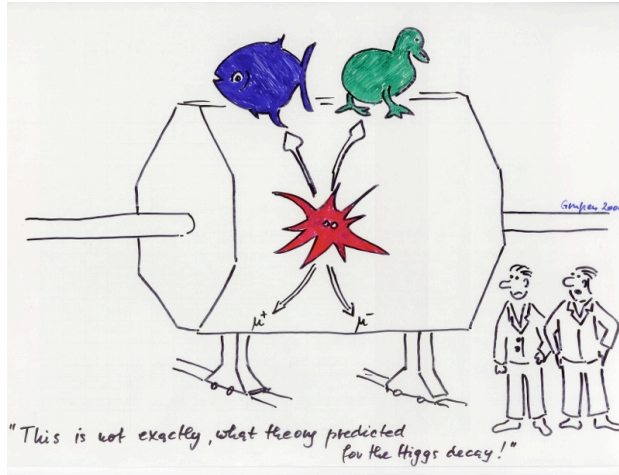


Physics 3: Particle Physics

Lectures 10. Beyond The Standard Model March 13th 2008



- * Neutrinos
- * Supersymmetry
- * Unification

1

Neutrinos

Pauli hypothesised the neutrino to explain the nuclear beta decay energy spectrum

- $n \rightarrow p e^- \bar{\nu}_e$

$$p_e = p_n - p_p - p_{\bar{\nu}}$$

The recoil of the nucleus is very small $\Rightarrow E_e \approx m_n c^2 - m_p c^2 - E_{\bar{\nu}}$

Neutrino means “little neutral one”. They have:

- no colour charge
- no electric charge
- only interact via the weak force: very small interaction rate.

Pauli: “I have done a terrible thing. I have postulated a particle that cannot be detected.”

Neutrinos can be detected via inverse beta decay e.g:

$$\begin{aligned} \nu_e + n &\rightarrow p + e^- \\ \bar{\nu}_e + p &\rightarrow n + e^+ \end{aligned}$$

Tiny cross section:
 $\sigma(\bar{\nu}_e + p \rightarrow n + e^+) \approx 5 \times 10^{-20} \left(\frac{E_{\bar{\nu}}}{\text{MeV}} \right)^2 \text{ barn}$

Mean free path of neutrinos in water is ~60 light years for a 1 MeV anti-neutrino!

2

Neutrino Sources & Detectors

Where can we find neutrinos?

The Sun

- Standard Solar Model predicts rates and energy spectra for ν_e

Cosmic rays

- $\pi^+ \rightarrow \mu \nu_\mu; \mu^+ \rightarrow e^+ \nu_e \bar{\nu}_\mu$

Nuclear Reactors

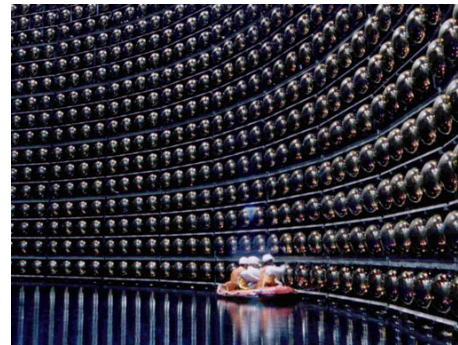
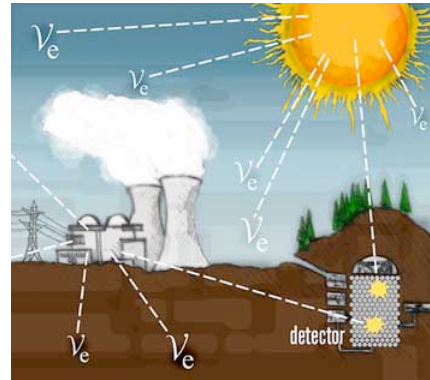
- e.g. $n + {}^{235}\text{U} \rightarrow {}^{144}\text{Ba} + {}^{89}\text{Kr} + 3n + 3\bar{\nu}_e$

Accelerators

- Decay of muons: $\mu^- \rightarrow e^- \bar{\nu}_e \nu_\mu$

To detect neutrinos:

- Get a lot of stuff
e.g. water, cleaning fluid, steel...
- Leave in a area of neutrino flux
- Be patient.



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Solar Neutrinos

Solar neutrinos were first detected during the 1970s in the Homestake mine, South Dakota

- 100,000 gallons of C_2Cl_4 (dry cleaning fluid)



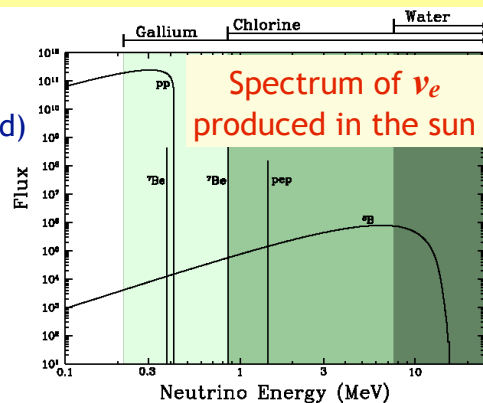
Standard Solar Model predicts the flux of neutrinos to produce 1.5 Ar atoms/day

- Only 0.5 Ar atoms/day found

Either:

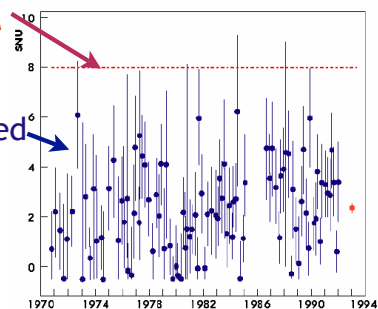
- Standard Solar Model is wrong
- we don't know the cross section for neutrino scattering
- electron neutrinos are going missing between creation in the sun and detection on the earth

Final solution now accepted:
 ν_e turn into ν_μ and ν_τ within sun!



Prediction of SSM

Measured



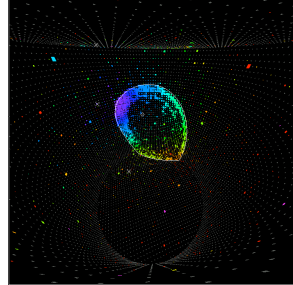
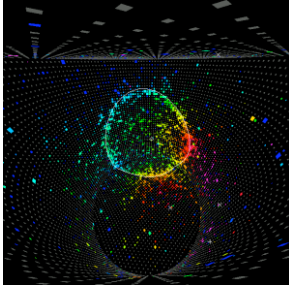
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Atmospheric Neutrinos

The Super-Kamiokande experiment in Japan detects atmospheric neutrinos. Reactions are:

$$\nu_e + N \rightarrow e^- + X$$

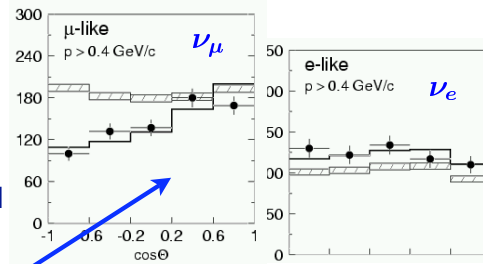
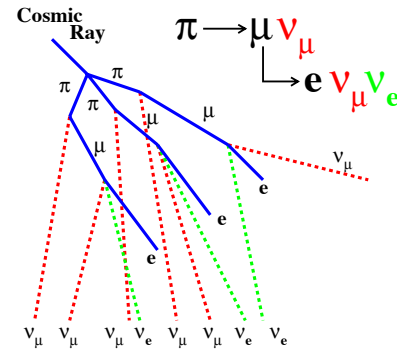
$$\nu_\mu + N \rightarrow \mu^- + X$$



As most neutrinos don't interact in the earth neutrinos produced in every θ direction around the earth are detected.

Observed: deficit of ν_μ from other side of earth.

- Interpreted as $\nu_\mu \rightarrow \nu_\tau$ OSCILLATIONS in the earth's crust.
- Different flavours of neutrinos mix and neutrinos have mass!



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Neutrino Masses

Both solar and atmospheric neutrino experiments provide evidence for neutrinos changing flavour as they travel.

(i.e. L_e, L_μ, L_τ not perfectly conserved!

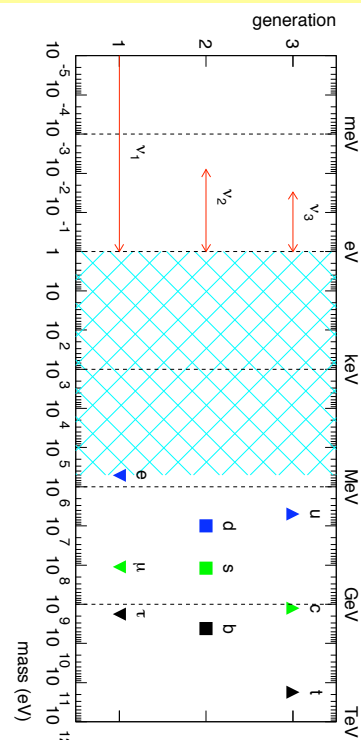
Net number of leptons *is* conserved.)

Neutrino flavour change is only possible if neutrinos have mass!

- Neutrino masses are not allowed in the Standard Model

Many experiments now underway with accelerator and reactor neutrinos. → Our understanding of neutrinos is rapidly evolving!

Neutrino masses provided the first evidence for particles physics beyond the Standard Model.



6

Supersymmetry

Particle physicists are hard to please...

Some of us would like extra fundamental particles, related to the observed particles by supersymmetry

Supersymmetry is symmetry between fermions and bosons:

- Every fundamental fermion has a boson partner. e.g.:

$$u \leftrightarrow \tilde{u}$$

up-quark $\leftrightarrow$ up squark

$$\tau \leftrightarrow \tilde{\tau}$$

tau-lepton $\leftrightarrow$ tau-slepton

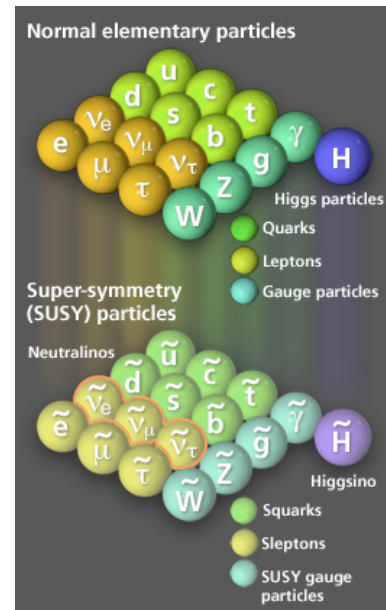
- Every boson (W, Z, γ , Higgs) has a fermion partner e.g.:

$$W \leftrightarrow \tilde{W}$$

W-boson $\leftrightarrow$ Wino

$$H \leftrightarrow \tilde{H}$$

Higgs $\leftrightarrow$ Higgsino



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Unification of the Forces

Maxwell equations' show that electric and magnetic interactions are manifestations of one interaction: **electromagnetism**.

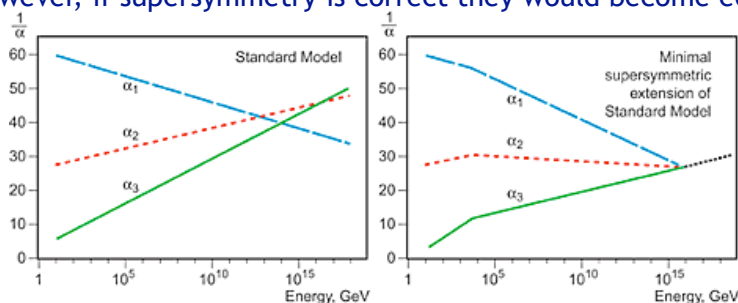
We've seen electromagnetic and weak interactions are manifestations of one underlying interaction: **the electroweak interaction**.

- Consequences of unification become apparent at high energies, $E \sim O(m_Z, m_W)$.

Maybe the electroweak and strong forces are a manifestation of a unified force? This would only become apparent at higher energies.

The couplings of the **electromagnetic**, **weak** and **strong** interactions change as a function of energy, but never become equal...

However, if supersymmetry is correct they would become equal



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And what about gravity?

The ultimate unification of the forces should include gravity.

- But gravity is very much weaker than the other forces...

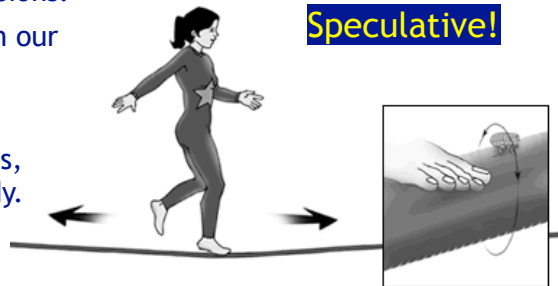
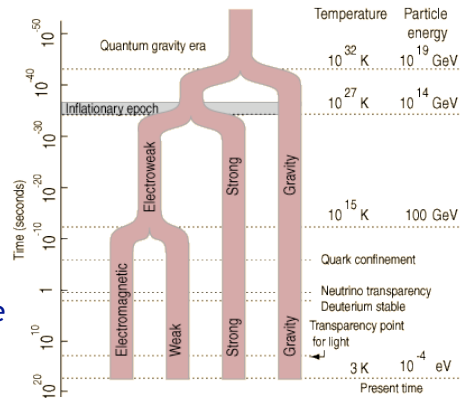
Many ideas proposed to explain this.

e.g. Extra dimensions

- Most particles (and us) can only travel in the regular 3 space + 1 time dimensions
- Gravitons - the bosons which propagate gravity - can travel in the extra dimensions.
- Strength of gravity is naturally weaker in our dimensions

$$F(r) = G \frac{m_1 m_2}{r^2} \rightarrow G_{\text{new}} \frac{m_1 m_2}{r^4}$$

- They have to be small extra dimensions, otherwise we'd have seen them already.
- If the dimensions are big enough we might see their effects at the LHC!



9

Summary: Known knowns and known unknowns

The Standard Model is a beautiful theory of (almost) all the measurements we see in particle physics.... But it isn't the whole picture.		"We can explain everything, but we understand (at a fundamental level) almost nothing!"
Neutrinos flavours mix: $\nu_e \leftrightarrow \nu_\mu \leftrightarrow \nu_\tau$ Only possible if neutrinos have mass!	Supersymmetry is one popular theory for physics beyond the Standard Model.	Supersymmetry provides a candidate particle for dark matter.
Ultimately we think the electroweak, strong and gravitational forces should be described by one underlying interaction.		Lots of fun to come!

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