

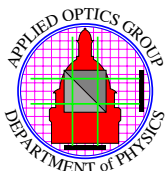


Topic 2: Scalar Diffraction

Aim: Covers Scalar Diffraction theory to derive Rayleigh-Sommerfeld diffraction. Take approximations to get Kirchhoff and Fresnel approximations.

Contents:

1. Preliminary Theory.
2. General propagation between two planes.
3. Kirchhoff Diffraction
4. Fresnel Diffraction
5. Summary





Scalar Wave Theory

Light is really a vector electromagnetic wave with \mathbf{E} and \mathbf{B} field linked by Maxwell's Equations.

Full solution only possible in limited cases, so we have to make assumptions and approximations.

Assume:

Light field can be approximated by a complex scalar potential. (amplitude)

Valid for:

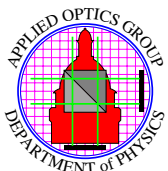
Apertures and objects $\gg \lambda$, (most optical systems).

NOT Valid for:

Very small apertures, Fibre Optics, Planar Wave Guides, Ignores Polarisation.

Also Assume:

Scalar potential is a *linear* super-position of monochromatic components. So theory only valid for Linear Systems, (refractive index does not depend on wavelength).



Scalar Potentials

For light in free space the \mathbf{E} and \mathbf{B} field are linked by

$$\begin{aligned}\nabla \cdot \mathbf{E} &= 0 \\ \nabla \cdot \mathbf{B} &= 0 \\ \nabla \wedge \mathbf{B} &= \epsilon_0 \mu_0 \frac{\partial \mathbf{E}}{\partial t} \\ \nabla \wedge \mathbf{E} &= -\frac{\partial \mathbf{B}}{\partial t}\end{aligned}$$

which, for free space, results in the “Wave Equation” given by

$$\nabla^2 \mathbf{E} \Leftrightarrow \frac{1}{c^2} \frac{\partial^2 \mathbf{E}}{\partial t^2} = 0$$

Assume light field represented by scalar potential $\Phi(\mathbf{r}, t)$ which MUST also obey the “Wave Equation”, so:

$$\nabla^2 \Phi \Leftrightarrow \frac{1}{c^2} \frac{\partial^2 \Phi}{\partial t^2} = 0$$

Write the Component of scalar potential with angular frequency ω as

$$\Phi(\mathbf{r}, t) = u(\mathbf{r}) \exp(i\omega t)$$

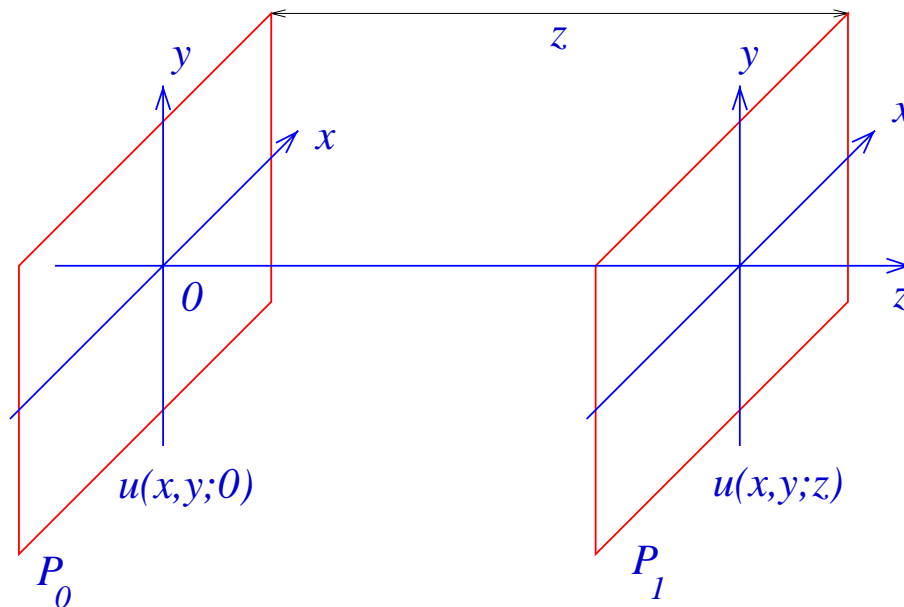
then substituting for Φ we get that,

$$[\nabla^2 + \kappa^2]u(\mathbf{r}) = 0$$

where $\kappa = 2\pi/\lambda$ or *wave number*.

So that $u(\mathbf{r})$ must obey Helmholtz Equation, (starting point for Scalar Wave Theory).

Diffraction Between Planes



In P_0 the **2-D** scalar potential is:

$$u(x, y; 0) = u_0(x, y)$$

where in plane P_1 the scalar potential is:

$$u(x, y; z)$$

where the planes are separated by distance z .

Problem: Given $u_0(x, y)$ in plane P_0 we want to calculate in $u(x, y; z)$ is any plane P_1 separated from P_0 by z .

Looking for a **2-D Scalar Solution** to the full wave equation.

Fourier Transform Approach

Take the **2-D Fourier Transform** in the plane, wrt x, y . so in plane P_1 we get

$$U(u, v; z) = \iint u(x, y; z) \exp(i2\pi(ux + vy)) dx dy$$

since we have that

$$u(x, y; z) = \iint U(u, v; z) \exp(i2\pi(ux + vy)) du dv$$

then if we know $U(u, v; z)$ in any plane, then we can easily find $u(x, y; z)$ the required amplitude distribution.

The amplitude $u(x, y; z)$ is a Linear Combination of term of

$$U(u, v; z) \exp(i2\pi(ux + vy))$$

These terms are **Orthogonal** (from Fourier Theory),

So **each** of these terms **must** individually obey the Helmholtz Equation.

Note: the term:

$$U(u, v; z) \exp(i2\pi(ux + vy))$$

in a **Plane Wave** with Amplitude $U(u, v; z)$ and Direction $(u/\kappa, v/\kappa)$, where $\kappa = 2\pi/\lambda$.

Propagation of a Plane Wave

From Helmholtz Equation we have that

$$[\nabla^2 + \kappa^2]U(u, v; z) \exp(i2\pi(ux + vy)) = 0$$

Noting that the $\exp()$ terms cancel, then we get that

$$\frac{\partial^2 U(u, v; z)}{\partial z^2} + 4\pi^2 \left(\frac{1}{\lambda^2} \Leftrightarrow u^2 \Leftrightarrow v^2 \right) U(u, v; z) = 0$$

so letting

$$\gamma^2 = \frac{1}{\lambda^2} \Leftrightarrow u^2 \Leftrightarrow v^2$$

then we get the relation that

$$\frac{\partial^2 U}{\partial z^2} + (2\pi\gamma)^2 U = 0 \quad \forall u, v$$

With the condition that

$$U(u, v; 0) = U_0(u, v) = F \{u_0(x, y)\}$$

we get the solution that

$$U(u, v; z) = U_0(u, v) \exp(i2\pi\gamma z)$$

This tells us how each component of the Fourier Transform propagates between plane P_0 and plane P_1 , so:

$$u(x, y; z) = \iint U_0(u, v) \exp(i2\pi\gamma z) \exp(i2\pi(ux + vy)) du dv$$

which is a general solution to the propagation problems valid for **all** z .

Free Space Propagation Function

Each Fourier Component (or Spatial Frequency) propagates as:

$$U(u, v; z) = U_0(u, v) \exp(i2\pi\gamma z)$$

Define: *Free Space Propagation Function*

$$H(u, v; z) = \exp(i\pi\gamma z)$$

so we can write:

$$U(u, v; z) = U_0(u, v)H(u, v; z)$$

Look at the form of $H(u, v; z)$.

$$u^2 + v^2 + \gamma^2 = \frac{1}{\lambda^2}$$

Case 1: If $u^2 + v^2 \leq 1/\lambda^2$ then γ is **REAL**

$$H(u, v) \text{ Phase Shift}$$

so **all** spatial frequency components passed with Phase Shift.

Case 2: If $u^2 + v^2 > 1/\lambda^2$ then γ is **IMAGINARY**

$$H(u, v) = \exp(\Leftrightarrow 2\pi|\gamma|z)$$

So Fourier Components of $U_0(u, v)$ with $u^2 + v^2 > 1/\lambda^2$ decay with Negative Exponential. (evanescent wave)

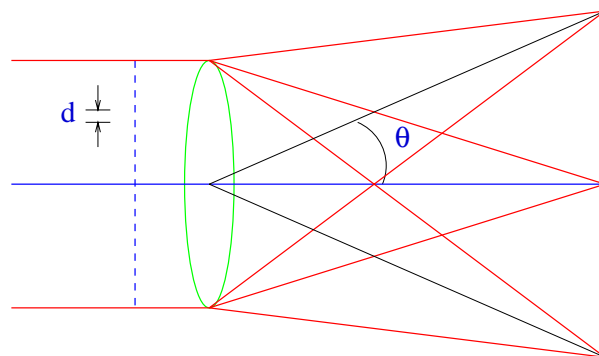
Frequency Limit for Propagation

In plane P_1 where $z \gg \lambda$ then the negative exponential decay will remove high frequency components.

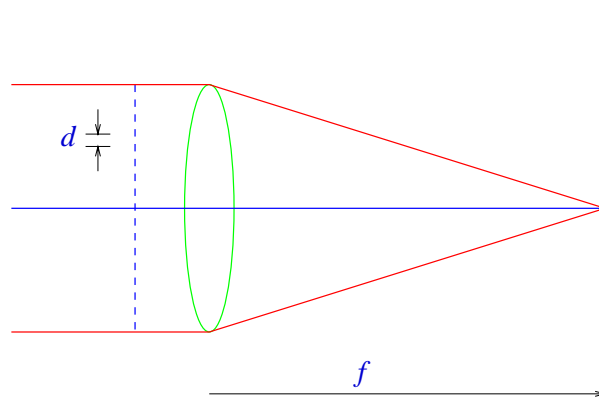
$$U(u, v; z) = 0 \quad u^2 + v^2 > 1/\lambda^2$$

so Fourier Transform of $u(x, y; z)$ is of limited extent.

Maximum spatial frequency when $u = 1/\lambda$, this corresponds to a grating with period λ .



$$d \sin \theta = n\lambda \quad \text{for } d > \lambda$$



No diffraction when $d < \lambda$. Information not transferred to plane P_1 .

Convolution Relation

We have that

$$U(u, v; z) = H(u, v; z) U_0(u, v; z)$$

so by the **Convolution Theorem**, we have that

$$u(x, y; z) = h(x, y; z) \odot u_0(x, y; z)$$

With $u(x, y; z)$ the distribution in P_1 due to $u_0(x, y; z)$ in P_0 .

We then have that

$$h(x, y; z) = \iint_{u^2+v^2 < 1/\lambda^2} \exp(i2\pi\gamma z) \exp(i2\pi(ux + vy)) \, dudv$$

where

$$\gamma = \sqrt{\frac{1}{\lambda^2} \Leftrightarrow u^2 + v^2}$$

“It-Can-Be-Shown” (with considerable difficulty), that

$$h(x, y; z) = \Leftrightarrow \frac{2\pi}{\lambda^2} \frac{\partial}{\partial z} \left(\frac{\exp(i\kappa r)}{\kappa r} \right)$$

where we have that

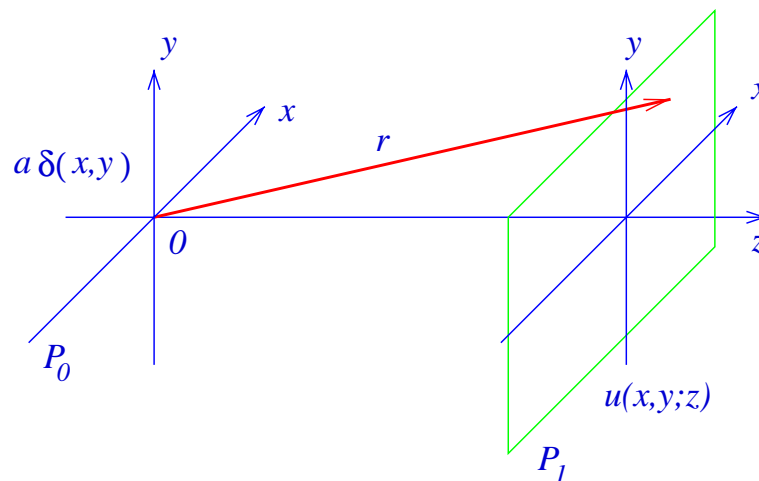
$$r^2 = x^2 + y^2 + z^2 \quad \text{and} \quad \kappa = \frac{2\pi}{\lambda}$$

we therefore get that

$$h(x, y; z) = \frac{1}{\lambda} \left[\frac{1}{\kappa r} \Leftrightarrow i \right] \left(\frac{z}{r} \right) \frac{\exp(i\kappa r)}{r}$$

which is known as “The Impulse Response Function for Free Space Propagation”

Point Object



In P_0 we have a Delta Function, so:

$$u_0(x,y) = a\delta(x,y)$$

So in plane P_1 we have

$$\begin{aligned} u(x,y;z) &= ah(x,y;z) \\ &= \frac{a}{\lambda} \left[\frac{1}{\kappa r} \Leftrightarrow i \right] \left(\frac{z}{r} \right) \frac{\exp(i\kappa r)}{r} \end{aligned}$$

where r is the distance from:

$$(0,0;0) \Rightarrow (x,y;z)$$

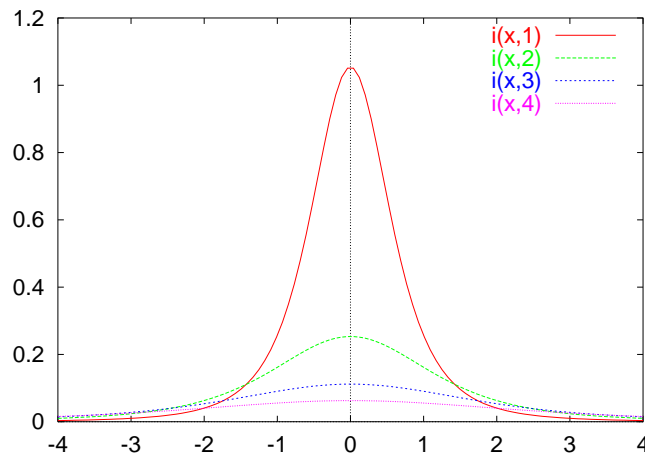
So the intensity in plane P_1 is given by:

$$i(x,y;z) = \frac{b^2}{r^2} \left[\frac{1}{\kappa^2 r^2} + 1 \right] \left(\frac{z}{r} \right)^2$$

where the λ^2 has been incorporated into the constant b .

Shape of Function

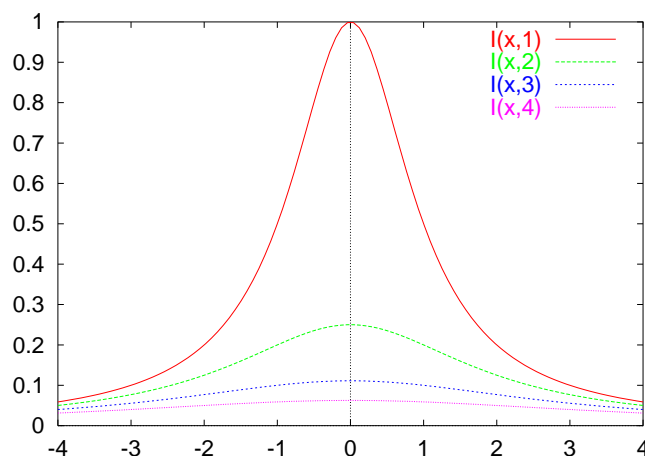
The shape of $i(x, 0; z)$ is shown below:



for $z = \lambda, 2\lambda, 3\lambda, 4\lambda$, x -scale in λ s.

Compare with a isolated 3-D point source, spherical expanding waves, so intensity in plane at distance z of:

$$I(x, y; z) = \frac{b^2}{r^2}$$



Note We have a 2-D Delta Function, (whole in a screen), and NOT an 3-D point source.

Full Expression

The full convolution expression is

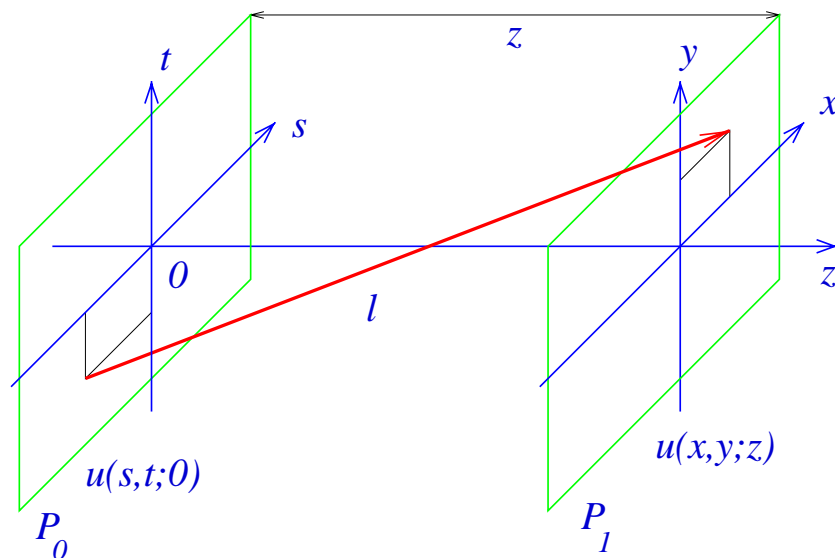
$$u(x, y; z) = \iint_{P_0} u_0(s, t) h(x \Leftrightarrow s, y \Leftrightarrow t; z) ds dt$$

where s, t are variables in plane P_0

The $h()$ term will contain terms of the form

$$(x \Leftrightarrow s)^2 + (y \Leftrightarrow t)^2 + z^2 = l^2$$

so l is Distance from $(s, t; 0) \Rightarrow (x, y; z)$



Full Expression is

$$u(x, y; z) = \frac{1}{\lambda} \iint_{P_0} u_0(s, t) \left[\frac{1}{\kappa l} \Leftrightarrow l \right] \frac{\exp(i\kappa l)}{l} \left(\frac{z}{l} \right) ds dt$$

Rayleigh-Sommerfeld Diffraction Equation

Kirchoff Diffraction

Look at Impulse Response Function:

$$h(x, y; z) = \frac{1}{\lambda} \left[\frac{1}{\kappa r} \Leftrightarrow i \right] \left(\frac{z}{r} \right) \frac{\exp(i\kappa r)}{r}$$

Most Practical cases, P_0 and P_1 separated by **MANY** wavelength,

$$\Rightarrow z \gg \lambda \quad \Rightarrow r \gg \lambda$$

Approximate the term:

$$\left[\frac{1}{\kappa r} \Leftrightarrow i \right] \approx \Leftrightarrow i$$

so that

$$h(x, y; z) \approx \frac{1}{i\lambda} \left(\frac{z}{r} \right) \frac{\exp(i\kappa r)}{r}$$

Look at Terms

$$\frac{\exp(i\kappa r)}{r}$$

Spherically expanding wave from point $(0, 0; 0)$

$$\frac{z}{r}$$

“**Obliquity Factor**” which forces propagation in z direction. (In Goodman’s notation the obliquity is written as a $\cos(\)$ term).

The other terms are just constants.



Note: For 3-D point source, get expanding spherical wave obliquity factor results from 2-D source in a plane.

Impulse Response Function: Spherical Wave with directional weighting term.

Model: Each point in P_0 acts a source of impulse response functions, that sum is P_1 .

$$h(x, y; z) = \text{Spherical Wave} \begin{bmatrix} z \\ r \end{bmatrix}$$

This is **Hygen's Secondary Wavelet** (weighted by obliquity factor).

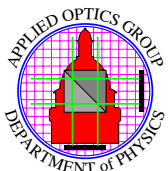
Hygen's Secondary Wavelets

Model: Each point on the wavefront gives rise to Spherical Waves,

$$h(x, y; z) = \frac{\exp(i\kappa r)}{r}$$

Add postulate that "wave propagate in positive z direction".

$$\text{Kirchhoff} \Leftrightarrow \text{Hygen's} \begin{bmatrix} z \\ r \end{bmatrix}$$



Kirchhoff Diffraction Integral

In Rayleigh-Sommerfeld integral,

$$z \gg \lambda \Rightarrow l \gg \lambda$$

make the approximation that

$$\left[\frac{1}{\kappa l} \Leftrightarrow l \right] \approx \Leftrightarrow$$

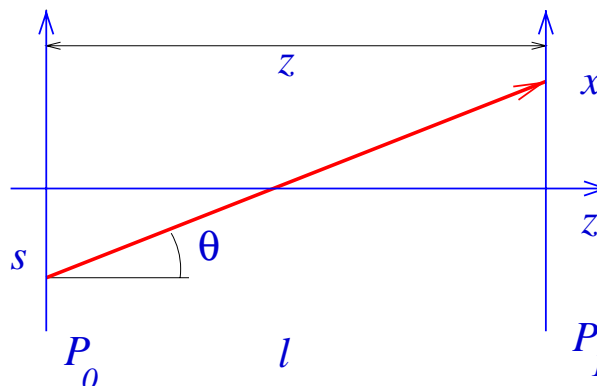
so we get that

$$u(x, y; z) = \frac{1}{i\lambda} \iint_{P_0} u_0(s, t) \frac{\exp(i\kappa l)}{l} \left(\frac{z}{l} \right) ds dt$$

which is valid (1%), $z > 20\lambda$.

Typical starting point for optical calculations.

Look at z/l factor:



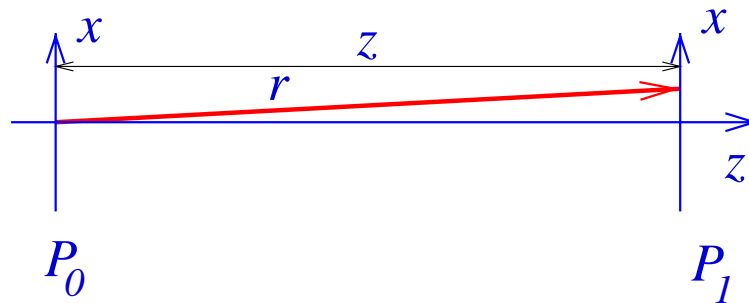
so that

$$\frac{z}{l} = \cos \theta$$

Same expression as in books [eg. Goodman page 52, equation (3-51)]

Fresnel Approximation

Assume extend of P_0 and $P_1 \ll z$



So we have that

$$|x| \& |y| \ll z$$

We have Kirchhoff impulse response

$$h(x, y; z) = \frac{1}{i\lambda} \left(\frac{z}{r}\right) \frac{\exp(i\kappa hr)}{r}$$

Which we can write as:

$$h(x, y; z) = A(x, y; z) \exp(i\kappa r)$$

where the amplitude

$$A(x, y; z) = \frac{1}{i\lambda} \frac{z}{r^2}$$

Since $x \& y \ll z$, we can expand r as

$$r = z \left(1 + \frac{x^2 + y^2}{z^2} \right)^{\frac{1}{2}} \approx z + \frac{x^2 + y^2}{2z}$$

Amplitude Term: First Order approx:

$$r \approx z \Rightarrow A(x, y; z) \approx \frac{1}{i\lambda z}$$

Phase Term: Second Order approx:

$$r \approx z + \frac{x^2 + y^2}{2z}$$

$$\exp(i\kappa r) \approx \exp(i\kappa z) \exp\left(i\frac{\kappa}{2z}(x^2 + y^2)\right)$$

So the Fresnel Approximation is that

$$h(x, y; z) \approx \frac{\exp(i\kappa z)}{i\lambda z} \exp\left(i\frac{\kappa}{2z}(x^2 + y^2)\right)$$

Fresnel Approximations

- 1) Replace Spherical waves by Parabolic Waves
- 2) Ignore Obliquity factor.

NOTE: only valid

$$|x| \text{ \& \ } |y| \ll z \Rightarrow \text{Small Objects}$$

This is also known as “Paraxial Approximation”. Useful in many practical systems.

Use in the rest of the course

Fresnel Diffraction

We have that

$$\begin{aligned}
 u(x, y; z) &= h(x, y; z) \odot u_0(x, y) \\
 &= \frac{\exp(i\kappa z)}{i\lambda z} \iint u_0(s, t) \exp \left[i \frac{\kappa}{2z} ((x \Leftrightarrow s)^2 + (y \Leftrightarrow t)^2) \right] ds dt
 \end{aligned}$$

This we can expand to get the

$$\underbrace{\frac{1}{i\lambda z} \exp(i\kappa z)}_1 \underbrace{\exp \left[i \frac{\kappa}{2z} (x^2 + y^2) \right]}_2 \times \underbrace{\iint u_0(s, t) \exp \left[i \frac{\kappa}{2z} (s^2 + t^2) \right]}_4 \underbrace{\exp \left[\Leftrightarrow i \frac{\kappa}{z} (sx + ty) \right]}_3 ds dt$$

Look the **FOUR** terms

1. Absolute amplitude and phase, depends only on z (constant which is not normally important).
2. Parabolic phase term, no effect on intensity.
3. Fourier Transform scaled by κ/z
4. Scalar distribution in P_0 weighted by parabolic phase term.

Fresnel Diffraction \Rightarrow Fourier Transform weighted by parabolic phase term [+ Extra phase and constants]

In Fourier Space

In Fourier Space we have

$$U(u, v; z) = H(u, v; z) U_0(u, v)$$

so we have that

$$H(u, v; z) = \frac{1}{i\lambda z} \exp(i\kappa z) \iint \exp\left[i\frac{\kappa}{2z}(x^2 + y^2)\right] \exp(i2\pi(ux + vy)) dx dy$$

This can be separated into 2 integrals in x and y , and with identity

$$\int_{-\infty}^{\infty} \exp(i\pi bx^2) \exp(i\pi ax) dx = \sqrt{\frac{\pi}{b}} \exp\left[i\pi \frac{a^2}{4b}\right]$$

It-Can-be-Shown that

$$H(u, v; z) = \exp(i\kappa z) \exp(i\pi\lambda z(u^2 + v^2))$$

which is again a parabolic term.

Fresnel diffraction \Rightarrow Multiplication in Fourier plane by parabolic phase term.



Summary

Scalar Wave solution for propagation between two plane:

Full Solution \Rightarrow **Rayleigh-Sommerfeld**

If distance between planes $>$ a “few wavelengths” then

Scalar Kirchoff Diffraction

This is Hygen’s Secondary Spherical wavelets plus obliquity factor due to 2-D whole in plane.

If distance between planes *Large* and planes are *Small* make small angle approximation to get

Fresnel Diffraction

which replaces Spherical waves with Parabolic and ignores obliquity factor.

Fresnel Diffraction will be used for the rest of this course.

