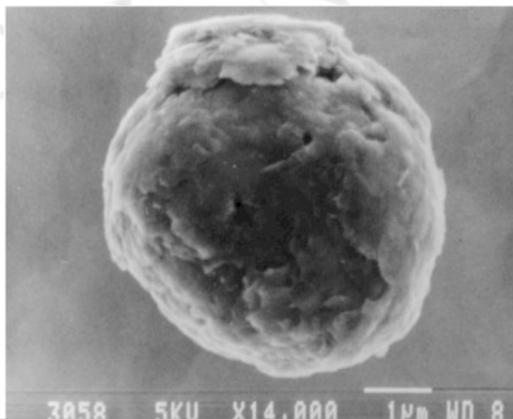


The Importance of ^{44}Ti

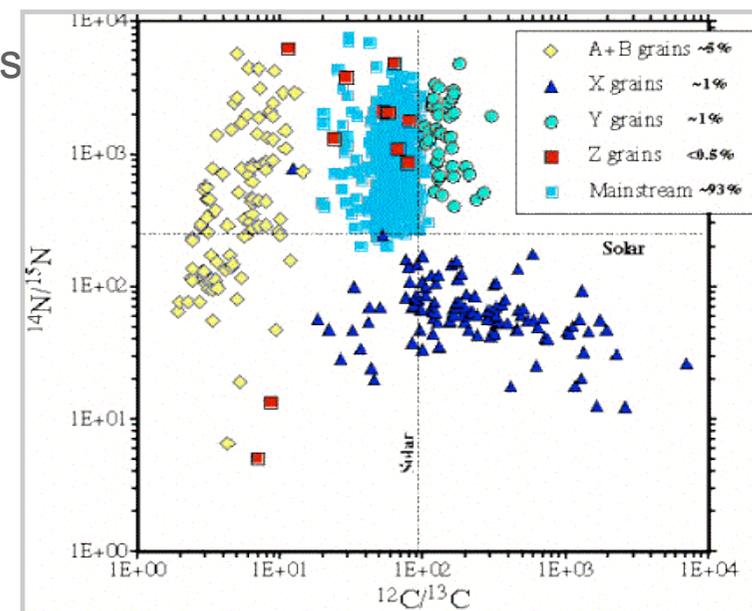
- ^{44}Ti that is ejected will become a γ -ray emitter
- Cassiopeia-A, Vela, not SN1987A
- $\tau=60$ yrs, $E_\gamma=1.157$ MeV
- 'Easily' observable
- INTEGRAL & other missions
- Also Meteoritic data
- Enrichment of ^{44}Ca in type X presolar grains



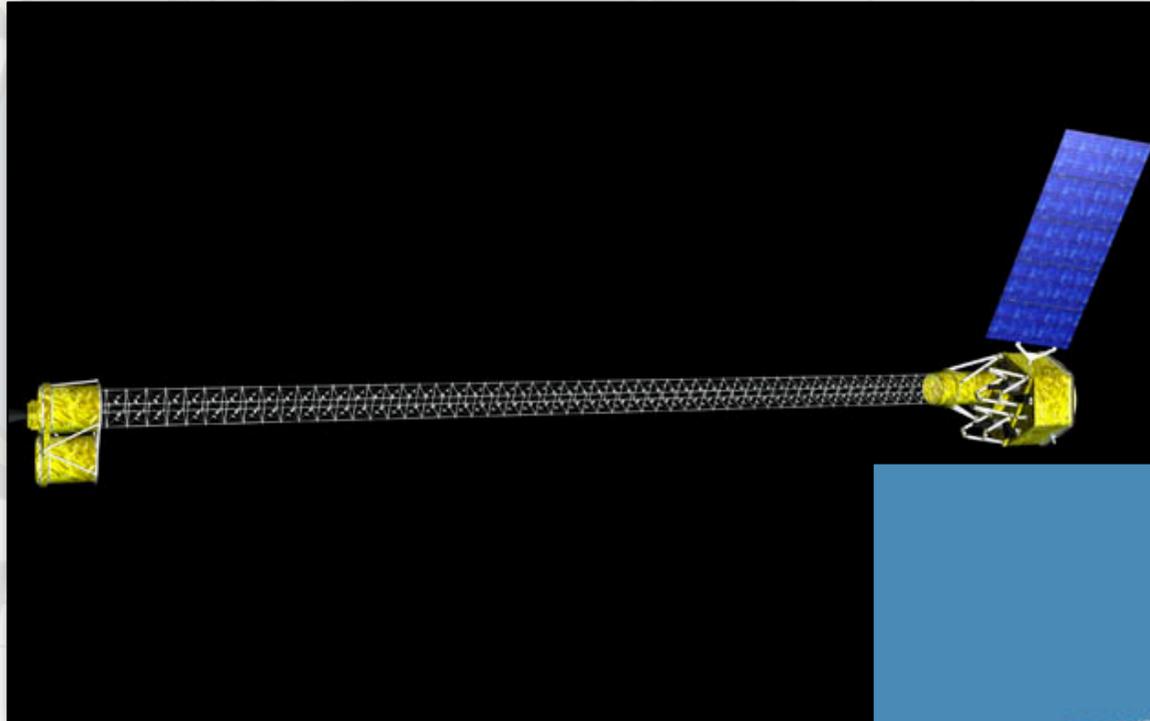
Integral



A grain from the Murchison Meteorite



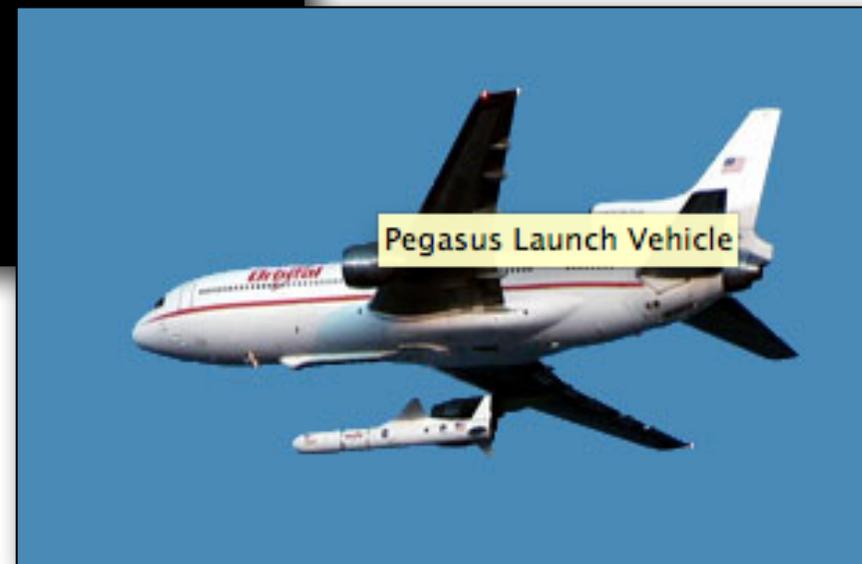
NuSTAR (the 'other' one)



Hard X-ray imager
6-80 keV

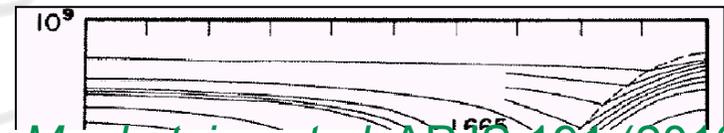
~NASA version of ESAs
GRI (UK: Tony Bird,
Bruce Swinyard, ASM)

LAUNCH:
14th March 2012, 8am

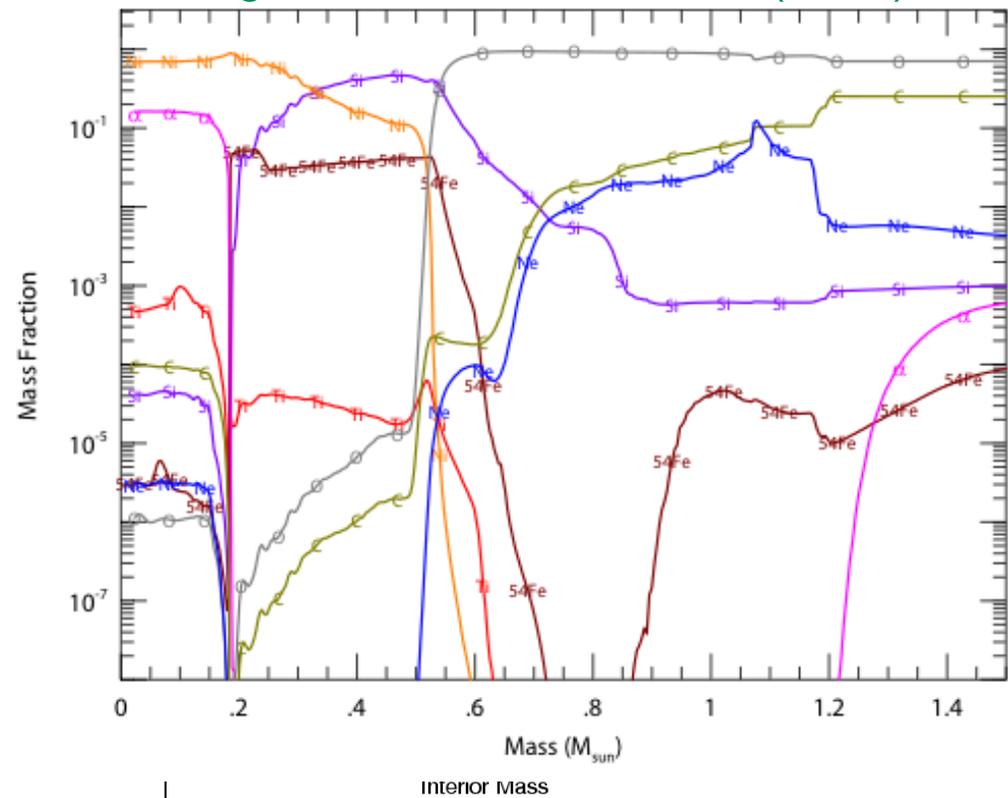


The Importance of ^{44}Ti

- Amount ejected sensitively depends on location of the 'mass cut'
- Material that 'falls back' is not available for detection
- ^{44}Ti yield a sensitive diagnostic of the explosion mechanism
- Thus, very useful for models to make comparisons against



Magkotsios et al: APJS 191 (2010) 66



Timmes et al. (1996)

The KEY Question

Could a measurement of ^{44}Ti gamma rays be a diagnostic of the underlying explosion mechanism?

Do the different explosion mechanisms affect the location of the mass cut?

- “there should be some interesting differences on average between the predictions/expectations of the various mechanisms (neutrino, acoustic, magnetic, etc.), we have yet truly to determine these.”

Adam Burrows, Princeton, Jan 2012

- *“By accident we noticed e.g. in simple parametrized explosions with the same explosion energy, that pistons (like used by Woosley and Heger) produce less ^{44}Ti than thermal bombs (like used by Nomoto, myself, Umeda). The point is just that in the innermost ejecta the entropies are higher.”*

Friedel Thielemann, Basel, Jan 2012

Key Reactions

Reaction rate sensitivity studies:

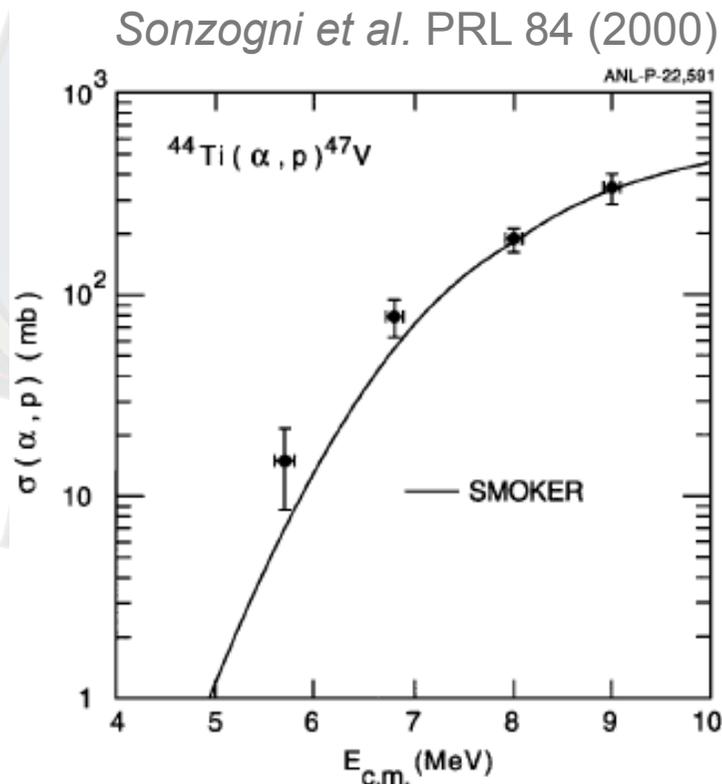
- *The et al: ApJ 504 (1998) 500*
- *Magkotsios et al: APJS 191 (2010) 66*
- Papers agree, $^{44}\text{Ti}(\alpha, p)$ most important reaction
- Importance stems from it being the bottle neck in reaction flow as material drops out of QSE

ORDER OF IMPORTANCE OF REACTIONS PRODUCING ^{44}Ti AT $\eta = 0^a$

Reaction	Slope
$^{44}\text{Ti}(\alpha, p)^{47}\text{V}$	-0.394
$\alpha(2\alpha, \gamma)^{12}\text{C}$	+0.386
$^{45}\text{V}(p, \gamma)^{46}\text{Cr}$	-0.361
$^{40}\text{Ca}(\alpha, \gamma)^{44}\text{Ti}$	+0.137
$^{57}\text{Co}(p, n)^{57}\text{Ni}$	+0.102
$^{36}\text{Ar}(\alpha, p)^{39}\text{K}$	+0.037
$^{44}\text{Ti}(\alpha, \gamma)^{48}\text{Cr}$	-0.024
$^{12}\text{C}(\alpha, \gamma)^{16}\text{O}$	-0.017
$^{57}\text{Ni}(p, \gamma)^{58}\text{Cu}$	+0.013
$^{58}\text{Cu}(p, \gamma)^{59}\text{Zn}$	+0.011
$^{36}\text{Ar}(\alpha, \gamma)^{40}\text{Ca}$	+0.008
$^{44}\text{Ti}(p, \gamma)^{45}\text{V}$	-0.005
$^{57}\text{Co}(p, \gamma)^{58}\text{Ni}$	+0.002
$^{57}\text{Ni}(n, \gamma)^{58}\text{Cu}$	+0.002
$^{54}\text{Fe}(\alpha, n)^{57}\text{Ni}$	+0.002
$^{40}\text{Ca}(\alpha, p)^{43}\text{Sc}$	-0.002

^a Order of importance of reactions producing ^{44}Ti at $\eta = 0$ according to the slope of $X(^{44}\text{Ti})$ near the standard reaction rates.

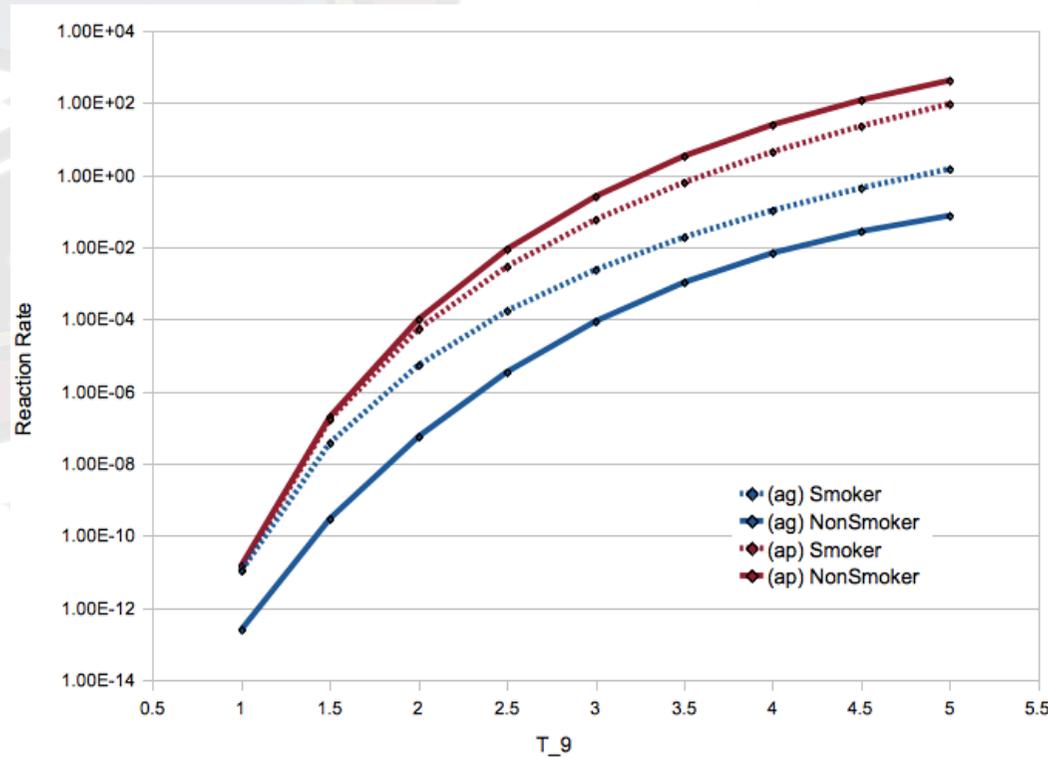
$^{44}\text{Ti}(\alpha, p)$ Present status



- Astrophysical region is 3-6 MeV
- Hoffman et al. APJ 715 (2010) 1383
- New evaluation of $^{44}\text{Ti}(\alpha, p)$ reaction rate
- Conclude that $^{44}\text{Ti}(\alpha, p)$ uncertainty has been underestimated (x3)
- Sonzogni compared to SMOKER
- NON-SMOKER provides significant update

Data achieved with $\sim 10^5$ pps on target

$^{44}\text{Ti}(\alpha, p)$ Reaction Rates



SMOKER → NONSMOKER: $^{44}\text{Ti}(\alpha, p)$ little effect
 $^{44}\text{Ti}(p, \gamma)$ major effect

$^{44}\text{Ti}(\alpha, p)$ Reaction Rates

- NON-SMOKER includes better treatment of isospin suppression for alpha-capture reactions on $N = Z$ nuclei
- Vockenhuber *et al.* states that $^{44}\text{Ti}(\alpha, p)$ NON-SMOKER rate is 100x smaller than SMOKER rate
[J. Phys G: Nucl. Part. Phys. 35(2008)]
- Rauscher [priv. comm.] says this is in error. Only a factor of 20.
- Consequences if $^{44}\text{Ti}(\alpha, \gamma)$ rate $>$ $^{44}\text{Ti}(\alpha, p)$ rate



Proposal

Direct measurement of $^{44}\text{Ti}(\alpha, p)$ at astrophysical energies

- Sonzogni made their own ^{44}Ti via $^{45}\text{Sc}(p, 2n)^{44}\text{Ti}$.
 - about 4×10^{15} atoms were used
- Priv. Comm.: Sonzogni approach has limited further capability
- Production of a ^{44}Ti target is viable: Daniel Bemmerer leads; Timescale 'few years'.
- Production of a ^{44}Ti ISOL beam is viable: beam development at GANIL, TRIUMF. Timescale is 'few years'
- Production of offline ^{44}Ti beam... this proposal

As of yesterday...

Beam request for ISOLDE 2012

April 2 to December 2, 2012

PLEASE COMPLETE ALL INFORMATION REQUESTED AND RETURN BY JANUARY 30, 2012

Experiment IS544

Measurement of the $^{44}\text{Ti}(\alpha, p)^{47}\text{V}$ reaction cross section, of relevance to gamma-ray observation of core collapse supernovae, using reclaimed ^{44}Ti

Spokesperson: A. Murphy

Total shifts approved: 28

Shifts used in 2011: 0

Remaining shifts: 28 (if shifts requested in January 2012 are approved)

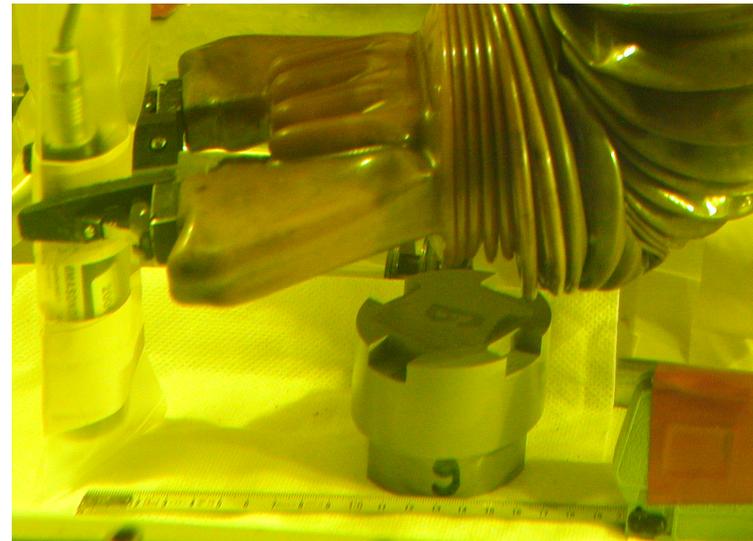
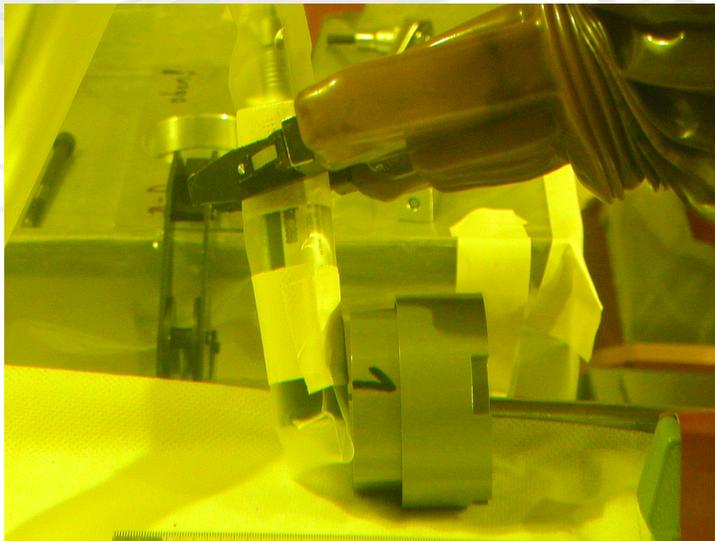
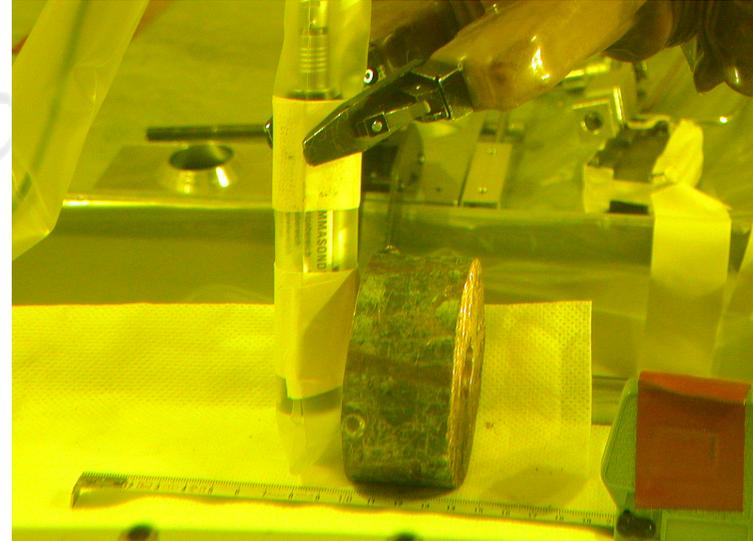
ERAWAST

Exotic Radionuclides from Accelerator Waste for Science and Technology

- A project to utilise long lived activity generated in PSI beam dumps
- Copper beam bumps, exposed to 1.5mA protons for ~12 years, dismantled ~15 years ago.
- ^{26}Al , ^{59}Ni , ^{53}Mn , ^{60}Fe or ^{44}Ti have been separated.
- SINQ facility material also available: other isotopes, e.g. ^{182}Hf



Applications: Nuclear physics, nuclear astrophysics, Geophysics, Radiopharmacy, AMS, RIMS,...



^{44}Ti availability

“The bottom line on the ^{44}Ti is that there is more of it than we are going to know what to do with. They have actually stopped processing the copper beam dump because they have been having a lot more luck with stainless steel test samples that were put in their SINQ facility (spallation neutron source). These samples have the added benefit that no ^{60}Co is produced, so they don't have to wait very long before they can handle the material.

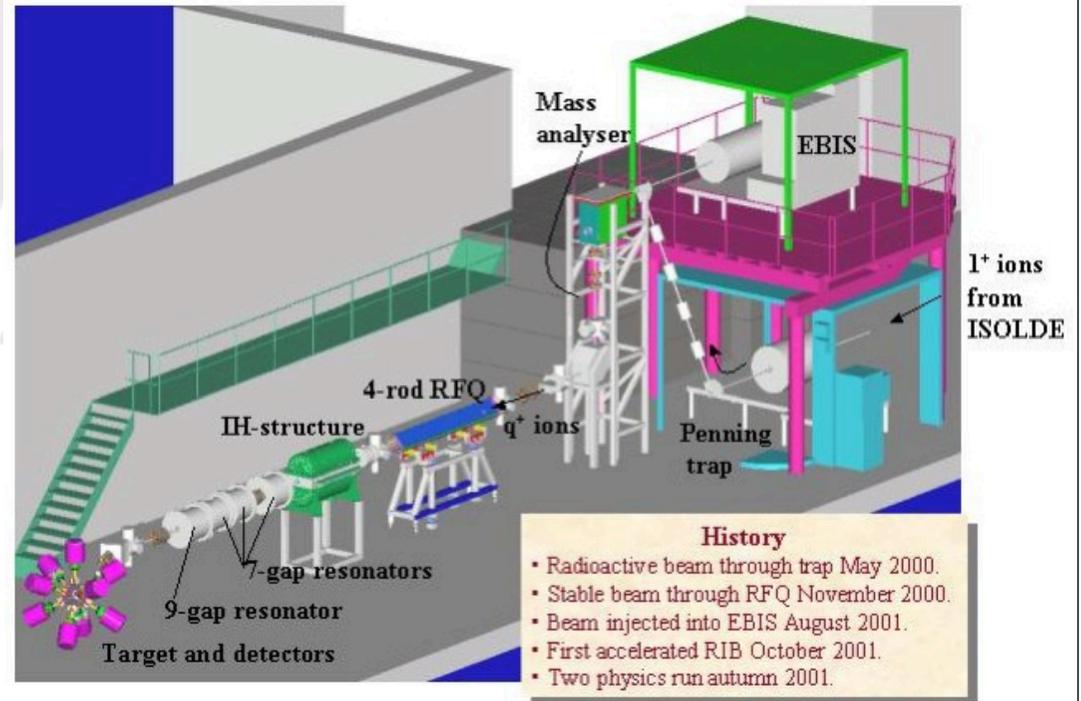
“They have determined that they have **300MBq** of ^{44}Ti in the samples produced to-date, and there are more experiments yet to be run. There are also tests on V-metal and V+Ti-metal and each of these samples is expected to contain **500MBq** of ^{44}Ti . There is plenty to share with the ^{44}Ti target experiment, and the medical people have moved on and are now producing their own. **The only limits for how much ^{44}Ti we can use to produce a beam will be set by the safety people.**”

Jennifer Fallis, 6/9/2011

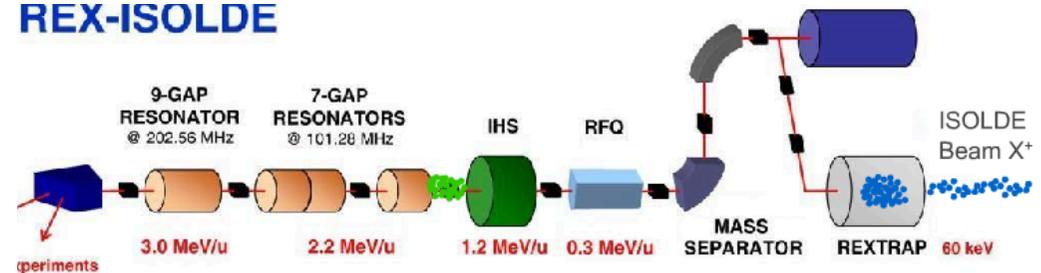
^{44}Ti Beam Production

- Development ongoing at ISOLDE, TRIUMF and GANIL
- 1 week at 10^5 pps is about 10^{12} ions
- With 10^{18} ions, 0.0001 total is efficiency required for a significantly improved measurement
- $^{44}\text{TiF}_4$ deposited on Ta foil
- FEBIAD ion source; extract TiF_3^+ molecules
- Accelerate, dissociate molecule, accelerate $^{44}\text{Ti}^{4+}$ ions
- Expected efficiency... ~0.1 - 1%

CERN REX - ISOLDE

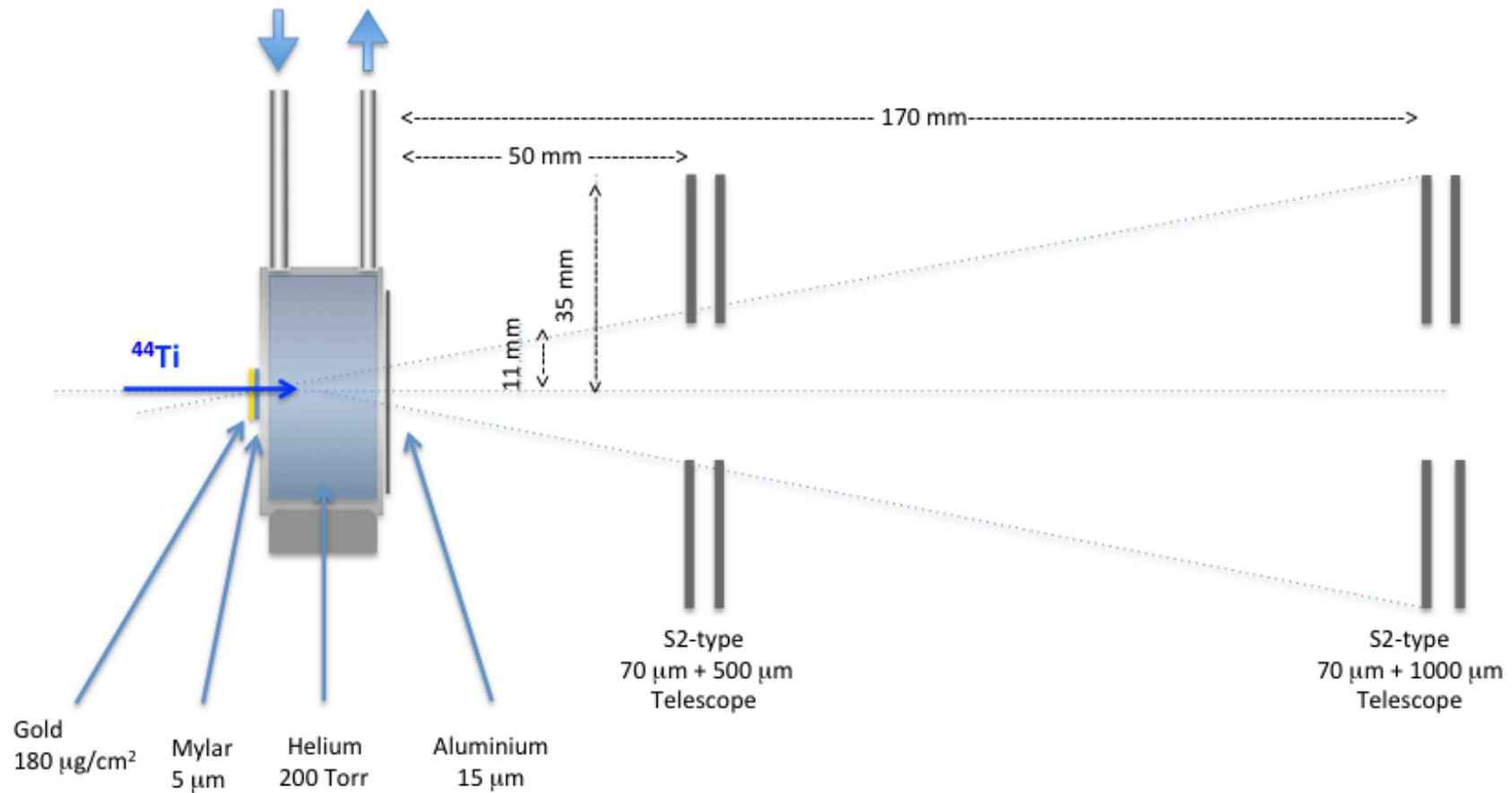


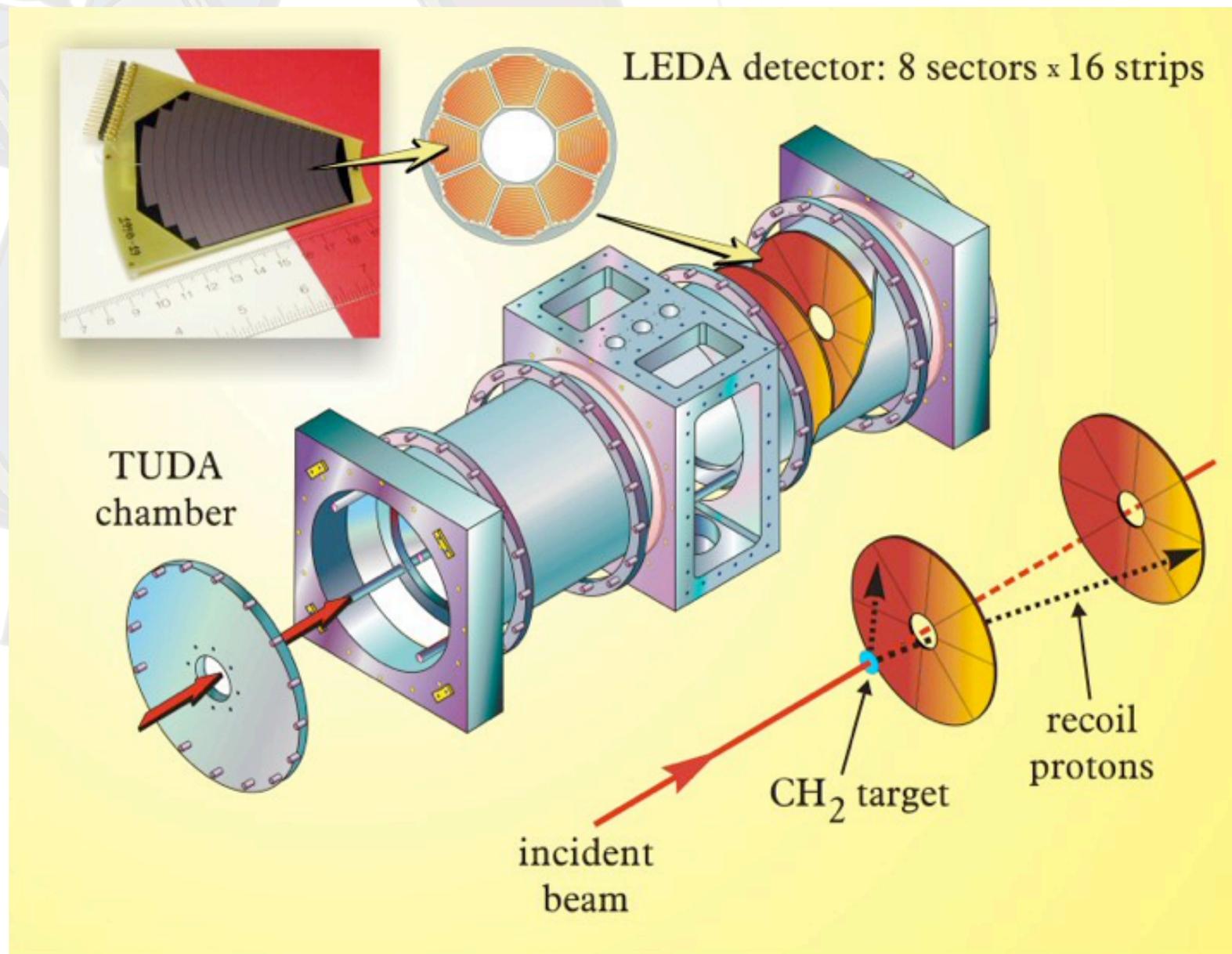
REX-ISOLDE

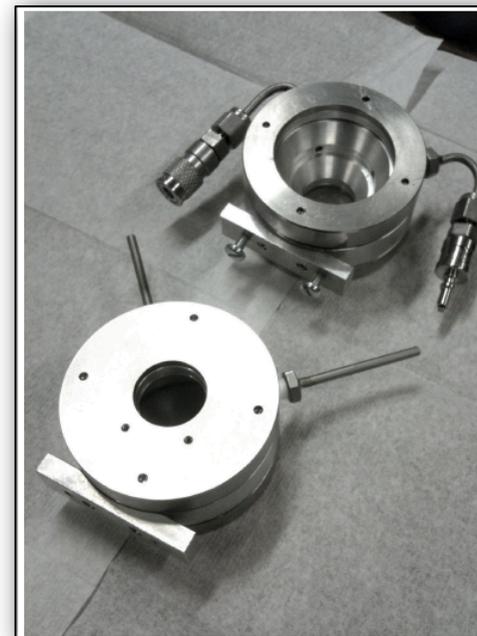
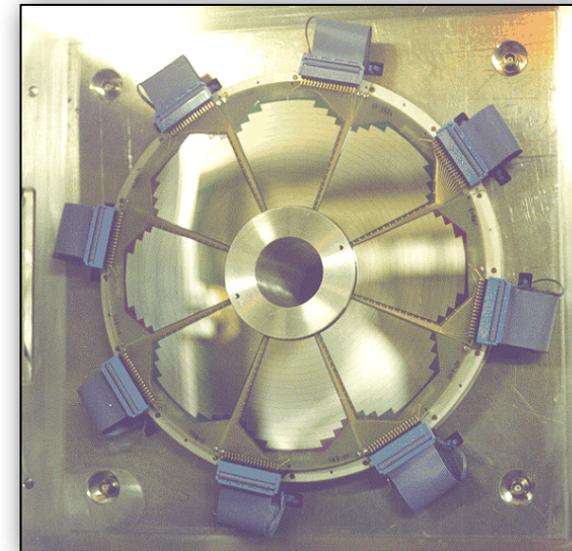
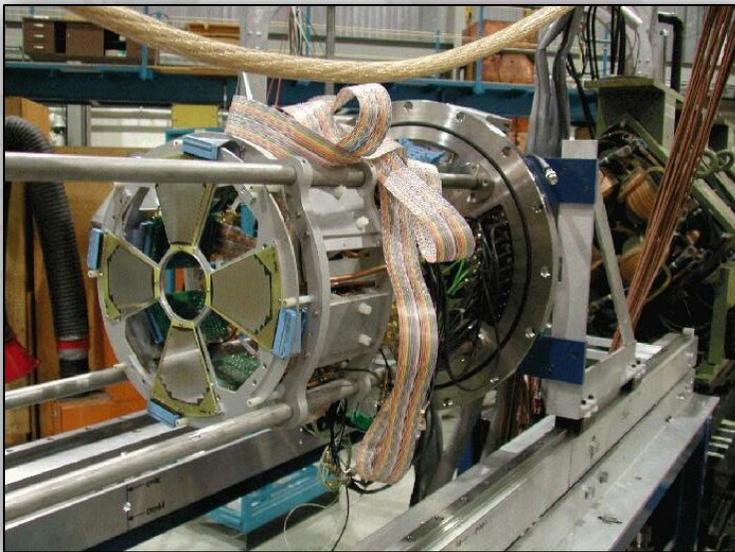


^{44}Ti Experiment configuration

Proposed Set-up for $^{44}\text{Ti}(\alpha, p)$ Experiment at ISOLDE





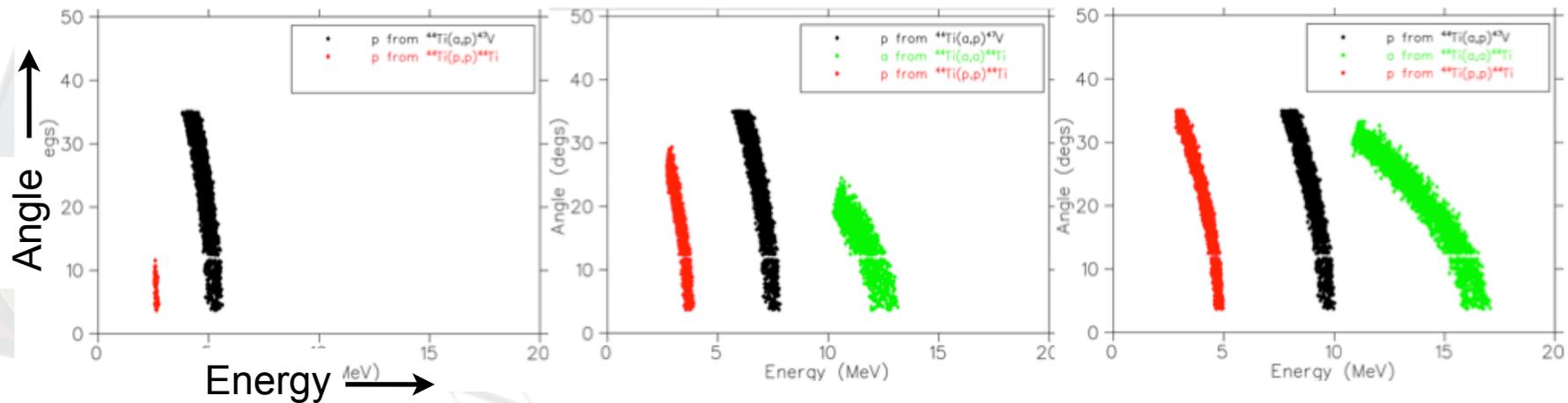


^{44}Ti Experiment Simulations

3 MeV

4 MeV

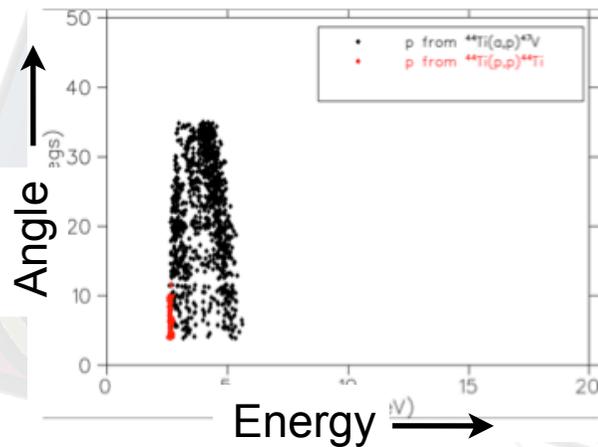
5 MeV



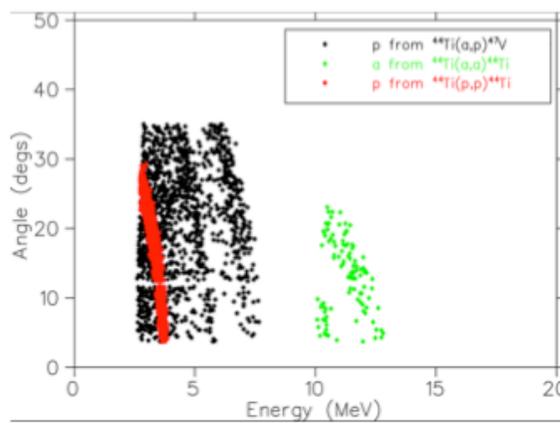
Clean separation

^{44}Ti Experiment Simulations

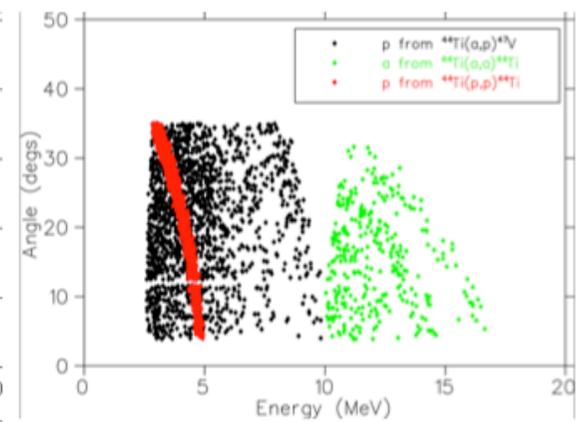
3 MeV



4 MeV



5 MeV



Very unrealistic population of excited states! 'Worst case'

^{44}Ti Event rate estimate

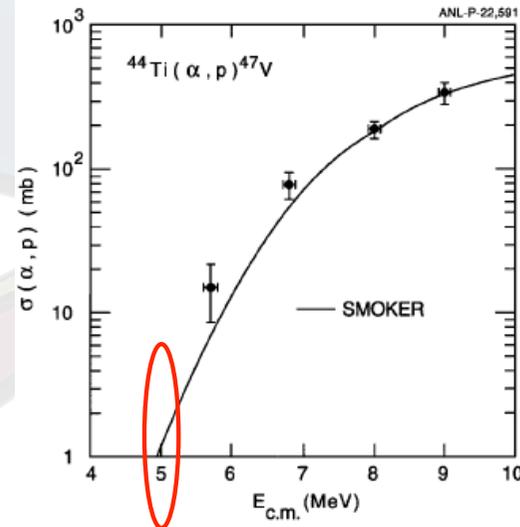


Table 1: Summary of expected rates of events.

Beam energy ¹ (MeV/u)	CM Energy ² (MeV)	Efficiency (%)	Cross section ³ (mb)	Rate (min ⁻¹)	Rate (hour ⁻¹)	Rate (day ⁻¹)
1.41	3.0	19.2	0.004	0.005	0.29	9
1.67	4.0	18.9	0.1	0.16	9.5	229
1.93	5.0	18.6	2.0	3.1	187	4500

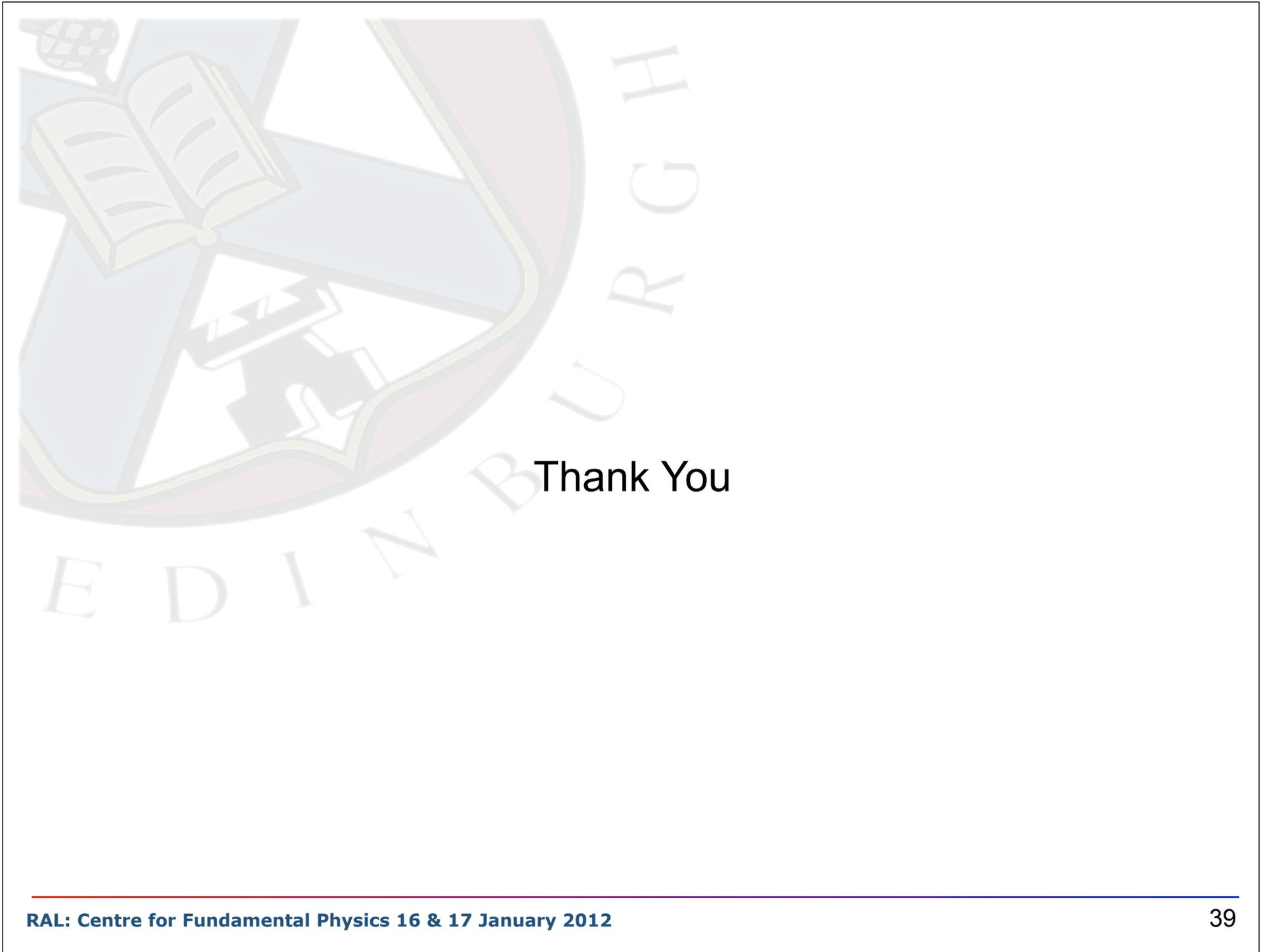
¹Beam energy to be supplied by ISOLDE.

²Centre of mass energy for $^{44}\text{Ti}(\alpha, p)$ reactions at the start of the gas cell.

³Estimated by extrapolation of Sonzogni *et al.*.

Scheduling

- ^{44}Ti ions are available 'now'
- Experimental apparatus (gas cell, silicon, DAQ) is available now
- Gas cell windows, vacuum chamber need a little work
- Monte Carlo optimisation is ongoing
- Experimental team is already sufficient (additional collaborators welcome)
- Safety situation?
- Assessment of measures required to prevent activation?
- Window of opportunity during CERN 'shutdown year'?
 - We have strong support of ISOLDE team



Thank You