

# Particle Physics - Measurements and Theory



## Outline

Natural Units

Relativistic Kinematics

Particle Physics Measurements

Lifetimes

Resonances and Widths

Scattering

Cross section

Collider and Fixed Target Experiments

Conservation Laws

Charge, Lepton and Baryon number,

Parity, Quark flavours

Theoretical Concepts

Quantum Field Theory

Klein-Gordon Equation

Anti-particles

Yukawa Potential

Scattering Amplitude - Fermi's Golden Rule

Matrix elements

# Particle Physics Units



## Particle Physics

is relativistic and quantum mechanical

→  $c = 299\,792\,458\text{ m/s}$

→  $\hbar = h/2\pi = 1.055 \cdot 10^{-34}\text{ Js}$

## Length

size of proton:  $1\text{ fm} = 10^{-15}\text{ m}$

## Lifetimes

as short as  $10^{-23}\text{ s}$

## Charge

$1\text{ e} = -1.60 \cdot 10^{-19}\text{ C}$

## Energy

Units:  $1\text{ GeV} = 10^9\text{ eV}$  --  $1\text{ eV} = 1.60 \cdot 10^{-19}\text{ J}$

use also MeV, keV

## Mass

in  $\text{GeV}/c^2$ , rest mass is  $E = mc^2$

## Natural Units

Set  $\hbar = c = 1$

→ Mass [ $\text{GeV}/c^2$ ], energy [ $\text{GeV}$ ]  
and momentum [ $\text{GeV}/c$ ] in  $\text{GeV}$

→ Time [ $(\text{GeV}/\hbar)^{-1}$ ], Length [ $(\text{GeV}/\hbar c)^{-1}$ ]  
in  $1/\text{GeV}$  area [ $(\text{GeV}/\hbar c)^{-2}$ ]

## Useful relations

$\hbar c = 197\text{ MeV fm}$

$\hbar = 6.582 \cdot 10^{-22}\text{ MeV s}$

# Particle Physics Measurements



How do we measure particle properties and interaction strengths?

## Static properties

Mass                      How do you weigh an electron?  
 Magnetic moment              couples to magnetic field  
 Spin, Parity

## Particle decays

Lifetimes	Force	Lifetimes
Resonances & Widths	Strong	$10^{-23}$ -- $10^{-20}$ s
Allowed/forbidden	El.mag.	$10^{-20}$ -- $10^{-16}$ s
Decays	Weak	$10^{-13}$ -- $10^3$ s
Conservation laws		

## Scattering

Elastic scattering               $e^- p \rightarrow e^- p$   
 Inelastic annihilation       $e^+ e^- \rightarrow \mu^+ \mu^-$

Cross section	Force	Cross sections
total $\sigma$	Strong	$O(10 \text{ mb})$
Differential $d\sigma/d\Omega$	El.mag.	$O(10^{-1} \text{ mb})$
Luminosity $L$	Weak	$O(10^{-1} \text{ pb})$
Particle flux		
Event rate $N$		

# Relativistic Kinematics



## Basics

4-momentum

Invariant mass

Four-vector notation

$$p^\mu = \left( \frac{E}{c}, p_x, p_y, p_z \right)$$

$$p^2 = p^\mu \cdot p_\mu = (E/c)^2 - \vec{p}^2 = m^2 c^2$$

## Useful Lorentz boosts relations

set  $\hbar = c = 1$

$$\gamma = E/mc^2 = E/m$$

$$\gamma\beta = pc/mc^2 = p/m$$

$$\beta = pc/E = p/E$$

invariant mass

$$m^2 = E^2 - p^2$$

$$\gamma = 1/\sqrt{1 - \beta^2}$$

$$\beta = \sqrt{1 - 1/\gamma^2}$$

## 2-body decays

$$P_0 \rightarrow P_1 P_2$$

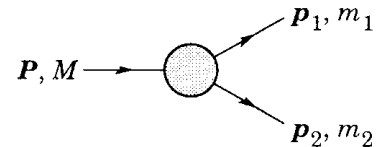
work in  $P_0$  rest frame

$$p^\mu = (m_0, \vec{0}) \quad p_1^\mu = (E_1, \vec{p}_1) \quad p_2^\mu = (E_2, \vec{p}_2)$$

$$p_2^\mu = (p^\mu - p_1^\mu) \Rightarrow p_2^2 = (p^\mu - p_1^\mu)^2 = p^2 + p_1^2 - 2p \cdot p_1$$

$$m_2^2 = m_0^2 + m_1^2 - 2m_0 E_1$$

$$E_1 = \frac{m_0^2 + m_1^2 - m_2^2}{2m_0} \quad |\vec{p}_1| = |\vec{p}_2|$$



Example:  $\pi^+ \rightarrow \mu^+ \nu_\mu$

work in  $\pi^+$  rest frame

use  $m_\nu^2 = 0$

$$p^\mu = (m_\pi, \vec{0}) \quad p_1^\mu = (E_\mu, \vec{p}_\mu) \quad p_2^\mu = (E_\nu, \vec{p}_\nu)$$

$$E_\mu = \frac{m_\pi^2 + m_\mu^2}{2m_\pi} = 109.8 \text{ MeV}$$

$$|\vec{p}_\mu| = \sqrt{E_\mu^2 - m_\mu^2} = 29.8 \text{ MeV}/c$$

# Lifetimes



## Decay time distribution

Mean lifetime

$$\tau = \langle d\Gamma/dt \rangle$$

aka proper time, eigen-time  
of a particle

$$\frac{d\Gamma}{dt} = \Gamma \exp\left(-\frac{t}{\tau}\right) \quad \Gamma = \frac{1}{\tau}$$

## Lifetime measurements

In laboratory frame

Decay Length

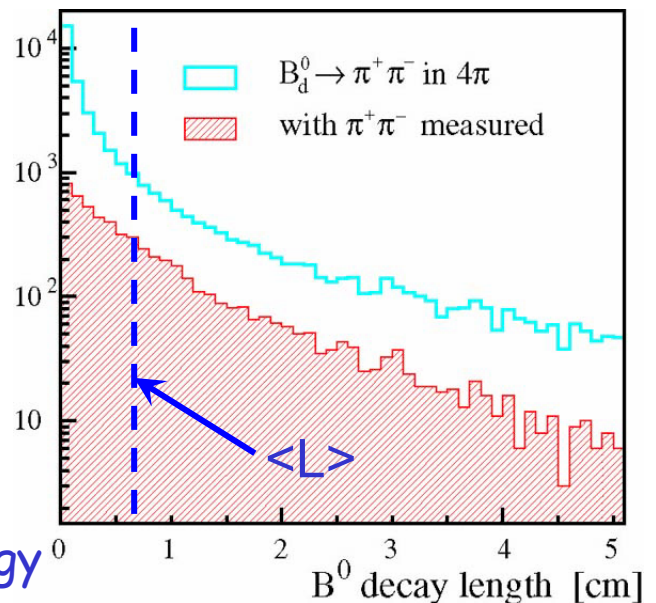
$$L = \gamma\beta c\tau$$

Example:  $B_d \rightarrow \pi^+\pi^-$   
in LHCb experiment  
 $\langle L \rangle \approx 7 \text{ mm}$

Average B meson energy

$$\langle E_B \rangle \approx 80 \text{ GeV}$$

$$\rightarrow \tau = 1.54 \text{ ps}$$



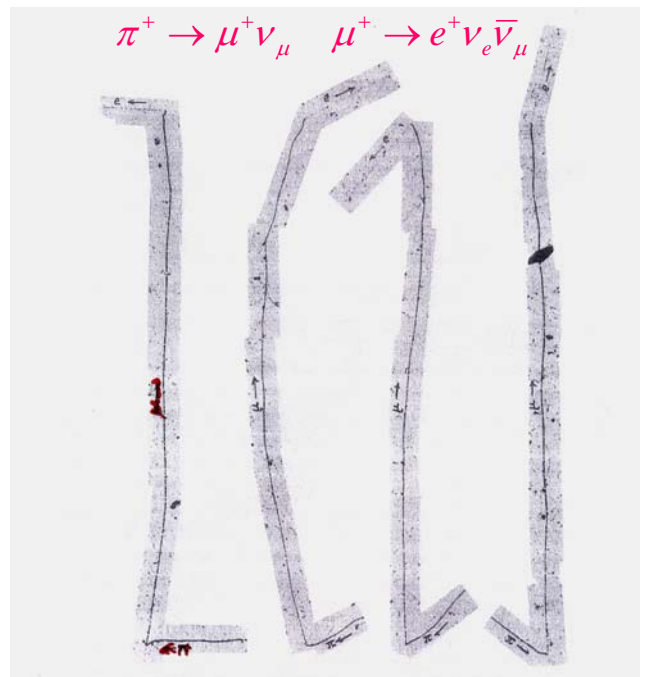
## Example: $\pi^+$ discovery

Decay sequence



Emulsions exposed to

Cosmic rays



# Resonances and Widths



## Strong Interactions

Production and decay of particles

**Lifetime**  $\tau \sim 10^{-23}$  s       $c\tau \sim O(10^{-15}$  m)

unmeasurable

## Heisenberg's Uncertainty Principle

$$\Delta E \Delta t \approx \hbar$$

Time and energy measurements are related

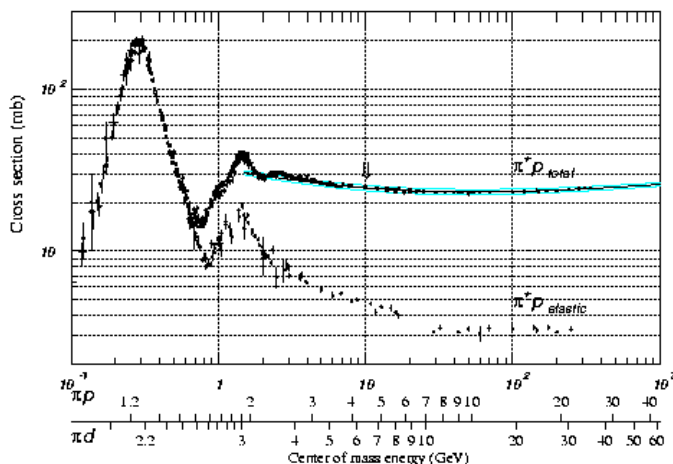
## Natural width

Energy width  $\Gamma$  and lifetime  $\tau$  of a particle

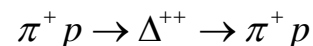
$$\Gamma = \hbar/\tau \rightarrow \text{Width } \Gamma = O(100 \text{ MeV})$$

measurable

## Example - Delta(1232) Resonance



Production



Peak at Energy

$E = 1.23 \text{ GeV}$

(Centre-of-Mass)

**Width**  $\Gamma = 120 \text{ MeV}$

**Lifetime**

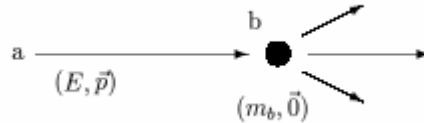
$$\tau = \hbar/\Gamma \approx 5 \cdot 10^{-24} \text{ s}$$

# Scattering



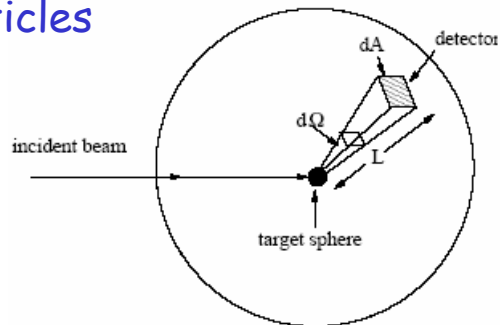
## Fixed Target Experiments

$$a + b \rightarrow c + d + \dots$$



- $n_a$  # of beam particles
- $v_a$  velocity of beam particles
- $n_b$  # of target particles per unit area

Incident flux  $F = n_a v_a$



## Cross Section

effective area of any scattering happening  
normalised per unit of incident flux  
depends on underlying physics

What you want to study

- $dN$  # of scattered particles in solid angle  $d\Omega$
- $d\sigma/d\Omega$  differential cross section in solid angle  $d\Omega$
- $\sigma$  total cross section

$$dN = n_a v_a n_b d\sigma = F n_b d\sigma = L d\sigma \Rightarrow \frac{d\sigma}{d\Omega} = \frac{1}{L} \frac{dN}{d\Omega}$$

$$L \text{ Luminosity} \quad N = \int \frac{d\sigma}{d\Omega} d\Omega \Rightarrow N = L \sigma$$

**Event Rate  $N$**

## Luminosity

Incident flux times number of targets

Depends on your experimental setup

1 barn = 1 b =  $10^{-24}$  cm<sup>2</sup>      Luminosity  $[L] = 10^{30...34}$  cm<sup>-2</sup>s<sup>-1</sup>

**Event Rate = Luminosity times Cross Section**

# Scattering



## Centre-of-Mass Energy

$$a + b \rightarrow c + d + \dots$$

Collision of two particles

$s$  is invariant quantity

Mandelstam variable

$$\begin{aligned} s &= (p_1^\mu + p_2^\mu)^2 = (E_1 + E_2)^2 - (\vec{p}_1 + \vec{p}_2)^2 \\ &= p_1^2 + p_2^2 + 2p_1 \cdot p_2 \\ &= m_1^2 + m_2^2 + 2(E_1 E_2 - |\vec{p}_1| |\vec{p}_2| \cos \theta) \end{aligned}$$

$$E_{\text{CoM}} = \sqrt{s} \quad \text{centre-of-mass energy}$$

Total available energy in centre-of-mass frame

$E_{\text{CoM}}$  is invariant in any frame, e.g. laboratory

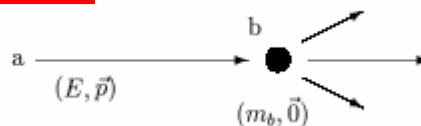
## Energy Threshold

for particle production

$$E_{\text{CoM}} = \sqrt{s} \geq \sum_{j=c,d,\dots} m_j$$

## Fixed Target Experiments

$$p_1^\mu = (E_{\text{lab}}, \vec{p}_1) \quad p_2^\mu = (m_2, \vec{0})$$



$$E_{\text{CoM}} = \sqrt{s} = \sqrt{m_1^2 + m_2^2 + 2E_{\text{lab}} m_2} \Rightarrow E_{\text{CoM}} \cong \sqrt{2E_{\text{lab}} m_2} \quad \text{if } E_{\text{lab}} \gg m_i$$

## Example:

100 GeV proton onto proton at rest

$$E_{\text{CoM}} = \sqrt{s} = \sqrt{(2E_p m_p)} = 14 \text{ GeV}$$

Most of beam energy goes into CoM momentum and is not available for interactions



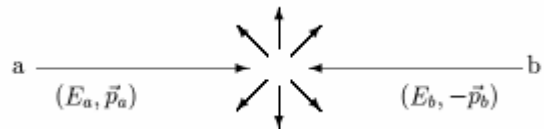
# Scattering



## Collider Experiments



Head-on collisions  
of two particles



$$\theta = 180^\circ$$

$$s = m_1^2 + m_2^2 + 2(E_1 E_2 - |\vec{p}_1| |\vec{p}_2| \cos \theta)$$

$$E_{\text{CoM}} = \sqrt{m_1^2 + m_2^2 + 2(E_1 E_2 + |\vec{p}_1| |\vec{p}_2|)} \Rightarrow E_{\text{CoM}} \cong \sqrt{4E_1 E_2} \text{ if } E_i \gg m_i$$

All of beam energy available for  
particle production

## Example

LEP - Large Electron Positron Collider at CERN

100 GeV  $e^-$  onto 100 GeV  $e^+$

Centre-of-mass energy

$$E_{\text{CoM}} = \sqrt{s} = 2E = 200 \text{ GeV}$$

Cross section

$$\sigma(e^+ e^- \rightarrow \mu^+ \mu^-) = 2.2 \text{ pb}$$

Luminosity

$$\int L dt = 400 \text{ pb}^{-1}$$

Number of recorded events  $N = \sigma \int L dt = 870$

# Conservation Laws



## Noether's Theorem

Every symmetry has associated with it a conservation law and vice-versa

## Energy and Momentum, Angular Momentum

conserved in all interactions

Symmetries - translations in space and time, rotations in space

## Charge conservation

Well established

$$|q_p + q_e| < 1.60 \cdot 10^{-21} e$$

Valid for all processes

Symmetry - gauge transformation

## Lepton and Baryon number (L and B)

$|L+B|$  conservation = matter conservation

Proton decay not observed (B violation)

Lepton family numbers  $L_e, L_\mu, L_\tau$  conserved

Symmetry - mystery

## Quark Flavours, Isospin, Parity

conserved in strong and electromagn processes

Violated in weak interactions

Symmetry - unknown

# Theoretical Concepts



## Standard Model of Particle Physics

### Quantum Field Theory (QFT)

Describes fundamental interactions of Elementary particles

Combines quantum mechanics and special relativity

Very small  
 $\Delta x \Delta p \approx \hbar c$

Very fast  
 $v \rightarrow c$

Classical Physics	Quantum mechanics
Special relativity	Quantum field theory

Natural explanation for antiparticles and for Pauli exclusion principle

Full QFT is beyond scope of this course

### Introduction to Major QFT concepts

Transition Rate

Matrix elements

Feynman Diagrams

Force mediated by exchange of bosons

# Klein-Gordon Equation



## Schrodinger Equation

For free particle

non-relativistic

1<sup>st</sup> order in time derivative

2<sup>nd</sup> order in space derivatives

not Lorentz-invariant

$$\frac{\hat{p}^2}{2m}\psi = \hat{E}\psi$$

$$-\frac{\hbar^2}{2m}\nabla^2\psi = i\hbar\frac{\partial}{\partial t}\psi$$

## Klein-Gordon (K-G) Equation

Start with relativistic equation

$$E^2 = p^2 + m^2 \quad (\hbar = c = 1) \quad E \rightarrow i\hbar\frac{\partial}{\partial t} \quad \vec{p} \rightarrow -i\hbar\vec{\nabla}$$

Apply quantum mechanical operators

$$\left(-\frac{\partial^2}{\partial t^2} + \vec{\nabla}^2\right)\psi = m^2\psi \quad \text{or} \quad \left(\frac{\partial^2}{\partial t^2} - \vec{\nabla}^2 + m^2\right)\psi = 0$$

2<sup>nd</sup> order in space and time derivatives

Lorentz invariant

Plane wave solutions of K-G equation

$$\psi(x^\mu) = N \exp(-ip_\nu x^\nu) \Rightarrow E = \pm\sqrt{p^2 + m^2}$$

→ negative energies ( $E < 0$ )

also negative probability densities ( $|\psi|^2 < 0$ )

## Negative Energy solutions

→ Dirac Equation, but -ve energies remain

→ Antimatter

# Klein-Gordon Equation



## Interpretation

K-G Equation is for **spinless particles**

Solutions are **wave-functions** for bosons

## Time-Independent Solution

Consider **static** case, i.e. no time derivative

$$\nabla^2 \psi = m \psi$$

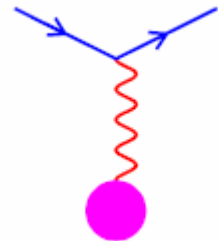
Solution is spherically symmetric

$$\psi(r) = -\frac{g^2}{4\pi r} \exp(-mr)$$

Interpretation - **Potential**

analogous to Coulomb potential

**Force is mediated by exchange of massive bosons**



## Yukawa Potential

Introduced to explain nuclear force

$$V(r) = -\frac{g^2}{4\pi r} \exp\left(-\frac{r}{R}\right) \quad R = \frac{\hbar}{mc}$$

**g** strength of force - "strong nuclear charge"

**m** mass of boson

**R** Range of force      see also nuclear physics

For  $m = 0$  and  $g = e \rightarrow$  Coulomb Potential

# Antiparticles



## Klein-Gordon & Dirac Equations

predict negative energy solutions

### Interpretation - Dirac

Vacuum filled with  $E < 0$  electrons

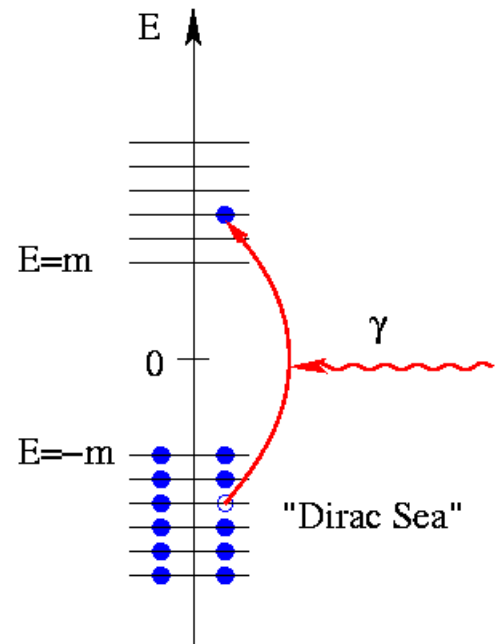
2 electrons with opposite spins per energy state - "Dirac Sea"

Hole of  $E < 0$ , -ve charge in Dirac sea  $\rightarrow$  antiparticle

$E > 0$ , +ve charge

$\rightarrow$  positron,  $e^+$  discovery (1931)

Predicts  $e^+e^-$  pair production and annihilation

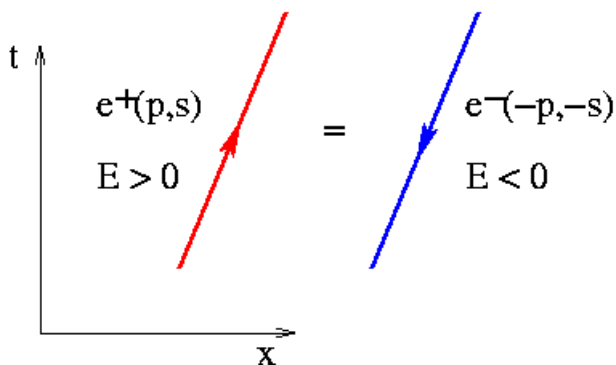


### Modern Interpretation - Feynman-Stueckelberg

$E < 0$  solutions: Negative energy particle

moving backwards in space and time correspond to

Feynman -Stueckelberg



### Antiparticles

Positive energy, opposite charge moving forward in space and time

$$\exp[-i((-E)(-t) - (-\vec{p}) \cdot (-\vec{x}))] = \exp[-i((Et - \vec{p} \cdot \vec{x}))]$$

# Scattering Amplitude



## Transition Rate $W$

Scattering reaction  $a + b \rightarrow c + d$

$$W = \sigma F$$

Interaction rate per target particle  
related to physics of reaction

## Fermi's Golden Rule

$$W = \frac{2\pi}{\hbar} |M_{fi}|^2 \rho_f$$

Matrix Element  $M_{fi}$   
scattering amplitude

Density  $\rho_f$   
# of possible final states  
"phase space"

non-relativistic

1<sup>st</sup> order time-dependent perturbation theory

see e.g. Halzen&Martin, p. 80, Quantum Physics

## Matrix Element

Contains all physics of the interaction

$$M_{fi} = \langle \psi_f | \hat{H} | \psi_i \rangle$$

Hamiltonian  $H$  is perturbation - 1<sup>st</sup> order

Incoming and outgoing plane waves

works if perturbation is small

Born  
Approximation

# Matrix Element



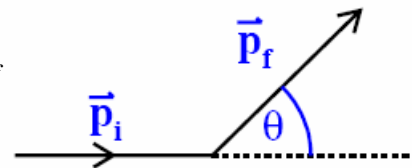
## Scattering in Potential

Example:  $e-p \rightarrow e-p$

Incoming and outgoing plane waves

Matrix element

Momentum transfer  $\vec{q} = \vec{p}_i - \vec{p}_f$



$$M_{fi} = \int \psi_f^* V(\vec{r}) \psi_i d^3\vec{r}$$

$$= \frac{1}{N^2} \int \exp(-i\vec{p}_f \cdot \vec{r}) V(\vec{r}) \exp(i\vec{p}_i \cdot \vec{r}) d^3\vec{r}$$

$$= \frac{1}{N^2} \int \exp(i\vec{q} \cdot \vec{r}) V(\vec{r}) d^3\vec{r} \quad \vec{q} = \vec{p}_i - \vec{p}_f$$

$M_{fi}(\vec{q})$  is Fourier transform of Potential  $V(\vec{r})$

## Scattering in Yukawa Potential $V(r) = -\frac{g^2}{4\pi r} \exp(-mr)$

$$M_{fi} = -\frac{g^2}{4\pi} \int_0^\infty \int_0^\pi \int_0^{2\pi} \exp(i|\vec{q}|r \cos \theta) \frac{\exp(-mr)}{r} r^2 dr \sin \theta d\theta d\phi$$

$$= -\frac{g^2}{2i|\vec{q}|} \int_0^\infty (\exp(i|\vec{q}|r) - \exp(-i|\vec{q}|r)) \exp(-mr) dr$$

$$M_{fi} = -\frac{g^2}{(m^2 + |\vec{q}|^2)}$$

**Propagator**

term in  $M_{fi}$   $1/(m^2 + q^2)$

**Cross section**  $\frac{d\sigma}{d\Omega} \propto |M|^2 \propto \frac{1}{(m^2 + |\vec{q}|^2)^2} \Rightarrow \frac{d\sigma}{d\Omega} \propto \frac{1}{q^4} \quad m=0$

Result still holds relativistically

4-momentum transfer  $q^\mu = (E_i - E_f, \vec{p}_i - \vec{p}_f)$